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Article:

Huang, X, Yang, S, Haywood, A orcid.org/0000-0001-7008-0534 et al. (6 more authors) (2023) Simulations reveal causes of inter-regional differences in Pliocene climatic periodicity. *Science Bulletin*, 68 (2). pp. 146-149. ISSN 2095-9273

<https://doi.org/10.1016/j.scib.2022.12.031>

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1 **Short Communication**

2 **Simulations reveal causes of inter-regional differences in Pliocene**
3 **climatic periodicity**

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16 Received 2022-09-07, Revised 2022-11-03, accepted 2022-11-04

17 The Pliocene Epoch (5.33 to 2.58 Ma) is the last period of sustained warmth
18 before the emergence of the large-scale Pleistocene glaciations. This period was
19 generally warm and wet and is often used as an analogue for near-future climate
20 conditions in terms of CO₂ levels (~400 ppm, 1 ppm=1 μmol/mol), and comparable
21 temperatures. Evaluating the climate system and its fluctuation patterns during the
22 Pliocene warm period provide scientists with the opportunity to understand a
23 warmer-than-present world, and are fundamental to our understanding and ability to
24 accurately project future climate and environmental change.

25 During the Pliocene, a dominant obliquity cycle (41 ka) was identified from
26 marine δ¹⁸O-based global ice volume reconstructions and high-latitude continental
27 records (Fig. 1; e.g., [1]). In contrast, dust and pollen records from low latitudes and
28 the Mediterranean region show a dominance of the precession cycle (21 ka) (Fig. 1;
29 e.g., [2]), and proxy records derived from aeolian deposition in the mid-latitude

30 monsoonal Asia display a mix of precession, obliquity, and eccentricity (100 ka)
31 cycles (Fig.1; e.g., [3]). However, the underlying mechanisms for these inter-regional
32 differences in climatic cycles remain to be clarified.

33 Numerical experiments have emerged as an efficient means of understanding
34 past climate over regional and global scales, and have been used to explore the
35 process and mechanism of Pliocene climate change [4]. Therefore, the dominant
36 orbital parameter drivers of the Pliocene climate variability can be studied using
37 simulations. In this study, we designed sensitivity experiments using a coupled
38 climate model to explore the effect of changes in orbital parameters on climate during
39 the Pliocene, with the aim of investigating the mechanism of inter-regional
40 differences in climatic periodicity.

41 The Hadley Centre coupled climate model version 3 (hereafter referred to as
42 HadCM3) was used for this study. Model description and detail boundary conditions
43 can be found in the supplementary material. Our simulations are based on Bragg et al.
44 [5] but used the MOSES 2 surface exchange scheme as described in Tindall et al. [6].
45 The other one difference between our simulations and Bragg's PlioMIP1 simulation is
46 that we use different orbital parameters than those applied in the PlioMIP1 simulation
47 which assigns modern orbital configuration [5].

48 We perform a set of idealized sensitivity experiments (Table 1) to isolate the
49 effect of variations in obliquity, precession, and eccentricity on climate during the
50 mid-Pliocene warm period. In briefly, the suite includes (1) maximum eccentricity,
51 maximum precession, and minimum obliquity (hereafter $E_{\max}P_{\max}O_{\min}$), (2) maximum
52 eccentricity, minimum precession, and minimum obliquity (hereafter $E_{\max}P_{\min}O_{\min}$),
53 (3) maximum eccentricity, maximum precession, and maximum obliquity (hereafter
54 $E_{\max}P_{\max}O_{\max}$), and (4) minimum eccentricity, maximum precession, and minimum
55 obliquity (hereafter $E_{\min}P_{\max}O_{\min}$). These eccentricity, obliquity and precession values
56 represent the extremes of their theoretical orbital variations obtained from Laskar et al.
57 [7] during the mid-Pliocene warm period. Except the orbital parameters, all other
58 boundary conditions in the sensitivity experiments are the same.

