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Persistence of the $Z = 28$ shell gap in $A = 75$ isobars: Identification of a possible $(1/2^-)$ μ s isomer in ^{75}Co and β decay to ^{75}Ni

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Background: The evolution of shell structure around doubly-magic exotic nuclei is of great interest in nuclear physics and astrophysics. In the ‘south-west’ region of ^{78}Ni , the development of deformation might trigger a major shift in our understanding of explosive nucleosynthesis. To this end, new spectroscopic information on key close-lying nuclei is very valuable.

Purpose: We intend to measure the isomeric and β decay of ^{75}Co , with one proton- and two neutron-holes relative to ^{78}Ni , to access new nuclear structure information in ^{75}Co and its β -decay daughters ^{75}Ni and ^{74}Ni .

Methods: The nucleus ^{75}Co is produced in relativistic in-flight fission reactions of ^{238}U at the Radioactive Ion Beam Factory (RIBF) in the RIKEN Nishina Center. Its isomeric and β decay are studied exploiting the BigRIPS and EURICA setups.

Results: We obtain partial β -decay spectra for ^{75}Ni and ^{74}Ni , and report a new isomeric transition in ^{75}Co . The energy ($E_\gamma = 1914(2)$ keV) and half-life ($t_{1/2} = 13(6)$ μ s) of the delayed γ ray lend support for the existence of a $J^\pi = (1/2^-)$ isomeric state at 1914(2) keV. A comparison with PFSDG-U shell-model calculations provides good account for the observed states in ^{75}Ni , but the first calculated $1/2^-$ level in ^{75}Co , a prolate $K = 1/2$ state, is predicted about 1 MeV below the observed $(1/2^-)$ level.

Conclusions: The spherical-like structure of the lowest-lying excited states in ^{75}Ni is proved. In the case of ^{75}Co , the results suggest that the dominance of the spherical configurations over the deformed ones might be stronger than expected below ^{78}Ni . Further experimental efforts to discern the nature of the $J^\pi = (1/2^-)$ isomer are necessary.

I. INTRODUCTION

The neutron-rich region approaching ^{78}Ni , with 28 protons and 50 neutrons, is in the spotlight of the most important radioactive-ion beam facilities as this nucleus is

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the most exotic doubly-magic one ever synthesized in laboratory [1–3]. As more access to new structural features is obtained in nearby nuclei [4–14], ^{78}Ni appears to be the last spherical system prior to the hitherto unattainable domain of the r -process reaction path [15]. Indeed, the most advanced theoretical calculations presently available [15, 16] predict the coexistence of spherical and prolate deformed shapes in ^{78}Ni , with a 0^+ deformed bandhead lying at about the same excitation energy of the first 2^+ state, or even below [15, 17]. Interestingly, first experimental fingerprints for the existence of the deformed configuration have recently been provided by R. Taniuchi *et al.* [17], who have proposed a 2.91-MeV deformed (2^+) candidate just above the spherical (2_1^+) state at 2.60 MeV. Just by adding a few neutrons or removing a few protons to the doubly-magic system, the prolate-deformed configuration is expected to drop below the spherical one and become yrast. Such an inversion, with ^{78}Ni as the doorway to the new ground-state deformation region, might have a substantial effect on the theoretical predictions on the location of the neutron drip line, as deformed systems are expected to be more tightly bound [18], hence making a difference to our understanding of the r -process nucleosynthesis pathways.

The robustness of the $Z = 28$ closed shell and the coexistence of deformed and spherical shapes in the neutron-rich $\nu g_{9/2}$ Ni isotopes have been a matter of debate in a number of recent experimental and theoretical works [11, 19–32]. Particularly important is the conservation of the seniority quantum number v —the number of protons or neutrons that are not coupled in pairs to $J = 0$ [33, 34]—as it is a good indicator of gap stability. In the Ni nuclei filling the $\nu g_{9/2}$ shell, the seniority is still a good quantum number for a sub-set of solvable eigenstates [35–40], although it is still unknown if the deformation-driving forces might induce mixing of seniorities in close-lying states with equal J [11, 41, 42].

The increase in collectivity as protons are removed from the $Z = 28$ closed-shell regime has also been deeply investigated [5, 6, 13, 26, 43–47], indicating the development of a new island of inversion around $N = 40$ that extends beyond the harmonic-oscillator shell. Theoretically, the shaping of deformation around $N = 40$ appears to be driven not only by the variation in the number of protons and neutrons as one moves away from stability, but by many particle-hole excitations across energy gaps eventually induced by the proton-neutron tensor component of the nuclear force [48–51], which causes a reduction of the $\pi f_{7/2} - \pi f_{5/2}$ spin-orbit splitting as neutrons occupy the $\nu g_{9/2}$ orbital [16, 22, 52].

The convergence of the $N = 40$ island of inversion with a newly predicted region of deformation around $N = 50$ has been theorized recently [15]. This phenomenon is comparable to the merging of the $N = 20$ and $N = 28$ closed shells, with similar underlying mechanisms driving the onset of deformation and the disappearance of the classical shell closures. Although the observation of a deformed candidate state in coexistence with the normal spherical shape in ^{78}Ni supports this prediction, more

experimental information on lighter $Z \leq 28$ nuclei towards $N = 50$ is needed to fully comprehend how the shell structure evolves in the neutron-rich region below $Z = 28$, and if there is a new $N = 50$ island of inversion in coalescence with the one at $N = 40$.

