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Morgado, E, Morgan, DJ orcid.org/0000-0002-7292-2536, Harvey, J orcid.org/0000-0002-0390-3438 et al. (6 more authors) (2022) The Magmatic Evolution and the Regional Context of the 1835 AD Osorno Volcano Products (41°06'S, Southern Chile). Journal of Petrology, 63 (11). egac105. ISSN 0022-3530

https://doi.org/10.1093/petrology/egac105

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AD Osorno volcano products (41°06´S, Southern Chile)

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23 ABSTRACT

24 Osorno volcano (41°06'S, 72°20'W) is a composite stratovolcano of the Central Southern 25 Volcanic Zone of the Chilean Andes. It is the southernmost member of a NE-SW trending 26 alignment of volcanic edifices including La Picada and Puntiagudo volcanoes and the Cordón 27 Cenizos chain. According to contemporary descriptions recorded by Charles Darwin in 1835, 28 two eruptive events occurred: the first during January-February, and the second during 29 November-December 1835 and January 1836. The volcano erupted basaltic andesite lavas 30 and tephra fall deposits (52.4 to 52.9 SiO₂ wt. %) which contain phenocrysts of olivine, 31 plagioclase, clinopyroxene, and spinel. The compositions of these phenocryst phases, 32 together with those of olivine-hosted melt inclusions, allowed us to constrain intensive 33 parameters for the pre-eruptive magmas. These varied from 1060 to 1140 °C, with an oxygen 34 fugacity buffer of $\sim \Delta QFM + 1.1$, dissolved water concentrations of up to 5.6 wt. % (average of ~4.2 wt. %) and maximum pressures equivalent to ~7 km depth. Textural relations, such 35 36 as crystal accumulations and clots, zoning in crystals and other indications of disequilibrium, 37 lead us to infer the involvement of a crystal mush, rich in individual crystals and clots of 38 crystals, which underwent a degree of disaggregation and entrainment into the transiting 39 magma prior to eruption. Comparison of trace element abundances, including rare earth 40 elements, fluid-mobile elements, and relatively fluid-immobile elements, combined with 41 ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd isotope ratios, allow us to consider variations in slab-derived fluid 42 input and the minor role of crustal contamination on the Osorno eruptive products and those 43 from neighbouring volcanic systems. Our results suggest both a greater contribution from 44 slab-derived fluid and a higher degree of partial melting in the systems supplying

45 stratovolcanoes (Osorno, Calbuco, and La Picada) relative to those supplying small eruptive46 centres built over the major regional Liquiñe-Ofqui Fault Zone.

47 INTRODUCTION

48 In subduction zones, geochemical features of relatively evolved volcanic products across-arc 49 may be preserved (Watt et al., 2013). This allows to contrast whole-rock chemical 50 compositions of products from different volcanoes to understand mantle- and slab-related 51 contributions. Specifically in Southern Andean volcanoes, several studies (e.g., Hickey-52 Vargas et al., 1989; 2002; López-Escobar et al., 1992; 1995a; Watt et al., 2013; Jacques et 53 al., 2014; Morgado et al., 2015; Rawson et al., 2016; McGee et al., 2017) have reported 54 chemical variations of erupted material over small spatial scales and have attributed crustal 55 and subcrustal processes as causes for the observed variations in magma composition. 56 However, studies combining major and trace elements, isotopic ratios, and mixing modellings of products from different volcanoes at the same latitude, are uncommon. 57 58 Besides whole-rock chemistry, recent articles (e.g., Ruth et al., 2018; Weller and Stern, 2018; 59 Morgado et al., 2019a; Couperthwaite et al., 2020; Tassara et al., 2020) have also considered 60 melt inclusion composition, mineral chemistry, and diffusion chronometry in volcanic rocks, 61 which have placed observed chemical variabilities in a geological context, permitting events 62 occurring during the magmatic evolution to be identified. Particularly, melt inclusion 63 compositions, which have been used to constrain pressure of the magmatic storage and water 64 dissolved in melt, could undergo significant reequilibration after recrystallisation in the rims 65 and volatile loss, leading to misinterpretations of magmatic processes.

66 In Southern Chile, from 1575 to 1835, Osorno volcano had an eruptive frequency of one 67 event per ~25 years, but its last eruption occurred in 1835 (Darwin, 1839a; Moraleda y 68 Montero, 1888; Martin, 1901; O'Compley, 1936; Petit-Breuilh, 1999). Despite this relatively 69 long quiescence, Osorno has shown an elevation of seismicity base level in recent years, 70 interpreted as a sign of volcanic unrest (SERNAGEOMIN-OVDAS 2017a; 2018a, b, c; 71 2019; 2022). Here, we report our findings on the last eruption cycle (in 1835 AD) of Osorno 72 volcano (Central Southern Volcanic Zone of the Southern Andes of Chile; Fig. 1); eruptions 73 that lack detailed studies of the magmatic conditions that preceded volcanic activity. We 74 describe crustal magmatic processes and assess the pre-eruptive physical magmatic 75 conditions via mineral chemistry and coupling melt inclusion compositional profiles and 76 diffusion chronometry. Finally, to place the Osorno volcano into a wider context with 77 neighbouring volcanic activity, we compare and contrast the nature of the chemical 78 enrichment/depletion of the erupted products across-arc from volcanoes at roughly the same 79 latitude (41°06'S) from W to E, i.e., from Calbuco, Osorno, La Picada and La Viguería 80 volcanoes. Comparison of volcanoes at the same latitude decreases the uncertainties related 81 to the mantle source and minimises the potential effects of latitudinal variability in the supply 82 of material associated with trench sediments (Watt et al., 2013) or other geodynamic 83 anomalies (e.g., subduction of oceanic fracture zones; Sellés et al., 2004).

84 **GEOLOGICAL SETTING**

85 Southern Volcanic Zone

86 The Southern Volcanic Zone (SVZ; Fig. 1) of the Andes is a volcanic chain produced by the87 subduction of the oceanic Nazca plate below the South American continental plate (e.g.,

88 Jarrard 1986; Dewey and Lamb 1992; DeMets et al., 1994; Angermann et al., 1999). The 89 SVZ comprises over 70 Pleistocene and Holocene stratovolcanoes and several large volcanic 90 fields (Stern et al., 2007). The whole volcanic segment is ~1400 km long (33° to 46°S) and 91 is accompanied by the NS regional Liquiñe-Ofqui Fault Zone (LOFZ), which is an intra-arc 92 fault that runs along-strike, just like the much of the arc volcanism between ~38° and ~46°S 93 (Cembrano and Lara, 2009; Sánchez et al., 2013; Pérez-Flores et al., 2016). The SVZ has 94 been subdivided into four segments on the basis of tectonic setting, geochemical and 95 petrologic considerations (Tormey et al., 1991; López-Escobar et al., 1995a; Stern, 2004; 96 Stern et al., 2007). These subdivisions are shown in Fig. 1 and comprise the Northern (NSVZ, 97 33.3 to 34.4°S), Transitional (TSVZ, 34.4 to 37°S), Central (CSVZ, 37 to 42°S), and Southern 98 (SSVZ, 42 to 46°S) zones.

99 The present study focuses on volcanic centres from the CSVZ, the most active region of the 100 SVZ (and the whole Andes; Stern et al., 2007), in particular on Osorno volcano considering 101 a geochemical comparison with products from other volcanoes at the latitude of ~41°S: La 102 Picada, and Calbuco stratovolcanoes and La Viguería cone. Osorno volcano is the 103 westernmost member of a NW alignment together with La Picada, Puntiagudo, and Cordón 104 Cenizos (Moreno et al., 2010; Vander Auwera et al., 2019; Fig. 2). The nearby Calbuco 105 volcano is not located directly upon any major regional structure (Sellés and Moreno, 2011). 106 Conversely, La Viguería cone (41°22'S) belongs to Cayutué-La Viguería small eruptive 107 centres built over the LOFZ master fault. According to the conceptual model of Cembrano 108 and Lara (2009), NW alignments are related to tension cracks at depth, which may reach the 109 surface and then provide pathways for magma ascent and, therefore, allow the building of either a stratovolcano or an elongated cluster of minor eruptive centres, depending on the 110

balance between strain rate and magma input (cf. Takada, 1994). Magmas erupted at small
eruptive centres can be shown to have ascended rapidly (e.g. Morgado et al., 2017), and with
no pronounced interaction with the crust (e.g., McGee and Smith, 2016; Hickey-Vargas et
al., 2016; Smith and Németh, 2017). This style of volcanism is usually related to extensional
domains (e.g., Nakamura, 1977; Takada, 1994; Piochi et al., 2005).

116 Osorno volcano: general features

Osorno is an active stratovolcano of the Central Southern Volcanic Zone of the Chilean Andes. Its maximum height is 2661 meters above sea level (m.a.s.l.) and its base level is between 50 and 250 m.a.s.l. with an estimated volume of 130 to 160 km³, covering an area of 250 km² (Moreno et al., 2010; Völker et al., 2011). It is located at 41°06'S, 72°20'W, and lies ~13 km west of the main trace of the LOFZ (Fig. 2). Osorno volcano is the southernmost member of an NE-SW alignment that includes La Picada volcano and the Puntiagudo– Cordón Cenizos volcanic chain (Fig. 2).

124 Osorno volcano has been active since the Middle Pleistocene (ca. 200 ka; Moreno et al., 125 2010). Its postglacial activity has been reported as summit crater eruptions, fissure-fed lava 126 flows, parasitic cones and dacitic domes on the flanks (Moreno et al., 2010). The basement 127 beneath Osorno volcano comprises Cretaceous granitoids (ca. 70 Ma; Moreno et al., 2010) 128 and Miocene plutonic rocks from the Northern Patagonian Batholith (López-Escobar et al., 129 1992; Adriasola et al., 2006). Vander Auwera et al (2019) found Miocene-age (ca. 9.6 Ma) 130 gabbroic xenolith blocks within the volcanic products of the adjacent La Picada volcano and 131 Hickey-Vargas et al. (1995) described gabbroic and granulitic xenoliths recovered from

- 132 volcanic material of the 1961 eruption of the neighbouring Calbuco volcano. All the xenoliths
- 133 from previous studies have been interpreted as being derived from the local basement.
- 134 The products of Osorno volcano vary from basalts to dacites (~50-70% SiO₂; e.g., Tagiri et
- 135 al., 1993; Moreno et al., 2010; Bechon et al., 2022) and have been classified into four units
- 136 (1, 2, 3, and 4) depending on the age of the deposits. In the unit Osorno 4, several historical
- 138 1834 to 1835 have been recognised and briefly described, whereas activity since 1834 has

eruptions during the years 1575, 1640, 1644(?), 1719, 1737, 1765, 1778 to 1779, 1790, and

- 138 1834 to 1835 have been recognised and briefly described, whereas activity since 1834 has
- 139 been confined to fumarolic episodes in 1852, 1882 to 1883, 1911 to 1915, 1963, 1969, 1985,
- 140 and 1993 (Petit-Breuilh, 1999).

