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Interpreting spatial variations in anisotropy: Insights into the Main Ethiopian Rift from SKS waveform modelling

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SUMMARY

Seismic anisotropy is a common feature in the upper mantle and measuring shear-wave splitting in core phases is a common approach in estimating its characteristics. Large lateral variations in estimated splitting parameters are observed over small spatial distances in many differing tectonic regions, including areas of continental break-up such as the Main Ethiopian Rift (MER). We investigate the ability of shear-wave splitting analysis to constrain spatial variations in anisotropy using a one-way wave equation modelling scheme to generate band-limited waveforms for a suite of models representing regions with rapidly changing anisotropy. We show that shear-wave splitting can identify lateral variation in anisotropy on the order of 20-50 km, where a change in fast direction demarcates the transition in anisotropy. Additionally, variation in the amount of splitting is complicated close to the transition, and is sensitive to the vertical thickness of anisotropy. We have used these modelling results to interpret shear-wave splitting measurements for the Main Ethiopian Rift. The model that best fits the observations has a 100 km wide rift zone with a fast direction of 30° outside and 20° inside the rift. The model has 9% anisotropy close to the western margin, with 7% anisotropy elsewhere. In all regions of the model we constrain the anisotropy to begin at a depth of 90 km. The depth of anisotropy is consistent with geochemical estimates of the depth of melt initiation beneath

18 the region. Also the elevated splitting beneath the western margin supports evidence of low
19 velocities and highly conductive zones from seismic tomography and magneto-tellurics, sug-
20 gesting melt is more focused along the western margin. This study shows how observations of
21 *SKS*-wave splitting from dense seismic networks can be used to map sharp lateral changes and
22 constrain the depth of the anisotropy.

23 **Key words:**

24 1 INTRODUCTION

25 Seismic anisotropy can be described as the variation of seismic wave speed with direction of prop-
26 agation. In most studies the main cause of anisotropy in the upper mantle is assumed to be the
27 lattice preferred orientation (LPO) of olivine where the olivine fast axis (a-axis) aligns in the di-
28 rection of upper-mantle flow (Babuska & Cara, 1991; Mainprice et al., 2000). This could be caused
29 by current mantle processes, or due to accumulated strain which has ‘frozen’ in an anisotropic sig-
30 nature from previous deformation events. Other mechanisms that cause upper mantle anisotropy
31 are fluid filled cracks (Crampin & Booth, 1985) or the preferred orientation of inclusions (e.g.,
32 oriented melt pockets (OMP), see Kendall, 1994; Blackman & Kendall, 1997) mechanisms that
33 can be very efficient at generating large amounts of anisotropy (Kendall, 2000).

34 A common way of constraining anisotropy in the upper mantle is shear-wave splitting analysis.
35 When a shear-wave enters an anisotropic medium it splits into two quasi-shear waves that are
36 polarised orthogonally to each other and propagate with different velocities. These split shear-
37 waves can be used to characterise anisotropy in terms of an apparent symmetry axis (typically
38 fast shear-wave direction, ϕ) and the time-lag between fast and slow shear waves (δt , a proxy for
39 amount, or extent of anisotropy).

40 Many studies investigating upper mantle anisotropy based on shear-wave splitting utilise *SKS*-
41 phases. This is a wave that travels as an *S*-wave through the mantle and a *P*-wave through the outer
42 core. It is advantageous to use this phase because it is a clear arrival over a range of epicentral
43 distances (85° - 120°), making it observable in most regions. Also, it is possible to ignore source
44 side anisotropy due to the fact that the seismic energy converts to a *P*-wave at the CMB, thus the

45 measured anisotropy can be attributed to the mantle beneath the station. Another benefit is that
46 the rays travel almost vertically in the uppermost mantle making *SKS*-wave splitting a useful tool
47 to distinguish lateral variations in anisotropy (see Silver, 1996; Savage, 1999; Kendall, 2000, for
48 examples). However, due to the near vertical raypaths it is very hard to place any constraints on the
49 depth extent of anisotropy, which has led to much debate on whether anisotropy can be attributed to
50 lithospheric fabric (fossil anisotropy (Silver, 1996), fluid filled cracks (Crampin & Booth, 1985)),
51 or asthenospheric processes (for example, flow at plate boundaries (Blackman & Kendall, 2002),
52 simple asthenospheric flow (Savage, 1999), density driven flow (Behn et al., 2004)).

53 Some attempt has been made to constrain the depth extent of anisotropy based on consideration
54 of *SKS* Fresnel zones (Alsina & Snieder, 1995; Rumpker & Ryberg, 2000). For instance, it can be
55 assumed that two different splitting results from nearby stations indicate that the *SKS*-waves are
56 sampling different anisotropic regions. Thus, the depth of origin of the anisotropy can be estimated
57 as anywhere shallower than the depth where the Fresnel zones beneath each station overlap. Favier
58 & Chevrot (2003) and Chevrot et al. (2004) apply a finite-frequency Frechet derivative approach
59 and calculate 3-D sensitivity kernels for shear-wave splitting intensity, which are similar to those
60 estimated from Fresnel zone estimates (Alsina & Snieder, 1995; Rumpker & Ryberg, 2000).

61 Other studies have utilised finite difference modelling schemes to investigate the ability of
62 shear-wave splitting to identify lateral and depth variations in anisotropy (e.g., Rumpker & Silver,
63 2000), and have placed some constraints on the distribution of anisotropy beneath transform faults
64 and shear zones (Rumpker et al., 2003; Chevrot, 2006; Chevrot & Monteiller, 2009) and a
65 plume setting (Rumpker & Silver, 2000). All these studies show that regions with laterally varying
66 anisotropy give rise to complicated splitting measurements.

67 In this paper we address the suitability of the *SKS*-wave splitting technique to constrain sharp
68 lateral variations in anisotropy, and further investigate the ability of this seismological technique
69 to constrain the depth extent of the anisotropy. This is done using a finite-frequency waveform
70 modelling technique (Angus et al., 2004). This study improves on previous modeling of laterally
71 varying anisotropy by providing some general guidelines that can be applied to shear-wave split-
72 ting observations in regions with sharp lateral changes in anisotropy. This study was motivated by

73 observations of *SKS*-wave splitting beneath the Main Ethiopian Rift (MER) (Kendall et al., 2005,
74 2006), and these results are used as a case study to highlight the utility of the modeling. This ap-
75 proach is applicable to any region that has sharp boundaries in anisotropy, such as transform faults
76 or suture zones.

77 **2 FINITE FREQUENCY WAVEFORM MODELLING**

78 We construct synthetic seismograms using the one-way wave equation modelling scheme of Angus
79 et al. (2004) and Angus & Thomson (2006), for waves propagating vertically through a medium
80 containing oriented melt pockets (Figure 2). The model is constructed by calculating the elastic
81 constants for vertically aligned melt pockets, using the approach of Hudson (1981) and applying
82 these elastic constants at each node. The elastic constants are calculated using *P*- and *S*-wave veloc-
83 ities of 7.8 km s^{-1} and 4.0 km s^{-1} (matrix material) and 2.5 km s^{-1} and 0.0 km s^{-1} (crack material)
84 and densities of 3.8 kg/m^3 and 2.7 kg/m^3 for the matrix material and crack material, respectively.
85 The cracks consist of penny shaped inclusions with an aspect ratio of 0.01. Crack density is then
86 varied to calculate elastic constants with varying magnitudes of anisotropy. We characterise this
87 anisotropy in terms of maximum shear-wave anisotropy (i.e. 10% anisotropy refers to a maximum
88 shear-wave anisotropy of 10%). Motivated by Kendall et al. (2005) a MER rift-like model is con-
89 structed (Figure 2) (e.g., a rotated horizontal symmetry axis in the rift zone). The simulations are
90 done in 3D, but with only 2D variations in anisotropy.