With these goals in mind, the isomeric and β decay of ^{75}Co , with one proton and two neutron holes relative to the ^{78}Ni doubly-magic core, were investigated following the in-flight fission of a relativistic ^{238}U beam on a thin natural Be target at the Radioactive Isotope Beam Factory (RIBF) at RIKEN, Japan. Despite being pinpointed as one of the $A \sim 78$ nuclei with a significant impact on r -process estimates [53], the only experimental information reported hitherto in the literature for ^{75}Co is limited to the half-life $t_{1/2}$ and an upper limit for the β -delayed one-neutron emission probability P_{1n} [2, 32, 54]. Here we provide an additional lower P_{1n} limit, of help to extend the experimental databases used by nuclear astrophysicists. For ^{75}Ni , four γ rays at 232 keV, 893 keV, 950(20) keV and 1100(20) keV have recently been reported [30, 32]. While the first two were observed following β decay of ^{75}Co in an in-flight fragmentation experiment in NSCL [32], the latter two were reported in an intermediate Coulomb excitation experiment carried out at RIKEN [30]. Of them, only the 232-keV and 950(20)-keV transitions have tentatively been placed in the level scheme of ^{75}Ni , decaying directly to the ground state from levels with proposed spins and parities $J^\pi = (7/2^+)$ and $(13/2^+)$, respectively. In both cases the J^π arises from the $\nu g_{9/2}^n$, $v = 3$ seniority configuration. In the present work, an extended experimental study with new γ -ray transitions in both ^{75}Co and ^{75}Ni is reported. The new spectroscopic information is compared to state-of-the-art large-scale shell-model calculations using the PFSDG-U interaction in the $pf - sdg$ valence space [15]. On the basis of the experimental and theoretical results presented here, we argue new spin assignments and discuss the evolution of the spherical and deformed configurations in the ‘south-west’ quadrant of ^{78}Ni .

II. EXPERIMENTAL DETAILS

The present data were obtained during the EURICA campaign at the RIBF, operated jointly by the RIKEN Nishina Center and the Center for Nuclear Study of the University of Tokyo. The primary beam of ^{238}U was delivered by the RIKEN accelerator complex, which consisted of a linear injector (RILAC2) and four ring cyclotrons (RRC-fRC-IRC-SRC). The beam energy was 345 MeV/nucleon, with an intensity of approximately 3×10^{10} pps. The nucleus ^{75}Co and other neutron-rich nuclides close to ^{78}Ni were produced by in-flight fission [55] on a 3-mm-thick foil of ^9Be . The secondary beam species of interest were separated in both the first and second stages of the BigRIPS magnetic spectrometer [56] using dipole magnets. The selected fission fragments were identified through the standard ΔE - $B\rho$ -TOF method in the second stage of BigRIPS. Beam-line detectors, as fast plas-

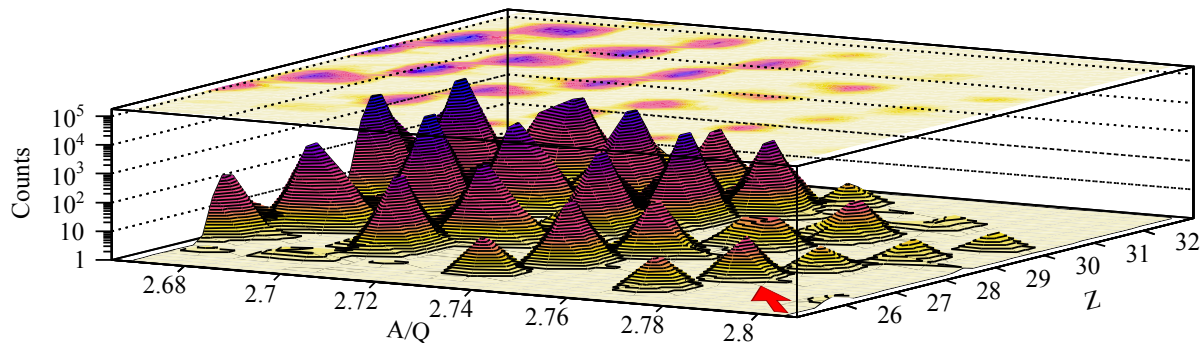


FIG. 1. (Color online) Three-dimensional cluster plot of the nuclei implanted in WAS3ABi as a function of their charge Z and mass-to-charge ratio A/Q . The cluster corresponding to ^{75}Co is indicated by the red arrow.

tic scintillators, parallel-plate avalanche counters and a multi-sampling ionization chamber, allowed for an event-by-event particle identification of the atomic number (Z) and the mass-to-charge ratio (A/Q) of the secondary-beam products.

The radioactive ion beam was conducted through the Zero-Degree Spectrometer (ZDS) [57] to the EURICA β -decay spectroscopy station, consisting of the active beam stopper WAS3ABi [58] and the γ -ray spectrometer EURICA [59]. Since the radioactive nuclei identified in BiGRIPS were very energetic, it was necessary to place a homogeneous aluminium degrader of variable thickness before WAS3ABi in order to adjust the range of the ions of interest within the implantation device.

The silicon array WAS3ABi was not only used to stop the radioactive nuclei but also to detect electrons and other charged particles emitted in their decay. WAS3ABi consisted of eight layers of 1-mm-thick double-sided silicon strips detectors (DSSSD) with an interspace of 0.5 mm between them. Each DSSSD had an active area of $60 \times 40 \text{ mm}^2$ and was segmented into sixty vertical strips and forty horizontal strips, providing a total of 2400 pixels of 1-mm pitch each. The WAS3ABi DAQ System recorded the pixel position, time and energy information of the implanted fission products and the emitted β electrons. Standard analog electronics were used to read the energy and time signals of each strip, optimized for the energy range of β particles. Meanwhile, in-flight fission fragments released around 1 GeV in the detector, overflowing the energy signals of the implantation strip and the neighbouring ones. The position of implantation was then defined by the X and Y strips with the fastest time signal [60].