137

141 The 1835 AD Osorno eruption: History

142 Charles Darwin (1839b; 1840) described two eruption stages of Osorno volcano during 1835. 143 Although he recognised release of "smoke" in November 1834 (Darwin, 1839a), the first stage of the 1835 Osorno eruption started on January 19th with moderate to vigorous 144 145 strombolian activity (VEI 2; according to the Pyle (2000) catalogue; Lara et al., 2012). The volcano was still erupting in this manner on February 18th, two days before the large-146 147 magnitude earthquake (M_w 8.0-8.3; Watt et al., 2009) of Concepción, Southern Chile 148 (Darwin, 1840). During the shock period, Darwin (1839b) described the emission of a dark 149 blue "smoke" column from the summit, following the eruption cessation, ending the first 150 period of eruption. Basaltic lavas are associated with this first period (Moreno et al., 2010), 151 which were emitted from an array of eruptive fissures and small pyroclastic cones (see Fig. 152 3). According to Darwin (1840), the second stage of eruptive activity, which started on 153 November 11th, resumed from the vents of the first stage (Lara et al., 2012). During December

154 5th, a fissure eruption occurred, associated with lava flows (Gillis, 1855) and lahars, which
155 burnt trees (Petit-Breuilh, 1999). The volcano was still erupting on the January 20th 1836
156 (Gillis, 1855).

157 SAMPLING AND ANALYTICAL METHODS

158 Sample collection

159 We collected samples from three volcanic systems: Calbuco, Osorno, and La Viguería.

160 *Calbuco volcano products.*

161 Calbuco samples correspond to four lapilli-sized pyroclasts from fall deposits and five bombs
162 from pyroclastic density current deposits from the 2015 eruption. The locations from which
163 they were taken, detailed quantification of the main petrographic features, and whole-rock
164 chemistry (major, minor, and trace elements) of Calbuco samples are available in Morgado
165 et al. (2019a).

166 *1835 AD Osorno volcano products.*

The collected samples from the 1835 AD Osorno eruption correspond to ten from lava flows erupted from the main crater (reaching lengths from 10 to 14 km long) from five locations (Table 1) and lapilli fall deposits erupted from parasitic cones in one location (Fig. 3; Table 1). Examination and quantification of the main petrographic features (e.g. crystallinity, vesicularity, and modal percentages of phenocryst phases) from the collected samples was done by optical microscope and image processing of pseudo-coloured images via the freeware JmicroVision, respectively (Roduit, 2005).

174 La Viguería cone.

The collected samples from La Viguería cone correspond to one from a lava flow (~1.2 km
length) and one from fall deposits (locations in Table 1). We performed the quantification of

- 177 the main petrographic features using the same methods described for Osorno samples.
- 178 Whole-rock major elements

179 Whole-rock major elements compositions of lapilli deposits of one location (Os-164), three 180 lava samples from three locations from the 1835 AD Osorno eruption (Os-144, Os-147, and 181 Os-154), and two samples (one lava and one bomb) from La Viguería monogenetic cone 182 (Fig. 2) were determined on fused glass discs by XRF (major elements) at ACME labs (now 183 Bureau Veritas Minerals, Canada) using OREAS-184 and SY-4 standards for all the major 184 elements except for FeO, for which CCU-1D, and OREAS-700 standards were used for FeO. 185 Accuracy for major elements in all standards is better than 3% (relative) except for TiO₂ in 186 SY-4 which is 4.5% (relative) and P₂O₅ in OREAS-184 and SY-4, which were 14% and 17% 187 (relative), respectively. The detection limits are reported in Table 2.

188 Whole-rock trace element abundances

The major and trace element chemistry of the 1835 Osorno volcano products and La Viguería 189 190 cone was obtained from bulk-rock powders. All trace element measurements were performed 191 on an ICP-MS Agilent 7500a at The Open University (Milton Keynes, UK). 0.1 g of finely 192 powdered sample was weighed out and 1 mL of Romil UpA HNO₃ and 4 mL of Romil UpA 193 HF were added. The samples were heated in sealed Teflon vials to 130 °C and agitated in an 194 ultrasonic bath for 20 minutes every 12 hours until they had completely dissolved. The 195 samples were then dried down to incipient dryness and the resulting gel brought up in 2 mL 196 of Romil UpA HNO₃ and agitated in an ultrasonic bath for 20 minutes then returned to the

hot plate overnight. After drying down for a second time the samples were redissolved in
concentrated UpA Romil HCl and returned to the hot plate overnight, before drying down a
third time. The residue was converted to a nitrate matrix using 1 mL of concentrated UpA
HNO₃, dried down and re-dissolved in 100 mL of 2 % UpA HNO₃ to give a final solution
with a thousandfold dilution, ready for mass spectrometry.

202 All element concentrations were corrected for blank effects (which resulted in corrections of 203 less than 1 % for all elements) and subsequently corrected offline for instrumental drift. 204 Where more than one analysis of a sample was undertaken during the same run the mean 205 value was taken prior to offline corrections were performed. Detection limits were estimated 206 to be five times the blank concentration for each element analysed and all data points with 207 concentrations lower than the detection limits were discarded. The precision was better than 208 8% (relative, 2σ) and accuracy in the trace elements of standards JB-2 (n = 5), BHVO-2 (n = 209 2) and AGV-1 (n = 2) is better than 10% (relative) except in Ni (21% relative, in AGV-1) 210 and Pb (13% relative, in BHVO-2). The uncertainties (2σ) of measurements are reported in 211 Table 2.

212 Strontium and neodymium isotope measurements

Strontium and neodymium isotopes were measured in volcanic products from Osorno volcano, La Viguería cone (see section 4.1), and six samples from 5 locations on the nearby Calbuco stratovolcano (Cal-149Ta, Cal-157, Cal-158, Cal-159, Cal-160; Morgado et al., 2019a). All ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd analyses in this study were conducted on a Thermo Scientific Triton Thermal Ionization Mass Spectrometer (TIMS) fitted with 10¹¹ Ohm amplifiers at the University of Leeds. Analyses were performed on ca. 100 mg of bulk rock powders digested using a combination of Romil-UpA HF, HNO₃ and HCl, using the same

220 method as described for the trace element determinations. Once in solution, samples were centrifuged prior to loading onto Eichrom Sr SpecTM columns with a bed volume of 221 222 approximately 100 µL. The REE fraction was collected before elution of the Sr fractions into 223 different vials. Prior to loading onto W filaments with a TaCl₅ activator, the Sr fraction was 224 passed through the Sr spec column a second time and dried with two drops of dilute HClO₄ 225 to remove any organic contaminants derived from the resin. The REE fraction was passed first through a TRU SpecTM column (ca. 100 µL bed volume) and finally a LN SpecTM column 226 227 (bed volume ca. 2 mL) before loading onto Re filaments. The USGS reference material 228 BHVO-2 was processed alongside the samples (Table 3), together with two total procedural blanks spiked with highly enriched ⁸⁴Sr, and ¹⁵⁰Nd solutions. Mean total procedural blanks 229 230 for Sr and Nd were 326 pg and 153 pg, respectively. Accuracy of the radiogenic isotope 231 analyses was assessed by analysing standard solutions of known composition alongside the 232 unknowns – NBS987 for Sr, La Jolla for Nd, and the reference material BHVO-2 for Sr and Nd, which yielded 87 Sr/ 86 Sr = 0.710241 ± 0.000006 for NBS987, 143 Nd/ 144 Nd = 0.511842 ± 233 0.000002 for La Jolla, 87 Sr/ 86 Sr = 0.703461 ± 0.000007 and 143 Nd/ 144 Nd = 0.512977 ± 234 235 0.000004 for BHVO-2, in agreement with published results (e.g., GeoReM and GEOROC 236 databases; Weis et al., 2005).

A static multicollection routine with standard cup configurations was used for all of the isotopic analyses. All measured ⁸⁷Sr/⁸⁶Sr ratios were corrected for mass bias relative to ⁸⁸Sr/⁸⁶Sr of 8.375209, using an exponential law. Data acquisition comprised the integration of two hundred 8 second measurements of masses 84, 85, 86, 87 and 88 for Sr isotope analyses. Mass 85 was measured to correct for any Rb interference on mass 87 assuming a natural ⁸⁷Rb/⁸⁵Rb of 0.386. For Nd isotope measurements, masses 143, 144, 146, 147, 148 and 150 were acquired in the same way as Sr isotopes. Mass 147 was measured to correct for
Sm interferences on mass 144, 148 and 150. Natural ¹⁴⁴Sm/¹⁴⁷Sm, ¹⁴⁴Sm/¹⁴⁸Sm, and
¹⁴⁴Sm/¹⁵⁰Sm of 0.2048, 0.7498 and 0.4923, respectively were assumed for this correction.

246 Mineral chemistry

247 The phenocryst and microlite phases were identified using Scanning Electron Microscopy 248 (FEI Quanta 650) at Leeds Electron Microscopy and Spectroscopy Centre at the University 249 of Leeds (LEMAS, Leeds, UK). The compositions of olivine, pyroxene, plagioclase, spinel, 250 ilmenite, titanomagnetite, and glass were measured using an electron probe microanalyser 251 (JEOL JXA8230) at LEMAS (Leeds, UK). Details of the analytical conditions of electron 252 microprobe measurements are available in Table 4. For ilmenite, pyroxene, and spinel the values of Fe³⁺ were obtained following the formulations of Droop (1987) and for 253 254 titanomagnetites, using the formulations of Stormer (1983).

255 Crystallographic orientation

Crystallographic orientations of olivine crystals were determined using electron back-scatter
diffraction in olivine crystals (EBSD; Prior et al. 1999) on a FEI Quanta 650 FEGSEM
equipped with a Nordlys EBSD camera at the School of Earth and Environment, University
of Leeds (Leeds, UK).

