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1 **What lies beneath: how can we study the inner workings of volcanoes and**
2 **their plumbing systems?**

3

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5

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9

10 **Abstract**

11 **Craig Magee** and **Chris Jackson** examine how geophysical seismic reflection data can be
12 used to image volcanoes and magma plumbing systems.

13

14 Volcanoes are awe-inspiring natural phenomena found throughout the Solar System. They
15 fascinate and enthuse children and adults alike, providing an important route into geoscience
16 for young adults. Beyond being photogenic, we study volcanoes for a variety of reasons (e.g.,
17 Martí and Ernst, 2008, Sigurdsson et al., 2015): (1) volcanoes provide important habitats for
18 life on Earth, whereas their eruption can threaten the very life they harbour; (2) volcanic
19 degassing, which may occur during major, catastrophic eruptions or during inter-eruption
20 periods of relative calm, can influence the global climate, and in the past have driven mass
21 extinctions; (3) the chemistry of igneous rocks, the products of magma (molten rock) moving
22 through the crust to be erupted at volcanoes, can provide crucial insight into plate tectonic

23 and deep Earth processes through time; and (4) the development of magma plumbing
24 systems, which modulate volcano growth and eruption, is intimately linked to the
25 accumulation of the many natural resources (e.g., minerals and metals) required to build a
26 sustainable future. We have also taken to examining volcanoes on other planets, such as
27 Mars, to see what they can tell us about planetary evolution. Given the importance of
28 volcanoes to human life and development, we have made great strides in understanding their
29 origins and formation.

30 We typically study active volcanoes by measuring their surface morphology and/or
31 examining the chemistry or mineralogy of erupted material (e.g., lavas or gas). These studies
32 help unravel how volcanoes have grown and evolved, and how magma is currently behaving
33 beneath them (e.g., Grosse et al., 2009, Grosse et al., 2014, Cashman and Sparks, 2013). By
34 coupling this information with studies of ancient volcanic centres, where erosion of the
35 volcano has allowed access to the network of pipes and conduits that deliver magma through
36 Earth's crust to the surface, we can identify tell-tale signs of magma movement (e.g.,
37 earthquakes and ground deformation); this helps us better understand their subsurface
38 structure and potential hazard (e.g., Sparks et al., 2012, Segall, 2013, Ebmeier et al., 2018,
39 Biggs et al., 2013, Biggs et al., 2011, Chouet and Matoza, 2013). Although we can go to the
40 field and study ancient magma conduits (intrusions), which typically have vertical (dykes) or
41 horizontal (sills) sheet-like geometries, we can only directly measure their intersection with
42 Earth's surface. This limitation in access to magma plumbing system structure led to the
43 paradigm that rock melts at depth towards the base of the crust, magma is transported
44 vertically upwards, and builds up in subsurface (magma) 'chambers' from which it is erupted
45 (e.g., Cashman and Sparks, 2013); this is what we call the 'textbook' scenario (Fig. 1A).
46 However, this simplistic view of vertically arranged conduits and voluminous reservoirs of
47 liquid is being replaced by a vision of transcrustal plumbing systems dominated by 'mush',

48 i.e. a crystal framework with pores containing melt, through which small volumes of magma
49 migrate and accumulate (Fig. 1B) (e.g., Cashman et al., 2017, Magee et al., 2018, Cruden and
50 Weinberg, 2018). Key to realising this new paradigm of magma plumbing systems has been
51 the integration of cutting-edge petrological and geochemical work with: (1) geophysical
52 techniques that can image zones of mush and magma within the crust (e.g., seismic
53 tomography); and (2) geodetic methods that can invert ground deformation patterns to model
54 the location, geometry, and dynamics of subsurface magma movement (see Magee et al.,
55 2018 and references therein). Here we discuss how we, and others, have used reflection
56 seismology, a technique traditionally applied to finding oil and gas, to contribute to the
57 evolving paradigm that describes how magma plumbing systems and volcanoes are
58 constructed. For example, we show how seismic reflection data, coupled with structural and
59 chemical analyses of sills exposed at Earth's surface, has helped demonstrate that volcanic
60 systems may not be vertically stacked, but instead comprise laterally extensive intrusion
61 networks (Fig. 1C) (Magee et al., 2016). We also take a brief look at two areas of reflection
62 seismology that look to advance our understanding of volcanism across the Solar System in
63 the near-future.