The EURICA array was set up surrounding the active beam stopper and was used to record the energy and time of γ rays during a time window of up to 110 μs after the detection of an implantation or β electron. As a result, the setup was sensitive to isomeric lifetimes ranging from several ns to several hundred μs . EURICA was made of 84 high-purity germanium (HPGe) crystals, arranged in 12 clusters of 7 crystals packed closely at an average distance of 22 cm from WAS3ABi. An absolute detection

efficiency of approximately 11% at 662 keV was achieved after applying a standard add-back routine [25].

III. ANALYSIS PROCEDURE

In the off-line analysis, implantation-like events were defined by an overflow energy signal in at least one X and one Y strip of WAS3ABi. These were requested to be in coincidence with a high-energy signal in the last fast plastic scintillator of ZDS and in anticoincidence with any signal above threshold registered in a β detector placed behind WAS3ABi [61]. In this way, secondary reaction products generated during the implantation process were rejected to a large extent. The DSSSD of implantation was then identified as the last one in which an X and a Y strip were overflowed. On the other hand, electron-like events were defined by non-overflow energy signals above β threshold ($\sim 50 \text{ keV}$) in anticoincidence with the last plastic scintillator of ZDS. Since a β electron typically fired several strips before leaving WAS3ABi, the total energy released in each DSSSD was obtained from the sum of the energies of adjacent strips within a 8- μs time gate. The (x,y) position of the β -like particles was then computed as the energy-weighted average of the x and y strips.

Once defined, implantations and β particles were correlated in position and time. In the present analysis, the spatial correlations were restricted to the DSSSD of implantation, and the maximum transverse distance was fixed to $\rho = \sqrt{(x-X)^2 + (y-Y)^2} = \sqrt{2}$ pixels. The time window was set to $t = 135 \text{ ms}$, corresponding to about five half-life periods of ^{75}Co [2]. Additional prompt-time correlations with γ rays were defined to explore the structure of implanted and descendant nuclei. These were set according to the expected nuclear half-lives, and had maximal time windows of 800 ns for non-isomeric β -delayed transitions and $\sim 50 \mu\text{s}$ for isomeric transitions. For the study of coincident γ rays, a 400-ns time window was set. Background contributions from randomly correlated events were evaluated by applying the so-called backward-time technique [62], which ex-

exploits correlations between implantations and preceding β electrons using the same conditions as for the normal correlations.

IV. RESULTS

The distribution of implantations as a function of Z and A/Q is shown in the three-dimensional plot of Fig. 1, where the red arrow pointing to the ^{75}Co nuclei illustrates the good quality of the particle identification. In total, $\sim 1.8 \times 10^4$ ions of ^{75}Co were implanted in WAS3ABi.

IV.1. Isomeric spectroscopy

The two-dimensional energy-time matrix of the γ rays detected within approximately 50 μs after the detection of ^{75}Co implantation events is shown in the left panel of Fig. 2. A newly identified isomeric γ transition can be seen at about 1900 keV. Matrix projections on the Y (energy) and X (time) axis are shown in the right top and bottom panels of the figure, respectively. The resulting energy for the isomeric transition is $E_\gamma = 1914(2)$ keV. Due to the scarce statistics, the half-life has been obtained from an unbinned maximum likelihood fit to a single exponential time distribution function describing the decay time behaviour of the γ events registered after the so-called *prompt flash* (which is visible at Time ~ 0). No background contributions have been considered in the fitting procedure due to the absence of random γ events in the region of interest of the energy-time matrix. The resulting half-life is $t_{1/2}(^{75}\text{Co}^*) = 13(6)$ μs . The reported uncertainty is only statistical and has been evaluated using the RooFit package [63] with the MINOS method for determination of error parameters [64].

IV.2. β -decay spectroscopy

The singles β -delayed γ -ray energy spectrum resulting from the analysis procedure described in Sec. III is shown in Fig. 3. There, the most intense transitions attributed to the β (^{75}Ni) and β_n (^{74}Ni) descendants are labelled in bold and italics, respectively. As an example, the coincidence spectra gated on the 232-, 1045- and 738-keV γ transitions are shown in the three panels of Fig. 4. While the first two are attributed to ^{75}Ni , the third one is assigned to the β_n daughter ^{74}Ni . In all cases, the background contributions have been evaluated and subtracted as described in Ref. [65].

The full list of γ -ray transitions, absolute γ intensities, and γ - γ coincidence relations observed following β decay of ^{75}Co is provided in Tables I and II. While Table I shows transitions placed in the level scheme of any of the descendant nuclei (^{74}Ni , ^{75}Ni or ^{75}Cu), Table II shows the list of γ rays attributed to ^{75}Ni that have not been placed in the level scheme due to the absence of coincident transitions. It is to note here the large intensity of

the γ rays at 1061.8(11) keV and 2458.8(15) keV.

The partial level schemes corresponding to the internal and β decay of ^{75}Co are shown in Fig. 5. The ordering of the transitions following β decay is proposed according to γ -ray intensity balances, γ - γ prompt coincidence relations, and γ -ray energy sum matchings according to the information displayed in Table I. Apparent β feedings and/or $\log ft$ values are indicated at the left of the level schemes. These can be considered as upper and lower limits, respectively, due to the large Pandemonium effect [66] expected in odd-mass nuclei with large Q_β values, as is the case of ^{75}Co , with $Q_\beta = 14380(580)$ keV [67]. In the figure, arrow widths are proportional to transition intensities, and tentative spins and parities are shown in parentheses on the left of each level.

