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# Measurement of the proton spin structure at long distances

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Measuring the spin structure of nucleons (protons and neutrons) extensively tests our understanding of how nucleons arise from quarks and gluons, the fundamental building blocks of nuclear matter. The nucleon spin structure is typically probed in scattering experiments using polarized beams and polarized nucleon targets, and the results are compared with predictions from Quantum Chromodynamics directly or with effective theories that describe the strong nuclear force. Here we report on new proton spin structure measurements with significantly better precision and improved coverage than previous data at low momentum transfer squared between 0.012 and 1.0 GeV<sup>2</sup>. This kinematic range provides unique tests of effective field theory predictions. Our results show that a complete description of the nucleon spin remains elusive. They call for further theoretical works that include the more fundamental lattice gauge method. Finally, our data agree with the Gerasimov-Drell-Hearn sum rule, a fundamental prediction of quantum field theory.

Understanding how hadronic matter arises from its fundamental constituents, quarks and gluons, is central to the study of nuclear and particle physics. Although the strong interaction is described by Quantum Chromodynamics (QCD), it remains the least understood force in the Standard Model. The difficulty arises because the QCD coupling constant  $\alpha_s$  becomes large at long distances [1], making traditional perturbative expansions in powers of  $\alpha_s$  infeasible. Consequently, complex phenomena like quark confinement are hard to understand quantitatively. The most fundamental approach to calculate QCD non-perturbatively is lattice gauge theory [2]. A second approach is provided by Effective Field Theories (EFT), which maintain rigorous, traceable connections to the underlying fundamental theory. A popular approach is chiral effective field theory ( $\chi$ EFT) [3, 4], which is constructed from hadronic degrees of freedom and incorporates the symmetries of QCD, including its approximate chiral symmetry. By making use of a perturbative expansion in small parameters,  $\chi$ EFT predicts experimental observables from a limited set of phenomenological inputs. Although generally successful,  $\chi$ EFT has been challenged by experimental data that depend explicitly on spin degrees of freedom [5, 6]. This is not unprecedented: other theoretical predictions had been thought to be robust until confronted with spin observables, including parity symmetry [7], the Ellis-Jaffe spin sum rule [8], the nucleon spin asymmetry  $A_1$  [9], and lattice QCD calculations of the nucleon axial charge [10]. Therefore, fully understanding QCD and nuclear matter requires an extensive set of spin observables.

We report on the measurements performed using a polarized electron beam to probe a polarized proton at the Thomas

Jefferson National Accelerator Facility (Jefferson Lab), in Virginia, USA. We measured spin-dependent cross sections in the nucleon resonance region at very low  $Q^2$ , i.e. at long distances. Here,  $Q^2$  is the square of the 4-momentum transferred from the electron to the proton and represents the inverse of the distance scale probed by the scattering. Polarized electrons with energies of 3.0, 2.3, 2.0, 1.3 and 1.1 GeV, produced by Jefferson Lab's Continuous Electron Beam Accelerator Facility (CEBAF), were scattered from a polarized proton target [11, 12]. The beam polarization ( $P_b$ ) was measured to be 85% with a total uncertainty of 2% using a Møller polarimeter [13]. The target contained granules of NH<sub>3</sub> that were dynamically polarized [11] at 1K in a 5 T magnetic field. The target polarization ( $P_t$ ) varied from 75% to 90%, as monitored by nuclear magnetic resonance polarimetry. As described below and in the Methods section, the product  $P_b P_t$  was measured to a relative precision of (2 – 5)%. The scattered electrons were identified using the CEBAF Large Acceptance Spectrometer (CLAS) [13], which was equipped with a multi-layer drift chamber detector for charged particle tracking, a scintillator hodoscope for particle time-of-flight measurement, an electromagnetic calorimeter and a Cherenkov Counter for discriminating scattered electrons from other background particles. The Cherenkov Counter in one of the six sectors of CLAS was modified specifically for this experiment to detect electron scattering at angles as low as 6°. Only this sector was used to collect the inclusive electron scattering data reported here.