64

65 **Reflection seismology**

66 Reflection seismology is a technique that involves firing acoustic energy (i.e. seismic waves)
67 from a source down into the subsurface (Fig. 2). When this acoustic energy encounters
68 boundaries between materials with different physical properties, some energy is transmitted
69 onwards deeper in the Earth, whilst some is reflected back to the surface (Fig. 2). By
70 measuring the arrival times (i.e. the travel-time) and amounts of reflected energy returned to
71 the surface, we can build an image of Earth's subsurface structure. Depending on the seismic

72 reflection survey design, these images may either: (1) constitute extensive 2D cross-sections
73 capable of imaging 10's of kilometres below Earth's surface across 100's of kilometres
74 distance; or (2) form part of a dense grid of seismic lines that can be interpolated to image the
75 Earth's subsurface structure in 3D, typically over a small area than 2D techniques.
76 Traditionally, these seismic reflection techniques have been used to explore for hydrocarbons
77 in basins comprising thick accumulations of sedimentary rocks. However, igneous rocks
78 typically have very different physical properties (e.g., densities of $>2.5 \text{ g/cm}^3$ and acoustic
79 velocities of $>4000 \text{ ms}^{-1}$) compared to surrounding sedimentary rocks (e.g., densities of 1.6–3
80 g/cm^3 and acoustic velocities of $>2000\text{--}5000 \text{ ms}^{-1}$) (Fig. 2). Due to these difference in the
81 physical properties of igneous and sedimentary rocks, volcanoes and their crystallised magma
82 plumbing systems are thus commonly well-imaged in seismic reflection data (e.g., Fig. 2).

83

84 **Seismic reflection imaging of volcanoes**

85 Most volcanoes studied using seismic reflection data are ancient and have long since been
86 buried by layers of younger sedimentary rock (e.g., Fig. 2). Importantly, these encasing
87 sedimentary strata protect the volcanic edifices from erosion and thereby preserve external
88 volcano morphologies (Jackson, 2012). By interpreting the top surface of volcanoes in
89 seismic reflection data we can create a topographic map of ancient volcanic fields (e.g., Fig.
90 3A), comparable to imaging modern volcanoes using satellite or bathymetry techniques.
91 From these maps we can extract and quantify the shape of ancient volcanoes (e.g., Sun et al.,
92 2020, Magee et al., 2013b, Reynolds et al., 2018). Such quantitative information on edifice
93 morphology has been used to reconstruct how modern volcanoes may grow, by assuming
94 small volcanoes become large volcanoes (e.g., Grosse et al., 2014, Grosse et al., 2009, Rossi,
95 1996). For example, Grosse and Kervyn (2018) compile data from a range of shield

96 volcanoes and developed a complicated framework to describe the controls on how they may
97 grow and erupt. A key limitation to these studies of modern volcanoes is that interpreted
98 growth models cannot be easily validated; i.e. we cannot see the inner geometry of modern
99 and commonly still-active volcanoes to test how they grew. Seismic reflection data
100 circumvents this lack of access because, in addition to imaging their external morphology,
101 internal layers can also be imaged (e.g., Figs 2 and 3B) (e.g., Sun et al., 2020, Magee et al.,
102 2013b, Reynolds et al., 2018). Several studies have used seismic reflection data to examine
103 how the external morphology of ancient volcanoes compares to their internal architecture,
104 showing that many appear to grow via a commensurate increase in summit height and basal
105 diameter (e.g., Fig. 3B) (e.g., Sun et al., 2020, Magee et al., 2013b, Reynolds et al., 2018).
106 Overall, seismic reflection data offer an opportunity to examine the external and internal
107 morphology of volcanoes for the first time.