91 In the models the fast direction is oriented 30° from north outside the rift, and north-south
92 inside the rift. An *SKS* like wave (i.e. small waveform curvature (4.2^{-12} s/m), 8 s period) is then
93 propagated through the 3D anisotropic model using the finite-difference one-way wave propagator
94 formulation (Angus et al., 2004; Angus & Thomson, 2006). There is a sharp transition between
95 these two regions (i.e. grid points on either side of the transition between rift and flank have
96 differing fast axis orientations and hence elastic constants). The one-way wave equation method
97 models transmitted waves (it ignores backscattering which is not an issue in modelling *SKS*-wave
98 arrivals), taking into account frequency dependent coupling effects due to, for instance, rapidly
99 rotating wave polarisation effects due to slowness surfaces (Crampin & Yedlin, 1981).

| Model number | Varying parameter | Range | Figure |
|--------------|---------------------------------------|----------|--------|
| 1 | Width of anisotropic zone | 5-40 km | 1 |
| 2 | Magnitude of anisotropy | 4-10% | 2 |
| 3 | Depth extent of anisotropy | 25-85 km | 3 |
| 4 | Dominant period of incoming wave | 0.5-10 s | 4 |
| 5 | Initial polarisation of incoming wave | 15-165° | 5 |

Table 1. Parameters tested in the various models

100 All waves are **Ricker wavelets (although wavelet type has no effect on the result)**, have a
101 dominant period of 8 s (except model 6, where frequency dependence is investigated), and initial
102 source polarisation of 45° (except model 7, where initial polarisation dependence is investigated).
103 The output waveforms are then analysed identically to the data (see Kendall et al., 2005), and the
104 apparent splitting is estimated for a profile spanning the ‘rift’ zone. **To estimate the splitting we**
105 **use the Teanby et al. (2004) cluster analysis method which is based on the Silver & Chan**
106 **(1991) method. This technique performs a grid search over δt and ϕ , rotating the horizontal**
107 **components by ϕ , and shifting their relative positions by δt . The values of δt and ϕ which**
108 **provide the most linear particle motion provides our estimate of the splitting. A statistical**
109 **F-test is used to assess the uniqueness of the result, thus providing an error estimate.** The ob-
110 served splitting depends on several parameters (outlined in Table 1) and each parameter is studied
111 in turn (Figures 3-7). Finally, based on the modelling results, the observed splitting parameters of
112 Kendall et al. (2005) (Figure 1) are modelled to place estimates on the anisotropic characteristics
113 beneath the MER.

114 2.1 Model class 1: Varying width of ‘rift’ zone

115 The first variable tested is rift width (all other variables are held constant: maximum S -wave
116 anisotropy=10%, depth of anisotropy=45 km (from surface), dominant period=8 s, initial polarisation=45°).
117 We vary the width of the rift zone between 100 km and **10 km** (Figure 3). A smooth transition in
118 ϕ is seen for all rift widths. The inflexion point in the ϕ profile marks the boundary between
119 anisotropic regions (Figure 3). However, δt shows considerably more variation across the rift
120 boundaries. The expected δt for this model is a constant value of 1.27 s, but we see large vari-

121 ations, similar to those seen for other studies of inhomogeneous anisotropic media (e.g., plume,
122 Rümpker & Silver (2000); transform fault, Rümpker et al. (2003)).

123 **2.2 Model class 2: Varying magnitude of anisotropy**

124 The second variable tested is the magnitude of anisotropy (all other parameters are held con-
125 stant: rift width=40 km, depth of anisotropy=45 km (from surface), dominant period=8 s, initial
126 polarisation=45°). All models within this class show a smooth transition in ϕ , with the inflexion
127 point showing the transition between anisotropic regions. The δt profile shows a similar trend for
128 all models. The magnitude of splitting depends on the amount of anisotropy, but the peaks and
129 troughs of the δt curve all lie in the same place (see Figure 4).

130 **2.3 Model class 3: Varying thickness of anisotropic zone**

131 The third variable tested is depth (i.e. thickness from the surface) of the anisotropic zone (all other
132 parameters are held constant: rift width=40 km, maximum *S*-wave anisotropy=10%, dominant pe-
133 riod=8 s, initial polarisation=45°). The variation of ϕ is smooth for nearly all models in this class,
134 with the inflexion point showing the transition between anisotropic regions. An exception occurs
135 in the model with an 85 km thick layer (Figure 5). The variation in ϕ seen in the 85 km thick layer
136 has some deviation at the boundary between anisotropic regions. This may be due to multipathing
137 effects; a result of the longer wavepath through a complex, highly anisotropic medium. The δt
138 curve shows similar variation to that seen in model 2. Although similar to model 2, where the
139 magnitude of splitting increases with anisotropic strength rather than increasing path length, there
140 is an observable moveout of the peaks with increasing thickness (Figure 5). This sensitivity with
141 depth can be used to interpret something about the depth to the anisotropic region. We discuss this
142 further in relation to the MER results in section 3.1.

143 **2.4 Model class 4: Varying frequency of propagating wave**

144 The fourth variable tested is the dominant period of the incoming wave (all other parameters are
145 held constant: rift width=40 km, maximum *S*-wave anisotropy=10%, depth of anisotropy=45 km,

146 initial polarisation= 45°). It is evident that varying the dominant period has a large effect on the
 147 variation observed in both δt and ϕ (Figure 6). For higher frequencies the curves match the input
 148 model well, with little deviation in δt and a sharp transition in ϕ . The observation that the inflexion
 149 points in ϕ describe the width of the rift zone still applies for all frequencies. For higher fre-
 150 quencies the peaks in δt are narrow, an effect of approaching the ray theoretical limit. This shows
 151 the importance of investigating shear-wave splitting using a finite-frequency approach, where the
 152 influence of frequency dependent shear-wave coupling is accounted for.

153 2.5 Model class 5: Varying initial polarisation of the incoming shear wave

154 The fifth variable tested is the initial polarisation of the incoming wave (all other parameters are
 155 held constant: rift width=40 km, maximum S -wave anisotropy=10%, depth of anisotropy=45 km,
 156 dominant period=8 s). The variation in ϕ is dependent on the initial polarisation, but the inflexion
 157 points still define the width of the rift zone (Figure 7). The transition in ϕ from 30° to 0° occurs
 158 over distances of ~ 20 -50 km. It is evident that δt is strongly dependent on the initial polarisation,
 159 with the peaks in δt occurring either side, and on top of the transition between anisotropic regions
 160 (Figure 7). This is similar to previous studies of inhomogeneous anisotropic structure which show
 161 that the measured splitting parameters are highly dependent on the initial polarisation of the shear-
 162 wave (Silver & Savage, 1994; Rumpker & Silver, 2000; Rumpker et al., 2003).

163 The variation in the δt profile depends on the relationship between the initial polarisation and
 164 the two anisotropic symmetry axes. For example, model 5 shows that for an initial polarisation
 165 of 105° the peak in the δt curve lies directly above the transition zone. This initial polarisation is
 166 oriented 75° from both the 30° fast direction outside the rift, and the 0° inside the rift ($105^\circ = -$
 167 75°). **However, when the initial polarisation is preferentially close to one of the symmetry axis**
 168 **(assuming it is not so close that a null measurement is recorded), it will induce an asymmetry**
 169 **in the observed δt measurements.** This result is consistent with the observation that the finite-
 170 frequency sensitivity kernel of the incoming shear-wave depends on the initial polarisation (Favier
 171 & Chevrot, 2003; Chevrot et al., 2004).”

3 DISCUSSION

For all the scenarios modelled in section 2 it is evident that sharp lateral changes in anisotropic fabric can be detected using shear-wave splitting. The sharpness of the transition in splitting parameters depends on three parameters, the thickness of the anisotropic layer (Figure 5), the frequency of the incoming wave (Figure 6) and the initial polarisation of the shear-wave (Figure 7). For example, for an 8 s wave (e.g., *SKS*), the transition in ϕ from 30° to 0° occurs over ~ 20 km (25 km thick layer) to ~ 40 km (85 km thick layer), assuming a constant initial polarisation. For a constant thickness of anisotropy (45 km) the transition in ϕ varies from ~ 20 -50 km, depending on the initial polarisation. In all model classes, except where the splitting is very large, the inflexion points in the ϕ curve define the transition in anisotropic fabric. For the MER this will indicate the rift width, but this phenomenon can also potentially be used to determine the location of transform faults and suture zones.

Having constrained the rift width from ϕ it is possible to use the variation in δt to place constraints on the thickness of the anisotropic layer. The position of the peaks and troughs in δt vary for differing rift widths, frequency content of the incoming wave, thickness of the anisotropic layer and initial polarisation. By assuming a dominant period of the incoming *SKS*-phase of 8 s, we can use the position of the peaks and troughs to estimate the thickness of the anisotropic layer. This can only be done for one initial polarisation (approximately the same as back-azimuth for an *SKS*-wave).

From studying the variations in ϕ and δt we can estimate the width and thickness of the anomalous anisotropic zone, and from this it is simple to estimate the magnitude of anisotropy. The magnitude of anisotropy is calculated by determining the amount needed to match the splitting results far from the rift. In our models we impose a vertical transition in anisotropic regions. We acknowledge that dipping boundaries may effect these results, and more modeling is need to constrain this.