The dominant scattering process is the one-photon exchange, in which the incident electron exchanges a single virtual photon with the nucleon of mass  $M$ , see Fig. 1. The 4-momentum transferred from the electron to the nucleon is  $q^\mu = k^\mu - k'^\mu = (\nu, \mathbf{q})$ , in which  $k^\mu$  and  $k'^\mu$  are the 4-momenta of the incident and the scattered electrons, respectively, and  $\nu$  is the energy transfer. In the following, we describe this process using the Lorentz-invariant vari-

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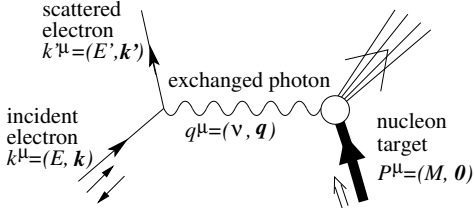


FIG. 1. The one-photon exchange process of polarized electron scattering off a polarized nucleon. The 4-momenta of the incident and the scattered electrons are  $k^\mu = (E, \mathbf{k})$  and  $k'^\mu = (E', \mathbf{k}')$ , respectively. The spin direction of the incident electron is indicated by the arrows  $\uparrow\downarrow$ . The nucleon, if at rest, has  $P = (M, \mathbf{0})$  and its spin is indicated by the outlined arrow  $\uparrow$ .

ables  $Q^2 = -q^2$ , and the Bjorken scaling variable  $x \equiv -q^2/(2P \cdot q)$  or the invariant mass of the photon-nucleon system  $W \equiv \sqrt{(P+q)^2} = \sqrt{P^2 + (1/x-1)Q^2}$ . The inclusive electron scattering cross section can be written as a linear combination of structure functions, of which  $F_1(x, Q^2)$  and  $F_2(x, Q^2)$  represent the spin-independent part of the cross section, and the spin structure functions  $g_1(x, Q^2)$  and  $g_2(x, Q^2)$  describe its dependence on the beam and target spin polarization. These structure functions encode the internal structure of the target. Alternatively, one can describe the spin-dependent part of the nucleon response in terms of virtual photo-absorption asymmetries  $A_1 = [g_1 - (Q^2/\nu^2)g_2]/F_1$  and  $A_2 = (\sqrt{Q^2}/\nu)(g_1 + g_2)/F_1$  [14]. The polarized cross section difference  $\Delta\sigma \equiv \sigma^{\downarrow\uparrow} - \sigma^{\uparrow\uparrow}$ , with  $\uparrow\downarrow$  representing the beam helicity state and  $\uparrow\downarrow$  the target spin orientation, is largely proportional to  $g_1$  (or equivalently  $A_1F_1$ ) with a small contribution from  $A_2F_1$ .

The proton spin structure function  $g_1$  and the product  $A_1F_1$  were extracted from the difference in the measured yield,  $N$ , of scattered electrons from a longitudinally polarized target between opposite beam helicity states:

$$\frac{N^{\downarrow\uparrow}}{Q_b^\downarrow} - \frac{N^{\uparrow\uparrow}}{Q_b^\uparrow} = \Delta\sigma(W, Q^2) \mathcal{L} P_b P_t a(W, Q^2), \quad (1)$$

where  $Q_b$  is the time-integrated beam current,  $\mathcal{L}$  is the areal density of polarized protons in the target, and  $a(W, Q^2)$  accounts for the detector acceptance and efficiency. The product  $\mathcal{L}P_bP_t$  was measured directly using elastic scattering on the proton and  $a(W, Q^2)$  was determined using a Monte Carlo simulation of the experiment; see the Methods section for details. Examples of our  $g_1$  results on the proton are shown in Fig. 2. Our results extend the measured  $Q^2$  range down to below the pion mass squared ( $m_\pi^2$ ), three times smaller than previous data [14–22], which makes it possible to rigorously test  $\chi$ EFT calculations for spin-dependent observables.