260 **RESULTS**

261 **Petrography**

262 Osorno 1835 AD products

263 The lava flow samples have a vesicularity of 4 to 17 vol. % and the lapilli samples have a

higher vesicularity of 31 to 61 vol. % (Table 5). The main phenocryst phases are plagioclase,

265 olivine, and clinopyroxene, as described for other Osorno materials in previous studies 266 (Tagiri et al., 1993; Bechon et al., 2022). The nature of the groundmass depends on the 267 sample; those collected from lava flows (Table 5) possess a glass-free microcrystalline 268 groundmass, and those corresponding to lapilli samples contain microlites and interstitial 269 The microlite phases comprise plagioclase, olivine, clinopyroxene glass. and 270 titanomagnetite. We also recognised several volcanic cognate xenoliths (up to ~1 cm long) 271 in the lapilli deposits at location Os-164 (Fig. 3). Their mineralogy comprises plagioclase, 272 orthopyroxene, olivine, ilmenite, and titanomagnetite as phenocrysts, with plagioclase, 273 clinopyroxene and titanomagnetite as microlites (details in Supplementary Data 1) in a glass-

- free groundmass.
- 275 La Viguería cone
- The lava flow sample has a vesicularity of ~12 vol. % and the lapilli sample has a higher vesicularity of ~43 vol. % (Table 5). The main phenocryst phases are olivine and clinopyroxene and the microlite phases identified using Scanning Electron Microscopy comprise plagioclase, olivine, and clinopyroxene (Table 5).
- 280 Mineral chemistry and textural features of Osorno 1835 products
- 281 The main features of the mineral phases from the samples of the 1835 AD Osorno eruption
- 282 are available in the Table 6.

283 *Plagioclase*.

284 Plagioclase crystals are found as isolated crystals, as inclusions in olivine and clinopyroxene,

and as crystal clots, together with olivine and clinopyroxene. They are characterized by

compositional zoning that can be categorized into three distinct zones: Zone 1 corresponds

to cores of phenocrysts with composition of An_{80-89} . Zone 2 corresponds to plagioclase surrounding Zone 1 or core composition where Zone 1 is absent, with compositions of An_{70} . 79. Zone 3 (An_{60-66}) corresponds to plagioclase rims, mainly observed in plagioclase phenocrysts from lava samples (Fig. 4). Plagioclase exhibit a composition equivalent to Zone 3 of plagioclase phenocrysts and display no zoning patterns. Plagioclase microlites have compositions of An_{52-59} .

Several plagioclase phenocrysts exhibit resorption textures in Zone 1 and Zone 2.
Additionally, the composition observed in plagioclase phenocryst rims and that filling the
spaces created because of resorption are the same and correspond to Zone 3 (Fig. 4).

296 *Olivine*

297 Olivine phenocrysts are present as both isolated crystals (e.g., Figure 5) and clots of crystals 298 (together with plagioclase, and clinopyroxene). There are two groups of olivine phenocrysts 299 regarding core compositions: Group 1 corresponds to crystals in the compositional range of 300 F076-79. Group 2 are crystals in the compositional range of F069-73. All the olivine phenocrysts 301 have normal zoning consisting of thin rims (< 15 μ m) of composition Fo₆₇₋₇₅. EBSD maps 302 (cf. Prior et al., 1999) show no measurable deformation in olivine phenocrysts (Fig. 6). 303 Olivine microphenocrysts exhibit the same compositions as Group 1 and Group 2 of olivine 304 phenocrysts (including ~3 µm thick rims with normal zonation). Olivine microlites have a 305 composition of Fo₅₂₋₆₀.

306 Pyroxene

Clinopyroxene phenocrysts occur as isolated crystals, inclusions in olivine, and as members
of crystal clots. The clinopyroxene compositional range corresponds to augite: En₄₃₋₅₂Wo₃₃₋₄₄Fs₈₋₁₆. The pyroxenes coupled with olivine grains from Group 1 in crystal clots have a
composition in the range of En₄₈₋₄₉Wo₄₀₋₄₂Fs₈₋₉ and those coupled with olivine grains from
Group 2 have a composition in the range of En₄₃₋₅₂Wo₃₃₋₄₄Fs₁₁₋₁₆.

312 Spinel

313 Spinel crystals are found as octahedral crystals typically preserved as inclusions in olivine 314 phenocrysts and rarely (< 5% of the Cr-spinel crystals) as isolated crystals. The Mg# and Cr# 315 $(Cr \#= molar Cr/(Fe^{3+}+Al+Cr))$ values yield ranges from 0.28 to 0.44 and from 0.27 to 0.36, 316 respectively. Spinel as inclusions in olivine phenocrysts from Group 1 have a Mg# ranging 317 from 0.36 to 0.44 and Cr# from 0.32 to 0.36, whereas those preserved as inclusions in olivine 318 phenocrysts from Group 2 have a Mg# from 0.28 to 0.29 and Cr# from 0.27 to 0.28. Measured 319 SiO₂ concentrations were low (< 0.2 wt. %) suggesting that analyses were not affected by 320 significant secondary fluorescence (Coogan et al., 2014) and that all the measured spinel 321 contents are reliable.

322 *Glass*

323 Glass occurs in the groundmass, interstitial between crystal clots, and as melt inclusions in 324 olivine and plagioclase phenocrysts (Fig. 4; 5). Olivine-hosted melt inclusions are generally 325 spherical with radii from 15 to 150 μ m, whereas plagioclase-hosted melt inclusions have 326 elliptical outlines with long-axis lengths from 40 to 160 μ m and short-axis widths from 20 to 327 50 μ m. Different compositional groups are recognised in glasses; glass from groundmass, 328 which has SiO₂ contents of 53.9 to 57.1 wt. %, versus interstitial glass between crystals from

- 329 clots and olivine-hosted melt inclusions, which both exhibit SiO₂ contents of 51.9 to 54.7 wt.
- 330 % (see Table 7, details in Supplementary Data 2).

331 Major and trace element geochemistry

332 The collected 1835 Osorno eruption samples are comparatively uniform with respect to their 333 major element concentrations. All are of basaltic andesite composition (52.4 to 52.9 SiO_2 wt. 334 %; Fig. 7) with high Al₂O₃ contents (19.6 to 20.7 wt. %), MgO contents from 4.09 to 4.55 wt. %, Mg# (Mg#= molar Mg/(Mg+Fe²⁺)) between 0.56 and 0.59, K₂O from 0.49 to 0.51 wt. 335 % and CaO contents from 10.2 to 10.9 wt. %. The 1835 Osorno products plot in the tholeiitic 336 337 trend (Fig. 8; Calbuco products in the tholeiitic trend and La Viguería products plot in the 338 calc-alkaline trend). Comparison with other nearby volcanic products erupted during the 339 Holocene can be seen in Fig. 7. Trace element concentrations are also similar between 340 samples, with invariant multi-element diagram and Rare Earth Element (REE) patterns and 341 a narrow range of La_N/Yb_N (2.1 to 2.2; Fig. 9, where the subscript N refers to values 342 normalised to Primitive Mantle concentrations; Sun and McDonough, 1989). The 1835 343 Osorno products show similar values and REE patterns to volcanic materials from the nearby 344 Calbuco and La Picada volcanoes, and the La Viguería cone (Fig. 9). In addition, the incompatible trace element diagram of the 1835 Osorno products shows similar trends in all 345 346 samples with positive Pb and Sr anomalies and negative Nb, P, and Ti anomalies, typical of 347 magmas from arc settings (Sun and McDonough, 1989; Fig. 9). The trends are consistent 348 with those reported by Bechon et al. (2022).

349 Sr-Nd isotopic compositions

350 Osorno (samples Os-144a, Os-147a, Os-154a, and Os-164a; see locations in Fig. 3) and 351 Calbuco (samples Cal-149Ta, Cal-149Tb, Cal-157a, Cal-158a, Cal-159a, and Cal-160; see 352 location in Morgado et al., 2019a) volcanic products each show a narrow range of ⁸⁷Sr/⁸⁶Sr 353 values (0.704309 to 0.704333, Osorno; 0.704384 to 0.704438, Calbuco), which though very close, do not overlap (Fig. 10). ¹⁴³Nd/¹⁴⁴Nd in Osorno and Calbuco volcanic products show 354 355 ranges of 0.512741 to 0.512784 (Osorno) and 0.512680 to 0.512835 (Calbuco). Two samples 356 from La Viguería (LV-152a and LV-153a), a volcanic cone overlying the LOFZ, exhibit lower ⁸⁷Sr/⁸⁶Sr (0.703705 to 0.703753), but ¹⁴³Nd/¹⁴⁴Nd (0.512763 to 0.512829) in the range 357 358 of those obtained in Osorno and Calbuco volcanic samples (Fig. 10). Non-volcanic crustal 359 rocks of the region are potential contaminants of Osorno and Calbuco products (granulites, 360 gabbros, and granitoids, Hickey-Vargas et al., 1995; Aragón et al., 2011) and most of their isotopic values differ from the ranges of ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd of Osorno and Calbuco 361 products (see Table 3). 362

363 **DISCUSSION**

364 Calculated groundmass composition

Oxygen fugacity modelling was performed using the compositions of both glass inclusions in olivine phenocrysts (Group 1 and Group 2) and the calculated anhydrous matrix composition (Os-144* composition, calculated from the composition of the sample Os-144, representing the more evolved residual melt, present immediately before eruption; see Table 8). Being low in the sequence, the sample Os-144 has experienced limited interaction with the atmosphere and most likely preserves a record of the oxygen fugacity most similar to pre371 eruptive conditions (Mollo et al., 2013). In addition, this sample has a representative whole-372 rock chemistry, typical mineral assemblages and displays textural features observed in all the 373 volcanic products analysed (see Tables 2 and 5), making Os-144 a representative effusive 374 product of the 1835 AD eruption. The fall deposits corresponding to the second eruptive 375 phase (Os-164) are excluded from this type of modelling because they are more prone to 376 rapid oxidation during the eruption than those from lava flows (Mollo et al., 2013). We 377 determined the groundmass composition (glass and microlites) by considering Os-144 (fresh 378 and with no evident signs of oxidation) as representative of the mean whole-rock composition 379 (Table 2). To obtain groundmass composition, we subtracted the composition of 380 representative phenocryst mineral phases, considering mass balance and modal abundances 381 (Table 5). From now on that composition is named as Os-144* (Table 8, details of calculation 382 are available in Supplementary Data 2). The calculated composition of Os-144* is similar to 383 those reported for several aphyric products of Osorno volcano from Unit 3 reported by Tagiri 384 et al. (1993) and Group 2 of measured glasses measured as part of this study.

385

Thermodynamic equilibrium

386 Chemical equilibrium between olivine and clinopyroxene was tested using equations from 387 Grove et al. (1997). According to these equations, we calculated a hypothetical melt 388 composition in equilbrium with each measured crystal composition. We did this by checking coefficient K_D^{Fe-Mg} , where $K_D^{Fe-Mg} = [X_{Fe}^{Mineral \, phase} \times X_{Mg}^{Melt}]/$ partition 389 their $\left[X_{Mq}^{Mineral \ phase} \times X_{Fe}^{Melt}\right]$ (Roeder and Emslie, 1970). If the phases are in equilibrium with 390 391 the same melt, then we considered the whole system, consisting of those solid phases to be 392 in equilibrium too (details in Supplementary Data 3).

