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Cross machine investigation of magnetic tokamak dust; morphological and elemental analysis

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Abstract

The presence of magnetic dust can be an important issue for future fusion reactors where plasma breakdown is critical. Magnetic dust has been collected from contemporary fusion devices (FTU, Alcator C-Mod, COMPASS and DIII-D) that feature different plasma facing components. The results of morphological and elemental analysis are presented. Magnetic dust is based on steel or nickel alloys and its magnetism is generated by intense plasma-material interactions. In spite of the strong similarities in terms of morphology and composition, X-ray diffraction analysis revealed differences in the structural evolution that leads to non-trivial magnetic responses.

1. Introduction

The presence of dust in fusion devices constitutes an important issue with safety and operational implications [1, 2, 3, 4]. Despite the fact that ferromagnetic dust has been detected in a number of machines [5, 6, 7, 8], magnetic dust has never been systematically investigated probably owing to the obvious absence of bulk magnetic materials in the vessel components. In particular, comprehensive dust collection activities in most tokamaks have never attempted a characterization of the magnetic properties of the sampled dust population [9, 10, 11, 12, 13, 14, 15].

It has been suggested that ferromagnetic dust can be generated in tokamaks by a non-equilibrium austenite-to-ferrite phase transformation that is induced by rapid temperature quenching in the presence of strong external magnetic fields [8]. Moreover, there is strong experimental evidence of pre-plasma remobilization of ferromagnetic dust in FTU that has been further supported by theoretical estimates [16]. This provided motivation for a re-evaluation of the old hypothesis concerning the possible impact of prematurely re-mobilized dust on the tokamak start-up [5, 16]. Accurate predictions of the extent of such impact require reliable input concerning the overall amount and the size distribution of the magnetic dust.

The present work reports the first cross-machine dust collection study that focuses on the magnetic properties of

tokamak dust and constitutes the first attempt to map the magnetic dust inventory of contemporary fusion devices. A morphological and elemental analysis of magnetic dust sampled from FTU, Alcator C-Mod, COMPASS and DIII-D is presented. These characteristics are compared to draw general conclusions concerning magnetic dust in tokamaks.

2. Magnetic dust collection and vessel composition

In the present work, the term *magnetic dust* refers to any particulate that can be isolated from a specimen under the effect of a strong permanent magnet, *i.e.* discussion is not confined to ferromagnetic dust but extends to strongly paramagnetic dust. The dust collection activities were carried out with different methods and in different locations for the tokamaks investigated. Magnetic dust from FTU, Alcator C-Mod and COMPASS was separated, by means of a strong magnet, from an entire dust batch collected by vacuuming. Magnetic dust from DIII-D was collected *in-situ* by means of a permanent magnet dragged around the vessel during a scheduled machine shut-down. The above difference naturally introduces limitations in the comparison between DIII-D and the other tokamaks.

The tokamaks under investigation are characterized by different plasma-facing material composition and different working toroidal magnetic field strengths. As summarized

Table I: Vessel material composition at the time of the dust collection together with the maximum strength of the toroidal magnetic field. TZM is a molybdenum dominated alloy (Ti 0.5wt%, Zr 0.1wt%, Mo balance); SS stands for stainless steel; Inconel is a Ni based super-alloy. Note also that, during the 2007 Alcator C-Mod experimental campaign, a toroidal row of W tiles was inserted in the strike point region [13].

Tokamak device	Year of dust collection	Wall material	Machine configuration	Limiter or divertor material	Maximum toroidal magnetic field
FTU	2013	SS 304LN	Limiter	TZM	8.0 T
Alcator C-Mod	2007	Mo	Divertor	Mo	8.0 T
COMPASS	2014/15	Inconel 625	Divertor	Graphite	2.1 T
DIII-D	2018	Graphite	Divertor	Graphite	2.2 T

in Table I, FTU and COMPASS feature plasma-facing materials that can directly seed magnetic dust (namely stainless steel and Inconel respectively), while Alcator C-Mod and DIII-D can only indirectly seed magnetic dust through diagnostic components or structural materials. Moreover, FTU and Alcator are compact machines with high toroidal magnetic field up to 8 T, whereas COMPASS and DIII-D have moderate field up to 2.2 T, implying a variable ability of magnetic dust production. Even though interest on plasma-facing materials is currently focused on the slightly paramagnetic tungsten and the slightly diamagnetic beryllium, as we shall discuss, the machine combination studied herein can be considered as representative of existing and future devices with respect to magnetic dust generation.