In our study, we utilize sum rules that relate integrals of structure functions to amplitudes calculable by lattice QCD [23, 24] or  $\chi$ EFT, or to known static properties of the target. One such relation is the Gerasimov-Drell-Hearn (GDH) sum rule [25, 26] for real photon absorption ( $Q^2 = 0$ ):

$$\int_{\nu_0}^{\infty} \Delta\sigma(\nu) \frac{d\nu}{\nu} = -\frac{2\pi^2\alpha}{M^2} \kappa^2, \quad (2)$$

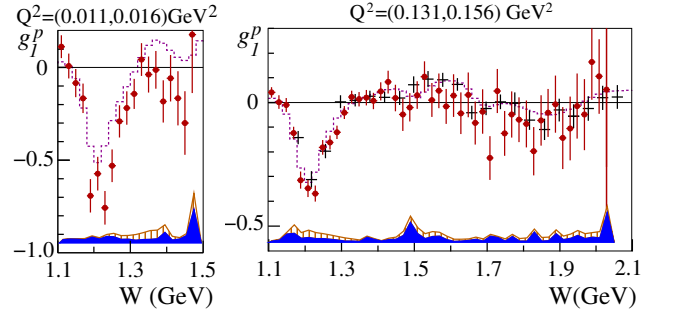


FIG. 2. Results on  $g_1$  of the proton (solid circles) vs invariant mass  $W$  for the lowest ( $0.011 \leq Q^2 \leq 0.016 \text{ GeV}^2$ ) bin and an intermediate ( $0.131 \leq Q^2 \leq 0.156 \text{ GeV}^2$ ) bin, compared to a parameterization of previous world data (dotted curve) [14]. The error bars are statistical. The solid and the vertically-hatched bands show the experimental and the parameterization uncertainties, respectively. Results from a previous experiment carried out in Jefferson Lab's Hall B [14] are shown when available (crosses).

with  $\kappa$  the anomalous magnetic moment of the target particle,  $\nu_0$  the inelastic threshold and  $\alpha$  the fine-structure constant. Theoretical arguments indicate that the divergence of the  $1/\nu$  factor is compensated by the fast decrease of  $\Delta\sigma$  with  $\nu$ . This is supported by experiments which have verified the GDH sum rule for the proton within about 7% accuracy [27, 28]. There exist several prescriptions that generalize the GDH sum rule to electron scattering in terms of moments of spin structure functions integrated over  $x$  (which is equal to  $Q^2/2M\nu$  in the laboratory frame). One often-used generalization is [29]:

$$\Gamma_1(Q^2) \equiv \int_0^{x_0} g_1(x, Q^2) dx = \frac{Q^2}{2M^2} I_1(Q^2), \quad (3)$$

where  $x_0 = Q^2/(W_{thr}^2 - M^2 + Q^2)$  corresponds to the electroproduction threshold  $W_{thr} = M + m_\pi = 1.073 \text{ GeV}$ . Equation (3) defines the integral  $I_1$ , which is related to the first polarized doubly-virtual Compton scattering (VVCS) amplitude that is calculable in the  $\nu \rightarrow 0$  limit with lattice QCD or  $\chi$ EFT [3, 4, 30–39]. The other prevailing generalization of the GDH integral is [40]:

$$I(Q^2) = \frac{2M^2}{Q^2} \int_0^{x_0} [A_1(x, Q^2)F_1(x, Q^2)] dx, \quad (4)$$

which can be calculated from both the first and the second spin-dependent VVCS amplitudes in the  $\nu \rightarrow 0$  limit. The  $I(Q^2)$  thus obtained can be extrapolated to  $Q^2 = 0$  to test the original GDH prediction  $I(0) = \kappa^2/4$ . In this work, we present results on both generalizations.

To form the spin structure integrals in Eqs. (3 & 4), the measured values of  $g_1$  or  $A_1F_1$  were used whenever available from our experiment up to a maximum  $x$  corresponding to  $W = 1.15 \text{ GeV}$ , which was chosen to limit the background from the elastic radiative tail (see Methods section) and down to a minimum  $x$  determined by the beam energy and the acceptance of CLAS. Contributions from regions at low  $x$  (down to  $x = 10^{-3}$ ) and at high  $x$  from  $W_{thr}$  to  $W = 1.15 \text{ GeV}$  were evaluated using a parameterization of previous data [14].