59 All experiments started from the end of the HadCM3 PlioMIP1 simulation, and

60 continued for another 160 ($E_{\max}P_{\max}O_{\min}$), 150 ($E_{\max}P_{\min}O_{\min}$), 240 ($E_{\max}P_{\max}O_{\max}$),
 61 and 200 ($E_{\min}P_{\max}O_{\min}$) model years until the model climate equilibrated to the
 62 boundary conditions. Climate statistics are based on time averages of the final 30
 63 years for each run. The results are presented as anomalies between the sensitivity
 64 experiments of the extremes of each orbital parameter, thereby allowing estimation
 65 each orbital parameter's effect on climate during the mid-Pliocene warm period.

66

67 *Table 1 Experimental design for the orbital extreme sensitivity experiments*

Experiment name	Eccentricity (E)	Precession (P)	Obliquity (O)	Boundary conditions	Years for analysis
$E_{\max}P_{\max}O_{\min}$	0.0607	Perihelion	22°	PRISM3D	30
$E_{\max}P_{\min}O_{\min}$	0.0607	Aphelion	22°	PRISM3D	30
$E_{\max}P_{\max}O_{\max}$	0.0607	Perihelion	24.5°	PRISM3D	30
$E_{\min}P_{\max}O_{\min}$	0.005	Perihelion	22°	PRISM3D	30

68

69 The effects of a change in Earth's obliquity (axial tilt) can be physically
 70 interpreted as a redistribution of the insolation by changing its meridional gradient,
 71 but at the same time keeping the total solar radiation incident upon the Earth the same
 72 [8]. An increase in the obliquity from 22.0° to 24.5° raises the annual mean air
 73 temperature (MAT) over most areas of the globe (Fig. 2a) but induces a small
 74 decrease at the equator (~ 1 °C). The corresponding zonally averaged annual MAT
 75 show a significant increase at high latitudes by more than 5 °C and a slight change by
 76 0–1 °C at the equator (Fig. 2b), which is consistent with the insolation changes (Fig.
 77 S1). The simulated annual MAT between precession minimum and maximum shows
 78 slight differences in most areas ($\sim \pm 2$ °C; Fig. 2c), and the corresponding zonally
 79 averaged annual MAT also shows a small change (± 1.5 °C) at most latitudes (Fig. 2d).
 80 For variation in eccentricity, the simulated annual MAT results demonstrate nearly
 81 pervasive warming over the globe (~ 1 –3 °C; Fig. 2e), with an ~ 2.5 °C increase in
 82 zonally averaged annual MAT at high latitudes and a warming of ~ 1 °C at low
 83 latitudes (Fig. 2f). Clearly, Variations in the Earth's orbital parameters (obliquity,

84 precession, and eccentricity) during the mid-Pliocene warm period have small effect
85 on the annual MAT consistently at most low latitudes, while obliquity change has the
86 most effect on the annual MAT at high latitudes.

87 The shift in obliquity from minimum to maximum values leads to a small
88 increase in mean annual precipitation (MAP; $\sim 0.5 \text{ mm d}^{-1}$) over most areas of the
89 globe (Fig. 2g) and a small decrease in MAP (~ -1 to 0 mm d^{-1}) at the equator (Fig. 2g,
90 h). The change in precession from aphelion to perihelion results in greatest
91 enhancement ($\sim 1.5\text{--}3 \text{ mm d}^{-1}$) and deficit ($\sim -2 \text{ mm d}^{-1}$) in MAP at low latitudes,
92 associated with small precipitation changes at high latitudes (Fig. 2i, j). The
93 eccentricity extreme-value experiments (Fig. 2k) show a spatial MAP pattern similar
94 to that observed in the obliquity experiments (Fig. 2g); i.e., a small precipitation
95 change ($\sim 0.5 \text{ mm d}^{-1}$) over the globe except a larger magnitude of change at low
96 latitudes ($\sim \pm 1.5 \text{ mm d}^{-1}$; Fig. 2l). Evidently, changes in the values of Earth's orbital
97 parameters (obliquity, precession, and eccentricity) during the mid-Pliocene warm
98 period have small effect on the MAP consistently at high latitudes, while precession
99 change has the most effect on the MAP at low latitudes.