393 Olivine-melt equilibrium and diffusive reequilibrium

394 The equilibrium conditions between melt inclusions and the olivine host (Fig. 11) are tested using the Toplis (2005) equation for K_D^{Fe-Mg} , considering Fe²⁺ and Fe³⁺ values in melt 395 396 following the methods of Maurel and Maurel (1982). This equilibrium filter is used prior to 397 the application of other methods that will be used to estimate intensive parameters. Most of 398 the olivine compositions (n=6) from Group 1 (Fo₇₆₋₇₉) and the two olivine crystals from 399 Group 2 (F_{069-73}) are in equilibrium with the hosted melt inclusions (Fig. 11). We only 400 consider glassy melt inclusions in olivine to ensure their compositions lie on the liquid line 401 of descent (Kent, 2008). Although inclusions with visible daughter and/or co-trapped 402 minerals exist, they are discarded from our attempt to obtain representative host melt contents (Sinton et al., 1993; Danyushevsky et al., 2002a; Kent, 2008). We also discarded melt 403 404 inclusions linked to the exterior of the host mineral. The host mineral (in this case, olivine) 405 crystallises in melt inclusions at the boundary between the melt inclusion and the host 406 mineral, which has a lower nucleation energy than homogeneous nucleation (cf. Roedder, 407 1979). According to Kent (2008), this growth on the inclusion wall may be apparent in 408 compositional profiles and also BSE images. This growth causes fractionation of the melt 409 within the inclusion, which can subsequently be modified by diffusion during magma 410 residence at high temperatures (Danyushevksy et al., 2000; 2002a). Larger melt inclusion 411 diameters are associated with longer re-equilibration timescales (Danyushevsky et al. 2000: 412 2002b). To study pre-eruptive conditions via melt inclusions, we examine the core 413 compositions of the largest melt inclusions (diameter > 50 μ m), which we interpret would 414 represent pre-eruptive composition before re-equilibration (cf. Danyushevsky et al., 2002b). 415 The only evidence of post-entrapment crystallisation is a thin rim ($< 1 \mu m$) of olivine growth.

416 Compositional traverses were measured in olivine-hosted melt inclusions and show MgO 417 variations because the growth of the olivine thin rims were balanced by MgO consumption 418 from the melt inclusion. We can be confident that the compositional profiles of olivine-hosted 419 melt inclusions represent diffusion and are not affected by convolution effects because the 420 host olivine phenocrysts have higher MgO contents than the melt inclusions. Hence, they can 421 be used to calculate timescales of re-equilibration. If secondary fluorescence is affecting the 422 measurements, then the compositional profiles within melt inclusions would show an 423 apparent increase in MgO content towards the rims. In contrast, we recognise a decrease in 424 MgO content (Fig. 12). In small melt inclusions (diameter $\leq 50 \ \mu m$) the compositional 425 profiles do not show the MgO content depletion towards the rim that we observed in > 50426 µm diameter melt inclusions. Compositional changes in melt inclusions from core to rim may 427 record the evolution of the melt close to the melt/mineral interface, influenced partially by a 428 high growth rate (Watson and Müller, 2009) and short timescales, hindering compositional 429 re-equilibrium. The calculated timescales of re-equilibration of six compositional traverses 430 for an unrealistically low temperature of 850 °C are only a few minutes (Fig. 12; Table 9) 431 and for more credible magmatic temperature of 1060 °C, are few seconds (Table 9). As such, 432 the compositions of the rims of olivine-hosted melt inclusions were modified only during 433 syn-eruptive process and the plateaus of uniform melt inclusion composition represent pre-434 eruptive composition. Then the core compositions of olivine-hosted melt inclusions would 435 not be significantly affected by the olivine growth in the olivine-melt interface. According to 436 the experiments of Newcombe et al. (2014), concentration profiles could affect the centres 437 of large melt inclusions in up to 100 minutes.

438 Using finite element methods, we modelled the variation of Mg concentration (C_{MgO} , in wt. 439 %) resulting from Mg diffusion in the olivine-hosted melt inclusions, of the radial component 440 of the time-dependent diffusion equation in spherical polar coordinates (Crank, 1975), 441 following the equation 1:

442
$$\left(\frac{\partial C_{Mg0}}{\partial t}\right) = D_{Mg0} \left(\frac{\partial^2 C_{Mg0}}{\partial r^2} + \frac{2}{r} \cdot \frac{\partial C_{Mg0}}{\partial r}\right) \qquad equation 1$$

443 where r is radius (in m), D_{MgO} is diffusivity (in m²/s), and t is time (in s). We calculated the 444 diffusivity of MgO in melt inclusions (D_{MgO}) following the equation of Chen and Zhang 445 (2008):

446
$$\ln(D_{MgO}) = -7.895 - \frac{26\,257}{T}$$
 equation 2

447 where T is in K. The modelled timescales, obtained for different input temperatures, are 448 available in Table 9. Details of all olivine-hosted melt inclusions can be found in 449 Supplementary Data 2 and example videos of modelled diffusion of MgO within melt 450 inclusions (and the corresponding MATLAB® codes) of the previously mentioned Osorno 451 samples are available in Supplementary Data 4. The uncertainties related to the Chen and 452 Zhang (2008) equation and glass chemistry measurement do not change significantly the 453 calculations performed. Therefore, the compositions of the core melt inclusions fit for 454 calculations of intensive variables.

455 Physical intensive variables

456 The propagations of uncertainties for the calculation of all the intensive conditions457 (temperature, pressure, oxygen fugacity, and dissolved magmatic water content) were carried

458 out assuming a Gaussian distribution of uncertainties and considering both internal
459 uncertainties (associated with the experimental method) and external uncertainties
460 (associated with precision and accuracy of electron microprobe analyses) uncertainties. A
461 summary of the intensive variables is in Table 10.

462 *Olivine-augite thermometry*

463 We used the olivine-augite Fe-Mg exchange geothermometer of Loucks (1996) in the cores

464 of olivine-augite equilibrated pairs, which yielded temperatures from 1100 to 1140 $^{\circ}C$ (± 15

465 °C), with an average of 1114 °C ($\sigma = 16$ °C) for olivine-augite pairs involving Group 1 olivine

466 grains (Fo₇₆₋₇₉) and similar temperatures from 1091 to 1166 °C (\pm 15 °C), with an average of

468 determined the thermodynamic equilibrium state between olivine and augite (n=30) via the

1125 °C ($\sigma = 24$ °C) for olivine-augite pairs involving Group 2 olivine grains (Fo₆₉₋₇₃). We

469 Loucks (1996) parametrisation (see section Supplementary Data 3 for details).

470 *Ca-in-olivine thermometry*

467

We used the Ca-in-olivine thermometer of Shejwalkar and Coogan (2013) in olivine cores (Fo₇₀₋₇₆), which yielded temperatures from 1094 to 1097 °C (\pm 22 °C), with an average of 1096 °C (σ = 2 °C) for Group 1 olivine phenocrysts (Fo₇₆₋₇₉) and temperatures from 1049 to 1173 °C (\pm 22 °C), with an average of 1129 °C (σ = 40 °C) for Group 2 olivine phenocrysts (Fo₆₉₋₇₃). We assume contemporary growth in clots of olivine together with clinopyroxene and plagioclase based on their textural relations.

477 *Olivine-spinel thermometry*

We used the olivine-spinel thermometer of Coogan et al. (2014) for spinel inclusions in olivine phenocryst cores (see Wan et al., 2008), which yielded temperatures from 1051 to 1070 °C (\pm 42 °C) with an average of 1061 °C (σ = 6 °C) for Group 1 olivine phenocrysts (Fo₇₆₋₇₉). We also obtained temperatures from 1104 to 1122 °C (\pm 42 °C) with an average of 1113 °C (σ = 13 °C) for Group 2 olivine phenocrysts (Fo₆₉₋₇₃).

483 Olivine-spinel oxybarometry

484 We used the olivine-spinel oxybarometer of Ballhaus et al. (1991) in spinels inclusions in 485 olivine phenocryst cores, which yielded an oxygen fugacity buffer of $\Delta QFM + 2.3 (\pm 0.7)$ 486 for Group 1 olivine phenocrysts (Fo₇₆₋₇₉) and Δ QFM + 2.7 (± 0.7) for Group 2 olivine 487 phenocrysts (Fo₆₉₋₇₃). We calculated the values of oxygen fugacity buffer considering the 488 temperatures obtained using the olivine-spinel thermometry (see Olivine-spinel thermometry 489 section) and a constant pressure of 2 kbar (results do not show significant variations with 490 pressure). The total uncertainty considers uncertainties of pressure, temperature, EPMA 491 measurements, and inherent errors of the method.

492 Spinel-melt oxybarometry in melt inclusions

493 For these calculations, we considered the same pairs of spinel inclusions and olivine 494 phenocrysts that we used for the determination of equilibrium olivine-spinel (following the 495 procedure of Kamenetsky et al. (2001). The equilibrium test of olivine-spinel pairs shows a 496 positive correlation of Mg# of spinel inclusions and the hosting forsterite olivine phenocryst 497 (details in Supplementary Data 3). Melt inclusion-olivine host equilibrium is determined as 498 explained in the Thermodynamic Equilibrium section. When the olivine host phenocryst is 499 in equilibrium with both the hosted melt inclusion and the hosted spinel inclusion, we use the Maurel and Maurel (1982) method to determine the Fe²⁺/Fe³⁺ ratio of olivine-hosted melt 500

501 inclusions considering the Fe^{2+}/Fe^{3+} of the olivine-hosted spinel composition (see Mineral 502 Chemistry section), which we assume equilibrated by transitivity. Finally, we applied the 503 Borisov et al. (2018) equation to calculate oxygen fugacity values recorded in olivine-hosted 504 melt inclusions:

505
$$log\left(\frac{FeO_{1.5}}{FeO}\right) = k \cdot log(f_{O_2}) + \frac{h}{T} + c + \sum d_i \cdot X_i + d_{SiAl} \cdot X_{SiO_2} \cdot X_{Al_2O_3} + d_{SiMg} \cdot X_{SiO_2}$$

506

$$\cdot X_{MqO}$$
 equation 3

where T is temperature (in K), X_i is the anhydrous mole fraction of melt components, and c, h, k, d_i , and d_{Xi} are regression coefficients (constant values as supplied in Supplementary Data 5).