Results on  $\Gamma_1(Q^2)$  and  $I(Q^2)$  are shown in Figs. 3 and 4.

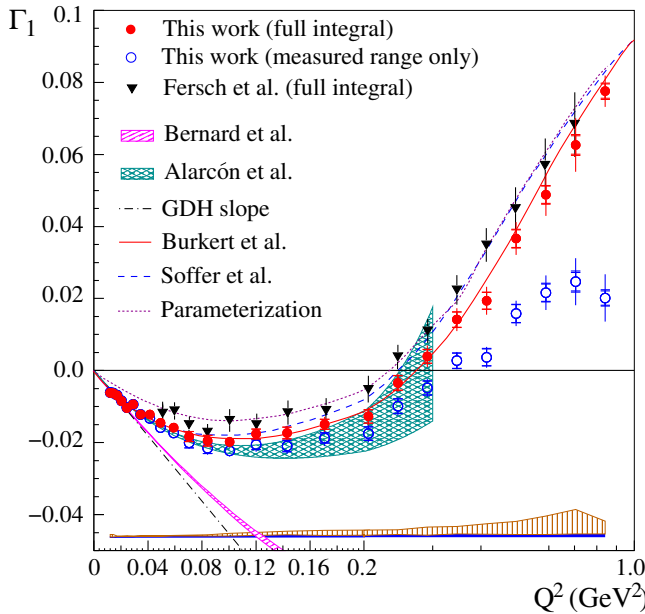


FIG. 3. Results on  $\Gamma_1(Q^2)$  for the proton. Integrals over the experimentally covered  $x$  range are shown as open circles. Full integrals are shown as solid circles. The inner and the outer error bars (sometimes too small to be seen) are for statistical and total uncertainties, respectively. Results from a previous experiment [14] are shown as solid triangles. The solid and the vertically-hatched bands show the experimental and the parameterization uncertainties, respectively. Also shown are the latest  $\chi$ EFT predictions by Bernard *et al.* [36] (diagonally hatched band) and Alarcón *et al.* [37] (cross-hatched band), phenomenological models by Burkert *et al.* [41] (solid curve) and Soffer *et al.* [42] (dashed curve), as well as our spin structure function parameterization [14] (dotted curve). The dash-dotted line is the slope predicted by the GDH sum rule as  $Q^2 \rightarrow 0$ .

To quantify the degree of agreement between our data and the recent  $\chi$ EFT predictions [36, 37], we computed the  $\chi^2$  per degree of freedom between these predictions and our results. We find that the predictions in [36] agree with our results only at the lowest few  $Q^2$  points, up to  $Q^2 = 0.024(0.014)$  GeV<sup>2</sup> for  $\Gamma_1$  ( $I$ ), if we require a  $\chi^2_{\text{reduced}} < 2$ . On the other hand, the predictions in [37] agree with our data over their full range, with  $\chi^2_{\text{reduced}} < 2$  up to  $Q^2 = 0.3$  GeV<sup>2</sup>. The phenomenological models [41, 42] agree well with our results for all  $Q^2$  values. The new results on  $\Gamma_1(Q^2)$  generally agree with a previous experiment [14] in the overlapping  $Q^2$  region. However, there exist visible differences between our results and the spin structure function parameterization [14], indicating that it can be improved with our new data. Extrapolating our results on  $I(Q^2)$  to  $Q^2 = 0$  yields

$$I^{\text{exp}}(0) = -0.798 \pm 0.042 \quad (5)$$

assuming the  $Q^2$ -dependence of  $I$  predicted by Alarcón *et al.* [37] within their quoted theoretical uncertainty (see details in the Methods section). This result is in good agreement with the GDH sum rule prediction  $I^{\text{GDH}} = -\kappa^2/4 = -0.804(0)$  for the proton and with the experimental photoproduction result  $-0.832 \pm 0.023(\text{stat}) \pm 0.063(\text{syst})$  [27, 28]. Our results provide, for the first time, a test of the GDH sum independent