108

109 **Seismic reflection imaging of magma plumbing systems**

110 A major benefit of reflection seismology is that not only can we image the external
111 morphology and internal architecture of volcanoes, but we can see their underlying magma
112 plumbing system (e.g., Fig. 2). Because seismic reflection data is compiled from acoustic
113 energy returned to the surface, this technique favourably images sub-horizontal-to-
114 moderately inclined magma conduits and related rock bodies (e.g., sill, inclined sheets, and
115 laccoliths; Smallwood and Maresh, 2002, Jackson et al., 2013, Magee et al., 2016, Eide et al.,
116 2018). Over the last two decades, seismic reflection-based studies have shown that many
117 intrusive components of magma plumbing systems within sedimentary basins have a saucer-
118 shaped morphology (e.g., Thomson and Schofield, 2008, Thomson and Hutton, 2004,
119 Schmiedel et al., 2017, Magee et al., 2013a, Planke et al., 2005); these are imaginatively

120 called ‘saucer-shaped sills’ and comprise a relatively flat inner base surrounded by inwardly
121 dipping sheets (Fig. 4A) (e.g., Malthe-Sørensen et al., 2004, Polteau et al., 2008). These
122 saucer-shaped sill reflections commonly contain discrete vertical offsets that are broadly
123 linear in map-view (Fig. 4A) (e.g., Schofield et al., 2017, Schofield et al., 2012a, Magee et
124 al., 2014, Magee et al., 2013c). By studying similar features in sills exposed in the field, these
125 vertical offsets have been interpreted to represent the seismic expression of ‘steps’, ‘bridges’,
126 or ‘magma fingers’ (Magee et al., 2019c, Schofield et al., 2012b, Hutton, 2009). These steps,
127 bridges, and fingers form during magma injection (e.g., Hutton, 2009, Pollard et al., 1975,
128 Schofield et al., 2012b, Schofield et al., 2010, Galland et al., 2019) and although they form in
129 response to different processes, their long axes all sub-parallel the magma propagation
130 direction (see Magee et al., 2019c and references therein). Mapping vertical offsets in sill
131 reflections related to step, bridge, or finger formation thus allows us to reconstruct how
132 magma once flowed through the host rock (e.g., Fig. 4A) (Schofield et al., 2017, Magee et al.,
133 2014, Schofield et al., 2012a). Critically, the interpretation of sills and magma flow patterns
134 across parts of sedimentary basins has shown that sills may be interconnected (i.e. forming a
135 sill-complex) and facilitate magma transport across 10’s kilometres vertically and 10–100’s
136 kilometres laterally (e.g., Fig. 1C) (Magee et al., 2016). The recognition that sill-complexes
137 can transport magma laterally over vast distances, which has been corroborated by studies of
138 sills exposed at Earth’s surface, questions long-held assumptions that volcanic systems are
139 vertically stacked and that eruption sites broadly overlie deep zones of melt generation (Fig.
140 1) (Magee et al., 2016).

141 In addition to being able to map intrusion geometries and magma flow patterns using
142 seismic reflection data, we can also image the deformation (folding and fracturing) of the
143 rock around sills that is generated to make space for magma injection. By examining host
144 rock deformation around intrusions, we can unravel when and how magma injection occurred

145 (e.g., Trude et al., 2003, Hansen and Cartwright, 2006). Typically, space for shallow-level
146 intrusions is generated by uplift of overlying rock and free surface (e.g., Koch et al., 1981,
147 Pollard and Johnson, 1973). Analysing such intrusion-related uplift of Earth's surface is key
148 to hazard assessment at active volcanoes as it allows us track magma movement and
149 accumulation in near real-time (e.g., Sparks et al., 2012, Segall, 2013, Ebmeier et al., 2018,
150 Biggs et al., 2013, Biggs et al., 2011). In particular, surface uplift can be inverted to model
151 underlying the shape and location of underlying intrusions, but this relies on assuming that
152 space for the intrusion is solely generated by uplift (e.g., Pritchard and Simons, 2004).
153 Seismic reflection-based studies of ancient intrusions, coupled with field analyses and
154 laboratory modelling, confirm that uplift of rock above sills can spatially accommodate their
155 intrusion (i.e. uplift produces structures called 'forced folds'; e.g., Fig. 4A) (e.g., Jackson et
156 al., 2013, Hansen and Cartwright, 2006, Magee et al., 2013a). However, studies have also
157 shown that other host rock deformation processes, such as collapse of pore spaces, can also
158 generate space for intruding magma; where these processes and uplift occur, the height of
159 forced folds can be significantly less than intrusion thickness (e.g., Magee et al., 2019b,
160 Magee et al., 2013a, Jackson et al., 2013, Galland, 2012, Morgan et al., 2008). This potential
161 difference between forced fold shape and size compared to that of the underlying intrusion
162 implies inversion of surface uplift patterns may underestimate magma volumes if no other
163 processes are considered to provide space for magma injection (e.g., Magee et al., 2018).