100 Climatic periodicity during the Pliocene has been studied extensively using
101 various proxies, enabling us to assemble a dataset that can be used to analyze global
102 climatic periodicity during the Pliocene (Fig. 1 and Table S1 online). This dataset
103 contains 76 records (Table S1 online), of which 24, 22, and 6 show a dominance of
104 obliquity, precession, and eccentricity, respectively. The remaining records display a
105 mixed signal of the two (17 records) or three (7 records) of the astronomical
106 parameters.

107 High-latitude terrestrial records and the deep-sea $\delta^{18}\text{O}$ records, which are
108 influenced mainly by ice volume at both northern and southern high latitudes [9],
109 exhibit a 41-ka cycle [1]. Our simulation results show that obliquity change has the
110 greatest effect on the high-latitude temperature, highlighting the strong influence of
111 obliquity on high-latitude ice volume, which is consistent with geological records (Fig.
112 1) and previous simulations [8]. It should be noted that some low-latitude deep-sea
113 $\delta^{18}\text{O}$ records also reveal a dominance of the obliquity cycle, as the abundances of ^{18}O

114 and ^{16}O in the ocean water are uniformly controlled by high-latitude ice volume.

115 Low-latitude and Mediterranean records, such as dust, pollen, Mediterranean
116 sapropel formation and lake levels (Fig. 1; e.g., [2]), appear to be directly linked to
117 monsoonal variations, showing a predominant 21 ka precession period. These records
118 are consistent with our simulated results and previous simulations [10], which show
119 that precession change has the most effect on the low-latitude precipitation. The
120 precession variations affect monsoon by changing the seasonal solar radiation at the
121 top of the atmosphere, thereby influencing the regional hydrological cycle. Of
122 particular note is that Mediterranean $\delta^{18}\text{O}$ records also reveal a clear precession cycle
123 (Fig. 1; e.g., [11]), which differs from the dominant obliquity cycle in other deep-sea
124 $\delta^{18}\text{O}$ records (Fig. 1; e.g., [1]). This may result from the combined effect of the
125 semi-isolated nature of the Mediterranean Sea and the rate of evaporation controlled
126 by precession-modulated low-latitude insolation.

127 It is striking that the geological records from the Chinese Loess Plateau display a
128 mix of astronomical signals (Fig. 1). Numerous studies have demonstrated that the
129 dust of the Loess Plateau is transported by the East Asian winter monsoon, while the
130 soil formation and biological processes are controlled by the East Asian summer
131 monsoon (e.g., [12]). The winter monsoon intensity is closely related to ice volume
132 oscillations at high northern latitudes [13], and the summer monsoon is directly linked
133 to precession variations [14]. Our simulations show that obliquity and precession have
134 the most effect on the high-latitude temperature and low-latitude precipitation,
135 respectively (Fig. 2a, i), which well explain the mix of precession and obliquity
136 signals over the Chinese Loess Plateau.

137 It should be pointed out that during the Pliocene warm period, the occurrence of
138 the eccentricity cycle (100 ka), which has dominated the loess and ocean records since
139 the Middle Pleistocene transition [15], is hard to be explained by the simulation
140 results. Previous studies have interpreted the eccentricity signal in loess as a result of
141 the asymmetric response to insolation forcing [3]; this issue will require systematic
142 future investigation based on reconstructions and simulations.

