

Ionosphere-Thermosphere-Mesosphere Variability imposed by Waves from Below in Future Climates

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The positions, experiences, and viewpoints expressed below are those of the authors as scientists in the space research community and are not the official positions of their employing institutions.

***Synopsis:** The structure and dynamics of the Ionosphere-Thermosphere-Mesosphere (ITM) system are significantly influenced by waves from the lower atmosphere. Understanding of the physical mechanisms at play and trends is impaired by data sparsity. This white paper reviews current knowledge in this area and highlights the importance and critical need for new, global, height-resolved observations of the ITM.*

1. INTRODUCTION

Over the past two decades, improved physical understanding of the Ionosphere-Thermosphere-Mesosphere (ITM) system, partly enabled by the development of new observational techniques, advances in the sophistication and computing power of models, and a new generation of tools allowing the relevant processes to be visualized and analyzed, demonstrated unambiguously that much of the terrestrial influences on the ITM is due to *internal atmospheric waves*. Mostly excited in the troposphere-stratosphere region (ca., 0-50 km), internal atmospheric waves play a key role in the transfer of energy and momentum, mixing of constituents, development of large-scale flows, and generation of ITM variability [e.g., *Liu, 2016; Forbes, 2021*]. The components of the atmospheric wave spectrum coupling terrestrial variability with ITM changes are gravity waves (GWs), atmospheric tides, Kelvin waves (KWs), and Rossby planetary waves (PWs), covering a spatial range from a few tens of kilometers to several thousand kilometers [e.g., *Oberheide et al., 2015*]. Understanding the impacts of vertically propagating waves on ITM variability is fundamental to understanding its thermal, dynamical, and compositional structure and long-term trends [e.g., *Qian et al., 2017; Rezac et al., 2018; Solomon et al., 2018, 2019*]. Recent research [see, e.g., reviews by *Sassi et al., 2019; Ward et al., 2021*] highlighted the importance of the manner in which waves produced in the troposphere-stratosphere region propagate vertically to influence the dynamics and transport of constituents in the ITM system, the character of the long-term trends in the ITM, and the changes taking place in the thermosphere that may affect the ITM structure, composition, and variability in *future climates*.

The satellite environment in near-Earth space depends on the complex ITM region as the maintenance of satellite orbits and space debris trajectories require the accurate prediction of neutral atmospheric densities over time and space [e.g., *Emmert, 2015*]. The accurate modeling of neutral atmospheric density requires an understanding of atmospheric waves and dynamics that modulate the neutral density spatially and temporally [e.g., *Leonard et al., 2012*]. Ionospheric plasma is also significantly influenced by atmospheric dynamics driven by waves [e.g., *Forbes and Zhang, 1997; Hocke and Schlegel, 1996; Laštovička et al., 2006; England et al., 2010; Forbes et al., 2018; Koucká Knižová et al., 2021*], with direct implications for satellite communication links and GPS-based navigation.

2. INTRA-SEASONAL AND INTER-ANNUAL VARIABILITY IN THE ITM

ITM neutral and electron density variability spanning timescales from months to years have been observed since the 1930s-1960s [e.g., *Appleton and Naismith, 1935; Berkner and Wells, 1938; Yonezawa, 1959; Paetzold and Zschorner, 1961*], yet significant uncertainty still exists on the physical mechanisms responsible for these variations. Attaining a better understanding and characterization of the various modes of variability in the ITM system from intra-seasonal (~30-90 days) to inter-annual timescales and their connections to terrestrial drivers is critical for achieving whole atmosphere predictability [*Sassi et al., 2019*] and for improving our ability to model changes associated with future climates [*Ward et al., 2021*]. Modes of intra-seasonal variability include the tropospheric Madden-Julian Oscillation (MJO; *Madden and Julian, 1971*); the stratospheric/mesospheric wind vacillations and global normal modes [*Smith, 1985*]; while modes of inter-seasonal and inter-annual variability include the tropospheric El Niño-Southern Oscillation (ENSO; *Trenberth, 1997*); the stratospheric quasi-biennial oscillation (QBO; *Baldwin et al., 2001*); the stratospheric/mesospheric annual (AO) and semiannual (SAO) oscillations [*Paetzold et al., 1961; Garcia et al., 1997*], and the solar cycle (SC) variation [*Oberheide et al.,*

2009]. Figure 1 (from *Sassi et al.*, 2019) illustrates the complexity of the coupling processes affecting the ITM based on their characteristic time scale and vertical domain.