164 To-date, a major limitation of seismic reflection data has been our relative inability to
165 seismically image buried, near-vertical sheet intrusions (dykes), which are expected to play a
166 dominant role in transporting magma through the crust to be erupted at the surface (e.g., Fig.
167 1). This problem occurs because near-vertical dykes reflect only a limited amount of acoustic
168 energy back to the surface and are thus typically poorly imaged in seismic reflection data
169 (Smallwood and Maresh, 2002, Planke et al., 2005, Thomson, 2007, Wall et al., 2010, Eide et

170 al., 2018). Where dykes have been rotated due to later deformation, they may be imaged in
171 seismic reflection data (e.g., Fig. 4B) (e.g., Phillips et al., 2018, Abdelmalak et al., 2015). For
172 example, Phillips et al. (2018) identify a swarm of dykes offshore southern Norway, which
173 were originally near-vertical but since their intrusion have been tilted (Fig. 4B). These data
174 revealed for the first time how dyke swarm geometry varies with depth, particularly showing
175 that only dykes within the centre of the swarm ascended to shallow levels (Fig. 4B) (Phillips
176 et al., 2018). A swarm of vertical dykes has just recently been discovered in seismic
177 reflection data from offshore North-west Australia, i.e. the Exmouth Dyke Swarm, where
178 individual dykes correspond to thin near-vertical zones where reflections from the flat-lying
179 sedimentary rocks are disrupted; i.e. the dykes are co-located with areas where less seismic
180 energy is returned from, such that they appear in seismic reflection data as zones of poor or
181 no imaging (Fig. 4C) (Magee and Jackson, 2020a). The Exmouth Dyke Swarm is perhaps up
182 to ~500 km long and ~300 km wide, and provides a unique opportunity to examine the 3D
183 geometry of a vertical dyke swarm (Fig. 4C) (Magee and Jackson, 2020a). In addition to the
184 dykes themselves, the seismic reflection data imaging the Exmouth Dyke Swarm also
185 captures overlying host rock deformation structures related to its emplacement (Fig. 5)
186 (Magee and Jackson, 2020a). These host rock deformation structures include (Fig. 5): (1)
187 dyke-induced faults, where rock has fractured and slipped along inclined surfaces (normal
188 faults) that extend along the length of and dip towards the dyke tops; and (2) pit craters, sub-
189 circular depressions that are underlain by a vertical pipe connected to dyke tops or dyke-
190 induced faults (Magee and Jackson, 2020a). Dyke-induced faults form because space for the
191 underlying dyke is generated by extension of the host rock, but because magma does not
192 reach the surface, the host rock extension above the dyke instead manifests as faults (e.g.,
193 Pollard et al., 1983, Rubin, 1992, Trippanera et al., 2015b). The origin of pit craters has
194 remained enigmatic (e.g., Wyrick et al., 2004), but seismic reflection data from offshore

195 North-west Australia provides the first conclusive evidence that they may extend down and
196 likely form in response to dyke intrusion (Fig. 5) (Magee and Jackson, 2020a).

197

198 **Ongoing and future developments**

199 There is a vast amount of seismic reflection data available that image volcanoes and their
200 magma plumbing systems, but relatively few scientists utilising these data. In addition to
201 progressing our understanding of the structures and processes described above, there are two
202 novel applications of seismic reflection data worth mentioning here. The first concerns how
203 we can use seismic reflection data to understand the subsurface structure of other planetary
204 bodies. For example, dyke-induced faults and pit craters have been recognised in magmatic
205 settings not only on Earth, but also on other planetary bodies such as Mars (e.g., Mastin and
206 Pollard, 1988, Rubin and Pollard, 1988, Pollard et al., 1983, Trippanera et al., 2015a, Whitten
207 and Martin, 2019, Frumkin and Naor, 2019, Okubo and Martel, 1998). Before their
208 recognition in seismic reflection data, we have typically only been able to access the surface
209 expression of these natural features and thus have had to infer their subsurface structure (e.g.,
210 Mastin and Pollard, 1988, Rubin and Pollard, 1988, Pollard et al., 1983, Trippanera et al.,
211 2015a, Whitten and Martin, 2019, Frumkin and Naor, 2019, Okubo and Martel, 1998). From
212 these inferences of subsurface structure, we have identified how the surface expression of
213 these features could be used to estimate subsurface properties; e.g., the measured distance
214 between and dip of a dyke-induced fault pair at the surface has been used to estimate depths
215 to dykes (e.g., Hjartardóttir et al., 2016, Wilson and Head, 2002, Trippanera et al., 2015b).
216 Seismic reflection data allow us, for the first time, to test these assumptions of how the
217 surface expression of natural dyke-induced faults and pit craters may reflect subsurface
218 structure (e.g., Fig. 4C and 5) (Magee and Jackson, 2020a, Magee et al., 2019a).