In the β -decay daughter ^{75}Ni , the levels at 1864.4(8) keV, 972.7(10) keV and 231.8(9) keV are established based on the observation of three independent γ -ray cascades to the ground state, namely the 891-973-, 891-741-232- and 1632-232-keV γ cascades, and the direct ground-state transition at 1865 keV. The location of the level at 1044.6(12) keV is based on two arguments: first, the 1045-keV transition is the most intense one in the set of coincident γ rays formed by the 1045-, 417-, 867-, and 1596-keV peaks (see Table I); and second, its energy could match with the tentative direct ground-state transition at 1100(20) keV reported by Ref. [30]. The placement of the states at 1461.1(17) keV, 1911.1(16) keV, and 3057(4) keV, built up on top of the 1044.6(12)-keV level, is proposed according to the coincidence relationships indicated in Table I. Meanwhile, the observed states at 1024.6(12), 1762.9(16), 2104.2(16), 2380(3) and 2606(3) keV in the β_n daughter ^{74}Ni were previously reported in the direct decay of ^{74}Co to ^{74}Ni [11] and identified as following the $(8_1^+) \rightarrow (6_1^+) \rightarrow (4_1^+) \rightarrow (2_1^+) \rightarrow 0^+$ and $(4_2^+) \rightarrow (2_1^+) \rightarrow 0^+$ γ cascades connecting states of seniorities $v = 2$ and $v = 4$. A lower limit for the β -delayed one-neutron emission probability of ^{75}Co has been deduced from the absolute intensity of the $(2_1^+) \rightarrow 0^+$ transition at 1025 keV, resulting in $P_{1n} \geq 14\%$. This value is in good agreement with the upper limit reported in the literature, $P_{1n} \leq 16\%$ [54], and points to a rather low ground-state feeding in the β_n decay $^{75}\text{Co} \rightarrow ^{74}\text{Ni}$.

V. DISCUSSION

V.1. β decay of ^{75}Co to ^{75}Ni

The ground state of the odd-even parent nucleus ^{75}Co is proposed to have a tentative $J^\pi = (7/2^-)$ based on the lowest-lying $\pi f_{7/2}^{-1}$ proton-hole configuration. This assignment is in accordance with the tentative $J^\pi = (7/2^-)$ attributed to the ground states of the lighter odd-even isotopes ^{71}Co and ^{73}Co [13, 21]. Meanwhile, the main contribution to the ground-state wave function in the daughter nucleus ^{75}Ni is expected from the unpaired neutron in the $\nu g_{9/2}$ shell, resulting in $J^\pi = (9/2^+)$. The assignment is supported by the systematics of lighter