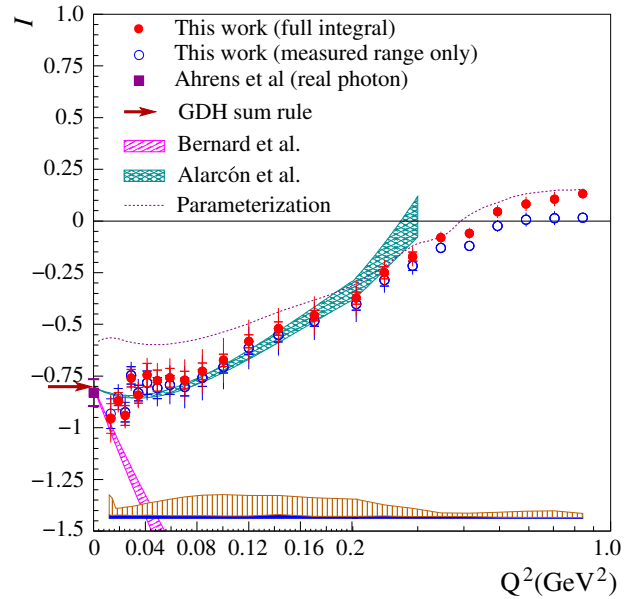


FIG. 4. Results on  $I$  for the proton, with symbols the same as in Fig. 3. The GDH value is shown by the arrow at  $I^{\text{GDH}} = -0.804$ . The experimental photoproduction result [27, 28] is shown by the solid square.

from exclusive photoproduction [27, 28].

Predictions from  $\chi$ EFT for  $I(Q^2)$  and  $\Gamma_1(Q^2)$  are constrained at  $Q^2 = 0$  by the GDH sum rule. No such constraint is available for  $\gamma_0(Q^2)$ , the generalized longitudinal spin polarizability, related by a sum rule to the integral of  $A_1 F_1$  [40, 43]:

$$\gamma_0(Q^2) = \frac{16\alpha M^2}{Q^6} \int_0^{x_0} x^2 A_1(x, Q^2) F_1(x, Q^2) dx. \quad (6)$$

This endows  $\gamma_0(Q^2)$  with additional resolving power to test the several theoretical predictions available. Furthermore, the  $x^2$  weighting in Eq. (6) suppresses the low- $x$  contribution. This is beneficial since the low- $x$  region is inaccessible experimentally and must be estimated using models, which introduces model uncertainty. The two integrals  $I$  and  $\gamma_0$  have different systematic uncertainties and therefore provide complementary tests of theoretical predictions.

Our results for  $\gamma_0(Q^2)$  are shown in Fig. 5. Neither of the new  $\chi$ EFT calculations describes the full data set well: The calculation from Ref. [36] agrees in magnitude (but not in slope) with our lowest  $Q^2$  results up to  $Q^2 \approx 0.025$  GeV<sup>2</sup>, while the calculation from Ref. [37] describes the shape of the data only marginally below that  $Q^2$  value. Together with the photoproduction data point [27, 28, 44], our data indicate a strong change in  $Q^2$  slope towards a value consistent with that predicted in Ref. [36] at very low  $Q^2$ . Classically,  $\gamma_0$  represents the distortion of the proton spin structure in response to the interference between various transverse electric and magnetic field components of the virtual photon shown in Fig. 1. In a hadronic picture,  $\gamma_0$  is principally due to the difference between the contribution from the pion cloud of the proton (positive) and the excitation of the  $\Delta$  (negative) [37]. The data thus indicate that the  $\Delta$  contribution dominates at

the photon point and becomes even more important for small- $Q^2$  virtual photons. This may be pictured intuitively from the extended size of the pion cloud whose contribution is quickly suppressed with increasing  $Q^2$ . However, at higher  $Q^2$ , the slope turns over since the polarizability is a global feature of the proton which must vanish as  $Q^2 \rightarrow \infty$ , as seen from the  $1/Q^6$  factor in Eq. (6).