219 Unfortunately, preliminary analyses suggest it may be more difficult to invert the surface
220 expression of dyke-induced faults and pit craters, to recover information on subsurface
221 structure, than previously anticipated (Magee and Jackson, 2020b).

222 Perhaps the most exciting application of reflection seismology to volcanology is the
223 move to generating seismic reflection images of active volcanoes. Along the East Pacific
224 Rise, a 3D seismic reflection survey was recently acquired over the Axial Seamount volcano
225 using a traditional acquisition design (Carbotte et al., 2020, Arnulf et al., 2014). These data
226 have imaged magma and mush in a series of sills beneath the Axial Seamount, and
227 highlighted the role of faults in directing magma flow pathways (e.g., Fig. 6A) (Carbotte et
228 al., 2020, Arnulf et al., 2014). Similarly, an active source seismic experiment at Mt St Helens,
229 conducted as part of the recent iMUSH (imaging Magma Under St Helens) project, has
230 imaged crustal reflections that illuminate its magma plumbing system (e.g., Levander and
231 Kiser, 2019). In addition to traditional surveying methods, reflections from passive (natural)
232 seismicity at Krafla volcano, Iceland, have been used to image a magma reservoir proven by
233 drilling data (e.g., Fig. 6B) (Kim et al., 2020). Acquiring seismic reflection data from active
234 volcanoes provides a new frontier in understanding the internal workings of volcanic
235 systems.

236

237 **Conclusions**

238 Reconstructing the emplacement and growth of volcanoes and their plumbing systems is
239 critical to understanding eruption locations and dynamics, which feed into hazard
240 assessments, and the distribution of magma-related critical raw materials resources. Here we
241 show the seismic reflection data provides unique imaging of the internal architecture of
242 volcanoes and illuminates magma plumbing systems. Although most volcanoes and intrusion

243 imaged in seismic reflection data are ancient, and their activity long-since ceased, we show
244 how we can peel away their layers, or map magma flow structures to reconstruct their
245 emplacement. We also demonstrate how seismic reflection data can be used to interrogate
246 host rock deformation processes that occur in order to generate space for intruding magma.

247

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254 we use (<http://www.ga.gov.au/nopims>) and to Schlumberger for access to their seismic
255 interpretation software.

256 **Figure captions**

257 Figure 1: The schools of thought on the structure of magma plumbing systems. (A)
258 Volcanoes generated above vertically stacked systems of dykes connecting magma reservoirs
259 and/or a transcrustal mush zone, occur directly above areas of melt production. (B) Sill
260 complexes can transport magma laterally and vertically through the crust, such that
261 associated volcanoes may be offset from areas of melt production.

262 Figure 2: Cartoon of Earth's subsurface structure depicting how seismic waves may reflect
263 from rock-rock boundaries, and how we can use this to image sedimentary layers, volcanoes,
264 sills, and associated host rock deformation (e.g., forced folds). Seismic section from the Bight
265 Basin, offshore S Australia.

266 Figure 3: (A) Oblique view of the Top Volcano surface mapped across part of the the Bight
267 Basin; note the volcanoes are all shield volcanoes but the surface is vertically exaggerated to
268 highlight the edifices. Vertical scale is presented in seconds two-way travel-time (TWT). (B)
269 Seismic sections through a volcano shown in (A), highlighting that internal layers within the
270 volcano can be mapped.