This modeling exercise has highlighted other features which may be observable in data to detect lateral variations in anisotropy. It is evident that stations close to the transition will show large variation in δt as a function of back-azimuth, whereas the same stations will show little variation

200 of ϕ . This is notably different from what is expected from two horizontal layers of splitting, which
 201 causes variations in both ϕ and δt (Silver & Savage, 1994). Another feature which can be observed
 202 is the variation of δt as a function of frequency, again close to the transition in anisotropic domains.

203 **3.1 Comparison with the Main Ethiopian Rift (MER)**

204 Kendall et al. (2005) observe a rotation in the splitting fast direction inside the Main Ethiopian
 205 rift valley, with the fast shear-wave aligning with the magmatic segments (Figure 1). This pattern
 206 could be caused by oriented melt pockets (OMP) or along-rift flow which causes a lattice preferred
 207 orientation (LPO) of olivine. A study of surface waves (**Kendall et al., 2006; Bastow et al., in
 208 review 2010**) addresses this ambiguity, due to the azimuthal variations expected for observed
 209 phase velocities on interstation paths being different for OMP or LPO. **Kendall et al. (2006) and
 210 Bastow et al. (in review 2010)** show that to satisfy both the *SKS*-wave splitting results and surface
 211 wave results an OMP source of anisotropy, down to depths of at least 70 km, must be present. Keir
 212 et al. (2005) analysed splitting in shear-waves from local earthquakes <20 km beneath the MER,
 213 and found fast directions similar to Kendall et al. (2005), aligning with the magmatic segments.
 214 They also found elevated splitting magnitudes above regions where melt has been inferred from
 215 wide-angle refraction (Mackenzie et al., 2005) and controlled source tomography (Keranen et al.,
 216 2004).

217 Other *SKS*-wave splitting results around Ethiopia show similar results. Ayele et al. (2004)
 218 observed splitting in Kenya, Ethiopia and Djibouti, and noticed that the magnitude of splitting
 219 increases with the amount of melt produced, inferred from a correlation between an increase in
 220 delay time and volcanism. Gashawbeza et al. (2004) performed shear-wave splitting on a network
 221 of wider aperture, but similar location to Kendall et al. (2005). They observed similar rift parallel
 222 trends in the fast directions, but argued that fossilised Precambrian anisotropy was the source of
 223 this splitting, with some more recent Neogene influence near the rift to explain the rotation in the
 224 fast directions.

225 Melt has been observed beneath the plateau in the form of highly conductive bodies in magneto-
 226 telluric surveys (Whaler & Hautot, 2006), and underplating is observed in wide-angle reflection

227 profiles (Mackenzie et al., 2005). It seems likely that a combination of both pre-existing fabric,
228 and melt induced anisotropy could be present beneath the plateau region. Further evidence of melt
229 beneath the MER comes from receiver function studies. High values of Poisson's ratio >0.3 for the
230 average crust, and underplating highlight the likelihood of melt beneath Afar (Dugda & Nyblade,
231 2006) and the MER (Stuart et al., 2006).

232 Using the criteria outlined in the previous section we can match the pattern of results observed
233 by Kendall et al. (2005) (Figure 1), and thus place more constraints on melt induced anisotropy
234 beneath the MER .

235 As was shown in the previous section, the variation of ϕ is relatively insensitive to all pa-
236 rameters (assuming an instantaneous change in anisotropic parameters at the transition), and the
237 observation that the 'rift' width can be defined by the inflexion points is robust, except in the pres-
238 ence of very high splitting. For the EAGLE dataset, this results in a rift width of ~ 100 km, based
239 on all the splitting results from different back-azimuths, with a fast direction of 30° outside the
240 rift, and 20° inside the rift (Figure 9).

241 We have shown that variation in δt is dependent on frequency, initial polarisation and vertical
242 thickness of the anisotropic zone. **To model this event we take a real SKS waveform from the**
243 **Ethiopian seismic station ADEE (Figure 8) (see Bastow et al., 2008, for station details), and**
244 **propagate this through the model . However, an 8 s Ricker wavelet, as used in the previous**
245 **sections, produces identical results.** We can not use all the data, as we did for measuring the
246 rift width, as they come from differing back-azimuths. To account for this we take results from
247 one very well constrained *SKS*-wave splitting event that was recorded across the whole array.
248 This event has a back-azimuth of 40° . We run a model, based on this information, to estimate the
249 depth of the anisotropy. The model has a 100 km wide rift zone, 10% anisotropy, a fast direction
250 of 30° outside the rift and 20° inside, an initial polarisation of 40° and a period of 8 s (Figure
251 10). It is evident that the peaks are moving out with anisotropic thickness, and if we plot the
252 peak-peak width as a function of depth it is evident that they lie on a straight line (Figure 10).
253 This observation is valid for initial polarisations which fall outside the null planes of either of the
254 anisotropic regions, and outside the null planes for effective splitting parameters observed near the

255 transition between anisotropic regions. For the event with back-azimuth of 40° the peak-peak width
256 is 133 km (Figure 9), which equates to a depth of 90 km. With this information we can estimate
257 the magnitude of anisotropy needed to generate the amount of splitting observed. However, the
258 peaks seen in δt vary in magnitude with the western plateau having elevated shear-wave splitting
259 compared to the eastern plateau. As a result we propose a model which has 9% anisotropy on the
260 western plateau and in the westernmost 30 km of the rift zone and 7% anisotropy on the eastern
261 plateau and easternmost 70 km of the rift zone (Figure 9). **Unfortunately, no other events were**
262 **suitably recorded across the whole array and thus comparisons of this model with splitting**
263 **results from other back-azimuths can not be performed.**

264 We estimate a region of anisotropy which extends to a depth of ~ 90 km beneath the MER, ex-
265 tending beneath the two margins and the rift valley. On the margins the anisotropy is oriented with
266 a fast direction of 30° , and beneath the rift valley the anisotropy is perturbed with an orientation of
267 20° . The anisotropy below the rift valley correlates with the magmatic segments, as described by
268 Kendall et al. (2005). The variation evident in δt can largely be explained as an effect of these two
269 different anisotropic regimes interacting. The variation in δt is highest on the western margin and
270 results in an estimate of 9% anisotropy beneath this region compared to 7% on the eastern margin.
271 This model is summarised in Figure 11. **This model was estimated using a trial and error ap-**
272 **proach. To fully constrain this model requires a more complete sampling of the whole model**
273 **space, which is computationally impractical. As a result we can not formally discuss errors**
274 **of the fit to the data here. To provide confidence in our models we can compare our results**
275 **with other geophysical data. To fully explore the model space, studies that invoke theoretical**
276 **sensitivity kernels may be suitable (Chevrot, 2006; Chevrot & Monteiller, 2009).**

277 If anisotropy is derived from oriented melt pockets, as suggested by Kendall et al. (2005) then
278 this suggests that melt present beneath the MER is aligned from a depth of 90 km, with an average
279 anisotropy of 7-9%. This equates to a melt fraction of 7-9%, assuming vertically oriented isolated
280 melt inclusions with an aspect ratio of 0.01, but the amount of melt needed to produce this amount
281 of splitting would be smaller for inclusions with a lower aspect ratio.

282 This estimate of the depth extent of melt induced anisotropy is supported by other studies.

283 Kendall et al. (2006) show that azimuthal anisotropy seen in surface wave results occurs at depths
284 between 50-70 km, and may extend further but resolution decreases below these depths. Based
285 on simple Fresnel zone estimates, Kendall et al. (2005) place the origin of the anisotropy seen
286 in the EAGLE data to be <100 km, a fairly accurate estimate based on these results. *P*- and *S*-
287 wave tomography show the lowest seismic velocities in the top 100 km (Bastow et al., 2005, 2008)
288 (Figure 11) and Ayele et al. (2004) image a discontinuity at a depth of ~ 90 km which they suggest
289 is the base of the lithosphere. Based on geochemical evidence, Rooney et al. (2005) suggest that
290 the base of the lithosphere is the origin of melt generation feeding the MER. Additionally, Keir
291 et al. (2009) suggest that partial melting of the lithosphere and subsequent magma injection causes
292 lower crustal earthquakes throughout the MER and western plateau.