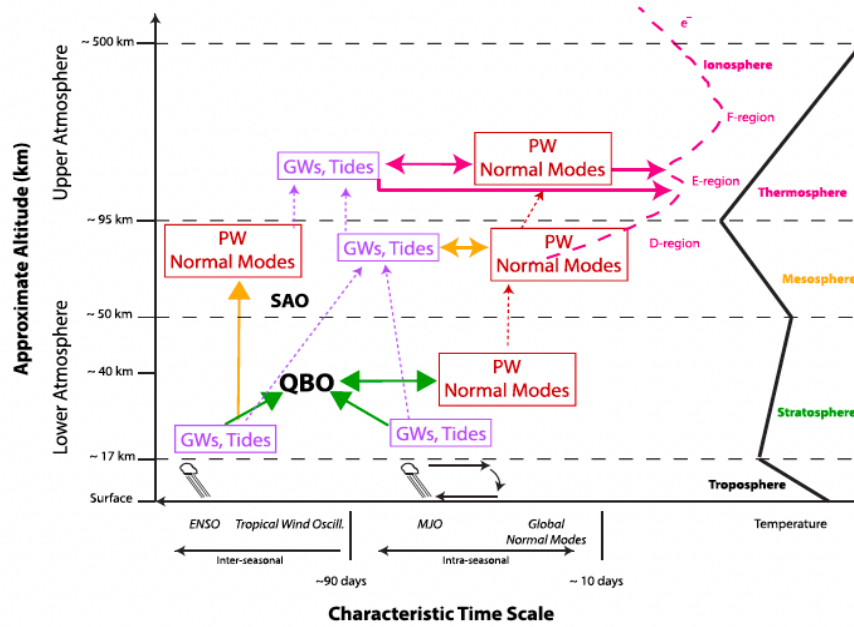


Figure 1 Illustration of the various processes affecting the ITM as a function of their characteristic time scale and vertical domain. Illustrations (not to scale) of typical temperature (thick black solid) and ionospheric electron density (purple dash) profiles are shown to the right. The solid arrows indicate interaction pathways, while the dashed arrows indicate the propagation directions in the vertical. [Figure from *Sassi et al.*, 2019].

3. INTRA-SEASONAL ITM VARIATIONS: CONNECTIONS TO THE MJO

The MJO is the dominant mode of intra-seasonal variability in tropical convection and circulation [Zhang, 2005] and is known to generate a whole spectrum of global-scale waves mainly through convective forcing [Wheeler and Kiladis, 1999]. The MJO was shown to modulate stratospheric GW, GW drag, and mean winds [Alexander et al., 2018] and lower and middle thermospheric tides and UFKWs [Kumari et al., 2020, 2021; Gasperini et al., 2020]. Li and Lu [2020] provided observational evidence that an ~15% peak-to-peak MJO-modulation of the GWs extends up to ~100 km altitude and into the extra-tropics, while Li and Lu [2021] found large MJO signals in MERRA-2 resolved GWs and parameterized GW drag at high northern latitudes during winter. Kumari et al. [2020] demonstrated that the MJO modulates the certain tidal amplitudes in the lower atmosphere by up to 25%, while Kumari et al. [2021] revealed that the modulation of tidal heating is comparatively more important than the modulation of background winds to impose the MJO signal on the low latitude E-region tides. The MJO-modulation of tides and UFKW has been shown to extend well into the thermosphere [Gasperini et al., 2017, 2020] with potential effects on the ionosphere either through E-region dynamo or direct upward propagation and/or composition changes. These recent studies suggest that the MJO may be responsible for an important fraction of the intra-seasonal variability observed in the ITM system, although work on this topic is at the very early stages.

4. INTER-SEASONAL ITM VARIABILITY: CONNECTIONS TO QBO, SAO, ENSO, SOLAR CYCLE

The QBO is the largest source of inter-annual variability in the tropical stratosphere [Baldwin *et al.*, 2001], and its influence extends to higher latitudes throughout the lower atmosphere [e.g., Anstey and Shepherd, 2014]. Ground- and space-based observations documented the presence of QBO-like oscillations in mesospheric winds over 25 years ago [e.g., Burrage *et al.*, 1996; Gurubaran and Rajaram, 1999; Vincent *et al.*, 1998]. There is now substantial observational and modeling evidence indicating that the QBO impacts inter-annual variability in lower thermosphere dynamics by affecting both tidal and GW breaking [Vincent *et al.*, 1998; Hagan *et al.*, 1999; Wu *et al.*, 2008a,b; Hibbins *et al.*, 2007; Pancheva *et al.*, 2009; Oberheide *et al.*, 2009; Xu *et al.*, 2009; Davis *et al.*, 2013; Gan *et al.*, 2014; Laskar *et al.*, 2016; Dhadly *et al.*, 2018]. Recent results by Yamazaki *et al.* [2017] suggest that the variation of atmospheric tides due to the stratospheric QBO could be an important source for inter-annual variability of the ionospheric wind dynamo. Yet, the question remains whether the QBO may have any measurable impact on the ionosphere.

The SAO in the tropical zonal wind is a mode affecting both the stratosphere and the mesosphere [Ern *et al.*, 2021; Smith *et al.*, 2017], with SAO amplitudes and phases varying between these two atmospheric regions. It has been shown that there is a relationship between the QBO and SAO and that at times their phases can align. There is also evidence of the role of the QBO influencing the mesospheric SAO; this is thought to be due to QBO filtering of large equatorial waves and GWs. These changes in the mesosphere map onto changes observed in the D-region ionosphere, where SAO signatures have been observed in Total Electron Content (TEC) and ion temperature observations [Silber *et al.*, 2016].