271 Figure 4: (A) A seismic section and 3D view of a saucer-shaped sill located offshore NW
272 Australia (modified from Magee et al., 2013a). (B) Dyke swarm imaged in seismic reflection
273 data, offshore S Norway (modified from Phillips et al., 2018). Inset: cartoon showing the
274 structure of the dyke swarm and explaining how basin flexure rotated the originally vertical
275 dykes to an inclined orientation. (C) Seismic section depicting vertical dykes, and overlying
276 dyke-induced faults, from the Exmouth Dyke Swarm, offshore NW Australia; the map shows
277 the spatial distribution of the dykes and the dataset used to study them (modified from Magee
278 and Jackson, 2020a).

279 Figure 5: 3D view of the top of a pit crater (i.e. a sub-circular depression) developed above
280 and linked to a dyke by a vertical pipe-like structure, in which reflections from sedimentary
281 rocks are offset downwards. Above the dyke, two dyke-induced faults are also observed that
282 parallel the dyke trend and dip towards the dyke top (modified from Magee and Jackson,
283 2020a).

284 Figure 6: (A) Seismic sections revealing the inner structure of Axial Seamount volcano, with
285 high-amplitude (bright) reflections interpreted to represent a magma reservoir comprising a
286 series of sills containing melt (Arnulf et al., 2014). (B) Seismic sections spanning the Krafla
287 volcano, Iceland, which images high-amplitude reflections shown by borehole data to
288 coincide with zones of magma (Kim et al., 2020).