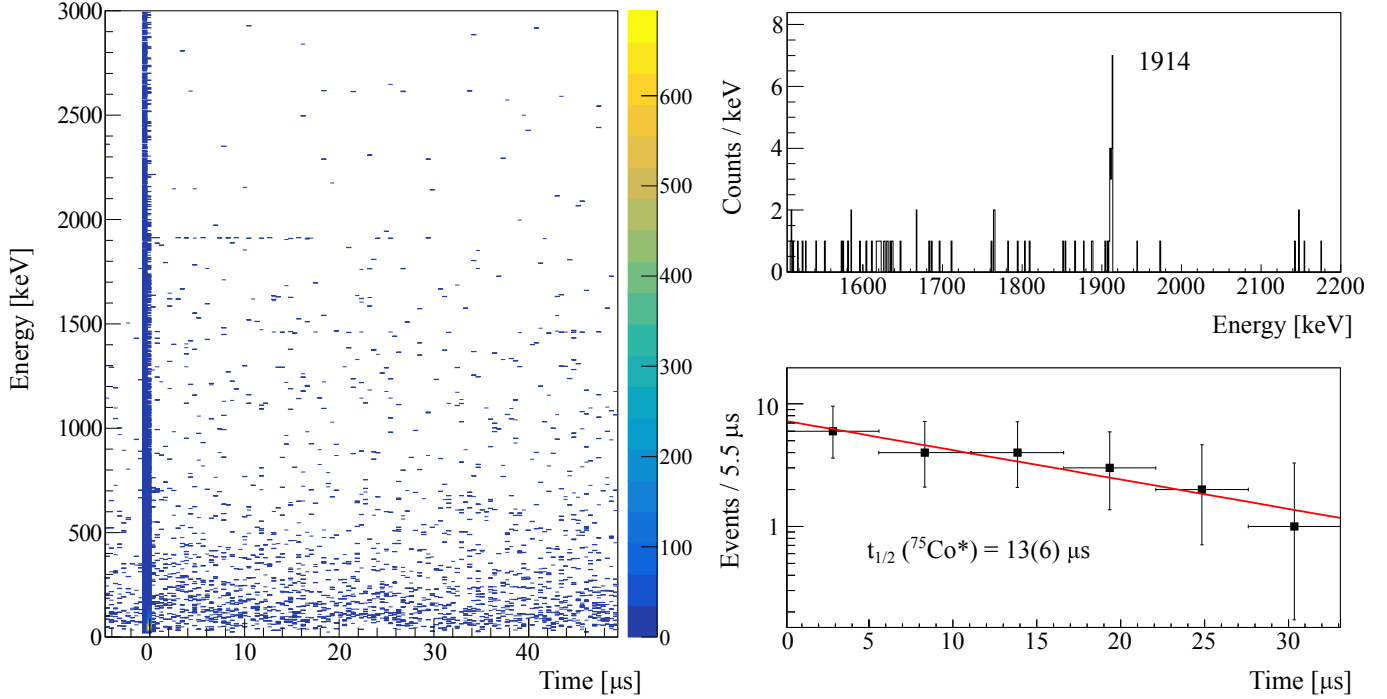


FIG. 2. (Color online) Left: Energy-time matrix showing the γ rays detected after implantation of ^{75}Co . The γ -ray energy is plotted against the ion- γ time difference. Top right: Projection of the matrix on the energy axis for a time window of ~ 35 μs . Bottom right: Projection of the matrix on the time axis, gated on the 1914-keV transition. The fit to a single exponential function is shown in red. The resulting half-life is indicated. See text for details.

TABLE I. List of γ transitions observed in the β decay of ^{75}Co and placed in the level scheme of any of the descendant nuclei. The nuclei to which the transitions are assigned, the γ -ray energies, the excitation energies of initial and final states, and the absolute γ -ray intensities are given. As well, the $\beta(\gamma\gamma)$ coincidence relations are indicated for each transition. An asterisk highlights those cases where the coincidence relation is established by only one observed count. For ^{75}Cu , the excited states are taken from Ref. [31].

Nucleus	E_γ (keV)	E_x^i (keV)	E_x^f (keV)	I_γ (%)	Coincident γ rays
^{75}Ni	231.8(9)	231.8(9)	0	41(4)	740.8, 891.4, 1632.2
^{75}Ni	416.6(13)	1461.1(17)	1044.6(12)	5.8(10)	1044.6, 1596
^{75}Ni	740.8(11)	972.7(10)	231.8(9)	7.1(13)	231.8, 891.4
^{75}Ni	866.5(12)	1911.1(16)	1044.6(12)	2.6(8)	1044.6
^{75}Ni	891.4(11)	1864.4(8)	972.7(10)	12.5(18)	231.8, 740.8, 972.9
^{75}Ni	972.9(12)	972.7(10)	0	7.0(13)	891.4
^{75}Ni	1044.6(12)	1044.6(12)	0	8.1(15)	416.6, 866.5, 1596*
^{75}Ni	1596(3)	3057(4)	1461.1(17)	4.0(11)	416.6, 1044.6*
^{75}Ni	1632.2(13)	1864.4(8)	231.8(9)	24(3)	231.8
^{75}Ni	1865.1(11)	1864.4(8)	0	1.6(7)	—
^{74}Ni	226.0(5)	2606(3)	2380(3)	1.5(5)	616.7, 738.3, 1024.6*
^{74}Ni	616.7(14)	2380(3)	1762.9(16)	5.8(11)	226.0, 738.3, 1024.6
^{74}Ni	738.3(11)	1762.9(16)	1024.6(12)	7.9(13)	226.0, 616.7, 1024.6
^{74}Ni	1024.6(12)	1024.6(12)	0	17(3)	226.0*, 616.7, 738.3, 1079.7
^{74}Ni	1079.7(11)	2104.2(16)	1024.6(12)	3.2(9)	1024.6
^{75}Cu	504.8(14)	1989	1483.5	2.3(7)	—
^{75}Cu	883.0(6)	949.7	66.2	1.8(7)	—
^{75}Cu	992.2(9)	992.2	0	3.4(9)	—
^{75}Cu	1483.0(15)	1483.5	0	2.4(8)	—

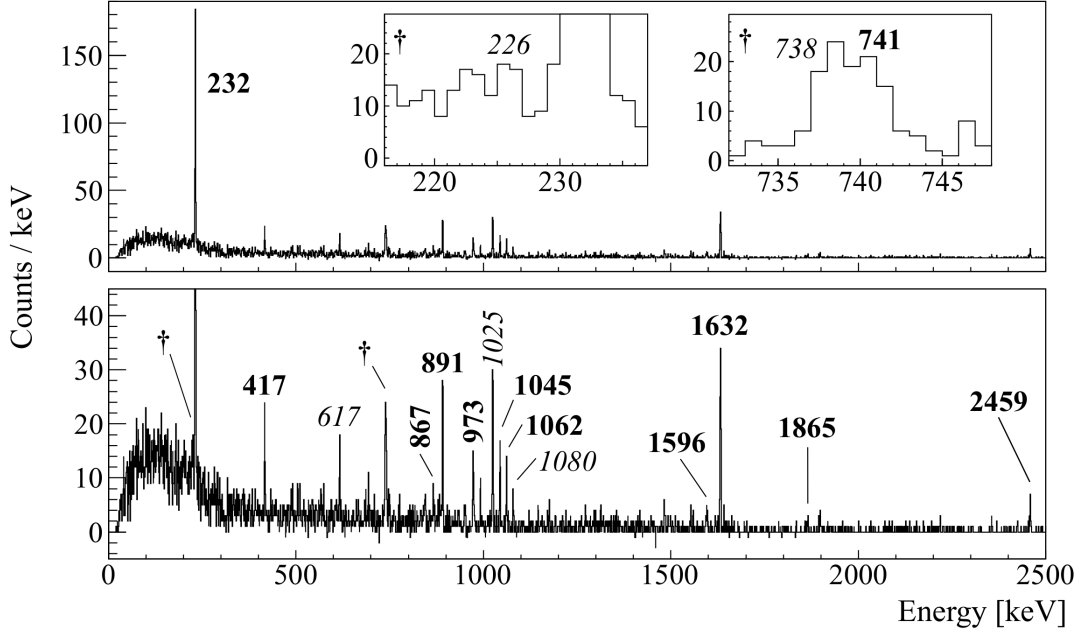


FIG. 3. The β -delayed γ -ray energy spectrum following implantations of ^{75}Co nuclei during a time interval of 135 ms. The panels present two different ranges of the Y axis to facilitate the observation of weak γ rays. The transitions assigned to the β (^{75}Ni) and β_n (^{74}Ni) daughters are marked in bold and italics, respectively. Expanded inset spectra are shown for the γ rays marked with a dagger.