293 Seismic tomography (Bastow et al., 2005, 2008) (Figure 11) and magneto-tellurics (Whaler &
294 Hautot, 2006) both show evidence for an asymmetry in melt production, with lower velocities and
295 higher conductivities present beneath the western plateau compared to the eastern plateau. This
296 is supported by our results which show elevated anisotropy in the west, mainly at the region we
297 define as the rift boundary. Holtzman & Kendall (in review) suggest that melt is concentrated in
298 regions of higher strain, and cite the elevated δt seen by Kendall et al. (2005) as evidence for this.
299 We show that an elevated δt can be explained by a simple variation in fast direction alone across
300 the region, but we still require elevated anisotropy close to the western margin, as suggested by
301 Holtzman & Kendall (in review), to explain the asymmetry seen in δt .

302 **4 CONCLUSIONS**

303 We have developed a modelling technique, using a one-way wave equation approach, to investigate
304 the effects of laterally varying anisotropy on shear-wave splitting. We have shown that:

- 305 (i) *SKS*-wave splitting can be used to identify changes in fast direction over lateral distances of
306 20-50 km (dependent on depth and initial polarisation).
- 307 (ii) The inflexion points in the ϕ profile demarcate the transition in anisotropy.

308 (iii) Variation in the position of the peaks and troughs seen in δt depend on anisotropic thickness
 309 from the surface, and for a given initial polarisation can be used to place depth constraints on the
 310 anisotropic region.

311 (iv) With information on the depth of the anisotropic region, estimates can be placed on the
 312 percentage of anisotropy in the region.

313 (v) At stations close to the transition δt varies as a function of back-azimuth and frequency,
 314 whereas ϕ shows little such variation. This can be used as an indicator of lateral changes in
 315 anisotropy

316 (vi) For higher frequencies the modelled splitting approaches the ray theoretical limit, and
 317 shows little variation in δt . Thus, a frequency dependence of δt could indicate a lateral transition in
 318 anisotropy. This also shows the importance of performing finite frequency waveform modelling as
 319 opposed to ray based approaches in regions where anisotropy varies over length scales comparable
 320 to the dominant seismic wavelength.

321 Determining the exact cause and symmetry of the anisotropy still requires analysis of other
 322 phases (e.g., joint shear-wave splitting and surface waves Brisbourne et al. (1999); Kendall et al.
 323 (2006)). However, we show that a simple rotation in the anisotropic characteristics in a 100 km
 324 wide ‘rift’ zone can explain much of the variation seen in the Kendall et al. (2005) Main Ethiopian
 325 Rift splitting results. The anisotropy across the model is confined to the uppermost 90 km. This re-
 326 gion of anisotropy coincides with regions with low velocities (Bastow et al., 2008), high anisotropy
 327 (Kendall et al., 2006), high conductivities (Whaler & Hautot, 2006), and the suggested base of the
 328 lithosphere (Ayele et al., 2004). It has also been suggested that this region is where melt is gener-
 329 ated that feeds the MER (Rooney et al., 2005). However, a simple rotation alone is not enough to
 330 reproduce these results. We require elevated anisotropy at the western margin to match the higher
 331 δt seen in the splitting results in this region. This coincides with the lowest velocities (Bastow
 332 et al., 2008) (Figure 11), highest conductivities (Whaler & Hautot, 2006) and regions of mag-
 333 matic underplate (Mackenzie et al., 2005; Cornwell et al., 2006). Holtzman & Kendall (in review)
 334 suggested that the elevated splitting is caused by focused melt along the margin, where strain is
 335 highest. The elevated anisotropy on the western margin required by our models supports this, with

336 little effect seen on the eastern margin where smaller evidence of melt related phenomena are ob-
337 served. These results show how observations of seismic anisotropy provides insights into the role
338 of magma in rifting.

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Figure 1. Average SKS-wave splitting results beneath the Main Ethiopian Rift (MER), adapted from Kendall et al. (2005). The orientation of the white bars shows the fast shear-wave direction, and the length is proportional to the amount of splitting. Heavy black lines indicate major border faults and magmatic segments are marked in red. The solid white line perpendicular to the rift indicates the profile used to construct the top panels. The inset plot shows the location of the EAGLE array in Ethiopia. The top panels show shear-wave splitting parameters as a function of distance from the rift axis. The red triangles show an interpolated fit to the data using a cubic B-spline interpolation with a knot spacing of 30 km. The shaded region shows the r.m.s. misfit of the data from the curve over a 30 km sliding window.

Figure 2. (Top) a schematic representation of the model used in the one-way wave equation modelling scheme. (Bottom) Hemispherical projections of the elastic constants applied at each node. The colour scheme shows the variation of magnitude of shear-wave anisotropy, and the black ticks show the fast direction of a wave propagating through the anisotropic medium, as a function of direction of propagation. Elastic constants are calculated for vertically aligned melt pockets using the approach of Hudson (1981), see text for details. For the modelling the symmetry plane on the rift flanks is oriented 30° from north, and in the rift segment it is oriented north-south. All other parameters (rift width, amount of anisotropy, depth extent of anisotropy, frequency of incoming wave and initial polarisation of incoming wave), are varied systematically (Table 1) and results are displayed in Figures 3-7.

Figure 3. Model 1: Varying rift width. Other parameters: maximum S-wave anisotropy=10%, anisotropic depth=45 km, Period=8 s, initial polarisation= 45° . Coloured dashed lines indicate the rift width for each model. Note smooth transition in ϕ , with inflexion points marking rift width, and complicated variation of δt .

Figure 4. Model 2: Varying maximum S-wave anisotropy. Other parameters: rift width=40 km, anisotropic depth=45 km, Period=8 s, initial polarisation= 45° . Black dashed line indicates the rift width for the model. Note similar curves to model 1, but with the δt varying proportionally to the amount of anisotropy. The peaks and troughs in the δt profile do not change with varying amounts of anisotropy.

Figure 5. Model 3: Varying anisotropic depth. Other parameters: rift width=40 km, maximum *S*-wave anisotropy =10%, Period=8 s, initial polarisation=45°. Black dashed line indicates the rift width for the model. Note similar curves to model 2, but now the peaks and troughs in the δt profile move out with increasing depth of the anisotropic layer. The flexure points still mark out the rift width, except for the 85 km case where multi-pathing effects cause a deviation in the ϕ profile.

Figure 6. Model 4: Varying frequency of incoming wave. Other parameters: rift width=40 km, maximum *S*-wave anisotropy=10%, anisotropic depth=45 km, initial polarisation=45°. Black dashed line indicates the rift width for the model. Note higher frequencies approach a ray based model, and show little deviation from the input model. inflexion points in the ϕ profile still map out the rift width.

Figure 7. Model 5: Varying initial polarisation. Other parameters: rift width=40 km, maximum *S*-wave anisotropy=10%, anisotropic depth=45 km, frequency=8 s. Black dashed line indicates the rift width for the model. Note a high variability in δt and ϕ profiles with initial polarisation.

Figure 8. An example of the Teanby et al. (2004) method of splitting for a) EAGLE station ADEE (Event information: 2001/12/02, 13:01:53, 39.40°N, 141.09°E, 123.8km, MW6.5), after Kendall et al. (2005). b) Synthetic data for an MER model (Figure 9). Both traces are located on the Eastern margin of the rift zone. (i and vii) Traces rotated into R and T directions before and after the anisotropy correction. R component is the initial shear-wave polarisation before entering the anisotropic region. T component is perpendicular to the R component. Energy on the corrected transverse trace should be minimised in the analysis window. (ii and viii) Uncorrected fast/slow shear waveforms. (iii and ix) Corrected fast/slow shear waveforms. (iv and x) Particle motion for uncorrected seismograms. (v and xi) Particle motion for corrected seismograms. A good result will show similar fast/slow waveforms and any elliptical particle motion will be linearised. (vi and xii) The results of the grid search over δt and ϕ . The method used minimises the second eigenvalue of the particle motion (i.e. the best result occurs where the particle motion is linear after removing the splitting). The optimum splitting parameters are represented by the cross, and the 1st surrounding contour denotes the 95% confidence interval.

Figure 9. Best fitting model to the results of Kendall et al. (2005). Rift width=100 km, maximum S -wave anisotropy=9% < -14 km, anisotropy=7% > -14 km, anisotropic depth=90 km, frequency=8 s, initial polarisation=40°. The δt model results are compared to splitting results from one event with a back azimuth of 40°, and the ϕ model results are compared with all splitting results.