The dust batches that were selected for analysis are listed in the following. (a) *FTU*. The dust was collected by gross vacuuming around the vessel floor and was meshed out into two different size groups [8]. The batch with sizes $< 210 \mu\text{m}$ was considered in the present investigation. (b) *Alcator C-Mod*. The dust was collected by vacuuming the floor, the upper areas of the lower outer divertor, the horizontal areas and beneath the “FG” sector of the outer divertor in 2007. (c) *COMPASS*. Three dust batches were collected by vacuuming in different locations: the batch Co2 was collected inside the east port duct, the batch Co3 was collected inside the south port duct and the batch Co4 was collected in the divertor at the Langmuir probe feed-through position. (d) *DIII-D*. Magnetic dust was collected by a permanent magnet dragged around bottom surfaces of an outboard midplane and bottom ports during an entry vent. In this case, the overall amount was minuscule.

Table II catalogues the dust batch names, weights and the analysis performed. It should be noted that dust from all the tokamak devices could be covered by boron as a result of boronisation activities. In addition, only for FTU, dust could be covered by lithium compounds courtesy of the liquid lithium limiter [8, 17]. Both elements are highly reactive with oxygen (O) and are difficult to detect by EDX analysis either due to the low sensitivity at low energies or due to partial overlapping with other elements.

3. Analysis of magnetic dust

The magnetic dust collected from the four aforementioned tokamaks is based either on stainless steel (SS) or on Nickel

Table II: Dust batches investigated. “N/A” suggests that the amount of dust is too scarce to be weighted, “XRD” stands for X-ray diffraction, “SEM” for scanning electron microscopy and “EDX” for energy-dispersive X-ray analysis. The fractions are in weight percentage.

Dust batch name	Magnetic amount / fraction	Non-magnetic amount / fraction	XRD	SEM-EDX
Ft1	12.57g/25.0%	50.28g/75.0%	yes	yes
All	0.074g/27.4%	0.196g/72.6%	yes	yes
Co4	0.042g/17.6%	0.196g/82.4%	yes	yes
Co3	N/A	N/A	no	yes
Co2	N/A	N/A	no	yes
D1	N/A	N/A	no	yes

(Ni) alloys that originate from different plasma facing components (PFCs) such as walls, divertor or limiter tiles, diagnostics or antennas. In terms of shape, the dust can be roughly categorized as flakes, splashes and spheroids.

Dust flakes. The majority of dust grains in all magnetic batches are flakes based on SS. Characteristic examples of flake-shaped dust that was collected from the four tokamaks are illustrated in figure 1. Their maximum dimension ranges from few tens up to $\sim 1500 \mu\text{m}$ depending on the fusion device of origin. EDX analysis revealed that the flakes have low oxygen content ($< 6 \text{at}\%$ on average) with the exception of FTU flakes which feature a much higher O concentration, probably due to the presence of Li. The average Fe/Cr ratio is $3.0 \div 5.0$, estimated on the grains analyzed from each device batch, to be compared with that of the standard SS AISI304/316 grade ($3.3 \div 4.0$). It should be noted that many debris in Alcator and COMPASS definitely originate from machining, performed either in-vessel or transported from outside. In fact, some debris have “accordion-like” structures typical of mechanical shavings, see figure 2(a,b). This has been verified by a morphological comparison with SS shaves obtained from workshop milling and lathing activities.