TABLE II. Transitions attributed to ^{75}Ni that have not been placed in the level scheme. The γ -ray energies and the absolute γ -ray intensities are shown.

Nucleus	E_γ (keV)	I_γ (%)
^{75}Ni	491.3(10)	1.5(5)
^{75}Ni	566.8(11)	1.5(5)
^{75}Ni	686.1(6)	1.4(5)
^{75}Ni	1061.8(11)	5.8(12)
^{75}Ni	1145.8(10)	1.8(7)
^{75}Ni	1175.9(13)	2.2(8)
^{75}Ni	1313.1(8)	1.4(6)
^{75}Ni	1559.2(8)	1.5(7)
^{75}Ni	2219.0(10)	2.0(8)
^{75}Ni	2355.7(13)	1.6(8)
^{75}Ni	2458.8(15)	5.6(15)

neutron-rich $\nu g_{9/2}$ even-odd Ni isotopes [21, 68] and the recent experimental studies of ^{75}Ni [30, 32].

The strong β feeding to the excited state at 1864.4(8) keV, $I_\beta = 38(7)\%$, and the corresponding $\log ft = 4.7(2)$, provide a robust proof for the occurrence of an allowed Gamow-Teller (GT) decay from the $\pi f_{7/2}^{-1}$ ground state of ^{75}Co . As the most energetic single-particle GT transition occurring in the region of ^{78}Ni transforms a neutron in the $\nu f_{5/2}$ orbital into a proton in the $\pi f_{7/2}$ shell, the associated wave function in the final 1864.4(8)-keV state of ^{75}Ni is expected to have a large $\nu f_{5/2}^{-1}$ contribution, resulting in a tentative spin and parity $J^\pi = (5/2^-)$. Further support for the $(5/2^-)$ assignment to this level

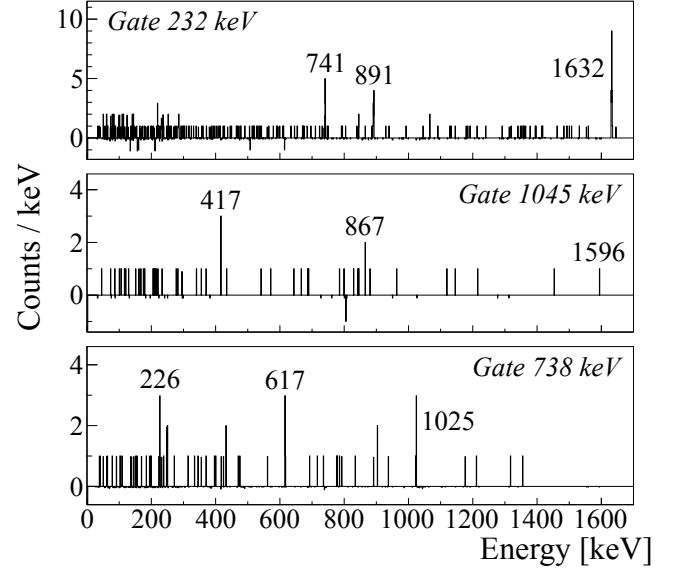


FIG. 4. From top to bottom, $\beta(\gamma\gamma)$ coincidence spectra gated on the 232-, 1045- and 738-keV γ transitions.

comes from the systematic comparison with the β -decay level schemes of lighter even-odd Ni isotopes (see Refs. [21, 68]), which shows that the most probable β -decay transition populates the yrast $(5/2^-)$ level. For the first excited state at 231.8(9) keV in ^{75}Ni , we propose $J^\pi = (7/2^+)$ despite an observed (apparent) feeding of $I_\beta < 15\%$. Our assignment is in agreement with the work of S. Go *et al.* [32] and it is equally based on the similarity

TABLE III. Possible spins and parities J^π for the state at 972.7(10) keV, multiplicities of the de-exciting transitions at 740.8(11) keV and 972.9(12) keV, and corresponding fractions of branching ratios R derived from single-particle estimates. See text for details.

J^π	Mult [$E_x(973 \text{ keV})$]	Mult [$E_\gamma(741 \text{ keV})$]	Mult [$E_\gamma(973 \text{ keV})$]	$R \left[\frac{BR_\gamma^{973}}{BR_\gamma^{741}} \right]$
$5/2^+$	M1	E2	E2	1.60×10^{-3}
$5/2^+$	E2	E2	E2	3.90
$3/2^+$	E2	M3	M3	5.00×10^{-7}

with the excitation energies of the $(7/2_1^+)$ states in ^{71}Ni (281 keV) and ^{73}Ni (239 keV) [21]. This spin and parity arises mainly from the coupling of the first (2^+) state in the ^{74}Ni core to the unpaired $\nu g_{9/2}$ neutron, and has as main configuration $\nu g_{9/2}^n$. Apart from this level, additional states with $J^\pi = 3/2^+, 5/2^+, 9/2^+, 11/2^+$ and $13/2^+$ are expected to arise from the $2_1^+(^{74}\text{Ni}) \otimes \nu g_{9/2}$ multiplet. Excepting the $9/2_2^+$ level, the rest of states are expected to lie at excitation energies between 1 MeV and 1.5 MeV (see Refs. [21, 30] and discussion in Sec. V.3).