289

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Figure 1

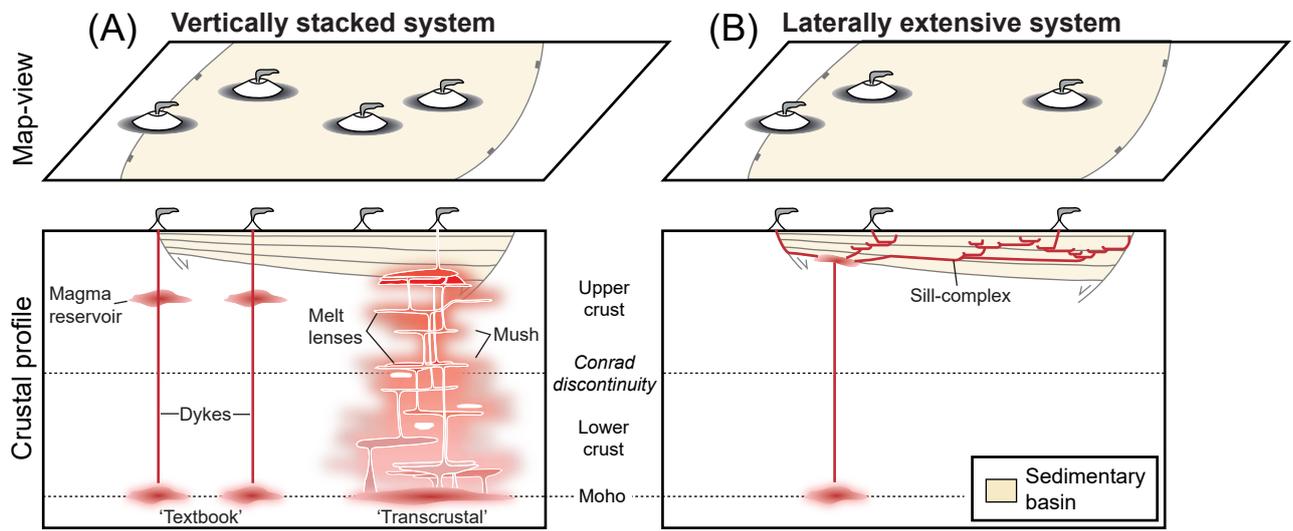


Figure 2

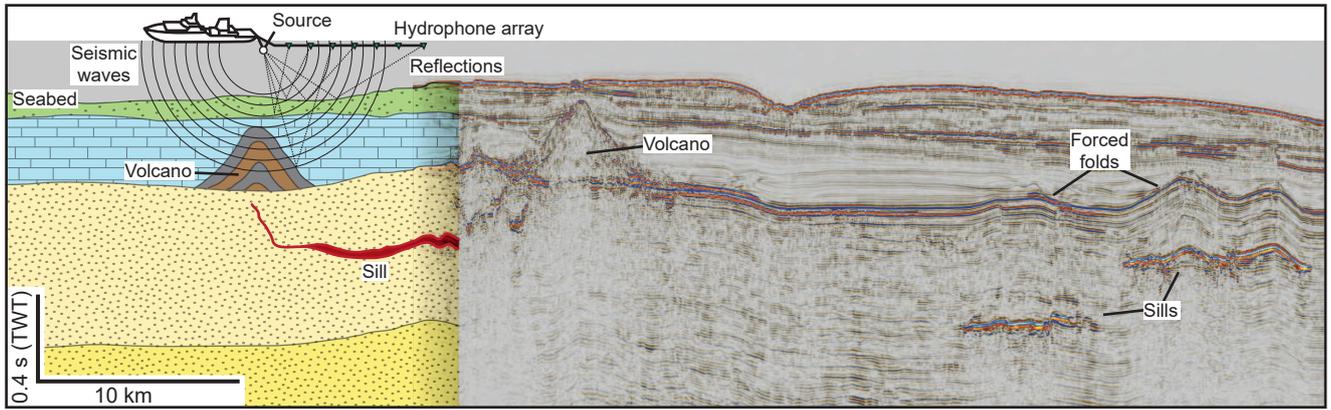