Splashed droplets. Splash-shaped magnetic dust has been found in all devices except DIII-D. This is expected owing to the absence of vacuuming in the present DIII-D collection as well as the enhanced adhesion of splashes [18] or of immobile dust with high temperature pre-history [19]. Splashes are based on SS or Ni alloys, see figure 3. Their maximum dimension ranges from 100 up to $\sim 500 \mu\text{m}$ and

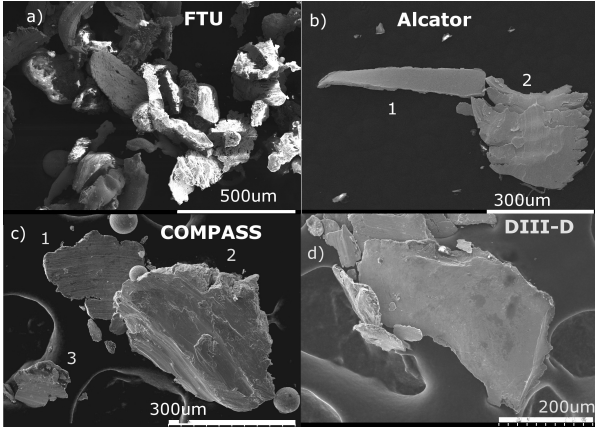


Figure 1: Dust flakes and their EDX analysis (at%). a) SS-based from FTU (C 3 ÷ 9, O 15 ÷ 40, Cr 1 ÷ 15, Fe 15 ÷ 53); b) Ni (1) and SS (2) based from Alcator (O 10.1-0, Cr 0-19.6, Fe 0.6-68.5, Ni 88.5-7.7, respectively); c) SS-based from the COMPASS Co3 batch (C 11 ÷ 33, O 0 ÷ 5.6, Cr 0.7 ÷ 16, Fe 45 ÷ 86, Ni 0 ÷ 6.6); d) SS-based from DIII-D (C 23.4, O 8.3, Cr 11.3, Fe 46.6, Ni 6.3).

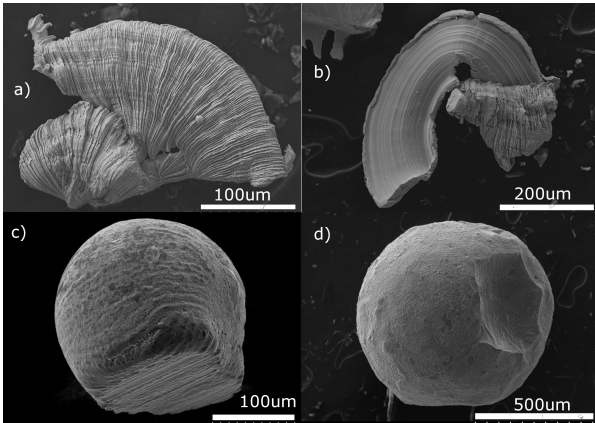


Figure 2: a-b) Chips in the Alcator magnetic batch probably due to machining. EDX analysis (at%): O 0.8-2.6, Cr 18.8-19, Fe 69.1-66.3 respectively. c-d) Spheroidal dust from in-lab Inconel studs welded on SS plates and EDX (at%); c) magnetic dust (O 63.7, Cr 18.6, Fe 1.4, Ni 0); d) non-magnetic dust (O 58.7, Cr 28.3, Fe 1.2, Ni 0.5).

they have a low O concentration (10 at% on average). The SS-based splashes possess a Fe/Cr ratio of 3.5 ÷ 3.9.