Figure 10. (Top) variation in δt profile with varying depth extent of anisotropy (similar to model 3). Blue dashed line shows the moveout with depth of the peaks in the δt profile. (Bottom) Least-squares fit to the location of the peaks in the δt profile with depth. Model parameters used in this model are rift width=100 km, maximum S -wave anisotropy=10%, variable anisotropic depth, frequency=8 s, initial polarisation=40°.

Figure 11. Map (left) and cross-section (right) view of the best fitting model from Figure 9, plotted over the P -wave tomographic images of Bastow et al. (2008). The red lines on the left plot highlight the modelled rift edges, and the dashed black line indicates the transition from 9% anisotropy to 7%. The white lines indicate the location of the tomographic cross-sections. There is a strong correlation between the slow velocity anomalies and the regions of 9% anisotropy. Also the slowest anomalies appear to be present in the top 100 km, similar to where we constrain the anisotropy to be.

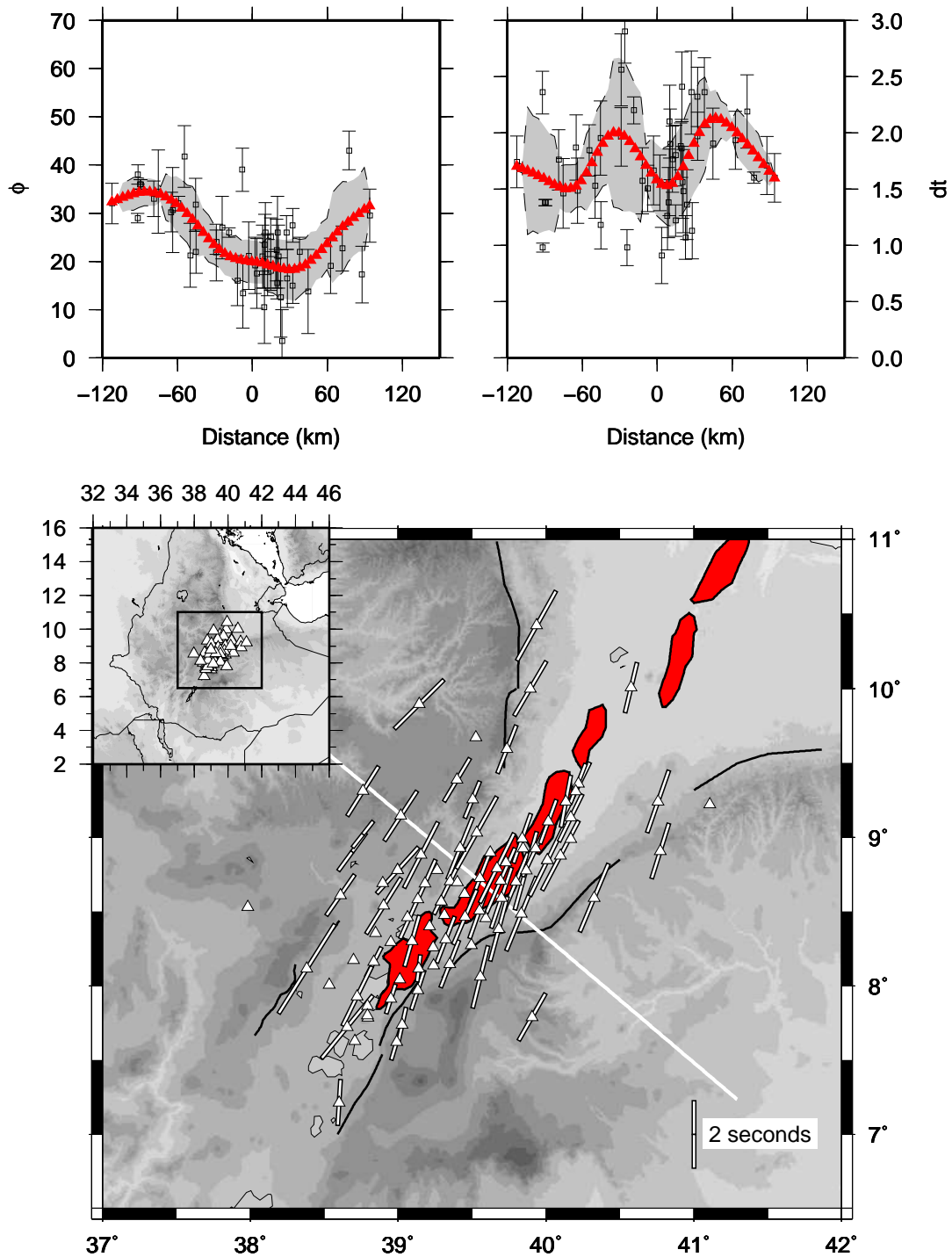


Figure 1

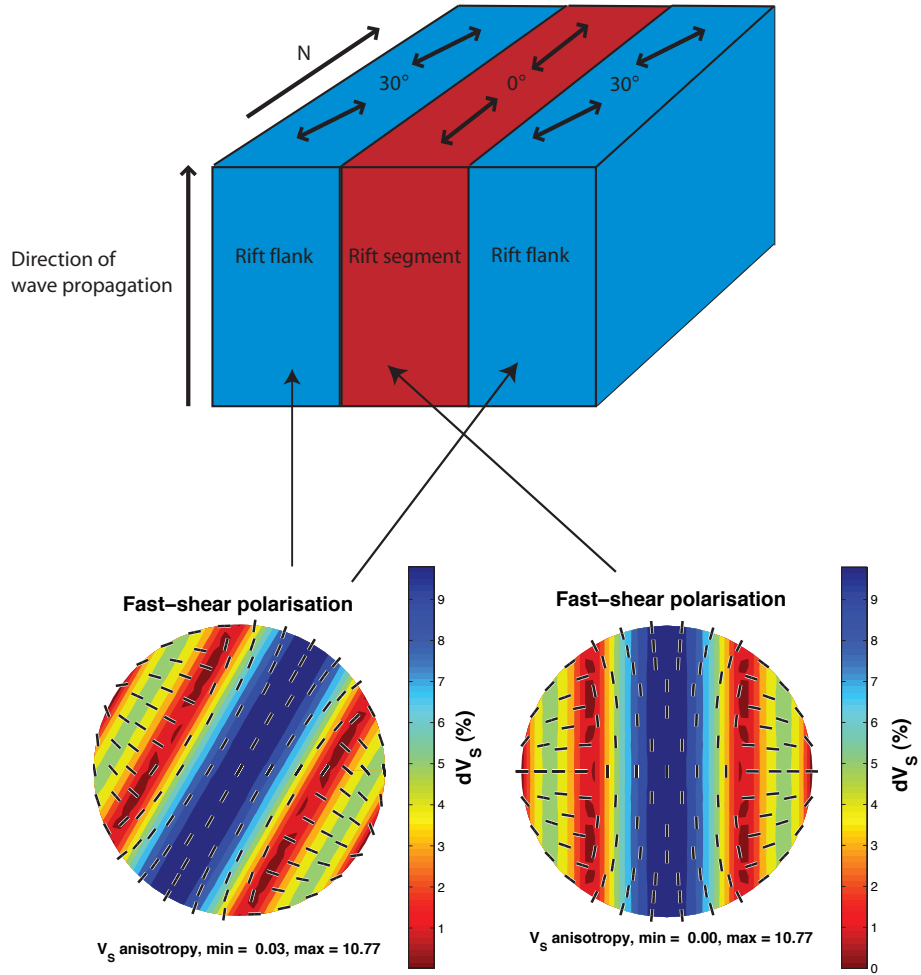


Figure 2

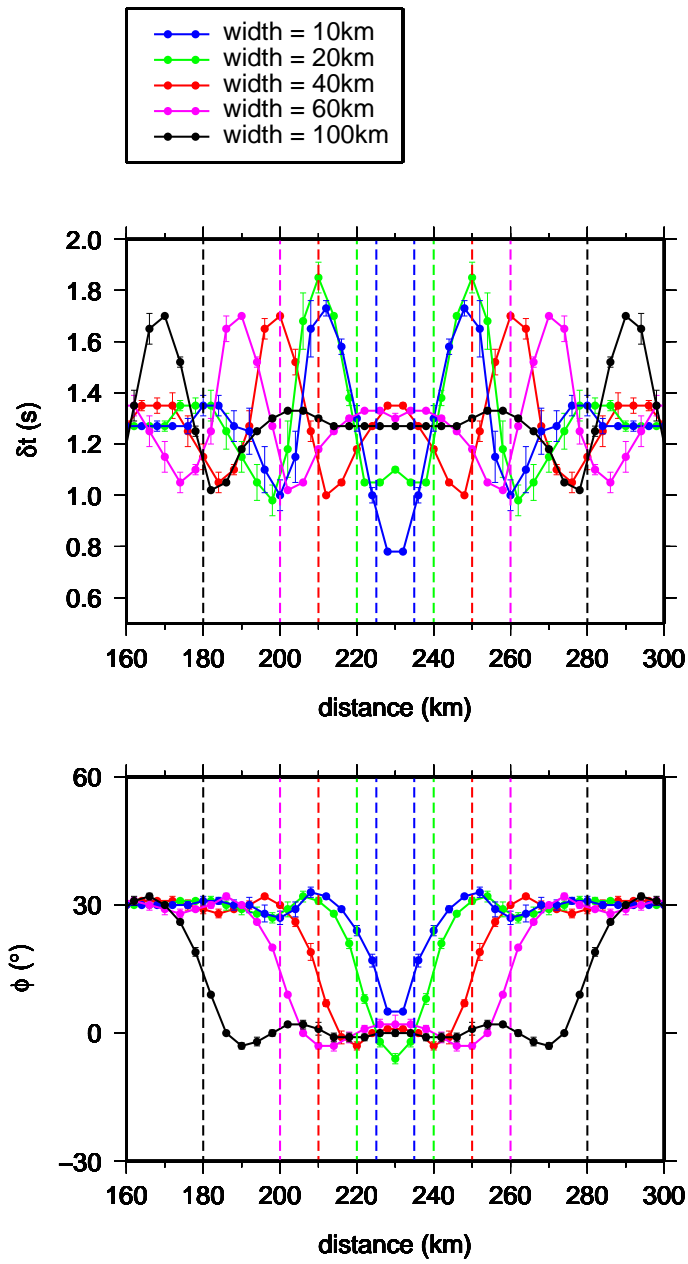


Figure 3