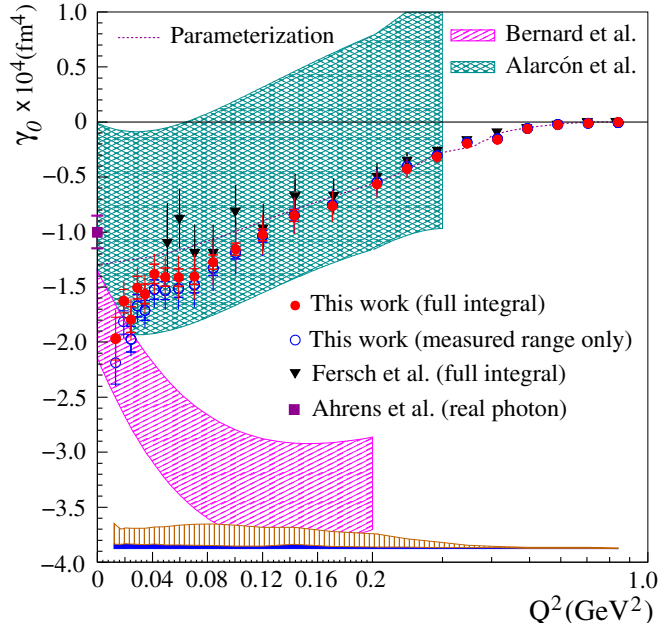


FIG. 5. Results on  $\gamma_0(Q^2)$  for the proton, with symbols the same as in Fig. 3. The photoproduction data point [27, 28, 44] is shown as the solid square.

Although the upper bound of the validity domain of  $\chi$ EFT is not known, the kinematic coverage of our data is well within its expected range between  $m_\pi^2 \approx 0.02$  GeV<sup>2</sup> and the chiral symmetry breaking scale,  $\Lambda_\chi^2 \approx 1$  GeV<sup>2</sup>. The actual validity range depends on the orders of the expansion parameter  $m_\pi/\Lambda_\chi$  at which the calculations are done, the expansion method, and the observable. One reason for the limited success of  $\chi$ EFT in describing our results may be coming from the difficulty to fully account for the  $\Delta$  resonance, the proton's first excited state. In fact, early  $\chi$ EFT calculations [30–32] did not explicitly include the  $\Delta$  excitation, which slows down the convergence of the  $\chi$ EFT perturbation series, or they included it phenomenologically [33–35]. This was thought to be the reason why many of the early nucleon spin structure function data [15–22] disagreed with calculations [30–35]. This disagreement prompted refined  $\chi$ EFT calculations [36–39] and a new experimental program at Jefferson Lab optimized to cover the  $\chi$ EFT domain [45, 46], including the measurement reported here. The latest calculations [36–39] both include the  $\Delta$  but differ in their expansion method to account for the  $\pi$ - $\Delta$  corrections. Ref. [36] treats the nucleon- $\Delta$  mass difference  $\delta M$  as a small parameter of the same order as  $m_\pi$ . Refs. [37–39] use  $\delta M$  as an intermediate scale such that  $\delta M/\Lambda_\chi \approx m_\pi/\delta M$  is used as the expansion parameter to account for the  $\Delta$ . In addition, the calculations [37–39] include empirical form factors in the relevant couplings to approxi-

mate the expected impact of high-order contributions. They make  $\gamma_0$  vanish at large  $Q^2$ , as observed, in contrast to calculation [36] which purely contains terms computed with  $\chi$ EFT and has no free-parameter that can be adjusted. For  $\gamma_0$ , which arises at third order in the  $\pi$ -N loops there are large cancellations between  $\pi$ -N loops and the  $\Delta$  contribution. Therefore, the calculations are very sensitive to the expansion and renormalization scheme, and the order of the expansion. This is why  $\gamma_0$  is especially well-suited to test  $\chi$ EFT. Finally, the integrals  $\Gamma_1$  and  $I$  contain Born terms in addition to the polarizability contributions calculated in  $\chi$ EFT. These terms are constrained by the GDH sum rule at  $Q^2 = 0$ . Refs [37–39] assume that their  $Q^2$ -dependence follows the corresponding proton form factors. This  $Q^2$ -dependence leads to the difference with Ref. [36] and the agreement with our data.