Spheroids. Sphere-like magnetic grains have been collected in all tokamaks under investigation as either SS- or Ni-based grains with diameters ranging from ~ 40 up to ~ 250 μm . Sphere-like dust can be divided into two classes; those with clear dendrite-like surface texture (see figure 4) and those without dendrite-like surface texture (see figure 5). The O content of the un-textured SS spheroids is lower than that of the textured SS spheroids (25 at% vs 45 at% on average), probably due to their higher surface porosity. The average Fe/Cr ratio is 2.5 ÷ 4 in both classes. The un-textured grains from DIII-D have a highly irregular surface but feature high oxygen content. The presence of a clear surface texture depends on the dynamic cooling rate after melting [20], which is not possible to estimate solely based on post-mortem evidence. The presence of both tex-

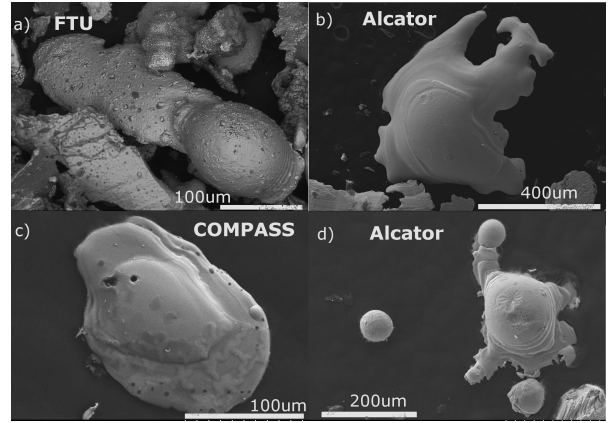


Figure 3: Dust splashes and their EDX analysis (at%). a) SS dust from FTU (O 10.1, Cr 14.2, Fe 61.3, Ni 8.0); b) SS dust from Alcator (O 1.1, Cr 18.8, Fe 72.2, Ni 6.6); c) Inconel dust from the COMPASS Co2 batch (B 24.6, O 4.8, Cr 14.3, Fe 0.9, Ni 37.7), no splashes in the Co3,4 batches; d) SS dust from Alcator (O 15.9, Cr 16.3, Fe 51.7, Ni 7.2).

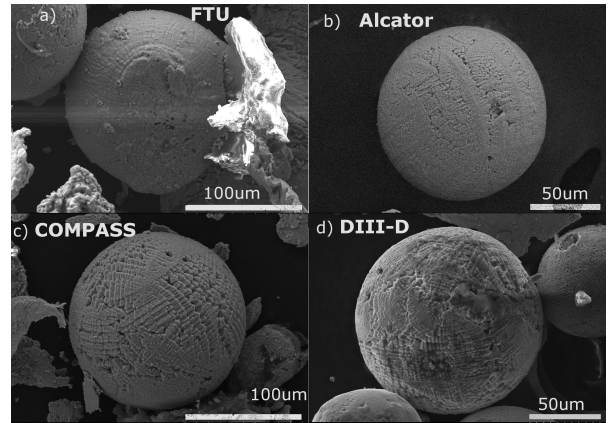


Figure 4: Spheroidal dust *with* clear dendrite-like textured surface and their EDX analysis (at%). a) Ni-based dust from FTU (O 15.4, Cr 12.7, Fe 10.5, Ni 51.24); b) SS dust from Alcator (O 53.7, Cr 8.0, Fe 32.8, Ni 4.8); c) SS dust from the COMPASS Co4 batch (O 48.3, Cr 8.7, Fe 32.1, Ni 3.3); d) SS dust from DIII-D (C 26.9, O 41.6, Cr 7.2, Fe 21.7, Ni 1.9).

tured and un-textured grains in the magnetic dust batches lead to the conclusion that the grain cooling rate is not a critical parameter for the austenite-to-ferrite phase transformation. Note that Ni-based sphere-like dust has been detected in the FTU, COMPASS Co2 and DIII-D batches.