143 In conclusion, the sensitivity of the climate to the extreme value of the orbital

144 parameters during the mid-Pliocene warm period was investigated using the HadCM3
145 model. Results show that the temperature changes induced by obliquity are >5 °C
146 over high latitudes and $\sim 0\text{--}2$ °C over low latitudes, while those induced by precession
147 or eccentricity are $\sim 0\text{--}3$ °C over high latitudes and $\sim 0\text{--}2$ °C over low latitudes. In
148 contrast, precipitation changes driven by obliquity, precession, and eccentricity are
149 consistently very small over high latitudes, while over low latitudes, the precession
150 induced precipitation changes are dramatic (>2 mm d⁻¹) and those induced by
151 obliquity or eccentricity are $\sim 0\text{--}1$ mm d⁻¹. Our results show that the most notable
152 effect of obliquity on climate occurs at high latitudes, whereas precession affects the
153 climate mainly over low latitudes. This well explains the regional differences in
154 climatic periodicity during the Pliocene: the predominant 41 ka climatic periodicity at
155 high latitudes and in marine oxygen isotope records regulated by high-latitude ice
156 volume, and the dominant 21 ka climatic cycle over low latitudes and the
157 Mediterranean region.

158 **Conflict of interest**

159 The authors declare that they have no conflict of interest.

160

161 **Acknowledgments**

162 This work was supported by the National Natural Science Foundation of China
163 (41725010 and 42107472), the Strategic Priority Research Program of Chinese
164 Academy of Sciences (XDB26000000 and XDB31000000), and the Key Research
165 Program of the Institute of Geology & Geophysics, Chinese Academy of Sciences
166 (IGGCAS-201905). We thank three anonymous reviewers for critical comments.

167 **Author contributions**

168 Xiaofang Huang, Shiling Yang, and Zhongli Ding designed the study; Xiaofang
169 Huang performed the analysis and wrote the draft; Yongda Wang, Minmin Sun, and
170 Shihao Zhang took part in the investigation; Xiaofang Huang, Alan Haywood, and
171 Julia Tindall designed the experiment; Shiling Yang and Zhongli Ding supervised the
172 project; Shiling Yang, Alan Haywood, Julia Tindall, Dabang Jiang revised the
173 manuscript.