510 The calculations yield an oxygen fugacity buffer range of $\Delta QFM + 1.1 (\pm 0.5)$ for Group 1 511 olivine phenocrysts and ΔQFM +1.4 (± 1) for Group 2 olivine phenocrysts. We performed 512 these calculations considering a temperature of 1060 °C and 1113 °C, respectively (± 42 °C; 513 see Olivine-spinel thermometry). The uncertainty of temperature depends on the uncertainty of both the Fe²⁺/Fe³⁺ ratio in olivine-hosted spinel inclusions and the method of Borisov et 514 515 al. (2018; Equation 3). These oxygen fugacity values are also consistent with those yielded 516 by the method of Jayasuriya et al. (2004): $\Delta QFM + 0.7 (\pm 0.3)$ and $\Delta QFM + 1.4 (\pm 0.3)$ for 517 Group 1 and Group 2 of olivine phenocrysts, respectively.

518 Water dissolved in melt inclusions

519 We calculated the water dissolved in magma in the olivine-hosted melt inclusions 520 (representing an early stage of magmatic evolution) using the Ca-in-olivine hygrometer of 521 Gavrilenko et al. (2016). The calculations yielded water contents from 3.4 to 4.8 wt. % (± 522 0.8 wt. %), with an average of 4.3 wt. % ($\sigma = 0.5$ wt. %) for Group 1 olivine phenocrysts and 523 from 3.4 to 5 wt. % (\pm 0.8 wt. %) H₂O, with an average of 4.2 wt. % (σ = 1.1 wt. %) for 524 Group 2 olivine phenocrysts. We applied the Gavrilenko et al. (2016) method only in those 525 cases when the melt inclusions are in equilibrium with the host olivine phenocryst (see Fig. 526 12). These results are consistent with those obtained using the Kelley and Cottrell (2009) 527 method for our olivine-hosted melt inclusions to determine water solubility in primitive 528 melts. Although Ca is sensitive to changes in olivine growth (e.g., Simkin and Smith, 1970; 529 Shejwalkar and Coogan, 2013), this effect is negligible for the studied olivine-melt inclusion 530 system because the timescales in which post-entrapment crystallisation could affect the Ca 531 in olivine composition would be relatively long, in contrast with the very short timescales 532 calculated for post-entrapment crystallisation in our samples. In consequence, the Ca in 533 olivine is not significantly affected by dissolution, reprecipitation and other processes. Our 534 calculated ranges of water dissolved in magmas are consistent with other values measured 535 for olivine-hosted melt inclusions from Llaima Volcano, in the SVZ (up to 4.4 wt.%; Bouvet 536 de Maisonneuve et al., 2012; Ruth et al., 2016). Also, our calculated ranges and averages of 537 water dissolved in magmas are consistent with those measured and reported for olivine-538 hosted melt inclusions from other arc geological settings, such as Marianas (e.g., Shaw et al., 539 2008; Kelley et al., 2010; Brounce et al., 2014), Central America (e.g., Roggensack et al., 540 1997; Sadofsky et al., 2008; Portnyagin et al., 2014), and Kamchatka (e.g., Auer et al., 2009; 541 Tobelko et al., 2019).

542 We also used the Putirka (2005; 2008) and Zhang et al. (2007) equations to determine H_2O 543 wt. % dissolved in melts and pressure iteratively, considering the plagioclase phenocrysts 544 composition (An₇₅, Zone 2) and the corresponding plagioclase-hosted melt inclusion 545 compositions. The calculations yielded water dissolved concentration between 2.1 and 4.2 546 wt. % and pressures up to 2.6 kbar (details in Supplementary Data 6). Those values are 547 consistent with our other water dissolved calculations. In addition, the "accurate volatiles by 548 difference" method of Hughes et al. (2019) yielded water dissolved contents from 0 to 3 wt. 549 % in olivine-hosted melt inclusions. Due to the slow diffusivity of Ca in olivine (Coogan et 550 al., 2005) compared to the high diffusivity of volatiles in melts (e.g., Zhang and Ni, 2010; 551 Zhang et al., 2010; Zhang and Gan, 2022), we infer that the method of Gavrilenko et al. 552 (2016) would yield water content representing an initial stage after melt inclusion 553 entrapment. Although CaO is very sensitive to olivine growth rate, the relatively constant 554 core CaO concentration in olivine phenocrysts suggest they did not change significantly in 555 time. In contrast, the procedure of Hughes et al. (2019) would likely give dissolved water 556 content in melts after diffusive hydrogen (and other volatiles) loss, which occurs rapidly in 557 response to changes in intensive conditions (e.g., Hauri, 2002; Chen et al., 2011; 2013; 558 Gaetani et al., 2012).

559 *Fe-Ti oxides thermometry and oxybarometry in xenoliths*

We used the Ghiorso and Evans (2008) thermometer and oxybarometer in ilmenitetitanomagnetite grain pairs from the volcanic xenoliths within the 1835 AD Osorno volcano products. The cores yielded temperatures from 751 to 841 °C (\pm 15 °C), with an average of 782 °C (σ = 33 °C) and oxygen fugacity buffer from Δ QFM-0.1 to Δ QFM+0.7 (\pm 0.1), with an average of Δ QFM+0.2 (σ = 0.3). The rims yielded higher temperatures from 751 to 983 °C (\pm 15 °C), with an average of 843 °C (σ = 90 °C), and an oxygen fugacity buffer ranging from Δ QFM-0.4 to Δ QFM+1.5 (\pm 0.1), with an average of Δ QFM+0.4 (σ = 0.8).

567 Thermodynamic modelling

568 We performed thermodynamic modelling using the freeware Petrolog3 (Danyushevsky and 569 Plechov, 2011) to obtain physical magmatic intensive conditions of two melt compositions: 570 olivine-hosted melt inclusions (both representing an early stage of magmatic evolution) in 571 Group 1 olivine phenocrysts (Fo_{76-79}) and Group 2 olivine phenocrysts (Fo_{69-73}). Via 572 Petrolog3, we determined the equilibrium between melt inclusions and their host olivine 573 phenocrysts by crystallisation modelling (for details see Supplementary Data 7). We 574 modelled the equilibrium at different oxygen fugacity buffer (from $\Delta QFM+0.4$ to 575 $\Delta QFM+2.4$) and pressure conditions for water-saturated systems. Our objective was to 576 construct a model that was consistent with the calculated water content dissolved in the melt 577 inclusions. Melt inclusions hosted by Group 1 olivine phenocrysts (Fo₇₆₋₇₉) reached 578 equilibrium with up to 1 wt. % crystallisation (see Supplementary Data 7) at pressures of up 579 to 2.5 kbar (equivalent to 7 km depth). Melt inclusions hosted by Group 2 olivine phenocrysts 580 (Fo₆₉₋₇₃) reached equilibrium with up to 4 wt. % crystallisation (see Supplementary Data 7) 581 at the same pressure conditions as inclusions hosted in Group 1 olivine phenocrysts (up to 582 2.5 kbar). Consequently, the melt inclusions were close to equilibrium with the host mineral. 583 These results confirm the validity of the chemical filter to determine the melt-olivine 584 equilibrium and suggest that the early stage of magmatic evolution occurred in the upper 585 crust, at depths up to 7 km depth (i.e., at pressures up to 2.5 kbar). These pressure conditions 586 yielded by thermodynamic modelling are consistent with those calculated, considering 587 measured dissolved water concentrations, in olivine-hosted melt inclusions from arc volcanic 588 products (e.g., Shaw et al., 2008; Auer et al., 2009; Werner et al., 2020; Rasmussen et al.,

589 2022) and those yielded by thermodynamic modelling for Osorno products (Rhyolite-590 MELTS; Bechon et al., 2022).

591 A shallow reservoir beneath Osorno volcano

592 We calculated the pre-eruptive temperature using several complementary methods. 593 Calculations via Fe-Mg exchange geothermometer (Loucks, 1996) yielded temperatures of 594 1114 °C and 1125 °C (± 15 °C) for olivine phenocrysts of composition in the range of Group 595 1 (F076-79) and Group 2 (F069-73), respectively. Ca-in-olivine thermometry (Shejwalkar and 596 Coogan, 2013) yielded slightly different temperatures of 1096 and 1129 (± 22 °C) for Group 597 1 and Group 2 olivine phenocrysts, respectively. Olivine-spinel thermometry (Coogan et al., 598 2014) yielded lower temperatures of 1061 °C and 1113 °C (± 42 °C) for Group 1 (Fo₇₆₋₇₉) and 599 Group 2 (Fo₆₉₋₇₃) olivine phenocrysts, respectively. Experiments by Müller et al. (2013) 600 suggest diffusion in clinopyroxene is limited. This suggests that olivine-augite thermometer 601 records temperatures that would be representative of the first stage of early magmatic 602 evolution for systems with up to 2 wt. % of water dissolved in melt (Loucks, 1996), which is 603 not the case for Osorno volcanic products. However, experiments performed in andesitic and 604 basaltic rocks from Santorini volcano (Andújar et al. 2015; 2016) yielded pressure (2-4 kbar), 605 temperature (1000-1040 °C) and dissolved water content in the melt (3-5 wt.%), which are 606 consistent with those we obtained from samples of the 1835 AD Osorno eruption.

According to Coogan et al. (2014), it seems likely that most of the Al in olivine phenocrysts
is due to reaction with the spinel Al component, and therefore olivine-spinel thermometry
would preserve a lower temperature more representative of the coexistence of olivine and
spinel rather than a liquidus temperature.

611 The oxygen fugacity of the early stage of evolution of the magmatic systems is represented 612 by the compositions of olivine-hosted melt inclusions. The melt inclusion oxybarometry 613 yielded an oxygen fugacity buffer from $\Delta QFM+1.1$ to $\Delta QFM+1.4$, suggesting that the 614 oxygen fugacity buffer remained relatively constant during the crystallisation of the olivine 615 phenocrysts from Group 1 and Group 2 of 1835 AD Osorno samples. The absence of 616 orthopyroxene in our samples and the calibration for spinel-bearing mantelic rocks of the 617 Ballhaus et al. (1991) olivine-orthopyroxene-spinel oxybarometer lead us to favour the 618 oxygen fugacity buffer values yielded by other methods.