Aiming to establish whether spheroidal magnetic dust can originate from in-vessel welding activities (a common practice inside the DIII-D vessel), dust has been collected after in-lab welding of Inconel 625 studs on Inconel 625 or SS AISI316 plates used in DIII-D. The analysis of the produced spheroidal grains revealed that only welding involving SS materials could generate magnetic grains, while pure Ni-based grains are not magnetic. Furthermore, the welding-produced magnetic SS dust seems to have been exposed to much higher heat loads and presents some irregular surface texture structure (though not well defined)

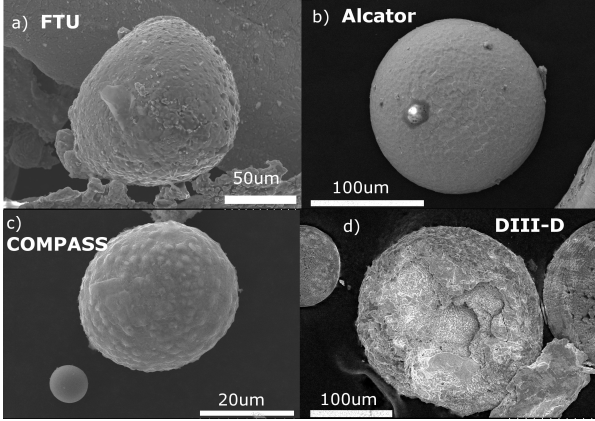


Figure 5: Spheroidal dust *without* any clear dendrite-like textured surface and their EDX analysis (at%). a) SS dust from FTU (C 19.6, O 15.7, Cr 9.6, Fe 47.2, Ni 7.6); b) SS dust from Alcator (C 0, O 0, Cr 21.2, Fe 70.0 Ni 8.0); c) SS dust from the COMPASS Co4 batch (C 50.4, O 9.1, Cr 7.9, Fe 27.4, Ni 4.4); d) SS dust from DIII-D (C 53.4, O 23, Cr 4.7, Fe 15.8, Ni 2.3).

compared to dust collected in tokamaks. Examples of magnetic and non-magnetic SS grains produced by in-lab welding activity are illustrated in figure 2c-d. Note the high O content and low Fe/Cr ratio ($\ll 1$).

X-ray Diffraction Analysis. XRD analysis was carried out on some magnetic dust batches (see table II). All batches analyzed were confirmed to be composed of SS and Ni alloys. Some of them, namely the ones from FTU and COMPASS, feature an increased ferrite peak compared to the peak of the standard untreated AISI304/316 SS material, see figure 6. This can be explained by the austenite-to-ferrite phase transformation that occurs in re-solidified spherical droplets and splashes when embedded in a strong magnetic field environment [8]. In addition, the FTU spectrum reveals the probable presence of martensite, typical of rapid cooling down of molten grains, a peculiarity that can be attributed to the fact that FTU is a cryogenic device [21, 22]. On the other hand, the XRD spectrum of Alcator magnetic dust (figure 6) is similar to the standard spectrum of untreated AISI316 steel, where the austenitic Fe- γ peak is much higher than the ferrite Fe- α one; Fe- α about 10% of Fe- γ . This can be explained by the fact that most Alcator magnetic grains were flakes, which as discussed above are believed to primarily originate from either in-vessel or ex-vessel machining. A more detailed analysis of the XRD results will be reported in a separate work.

4. Discussion and conclusions

This investigation catalogues the magnetic dust inventory of fusion devices that are equipped with different plasma-facing materials. Concerning its surface morphology, magnetic dust is encountered in the shape of flakes, splashes or spheroids with dimensions that range from few microns up to several hundreds of microns.