The almost non-existing β feedings to the levels at 972.7(10) keV and 1044.6(12) keV indicate that, more likely, they are not fed by allowed GT transitions but through internal γ feeding, in accordance with the forbiddenness of a $\nu g_{9/2} \rightarrow \pi f_{7/2}$ single-particle β transition. In the case of the 972.7(10)-keV level, the energy matches well with that of the reported $(13/2^+)$ state at 950(20) keV [30]. The observation of a prompt, strong γ ray at 891 keV connecting the $(5/2^-)$ level with this state, though, rules out spin assignments higher than $9/2$. Hence, the only positive-parity states remaining at this excitation energy are $J^\pi = 3/2^+$ and $5/2^+$. In Table III, the quotients R between the single-particle branching ratios of the γ rays decaying from the 972.7(10)-keV level are shown for each possible J^π and transition multipolarities. The γ -ray energies are 740.8(11) keV and 972.9(12) keV. A comparison with the experimental quotient of these two transitions, $R = 1.0(3)$, provides support for a $J^\pi = (5/2^+)$ assignment. Based on this, and given the previous spectroscopic information from relativistic Coulomb excitation (see Fig. 2 of Ref. [30]), we propose the 1044.6(12)-keV state to have $J^\pi = (11/2^+)$ or $(13/2^+)$.

The strong 1061.8(11)-keV transition observed in the singles γ spectrum of Fig. 3 could very likely de-excite the remaining $J^\pi = (13/2^+)$ or $(11/2^+)$ level to the ground state, as its energy matches well with those of the calculated $11/2^+$ and $13/2^+$ levels (see Fig. 5). However, as such high positive-parity spin states cannot be directly fed from β decay of the ^{75}Co ground state and no coincident γ rays have been observed for the 1061.8(11)-keV transition, it has not been placed in the level scheme shown in Fig. 5.

V.2. Isomeric decay of ^{75}Co

The absence of transitions in coincidence with the delayed 1914(2)-keV γ ray in ^{75}Co (see Fig. 2) leads to two possible scenarios. In the first one, the observed transition directly connects an isomeric level at 1914(2) keV with the $J^\pi = (7/2^-)$ ground state. In such a case, the measured γ -ray lifetime, $t_{1/2}^\gamma = 13(6) \mu\text{s}$, suggests an M3 or E4 character. An E4 nature can be rejected on the basis that the spin and parity of the initial level then would have to be $J^\pi = 15/2^-$, and a faster decay path would be opened through the $11/2^-$ state that is expected at about 1 MeV from systematics of the lighter $\nu g_{9/2}$ odd-even Co isotopes [13] and the present PFSDG-U calculations (see Fig. 5). For an M3 character, $J^\pi = 13/2^-$ or $1/2^-$ are possible. Of these, $J^\pi = 13/2^-$ would as well find a faster decay path through the $11/2^-$ or $9/2^-$ levels. Hence, only a $J^\pi = 1/2^-$ state could result in an isomeric decay to the $J^\pi = (7/2^-)$ ground state. This is the first option presented for the experimental level scheme of ^{75}Co in Fig. 5.

In the second scenario, the delayed transition may remain unobserved if it is of low energy and has a high conversion coefficient. With the current setup, this is more likely to happen for E2 or M2 transitions with energies below 50-60 keV [69, 70]. Then, the observed 1914(2)-keV γ ray would follow in the subsequent decay to the ground state. Taking a look at the lowest-lying states expected by the PFSDG-U calculations, a possible decay sequence would be $(1/2^-) \rightarrow (5/2^\pm) \rightarrow (7/2_{g.s.}^-)$. This option is also indicated in Fig. 5. In this latter case, given the large energy difference of the E2/M2 and M3 transitions, one would expect to detect as well a competing M3 branch to the ground state. Therefore, the non-observation of two close-lying γ rays in Fig. 2 lends support for the first interpretation, i.e., that the 1914(2)-keV transition more likely connects the first $(1/2^-)$ level with the $(7/2^-)$ ground state.

V.3. Comparison with shell-model calculations

The theoretical level schemes shown in Fig. 5 for ^{75}Co and ^{75}Ni have been obtained with shell-model (SM) calculations using the PFSDG-U interaction in a valence space consisting of the full pf shell for protons and the full $pf - sdg$ shell for neutrons [15]. In general, we find a good agreement between the observed yrast levels in ^{75}Ni and their calculated counterparts, with an accuracy below 200 keV. Regarding the 972.7(10)-keV level, the calculations also support a $(5/2_1^+)$ assignment from comparison of the reduced transition strengths, with $B(E2; 5/2_1^+ \rightarrow 9/2_1^+) \approx 48 e^2 fm^4$, $B(E2; 5/2_1^+ \rightarrow 7/2_1^+) \approx 1.4 e^2 fm^4$, and $B(M1; 5/2_1^+ \rightarrow 7/2_1^+) \approx 0,007 \mu_N^2$. In the present calculations, the neutron and proton effective charges used for the electric quadrupole operator $E2$ are $\varepsilon_n = 0.46$ and $\varepsilon_p = 1.31$ [15]. With them, the $B(E2)$ strengths predicted for the $11/2^+$ and $13/2^+$ levels are $B(E2; 11/2_1^+ \rightarrow 9/2_1^+) \approx 42 e^2 fm^4$ and