619 In the cognate xenoliths, the core of the Fe-Ti oxide grains yielded a temperature range of 620 751 to 841 °C and oxygen fugacity conditions up to $\Delta QFM+1.5$, which correspond to 621 subsolidus conditions, according to Rhyolite-MELTS modelling (Ghiorso and Sack, 1995; 622 Gualda et al., 2012). This suggests that Fe-Ti oxide crystals recorded a compositional 623 equilibrium after a quenching event, which does not represent pre-eruptive magmatic 624 conditions. This re-equilibrium at low temperature occurs because elements in Fe-Ti oxide 625 grains show much higher interdiffusivity than in silicate phases (Aragon et al. 1984; 626 Aggarwal and Dieckmann 2002a, b). Among the six measured traverses in ilmenite-627 titanomagnetite pairs, we recognised a heating event in only three ilmenite-titanomagnetite 628 pairs (details in Supplementary Data 8). This heating recorded in the grain boundaries 629 suggests the xenoliths were incorporated during the magma ascent or shortly before it. We 630 could not determine the timescales from heating to eruption with a spacing of $\sim 2 \,\mu m$ between 631 single measurements by EPMA.

632 Petrolog3 modelling of olivine-hosted melt inclusion compositions suggests that the early633 stage of magmatic evolution occurred in the upper crust, which is consistent with the shallow

634 magmatic reservoir beneath Osorno volcano inferred by Díaz et al. (2020) at 4-8 km depth 635 and 2 km to the E of the summit of Osorno volcano. This is also corroborated by 636 SERNAGEOMIN–OVDAS (2017b; 2018a, b; 2019), who reported recent volcano-tectonic 637 seismic activity, in which the associated sources correspond to 2.6-3.9 km depth. In 638 conclusion, all approaches that estimate magmatic pressure conditions suggest the existence 639 of a magmatic reservoir in the upper crust, which seems to be a common feature in the 640 plumbing systems of the SVZ of the Chilean Andes (e.g., Calbuco volcano, Nikkhoo et al., 641 2016; Delgado et al., 2017; Morgado et al., 2019a; Arzilli et al., 2019; Llaima volcano, 642 Bouvet de Maisonneuve et al., 2012; Villarrica volcano, Lohmar et al., 2012, Morgado et al., 643 2015; Quetrupillán, Brahm et al., 2018). Moreover, Díaz et al. (2020) also reported (based 644 on magnetotellurics) the presence of a deep (7-18 km depth) magnetic reservoir in the 645 Osorno volcano system.

646 Several prior studies have suggested that relatively reducing magmatic conditions produce 647 tholeiitic magmas, whereas oxidizing magmatic conditions produce calc-alkaline magmas 648 (e.g., Druitt et al., 1999; Andújar et al., 2015; 2016). However, Tatsumi et al. (2008) reported 649 that calc-alkaline and tholeiitic magmas could coexist in temporal and spatial proximity and 650 inferred the existence of both types of basaltic magmas beneath Zoa volcano (Japan). 651 Moreover, Elburg (2010) show that oxybarometers can yield lower oxygen fugacity 652 conditions for calc-alkaline compared to tholeiitic magmas, suggesting fO_2 is not the only 653 variable controlling the differentiation trend. Tholeiitic trends in volcanic products are 654 usually associated with relatively lower concentrations of water dissolved in melts (< 4 wt. 655 %) than calc-alkaline, which would play a role in the tholeiitic trend (e.g., Zimmer et al., 656 2010; Parman et al., 2011; Hamada et al., 2014). However, recent experiments (Firth et al., 657 2019) have shown that other reasons also could be invoked to explain both calc-alkaline and 658 tholeiitic trends in volcanic products: contrasting conditions for magmatic fractionation or 659 differences in the parental magma. We suggest the tholeiitic trend recognised in whole-rock 660 and melt inclusion compositions (considering an enrichment of total Fe from melt inclusions 661 to groundmass composition) from the 1835 AD Osorno eruption could initially evolve in the 662 lower crust, in a deep reservoir beneath Osorno volcano (7-15 km depth; Díaz et al., 2021), 663 generating a tholeiitic trend (as in Santorini volcanic products; Andújar et al., 2016). This 664 magma did not subsequently deviate from the tholeiitic trend during subsequent evolution in 665 the shallow magmatic reservoir (Bechon et al., 2022) because the composition of the melt is 666 not controlled solely by fO_2 , but also by dissolved water content, fractionation, and mixing 667 (Elburg, 2010).

668 Volatile exsolution between early and late evolution stages

669 We determined the water dissolved in magmas considering olivine-hosted melt inclusions 670 composition representing an early stage of magmatic evolution, which yielded values of up 671 to 5.6 wt. % (following the method of Kelley and Cottrell, 2009) and ~4.2 wt. % average 672 (following the method of Gavrilenko et al., 2016). Other methods to determine water 673 dissolved concentration are related to higher uncertainties: the iterative use of Putirka (2005; 674 2008) and Zhang et al. (2007) equations (which yielded concentrations between 2.1 and 4.2 675 wt. %) and the Hughes et al. (2019) method (up to 3 wt. % concentration), suggest volatile 676 exsolution after the early stage of magmatic evolution before the 1835 AD Osorno eruption. 677 This likely volatile exsolution process could be associated with an accumulation of gas 678 phases in the magma reservoir before the eruption (presumably at the top of the system), the 679 release

680

release of which may have been described by Darwin (1839a) as the "smoke" emitted around two months before the onset of the 1835 AD eruption of Osorno.

681 Crystal-mush within the magma reservoir and crystallisation events

682 The existence of a crystal mush in the magmatic reservoir can be inferred using several lines 683 of evidence: (i) glomerocrysts, (ii) interstitial glass between grains of crystal clots with 684 different composition from the outer glass (Fig. 6, Table 7), (iii) disequilibrium textures 685 (zoning and resorption) in isolated crystals as well as crystals from clots, and (iv) high 686 degrees of crystallinity (25 to 45 vol. %, Table 5). Based on textural features and the narrow 687 temperature crystallisation range (considering overlapping uncertainties), we infer that the 688 magma was multiply-saturated with the minerals observed throughout the crystallisation 689 sequence. However, the zoning patterns preserved in plagioclase phenocrysts suggest a 690 complex crystallisation history for that phase. The resorption observed in several plagioclase 691 phenocrysts, affecting plagioclase Zones 1 and 2, suggests the occurrence of plagioclase 692 destabilisation after initial crystallisation events. We interpret this destabilisation as relating 693 to processes that caused mush disaggregation. Several authors have given different 694 interpretations of resorption and sieve textures – typical features of disequilibrium between 695 plagioclase and melt – such as fluid-absent rapid decompression (Nelson and Montana, 1992; 696 Annen et al., 2006) or magma mixing (thermal and compositional mixing) usually related to 697 heating (Johannes et al., 1994; Ginibre and Wörner, 2007; Pietranik et al., 2006; Lohmar et 698 al., 2012). Additionally, the existence of patchy zoning with the same composition as 699 plagioclase phenocryst rims and microphenocrysts, equivalent to Zone 3 composition, is 700 consistent with a late pre-eruptive crystallisation episode, after mush disaggregation (Fig. 701 13). Tagiri et al. (1993) inferred that the porphyritic products from Osorno volcano, similar

to the volcanic products studied in this work, resulted from the addition of calcic plagioclase
to the aphyric products. The most reasonable explanation for this is melting of ~ 20 wt. % of
XAn₇₀₋₈₀, similar to Zone 2 plagioclase compositions. Compositions of aphyric materials
(Tagiri et al., 1993) are similar to Os-144* (groundmass composition), suggesting that
aphyric products from Units 1 and 3 from Osorno volcano also could be explained as

707 interstitial melt extracted from a crystal-mush magma reservoir.

708 Group 1 and Group 2 olivine phenocrysts exhibit fairly constant (unzoned) compositions 709 with only syn-eruptive zoning patterns in the crystal rims. This absence of pre-eruptive 710 zoning patterns suggests that there were either at least two crystal mushes or a zoned crystal 711 mush, where Group 1 and Group 2 olivine phenocrysts grew independently. For this, we 712 conclude that there could not have been any protracted interaction between Groups 1 and 2 713 olivine phenocrysts with a common melt. Moreover, the difference in compositions of 714 pyroxene coupled with Group 1 and those coupled with Group 2 in crystal clots, also suggests 715 the olivine phenocrysts Groups 1 and 2 grew independently of each other. Recent studies 716 (e.g., Bergantz et al., 2015; 2017; Schleicher et al., 2016; Morgado et al., 2019b) have shown 717 that crystal mushes within a magma reservoir interact with a magma intrusion in a "mixing 718 bowl", above the location of the intrusion itself. This "mixing bowl" corresponds to the lower 719 parts of the crystal mush where thermo-mechanical effects are at a maximum, whereas those 720 places located relatively far from it could form and maintain the distinct identities of the 721 unzoned Group 1 and Group 2 olivine crystals observed in this study of a magma reservoir 722 beneath Osorno volcano (e.g., Morgado et al., 2019b; Cheng et al., 2020). Although we 723 favour the "zoned crystal mush" alternative, Díaz et al. (2020) inferred at least two magma 724 reservoirs below Osorno volcano; one at upper crustal depths (between 4 and 8 km depth)

725 and a second at lower in the crust (between 7 and 15 km). The existence of those reservoirs 726 could explain two crystal mushes that evolved separately and interacted just before the 727 eruption, rather than one single "zoned crystal mush". The Cr-Al-Fe³⁺ composition of spinel inclusion in olivine phenocryst follow the oxidation trend in cumulus: increasing Fe^{3+} at 728 729 constant Cr composition (e.g., Simakin et al., 2021; details in Supplementary Data 3). The 730 composition of those spinels would suggest that the interstitial melt is reduced during olivine 731 oxidation (e.g., Knafelc et al., 2019; Simakin et al., 2021). Moreover, embayments in some 732 olivine phenocrysts suggest a disequilibrium event (presumably a heating event), which 733 could be related to crystal mush disaggregation with interstitial melt close to water saturation 734 (Alidibirov and Dingwell, 1996) prior the 1835 AD Osorno eruption, but it did not generate 735 zoning patterns at the upper levels of the zoned crystal mush. The very thin rims of olivine 736 phenocrysts would represent a syn-eruptive quenching overgrowth (e.g., Pankhurst et al., 737 2018; Couperthwaite et al., 2020).

738 Regional geochemical heterogeneity: Osorno in context

In volcanic rocks of relatively primitive composition, whole-rock chemistry can give insights into other pre-eruptive processes, such as mantle source and crustal contamination, which complement the information yielded by mineral chemistry and allow comparison with different nearby volcanic systems. The composition of the 1835 AD Osorno products shows similar major element chemistry and trace-element patterns to those reported by Tagiri et al. (1993), Moreno et al. (2010), and Bechon et al. (2022) for Units 4 from the Osorno volcano (Figure 14). Therefore, we consider the volcanic products of the 1835 AD Osorno eruption

746 as representatives of Holocene Osorno volcanic products in general.