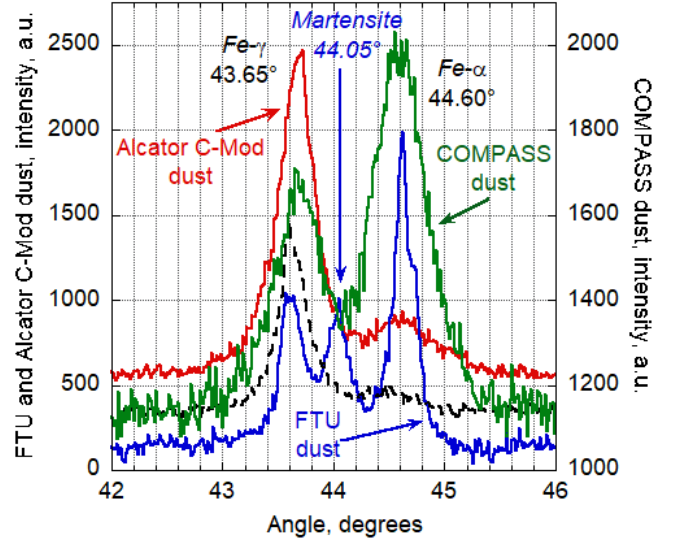


Figure 6: XRD spectra of the Alcator C-Mod (red), FTU (blue) and COMPASS Co4 (green) magnetic dust batches. The spectrum of untreated AISI316 steel is also shown with the black dashed line.

As far as the chemical composition is concerned, magnetic dust is based on steel or nickel alloys. The low surface content of oxygen on dust splashes and some spheroidal grains suggests that they originated from plasma-material interactions and not from in-vessel welding activities that can be a common practice. The latter possibility can be excluded for the majority of grains because of the morphological and chemical differences between plasma-generated and welding-generated dust, namely the high oxygen concentration and the Fe/Cr < 1 ratio in dust that is produced by in-lab welding. This is the so-called *Cr surface segregation phenomenon* that typically accompanies very high temperature treatment [23], as in welding, and leads to a higher O content owing to the formation of Cr oxides. The collection of Ni-based spheroidal dust (in DIII-D) further supports the claim that such grains do not originate from welding activities, as discussed in section 3.

X-ray Diffraction analysis has demonstrated the different nature of dust magnetism depending on the fusion device. In particular, in magnetic dust collected from FTU and COMPASS, the presence of ferrite and / or martensite structures is revealed that constitutes evidence of SS dust magnetic phase transition under intense plasma heat loads [8]. The presence of an ambient magnetic field plays a role in magnetic phase retention during dust resolidification, as observed by some authors [24, 25], but the magnetic field strength does not appear to be a key point, since magnetic phase transitions were detected in devices with high and moderate toroidal magnetic field. Probably, this stems from the fact that magnetic fields of the order of few Tesla suffice to saturate the magnetic domains of ferritic grains. On the other hand, the large quantity of SS flakes collected in Alcator C-Mod is revealed to be in the native austenite state.

In the case of ferromagnetic spherical dust, it has already been demonstrated that the magnetic moment force owing to the toroidal field can overcome both the gravitational and the adhesion force leading to pre-plasma remobilization [16]. In the case of strongly paramagnetic dust, a simple evaluation of the magnetic force that is exerted on a spherical grain in presence of toroidal field gradients suggests that it is several times larger than gravity. This indicates that, in principle, any SS grain could be classified as magnetic dust. The complications due to the non-spherical morphology of the flakes and a comparison with adhesion forces [26, 27] will be addressed in future work.

Overall, a non-negligible amount of magnetic dust has been collected from the tokamaks considered in this work. This holds for FTU and COMPASS that can directly seed magnetic dust from the first wall, but also holds for Alcator C-Mod and DIII-D that can only indirectly seed magnetic dust from their diagnostics and antennas. Taking into account that the ITER diagnostic first wall is composed of stainless steel [28], it is possible that great amounts of magnetic dust will be produced in ITER, where plasma breakdown is critical. This is confirmed by the latest projections for dust generation in ITER, which anticipate comparable amounts of stainless steel and tungsten dust by the end of the fusion power operation phase [29]. In particular, the lower estimate is 8 kg W and 9 kg SS dust, whereas the upper estimate is 37 kg W dust and 23 kg SS dust [29]. In the prospective of the possible extended utilization of reduced-activation ferritic-martensitic steels in DEMO [30], it is possible that large amounts of magnetic dust will also be generated beyond ITER. The understanding of the nature and origin of magnetic dust is an essential step towards the reliable modelling of its pre-plasma remobilization and of its interference with the start-up of future fusion devices.

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