In summary, the proton polarized structure functions  $g_1$  and  $A_1F_1$  and their integrals  $\Gamma_1$ ,  $I$  and  $\gamma_0$  have been measured in the very low  $Q^2$  region, down to 0.012 GeV<sup>2</sup>. Our results on  $I(Q^2)$  extrapolated to  $Q^2 = 0$  agree well with the original GDH sum rule. For  $Q^2 > 0$ , they provide precise tests of predictions from  $\chi$ EFT, the leading effective theory for the strong interaction. These tests use for the first time the proton spin degrees of freedom in the  $Q^2$  region where  $\chi$ EFT should be the most applicable. Although it is essential to understand the fundamental forces of nature from first principles, such descriptions are often impossible and one must use effective theories based on the new degrees of freedom that emerge from complexity [47]. Our data show that it remains difficult for  $\chi$ EFT to precisely describe *all* observables in which spin degrees of freedom are explicit. They provide strong incentive for future improvements of calculations using  $\chi$ EFT, the leading approach to the effective theory emerging directly from QCD, and for extending the more fundamental lattice QCD calculations to the spin-dependent structure of the nucleon.

## Methods

We measured the spin difference yields on the l.h.s. of Eq. (1) and solved that equation for  $\Delta\sigma(W, Q^2)$ , from which we extracted  $g_1$  and  $A_1F_1$  as functions of  $W$  and  $Q^2$ . We relied on the standard CLAS GEANT-3 Monte Carlo simulation package to fully simulate the spin-dependent yields, including all radiative effects and detector responses. The efficiency of the modified Cherenkov Counter was determined by comparing data taken with only the Electromagnetic Calorimeter in the trigger to those taken with the standard trigger that requires a coincidence between both detectors. The ratio of the latter to the former gave the Cherenkov efficiency. We selected only detector regions of well-understood acceptance in both the data and the simulation. This process fully determined the function  $a(W, Q^2)$  in Eq. (1). The same Eq. (1) was also used to extract the product  $\mathcal{L}P_bP_t$  by comparing the measured yield difference (l.h.s. of Eq. (1)), integrated over the elastic peak region  $0.85$  GeV  $< W < 1.0$  GeV, to the elastic cross section difference  $\Delta\sigma(W = M, Q^2)$  which can be calculated from the known electromagnetic form factors of the proton [48]. The polarized cross section  $\Delta\sigma(W, Q^2)$  in the simulation was calculated using an event generator for inclusive electron scattering [49] with up-to-date models of

structure functions and asymmetries, including near-final data from JLab experiment E08-027. We extracted our results on  $g_1$  and  $A_1F_1$  by varying our input parameterization for these quantities and finding the required values to make our simulation for the polarized yield agree with data. Corrections for higher-order quantum electromagnetic effects (radiative corrections) were applied in the simulation, of which one effect is the high-energy tail from elastic scattering (elastic radiative tail).

We propagated the uncertainties on the polarized yields to the final values for  $g_1$  and  $A_1F_1$ . Systematic uncertainties were studied by changing model parameters, or other inputs, and re-running the simulation. The overall uncertainty on the normalization factor  $\mathcal{L}P_bP_t$  for each beam energy varied from 2% to 5%, dominated by the statistics of the measured elastic peak and, to a lesser extent, the accuracy of the proton elastic form factors [48] that enter into  $\Delta\sigma(W = M, Q^2)$  and hence into our determination of that factor. Smaller contributions, all less than 1%, came from  $\pi^-$  and  $e^+e^-$  backgrounds, and scattering off the slightly polarized  $^{15}\text{N}$  in the target. The reconstruction of  $W$  has an uncertainty of less than 2 MeV, which was studied by shifting the simulated  $W$  spectrum and repeating the extraction. Uncertainties due to trigger and particle reconstruction and identification inefficiencies, as well as parameterizations for the structure functions,  $F_{1,2}$  and  $A_{1,2}$ , were studied by varying them in the simulation. Uncertainties in the radiative corrections were estimated by varying the amount of material the electron passed through in the simulation, and by adjusting the elastic radiative tail within reasonable limits. In all, the total experimental uncertainty is dominated by statistics.