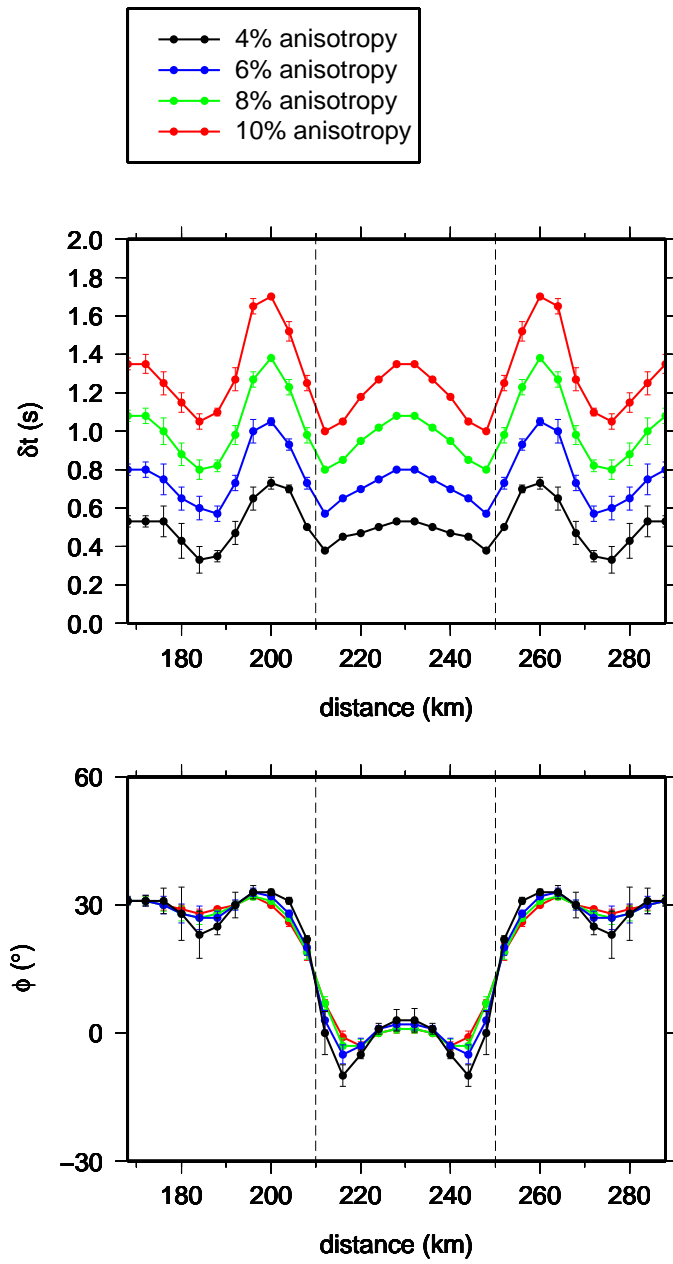


Figure 4

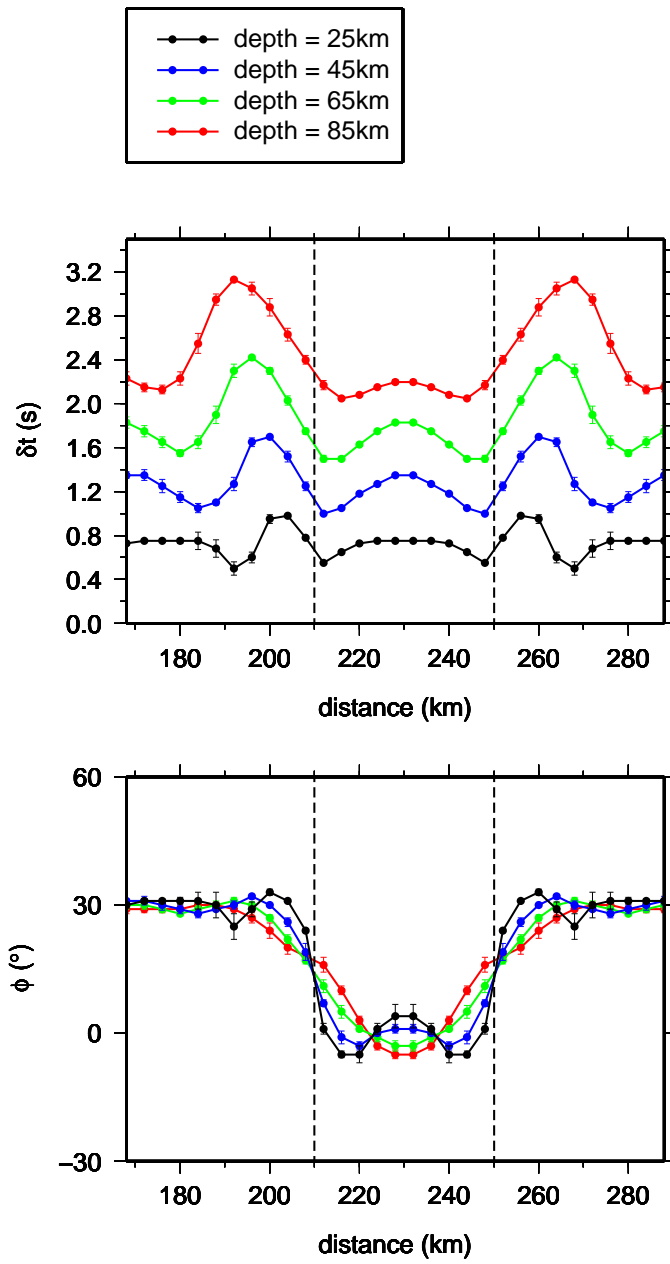


Figure 5

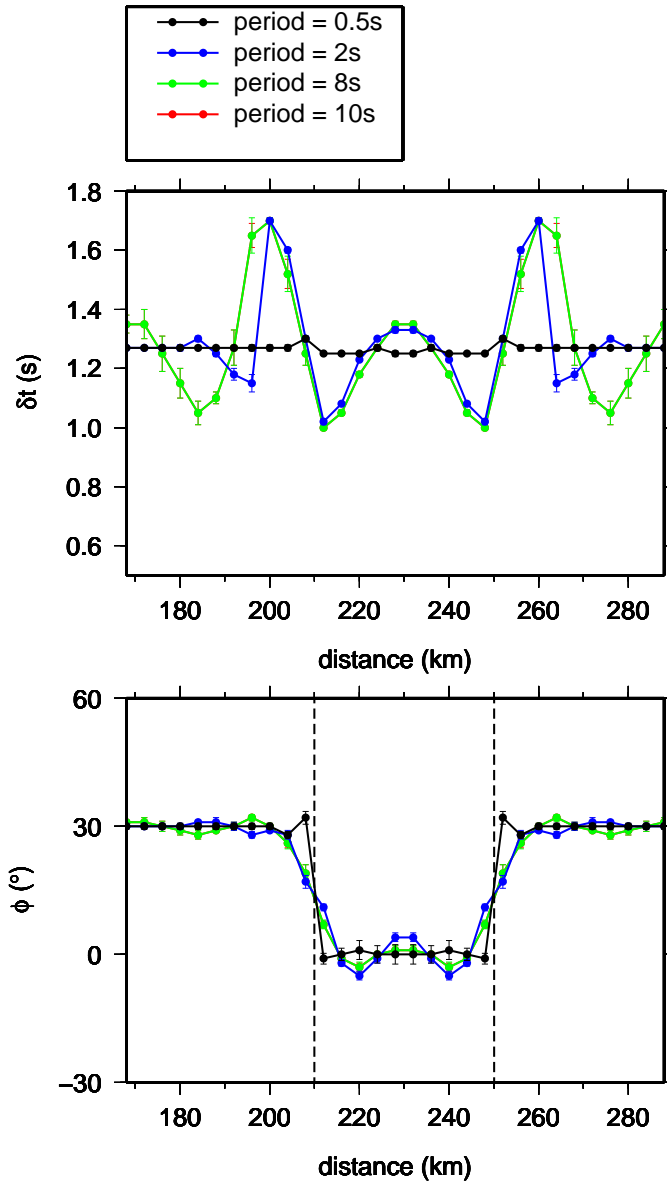


Figure 6

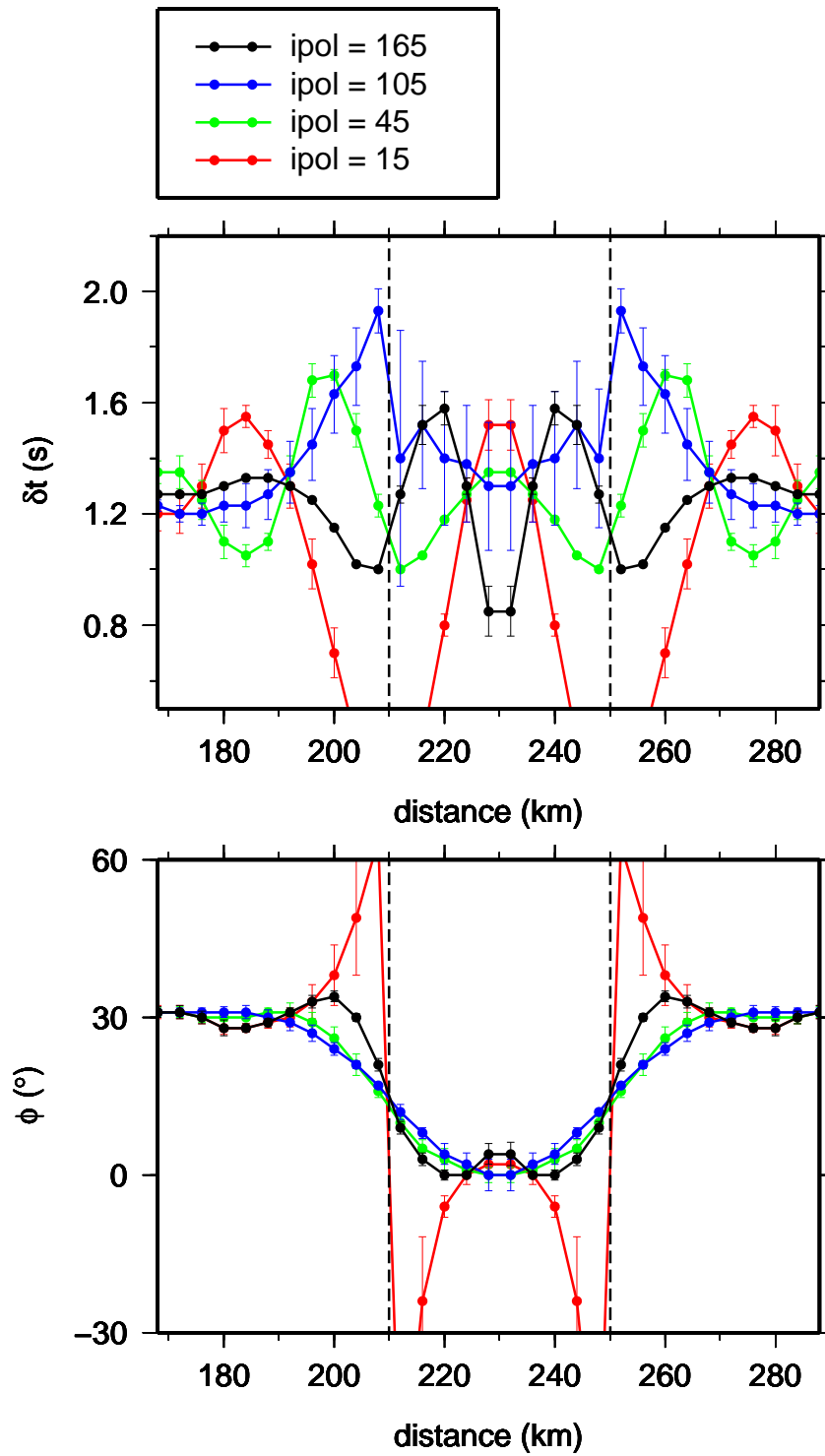


Figure 7

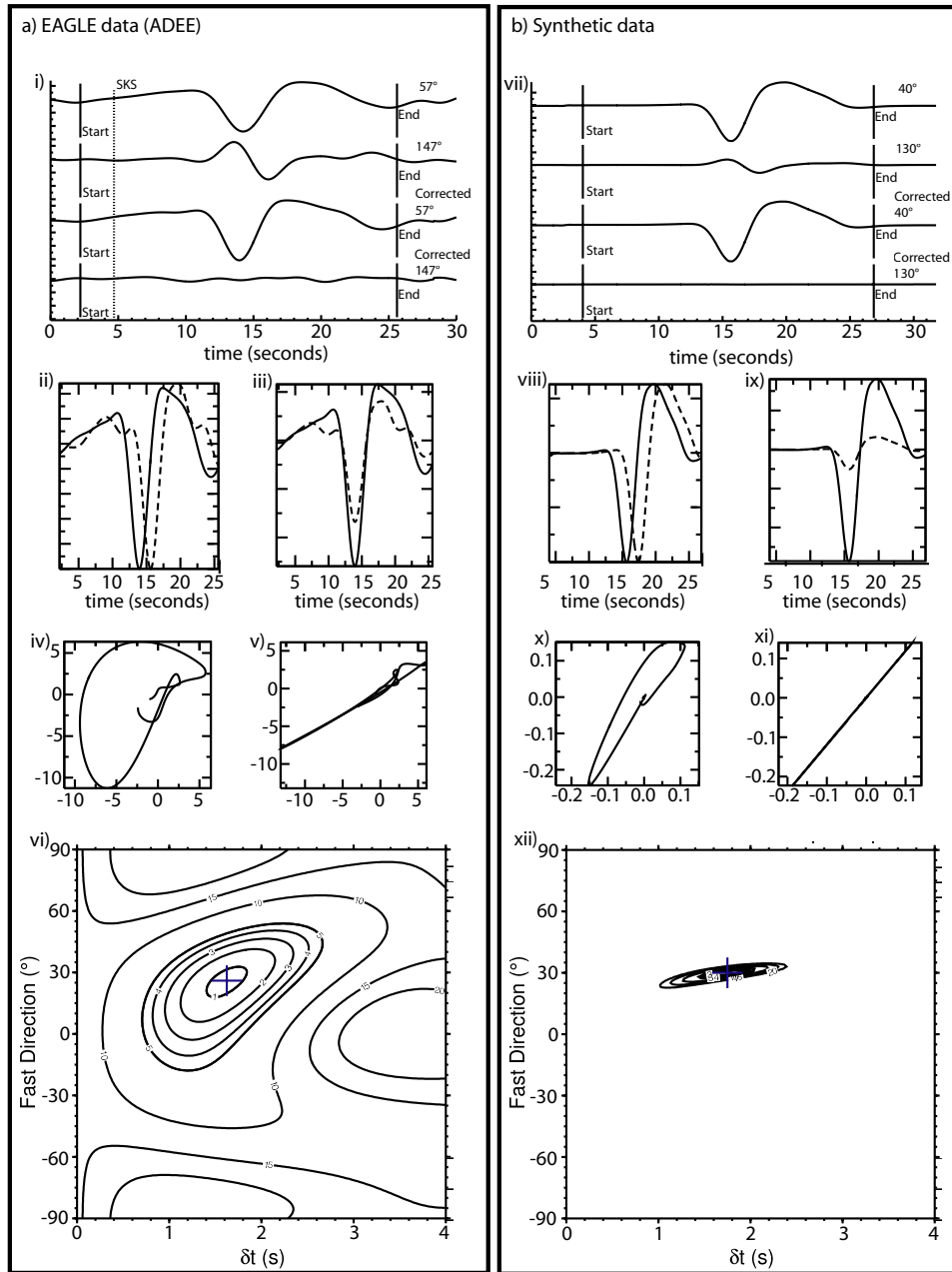


Figure 8

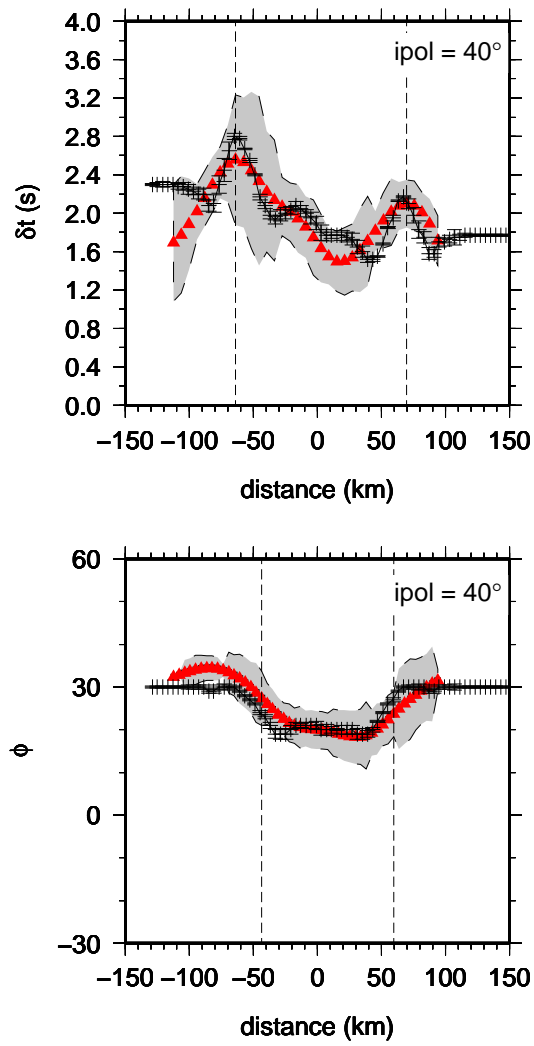


Figure 9

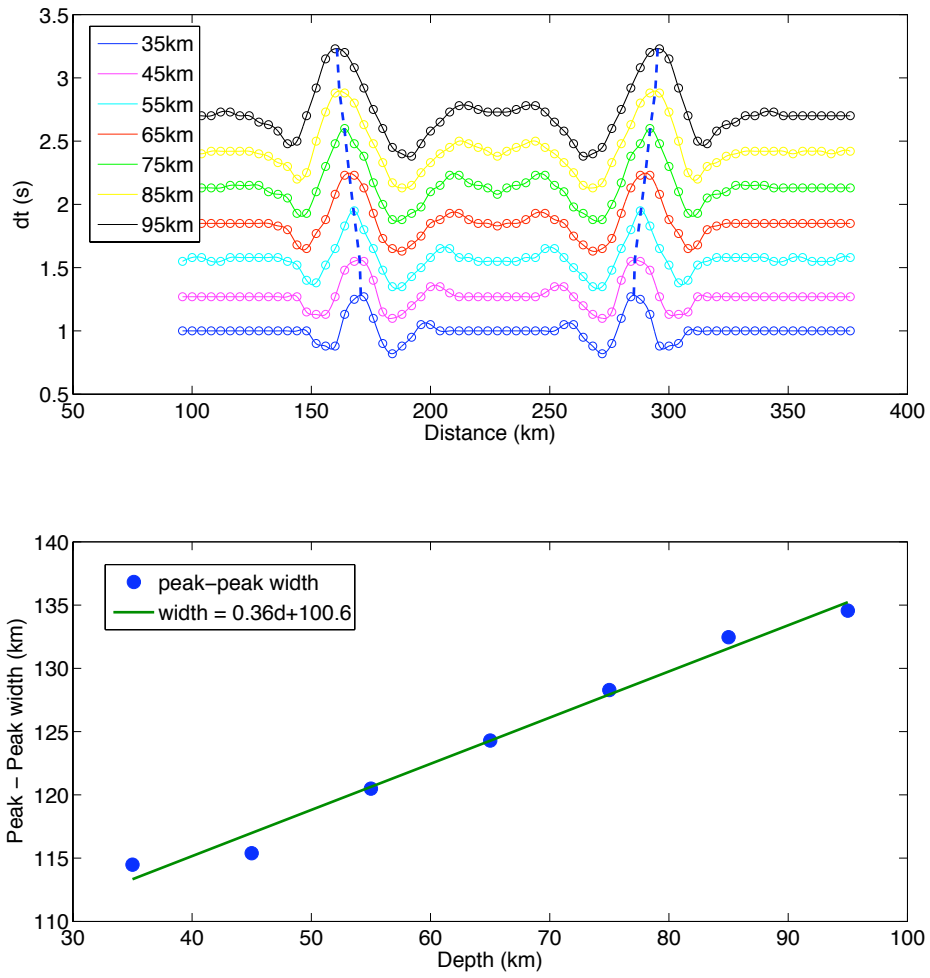


Figure 10

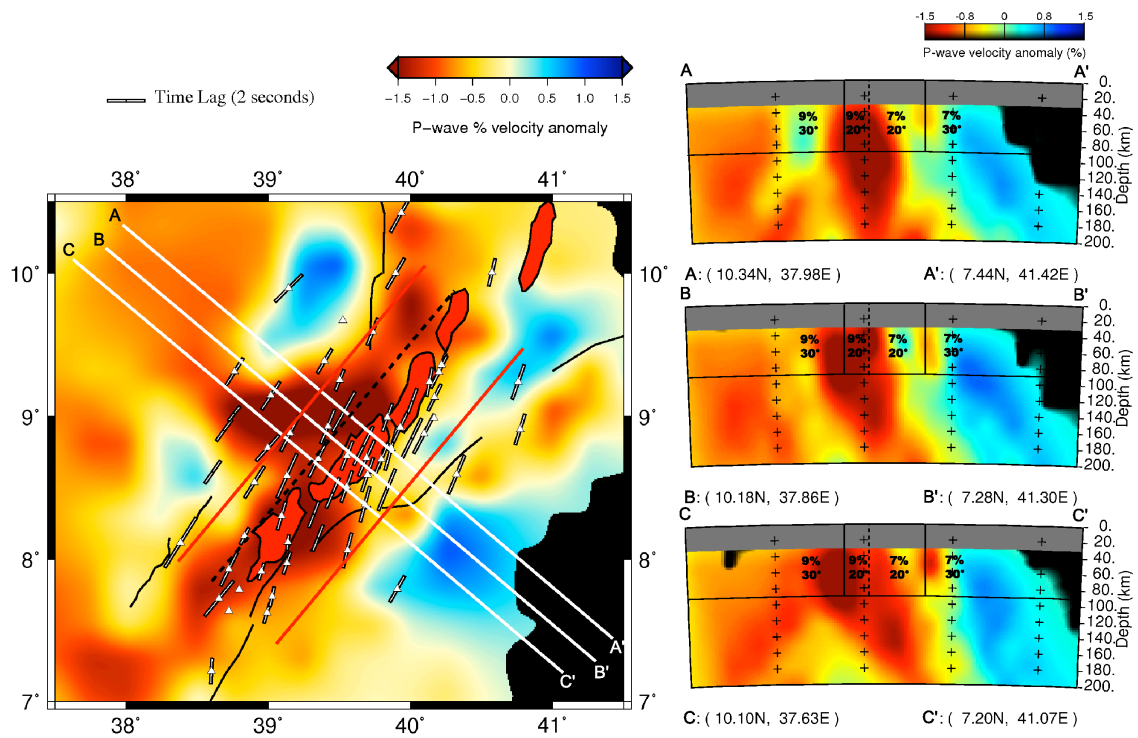


Figure 11