747 Volcanic products from Calbuco, Osorno, and La Picada - a stratovolcano on the same NE-748 SW lineament as Osorno – possess higher whole-rock fluid-mobile to fluid-immobile trace 749 elemental ratios (e.g., Ba/La vs Rb/Nd, Fig. 14) than La Viguería, a nearby small eruptive 750 centre built over the main trace of the LOFZ. We propose that the higher fluid-mobile to 751 fluid-immobile element ratios can be linked with a higher slab-derived fluid component in 752 the mantle source beneath Calbuco, Osorno and La Picada than La Viguería. Moreover, 753 Calbuco, Osorno, and La Picada volcanic products also have lower La/Yb ratios than La 754 Viguería volcanic material, with the lower La/Yb ratios being associated with either a higher 755 degree of partial melting (Fig. 14c), or melting of a previously melt-depleted source. 756 Although La Viguería samples are the most primitive of the studied region (lowest SiO_2 wt. 757 % and highest Mg#, and MgO wt. %; see Fig. 8 and Fig. 14), they also contain the highest 758 K₂O concentrations. Higher K₂O concentrations and a decrease in H₂O content in the 759 volcanic products from La Viguería could be derived from slab sediment melting, resulting 760 from a temperature increase of the slab surface (Hermann and Spandler, 2008). Temperature 761 increases at higher slab depth (Molnar and England, 1995) yielding water-poor sub-arc 762 mantle melts, lower H₂O supply to the mantle wedge and, in consequence, lower degrees of 763 partial melting recorded in volcanic products eastward from the Peru-Chile trench. This is 764 consistent with the conclusions of McGee et al. (2017) and Hickey-Vargas et al. (1989) for 765 the volcanic material from Caburgua-Huelemolle small eruptive centres, located eastward 766 from the nearby Villarrica volcano (39.3°S, Southern Andes), which is, like La Viguería, 767 built over the main trace of the LOFZ. Watt et al. (2013) also came to similar conclusions in 768 their study of the nearby Apagado and Hornopirén volcanoes at the latitude of 42°S.

A comparison of the REE systematics of the volcanic products from Calbuco, Osorno, La
Picada, and La Viguería (Fig. 9) suggests all these volcanic systems have a similar mantle
source. La Viguería cone samples (particularly the sample LV-152) show high Ni, Cr, and
low Rb values (Fig. 8 and 14), also suggesting their products are the most primitive of the
region.

774 Sources of volcanic products composition at latitude ~41°06`S

⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd isotope ratios can be used to identify the possible involvement of 775 776 subduction-related fluids and crustal contamination in the generation of the volcanic products 777 from Osorno, Calbuco, and La Viguería (we do not possess isotopic data from the other 778 nearby volcanic centres). We used the compositions of several representative end-member 779 rocks, whose chemistry could influence the Calbuco and Osorno whole-rock chemistry, 780 should the original magma interact with them in any significant way. The mantle source for magmas supplying these volcanoes lies close to a mixing line in ⁸⁷Sr/⁸⁶Sr vs. ¹⁴³Nd/¹⁴⁴Nd 781 782 space between a representative depleted MORB mantle (DMM; Workman and Hart, 2005) 783 and a representative enriched mantle end member EM-I (enriched mantle, Zindler and Hart, 784 1986; Garapić et al., 2015; Fig. 15). The most primitive samples from La Viguería plot on 785 this mixing line, whereas those from Osorno and Calbuco have, by comparison, elevated ⁸⁷Sr/⁸⁶Sr for a given ¹⁴³Nd/¹⁴⁴Nd (Fig. 15). This primitive composition of La Viguería 786 787 volcanic products could be related to the LOFZ, which would facilitate a more rapid 788 magmatic ascent, minimising the interaction with the crust (e.g., López-Escobar et al., 1995a; 789 Cembrano and Lara, 2009; Segovia et al., 2021). Figure 15 also shows a mixing line between 790 altered oceanic crust (AOC; Jacques et al., 2014, via Arc Basalt Simulator; Kimura et al., 791 2009) and CSVZ trench sediments (Kilian and Behrmann, 2003); both are likely potential

contaminants during the production of magma in this region. Possible sources of crustal
contamination are also shown on Figure 15 (details on Table 3). These include lower crustal
granulites, gabbros (Hickey-Vargas et al., 1995), and upper crustal granodiorites (Aragón et
al., 2011), all of which have been identified in the basement of this region.

796 In terms of potential fluid inputs, Calbuco, Osorno, and La Viguería volcanic products have lower ⁸⁷Sr/⁸⁶Sr than AOC and sediments. Also, in terms of ¹⁴³Nd/¹⁴⁴Nd, the studied samples 797 798 of volcanic products plot in between AOC (higher ratio) and sediments (lower ratio). Using 799 a simple mixing model (Fig. 15; details in Supplementary Data 9), the composition of a fluid 800 that is likely to have influenced the composition of the magmas investigated in this study 801 likely consists of contributions from both AOC and trench sediments. Using the composition 802 of La Viguería volcanic products as the end-member best representing the mantle source, in 803 order to reach the composition of Calbuco and Osorno volcanic products, the contributions 804 of the fluid component should have sediment: AOC ratios ranging from 0.15 to 9 (details of 805 calculations are available as Supplementary Data 9). These significant variations probably 806 occur because the sediments from the trench of CSVZ (Kilian and Behrmann, 2003) can be extremely heterogeneous in terms of its ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd signature (Kilian and 807 808 Behrmann, 2003; Fig. 15). Although it is reasonable to consider some contributions from 809 sediments in the slab-derived fluid, calculations of the fractions are unlikely to be conclusive. 810 The deviation of the Calbuco and Osorno products from the simple mixing line between the 811 representatives DMM and EM1 values likely reflects the influence of these slab-derived 812 fluids. This model does not consider mantle heterogeneity (as suggested for the CSVZ, at 813 ~39°30'S latitude by McGee et al., 2017), primarily because the similarities in the non fluid 814 mobile elements systematics of these three volcanoes preclude significant heterogeneity in

the mantle source. Chemical variations of slab-derived fluid input have also been suggested for the CSVZ, at ~39°30'S latitude by Hickey-Vargas et al., 1989), but are not accounted for in this first-order model. While chemical heterogeneity has been suggested, we are unable to determine the effect, if any, that different elemental concentrations would likely have on the final isotopic composition of the slab fluids, so we are reluctant to speculate and complicate our model unnecessarily.

821 In terms of assimilation, the presence of both granulite and gabbro xenoliths (Hickey-Vargas 822 et al., 1995) in the volcanic products of the 1961 eruption of Calbuco and the presence of the 823 North-Patagonian batholith in the region (Adriasola et al., 2006; Aragón et al., 2011), infer 824 the possibility of interaction between those crustal rocks and the ascending magma. 825 Following the mixing lines described above for fluid input, we modelled granodiorite, granulite, and gabbro assimilation to assess the ¹⁴³Nd/¹⁴⁴Nd variations observed in the 826 827 Calbuco and Osorno products (details in Supplementary Data 9). In order to shift the Osorno 828 volcanic products from the mixing line between DMM and EM1, the assimilation of 40% of granulite would be necessary to reproduce the ¹⁴³Nd/¹⁴⁴Nd range observed, but this value is 829 830 unrealistic (Taylor Jr., 1980; DePaolo, 1981). In constrast, less than 10 % assimilation of gabbro and granodiorite is required to reproduce the ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd ranges 831 832 observed. For Calbuco samples, granulite, gabbro, and granodiorite mixing models cannot reproduce the ¹⁴³Nd/¹⁴⁴Nd range observed and they are not considered as significant 833 834 contaminants.

These models suggest the most important factors controlling the Sr-Nd isotopic ratios of the studied volcanic products (at ~41°S) are the nature of the mantle source and fluid input from the subduction zone rather than crustal assimilation. This is consistent with the conclusions of Jacques et al. (2014), who investigated the compositions of volcanic products from the
whole CSVZ. Similarly, the ⁸⁷Sr/⁸⁶Sr of the volcanic material interpreted as xenoliths in 1835
Osorno volcanic products from the volcanic vent would not modify considerably the ⁸⁷Sr/⁸⁶Sr
ratios because the volcanic units 1, 3 and 4 of Osorno (unit 2 data are not reported) exhibit
similar values to those we present in this study (Tagiri et al., 1993; Fig. 10).

843 CONCLUSIONS

The last eruption of Osorno volcano occurred in 1835-1836 in two eruptive stages, with similar erupted products of basaltic andesite lavas and fall deposits, with plagioclase, olivine, and clinopyroxene as phenocrysts and spinels as inclusions in olivine phenocrysts. We used the composition of mineral phases, together with olivine-hosted melt inclusions to infer the following pre-eruptive features for the Osorno 1835 AD eruption (representative of the Osorno Unit 4):

850 There is at least one upper-crustal magma reservoir beneath Osorno volcano at depths 851 up to ~7 km (equivalent to 2.6 kbar, based on thermodynamic modelling and the use of the plagioclase-melt hygrometer-barometer equations iteratively), consistent with 852 853 geophysical constraints. Calculated pre-eruptive magma temperatures derived from 854 olivine-hosted spinel inclusions range from 1060 °C to 1113 °C with oxygen fugacity 855 buffer of $\sim \Delta QFM$ +1.1. Dissolved water concentrations in the olivine-hosted melt 856 inclusions are estimated to have been up to 5.6 wt. %, with an average of ~4.2 wt. %. 857 Disequilibrium textures in combination with evidence of local crystal accumulations _ 858 make it likely that there was interaction of the erupted magmas with a pre-existing 859 crystal mush. in erupted products (Fig. 16). This material underwent disaggregation 860 due to either increased temperatures or volatile concentrations (Fig.16).

Major and trace element concentrations in volcanic products from the latitude of
Osorno volcano (41°06'S) indicate a higher fluid contribution to their mantle sources
combined with higher degrees of source melting beneath the composite volcanoes
(Calbuco, Osorno, and La Picada) than for the smaller calc-alkaline eruptive centres
(represented by La Viguería cone), which are built over the LOFZ at the same
latitude, but several km eastward of the main volcanic arc.

- Mixing models suggest that the most dominant factors controlling the composition of volcanic products at 41°06'S latitude are mantle source (depleted peridotite) and slabderived fluid (with its sediment contribution). The contribution of crustal assimilates
 (represented by granites and gabbros) are restricted to < 10 wt.%. Our results have
 implications for the prediction of eruption characteristics in future eruptions of
 Osorno and neighbouring volcanoes.
- 873 The approach we employ here, using olivine-hosted melt inclusions composition
- 874 coupled to diffusion chronometry is adequate to study pre-eruptive and could be used
- 875 for future studies, together with whole-rock chemistry and geophysical techniques to
- assess magmatic and volcanic systems holistically.