To extrapolate our results on  $I(Q^2)$  to  $Q^2 = 0$ , we fit our data with a form obeying the  $Q^2$ -dependence of the Alarcón *et al.*  $\chi\text{EFT}$  calculation [37]. We chose this calculation because its  $Q^2$ -dependence agrees well with our data over a wide  $Q^2$  range. We found the intercept of our fit with the  $Q^2 = 0$  axis to be  $I^{exp}(0) = -0.798 \pm 0.013(\text{uncor}) \pm 0.040(\text{cor}) \pm 0.003(\text{range}) \pm 0.003(\text{form})$ , with  $\chi^2_{\text{reduced}} = 2.20$  determined with the “uncor” uncertainty. Here, “uncor” and “cor” refer to the experiment point-to-point uncorrelated and correlated uncertainties, respectively; “range” refers to the uncertainty due to the  $Q^2$  range ( $Q^2 \leq 0.1 \text{ GeV}^2$ ) used for the fit. The last contribution, “form”, is the uncertainty on the  $Q^2$ -dependence used for the fit. It is calculated from the uncertainty band given by the  $\chi\text{EFT}$  calculation [37]. Since the various uncertainties are largely independent, they are added quadratically, giving a total uncertainty of  $\pm 0.042$ . This is about twice smaller than that from photoproduction measurements of  $I(0)$  because the  $Q^2 \rightarrow 0$  extrapolation uncertainty calculated using [37] is negligible and because inclusive electroproduction automatically sums over all reaction channels, thereby removing uncertainties associated with the detection of final states needed in photoproduction. On the other hand, the extrapolation uncertainty is calculated from [37], which disagrees with [36]. This indicates that the uncertainty bands provided in the calculations may not reflect the full theoretical uncertainties. Extrapolating using the  $Q^2$ -dependence from [36] yields  $I^{exp}(0) = -0.625 \pm$

$0.022(\text{uncor}) \pm 0.039(\text{cor}) \pm 0.069(\text{range}) \pm 0.056(\text{form})$ , with  $\chi^2_{\text{reduced}} = 2.23$  determined with the “uncor” uncertainty. The “uncor” value here is larger because the fit is limited to very few data points ( $Q^2 \leq 0.024 \text{ GeV}^2$ ). This result differs notably from our main result, as expected from the very different slope of [36]. This discrepancy exemplifies the importance of testing and improving  $\chi\text{EFT}$  calculations, since well-controlled predictions would make electroproduction data very competitive for verifying the GDH sum rule and other real photon observables.

**Data availability** Experimental data that support the findings of this study will be posted on the CLAS database, <https://clasweb.jlab.org/physicsdb/>, or are available from X. Zheng upon request.

**Code availability** The computer codes that support the plots within this paper and the findings of this study are available from X. Zheng upon request.

**Author contributions** The members of the Jefferson Lab CLAS Collaboration constructed and operated the experimental equipment used in this experiment. A large number of collaboration members participated in the data collection. The following authors provided various contributions to the experiment design and commissioning, data processing, data analysis and Monte Carlo simulations: M. Battaglieri, R. De Vita, V. Drozdov, L. El Fassi, H. Kang, K. Kovacs, E. Long, M. Osipenko, S. Phillips, K. Slifer. The authors who performed the final data analysis and Monte Carlo simulations were A. Deur, S.E. Kuhn, M. Ripani, J. Zhang, and X. Zheng. The manuscript was reviewed by the entire CLAS collaboration before publication, and all authors approved the final version of the manuscript.

**Competing interests** The authors declare no competing interests.

#### Additional information

Supplementary information are available online that includes all numerical results reported here.

Correspondence and requests for materials should be addressed to A. Deur.

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