Figure 3

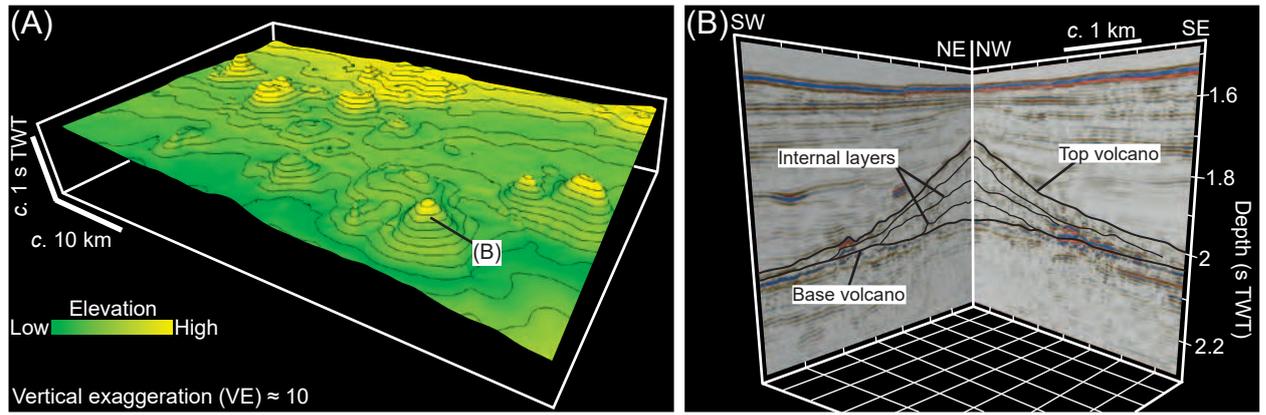


Figure 4

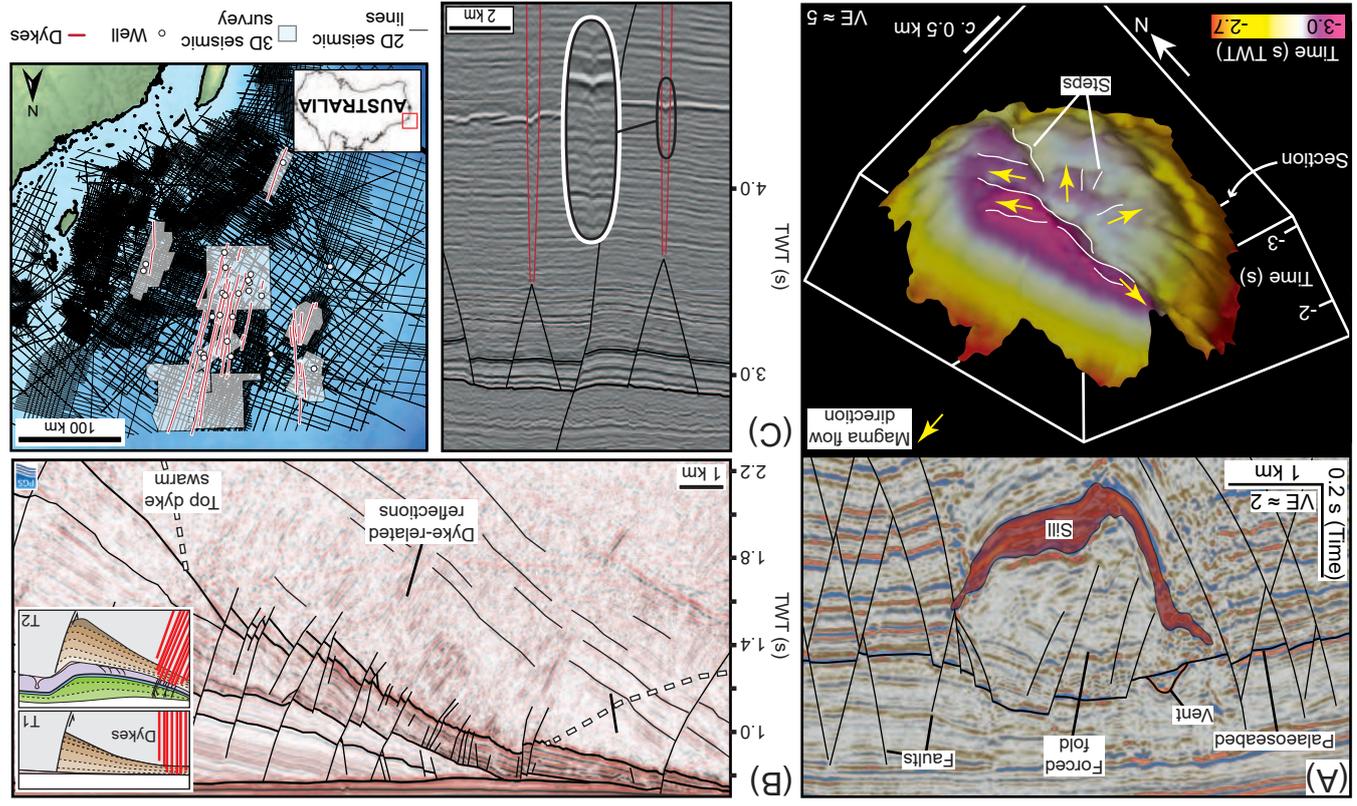


Figure 5

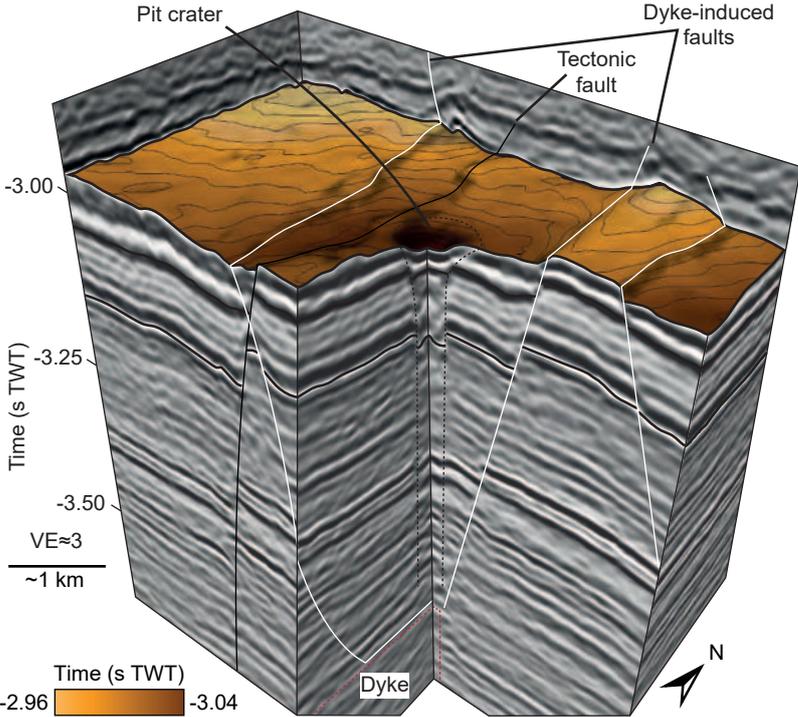


Figure 6