877 ACKNOWLEDGEMENTS

We acknowledge the help in the field of Marcela Vollmer. The financial support through
FONDAP project 15090013 (Centro de Excelencia en Geotermia de los Andes, CEGA) and
CONICYT PhD fellowship (72160268, EM), hosted at Leeds by DJM are acknowledged.
We thank Felix Boschetty for his contributions during the editing process. We would also
like especially thank Thomas Müller, David Pyle, Joan Andújar, Jan Lindsey, and Maxim
Gavrilenko for comments and suggestions on reading an earlier version of this manuscript

- and John Adam and Alexander Simakin for their constructive reviews. Editorial handling and
- 885 constructive comments of Gerhard Wörner are greatly appreciated.

886 Supplementary Data

887 Supplementary Data are available at Journal of Petrology online.

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Figure 1. a) Location of the different volcanic zones of the Andes. b) Location of the four
subdivisions of the Southern Volcanic Zone. c) The Central Southern Volcanic Zone and
location of stratovolcanoes and the Liquiñe-Ofqui Fault Zone (LOFZ). Location of the LOFZ
(solid and dashed lines) inferred from Cembrano et al. (1996) and Cembrano and Lara (2009).

1366 Figure 2. (a) Location of the Osorno volcano in Chile. (b) Location of Osorno, La Picada, 1367 Puntiagudo, and Calbuco volcanoes and Cordón Cenizos volcanic chain, La Viguería cone, 1368 and the Liquiñe-Ofqui Fault Zone (LOFZ, white dashed line). The alignment of Osorno, La 1369 Picada, Puntiagudo and Cordón Cenizos is represented by a yellow, dashed line. Advanced 1370 Spaceborne Thermal Emission and Reflection Radiometer (ASTER) Global Digital 1371 Elevation Model (GDEM) image was obtained via EarthExplorer, USGS 1372 (http://earthexplorer.usgs.gov).

1373 Figure 3. Osorno volcano and the deposits of the first and second stages of the 1835 eruption.

1374 Yellow stars represent locations where samples are collected from a lava flow, whereas lapilli

1375 sample locations are represented by red stars. Locations and dates of the lava flows and fall

1376 deposits are based on Moreno et al. (2010) and Lara et al. (2012). ASTER-GDEM image

1377 from EarthExplorer, USGS (http://earthexplorer.usgs.gov).

1378 Figure 4. (a) Backscatter electron (BSE) images of an isolated plagioclase phenocryst where 1379 different compositional zones are found. Brighter greyscale colours are correlated to higher 1380 density. (b) The arrow represents the profile measured by electron probe micro-analysis 1381 (EPMA) and its direction. (c) Relation and coefficient of determination (r^2) between BSE 1382 greyscale profile and the measured anorthite content ($An = 100 \times Ca/(Ca+Na+K)$; in molar 1383 proportions). "An" is a representative description of the composition of plagioclase because 1384 in all crystals the K content remains constant throughout (details in Supplementary Data 2). 1385 (d) Anorthite profile composition (An) by EPMA of the measured profile (circles) coupled 1386 with the BSE profile based on greyscale values calibrated with the composition measured by 1387 electron microprobe (solid line).

Figure 5. (a) Olivine-hosted melt inclusion of ~110 μm diameter. (b) Resorption feature
(embayment) in olivine phenocryst with glass and plagioclase microlites. The yellow dashed
line represents the olivine grain boundary.

Figure 6. (a) BSE image of a crystal clot of olivine grains. (b) Zoom of the BSE image,
which shows interstitial glass between the crystals constituting the clot. (c) EBSD map
represents the diversity of olivine crystal orientations, consistent with the observation that
the glasses are interstitial and are not melt inclusions.

Figure 7. Total alkali versus silica (Le Bas et al., 1986) plots of the Osorno 1835 volcanic
products (grey squares), the modified composition of Os-144 (Os-144*, yellow star), melt
inclusions (Group 1 of glass, red field), and glass from groundmass (Group 2 glass, green

field), La Viguería cone products (pink triangles), La Picada volcanic products (Vander
Auwera et al., 2019, blue field), and Calbuco 2015 volcanic products (Morgado et al., 2019a,
yellow squares).

1401Figure 8. (a) Variations of Ni ($\mu g \cdot g^{-1}$) versus SiO2 (wt. %), (b) K2O (wt. %) versus SiO2 (wt.1402%), (c) FeO*/MgO versus SiO2 (wt. %), (d) K2O (wt. %) versus MgO (wt. %), (e) CaO (wt.1403%) versus MgO (wt. %), and (f) Al2O3 (wt. %) versus MgO (wt. %) of Holocene erupted1404products from Calbuco volcano (Castruccio et al., 2016; Morgado et al., 2019a; Arzilli et al.,14052019; Namur et al., 2020), Osorno volcano (Tagiri et al., 1993; Moreno et al., 2010; Bechon1406et al., 2022; this study), and Cayutué-La Viguería field (López-Escobar et al., 1995a; this1407study).

- Figure 9. (a) Primitive mantle-normalised (Sun and McDonough, 1989) incompatible trace
 element diagram for Calbuco erupted products, taken from López-Escobar et al. (1995b)
 Morgado et al. (2019a). Major and trace element concentrations from Osorno and La
 Viguería are available in Table 2. (b) REE patterns of bulk rock samples from volcanic
 systems of the region (La Viguería, Osorno, Calbuco, and La Picada volcanoes).
- **Figure 10.** (a) ¹⁴⁴Nd/¹⁴³Nd versus ⁸⁷Sr/⁸⁶Sr of volcanic products of Osorno, Calbuco, sediments from the CSVZ trench, granulite xenoliths, gabbro xenoliths and from the basement, and granitoids from basement. (b) Comparison of ¹⁴⁴Nd/¹⁴³Nd versus ⁸⁷Sr/⁸⁶Sr with other volcanic zones from the SVZ (from Hickey-Vargas et al., 2016 and references therein).

1418Figure 11. Olivine core compositions $(X_{ol}(Fe^{2+}/Mg) vs X_{melt}(Fe^{2+}/Mg))$ for products of the14191835 Osorno eruption. Most of the olivine compositions from group 1 (Fo₇₆₋₇₉) are in1420equilibrium with the hosted melt inclusions, whereas the two olivine crystals from group 21421(Fo₆₉₋₇₃) are in equilibrium with the hosted melt inclusions. The Fe²⁺ in the melt is calculated1422using the olivine-hosted spinel inclusions Fe²⁺/Fe³⁺. We calculate the equilibrium lines as1423 $X_{ol}(Fe^{2+}/Mg) = K_D \times X_{melt}(Fe^{2+}/Mg)$, where K_D is calculated via the Toplis (2005) procedure.1424Dashed lines represent the uncertainty of the Toplis (2005) method.

- Figure 12. (a) BSE image of an olivine phenocryst and melt inclusions. (b) Zoom of the BSE
 image, which shows the EPMA profiles in olivine and large melt inclusion. (c) Mg#
 composition of olivine phenocryst traverse, which is flat (showing the absence of
 disequilibrium), (d) MgO (wt. %) composition profiles of the melt inclusion shown in b),
 which show depletion towards the rim. That depletion represents diffusion and the related
 timescales are a few minutes.
- Figure 13. Evolution of plagioclase phenocrysts zoning over time: first event (I) is nucleation
 and growth of plagioclase phenocryst cores (Zone 1 composition); after that (II), growth of
 plagioclase phenocryst rims and nucleation and growth of new plagioclase phenocrysts

1434 occurred (Zone 2 composition). These two first crystallization events occurred during the
1435 crystal mush formation. After the crystal mush building, another process generating
1436 plagioclase-melt disequilibrium and resorption (we suggest volatile additions or heating)
1437 occurred (III), and finally (IV), a last growth event occurred. We infer that the last growth
1438 event occurred before or during the eruption triggering.

1439 Figure 14. Plots of fluid mobile/immobile elements of Holocene volcanic samples of the 1440 region: (a) Ba/La versus Rb/La and (b) plot of Ba/Nd versus Rb/Nd. (c) plot of La/Yb versus SiO₂. (d) plot of Ni ($\mu g \cdot g^{-1}$) versus Cr ($\mu g \cdot g^{-1}$). (e) plot of Rb ($\mu g \cdot g^{-1}$) versus Mg#. In a) and 1441 1442 b) the arrows represent the direction in which the fluid component in the source is increasing. 1443 In c) the arrow represents the direction in which partial melting degree of the source is 1444 increasing. The arrows in d) and e) represent the direction in which the products are more 1445 primitive. We calculated Mg# assuming $Fe^{2+}/Fe^{3+} \approx 3$ (the same ratio we reported in this 1446 article via titration) in those samples from studies in which only Fe₂O_{3(t)} is reported (Moreno 1447 et al., 2010; Bechon et al., 2022).

Figure 15. ¹⁴³Nd/¹⁴⁴Nd vs. ⁸⁷Sr/⁸⁶Sr of whole-rock compositions of the samples from the 1448 1449 regional crustal rocks (granulite, gabbro, and granitoids; Hickey-Vargas et al., 1995; Aragón et al., 2011), Calbuco and Osorno stratovolcanoes, La Viguería small eruptive centre, and 1450 1451 samples representing mantle compositions (DMM and EM-I; Workman and Hart, 2005; 1452 Garapić et al., 2015) and mixing lines showing possible fluid composition supplies considering the ranges of ⁸⁷Sr/⁸⁶Sr ratio and ¹⁴³Nd/¹⁴⁴Nd for trench sediments in Kilian and 1453 1454 Behrmann (2003). The sediment compositional end-members are modelled in a) and b), 1455 respectively. The mixing lines are built according to the isotopic values presented in Table 1456 3.

1457 Figure 16. Schematic representation of the evolution of the shallow reservoir beneath Osorno 1458 volcano before the 1835 eruption. Figure (a) shows the initial crystal mush, which is 1459 perturbed (by heating and/or volatile addition), as represented in Figure (b). Figure (c) shows 1460 how crystals are disaggregated from the crystal mush and incorporated to the eruptible 1461 magma. Figure (d) shows when the eruption is triggered, it occurs in the main crater as well 1462 as the parasitic cones (image not to scale). The main crater and parasitic cones erupt lava 1463 flows and fall deposits. The sequence occurred for both eruptive events: January-February 1464 1835 and November 1835-January 1836. Volatile accumulation events during magma 1465 evolution and their subsequent release could have triggered the relatively explosive first 1466 event in January 1985.