The ENSO ocean-atmosphere coupling phenomenon is recognized as an important source of inter-annual variability in the lower thermosphere [Gurubaran *et al.*, 2005; Lieberman *et al.*, 2007; Pedatella and Liu, 2012, 2013; Warner and Oberheide, 2014; Liu, 2016; Liu *et al.*, 2017; Sun *et al.*, 2018]. The ENSO-related variation in diurnal tides in the lower thermosphere can modulate the ionospheric wind dynamo, coupling ENSO variability to the ITM system. Studies on the ENSO signature in the ionosphere are rare and challenging [Pedatella and Forbes, 2009; Pedatella and Liu, 2013; Chang *et al.*, 2018]. Pedatella and Liu [2013] simulated the influence of the ENSO-driven tidal variability on the low-latitude ionosphere and showed that the interannual tidal variability in the lower thermosphere can introduce 10-15% variability in the $E \times B$ vertical drift velocity and ionosphere peak density. The ENSO can also modulate the stratospheric QBO [e.g., Taguchi, 2010] as well as the QBO components in the temperature DW1 and DE3 in the lower thermosphere [e.g., Sun *et al.*, 2018]. Despite this recent progress, connections between ENSO and the ITM system are not as well established.

Solar cycle effects on tidal dissipation have been demonstrated in various studies [e.g., Oberheide *et al.*, 2009, Jones *et al.*, 2016]. Oberheide *et al.* [2009] (and references therein) showed that tidal dissipation becomes more important as solar activity increases. Jones *et al.* [2016] showed that the solar cycle variability in tidal-induced zonal-mean temperature changes results from tidally driven increases in nitric oxide (NO) infrared radiative (IR) cooling and that tidal modifications to the ionosphere are quite substantial through tidal-induced temperature and constituent changes. Mlynczak *et al.* [2010, 2014] reported strong solar cycle dependence in NO and carbon dioxide (CO₂) IR cooling in the thermosphere implying that the tidal effects on NO and CO₂ IR cooling

could modulate its inherent solar cycle behavior with implications for the whole ITM system. More modeling- and observational-based studies are needed to address the physical mechanisms at play.

5. OPEN QUESTIONS

It is now well established that internal waves play a leading role in impressing their long-term variability from the lower and middle atmosphere into the lower thermosphere and ionosphere. A growing body of recent evidence suggests that upward propagating waves may be the leading driver of long-term variability in the whole ITM system. Variability in the wave spectrum can be ascribed to changes in sources associated with tropospheric weather, variable propagation conditions, and nonlinear interactions between different parts of the wave spectrum [Yigit and Medvedev, 2015; Liu, 2016; Sassi *et al.*, 2019]. As their generation and propagation conditions may be undergoing modifications in a changing climate [Ward *et al.*, 2021], it is critical that we attain a better understanding of the physical mechanisms at play for improving our modeling and predictive capabilities.

Despite recent improved observational capabilities afforded by NASA's Ionospheric Connection Explorer (ICON, Immel *et al.*, 2018) and Global-scale Observations of the Limb and Disk (GOLD, Eastes *et al.*, 2017) missions, our ability to attain a comprehensive physical understanding of the processes at play is significantly impaired by the data sparsity in the ITM. Without global measurements with sufficient temporal and spatial resolution, physics-based models cannot be validated, and data assimilation for these heights remains a tentative venture. The establishment and maintenance of suitable observing capabilities are thus critical to allow for the dynamical conditions to be monitored.

The upcoming Geospace Dynamics Constellation (GDC) mission shall provide critical observations to better understanding long-term ITM coupling processes. Observations from the Dynamical Neutral Atmosphere-Ionosphere Coupling (DYNAMIC) mission will be particularly helpful by providing critical information on the height evolution of the wave spectrum in the thermosphere. GDC and DYNAMIC are expected to provide much-needed day/night wind, temperature, and composition observations throughout the thermosphere and ionosphere that will enable the investigation of wave-mean flow interactions, ion-neutral interactions, and dynamo processes critical for the study of long-term ITM variability in current and future climates.

A few, but not all, of the important open questions that need to be addressed in this area of research include:

1. *What are the physical mechanisms that transmit intra-seasonal, inter-seasonal, and inter-annual variability from the lower and middle atmosphere into the ITM system, and what is their relative importance?*
2. *What are the influences of lower atmospheric waves on the long-term trends of the ITM system? In particular, what are their impacts on eddy diffusion and CO₂ cooling rates that drive variations in O/N₂ and thus thermospheric density?*
3. *How are the influences of lower atmospheric waves on the long-term trends of the ITM system changing with respect to a changing solar energy budget on longer time scales?*

Simultaneous measurements from both GDC and DYNAMIC are critical to answering these outstanding questions along with many others.

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