174

175 **Supporting Information:**

176 Supporting information table S1

177

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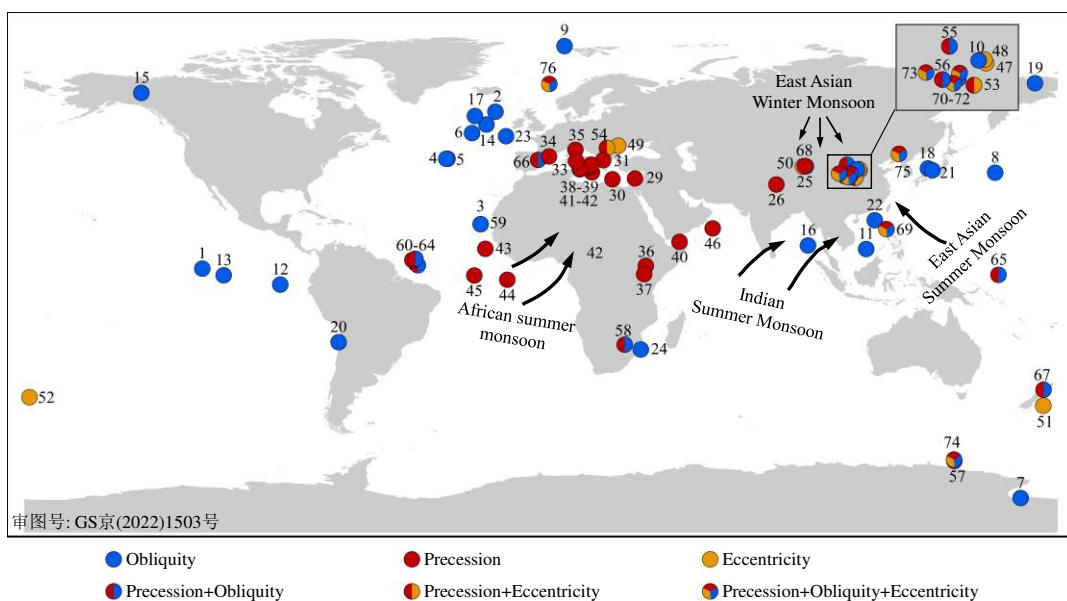
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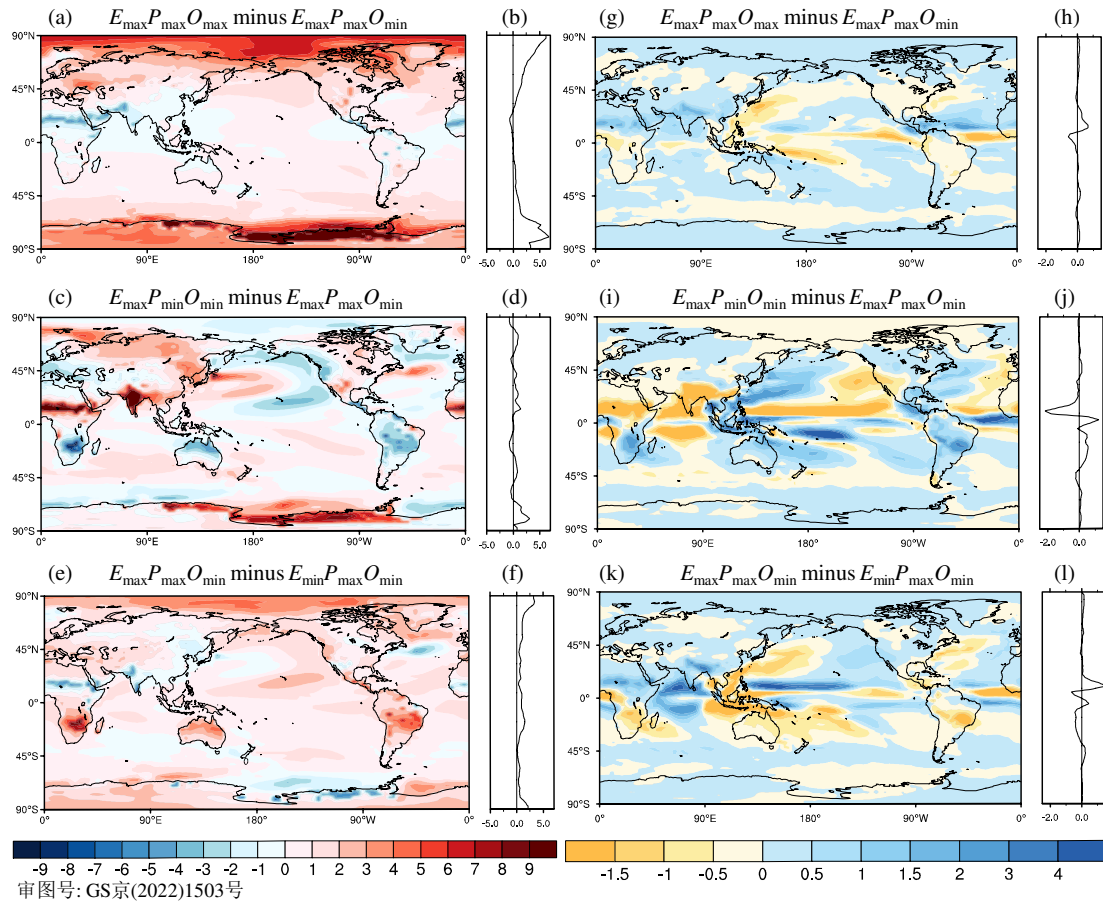
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223

224 Fig. 1. Paleoclimatic records for the climatic periodicity during the mid-Pliocene.

225 Numbers refer to sites listed in supplementary Table S1 (online). Circles with
 226 different colors represent different astronomical signals (see legend). Circles with two
 227 or three colors denote a mixed signal of two or three orbital parameters, respectively.
 228



229 审图号: GS京(2022)1503号
 230 Fig. 2. Annual mean surface air temperature (units: °C) and precipitation anomalies
 231 (units: mm d⁻¹) between the obliquity maximum and minimum (a, g); the precession
 232 maximum and minimum (c, i); and the eccentricity maximum and minimum
 233 variations during the mid-Pliocene warm period (e, k). Zonally averaged temperature
 234 (b, d, f) and precipitation (h, j, l) changes are plotted in the side panels.