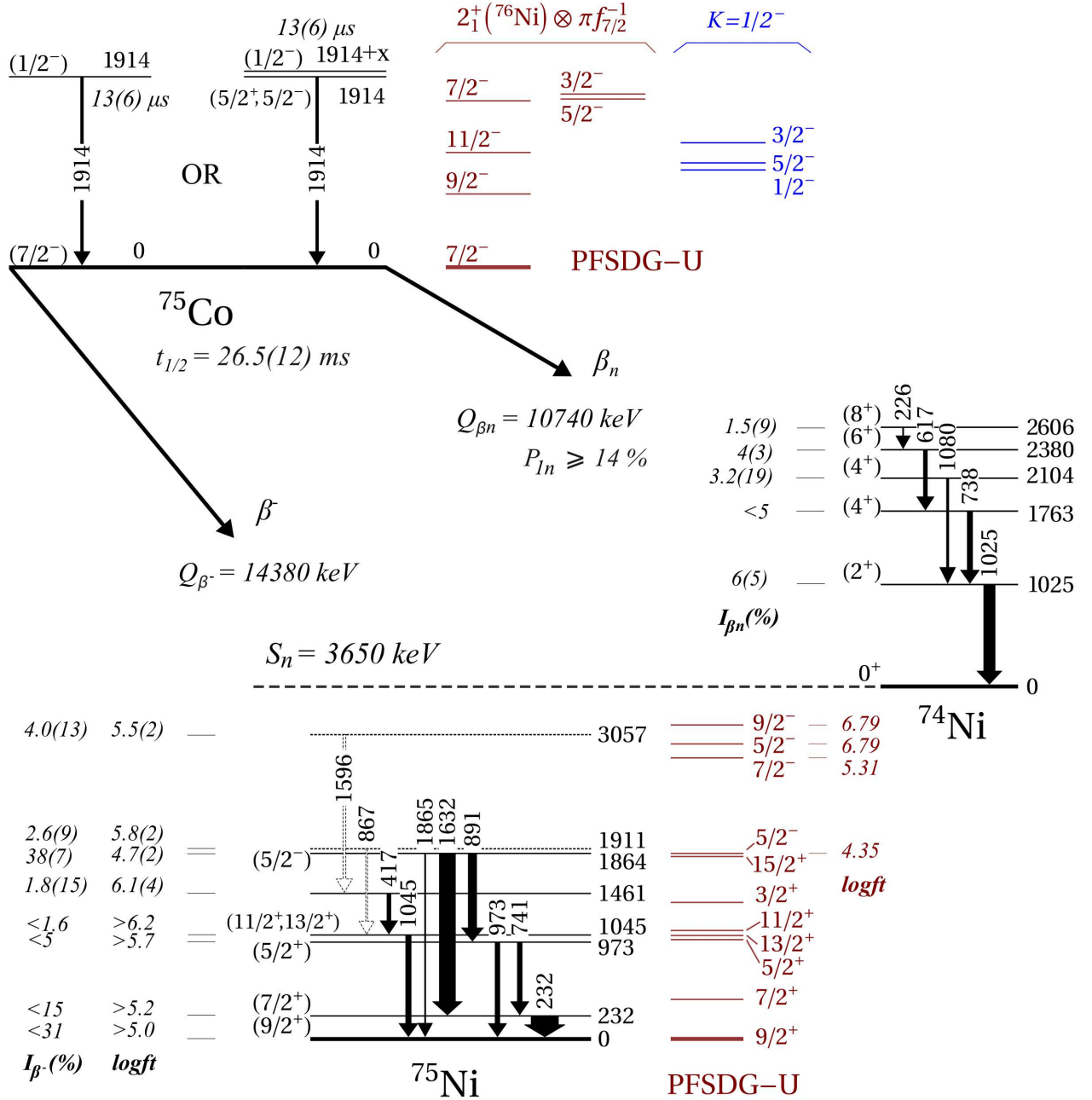


FIG. 5. (Color online) Experimental isomeric and β -decay level schemes of ^{75}Co extracted from the present work. The energies of the levels are given in keV. The thickness of the arrows is proportional to the intensities of the transitions connecting the states. Spins and parities, apparent β feedings and $\log ft$ values are indicated at the sides of the levels. Theoretical states, obtained with the PFSDG-U interaction [15], are indicated in red and blue. See text for details.

$B(E2; 13/2_1^+ \rightarrow 9/2_1^+) \approx 56 e^2 fm^4$, respectively. These theoretical results are of the same order than the calculations named as SM1 and SM2 in Ref. [30], and are similarly systematically lower than the recently measured $B(E2)$ values in ^{75}Ni [30]. The origin of the discrepancies might be due to the population of the $(5/2^+)$ state in the intermediate Coulomb excitation experiment, as it lies close in energy to the $(13/2^+)$ and $(11/2^+)$ candidates proposed by Ref. [30].

In addition, we have computed theoretical $B(GT)$ strengths for the $^{75}\text{Co} \rightarrow ^{75}\text{Ni}$ decay with the PFSDG-U SM calculations. The corresponding $\log ft$ values for the lowest allowed states are given in Fig. 5. These have been obtained assuming a standard quenching factor of 0.7. It is to note that the valence space employed is only suited for levels of natural (positive) parity; hence, the negative-parity states derived from the $B(GT)$ computation have been placed by assuming that the excitation

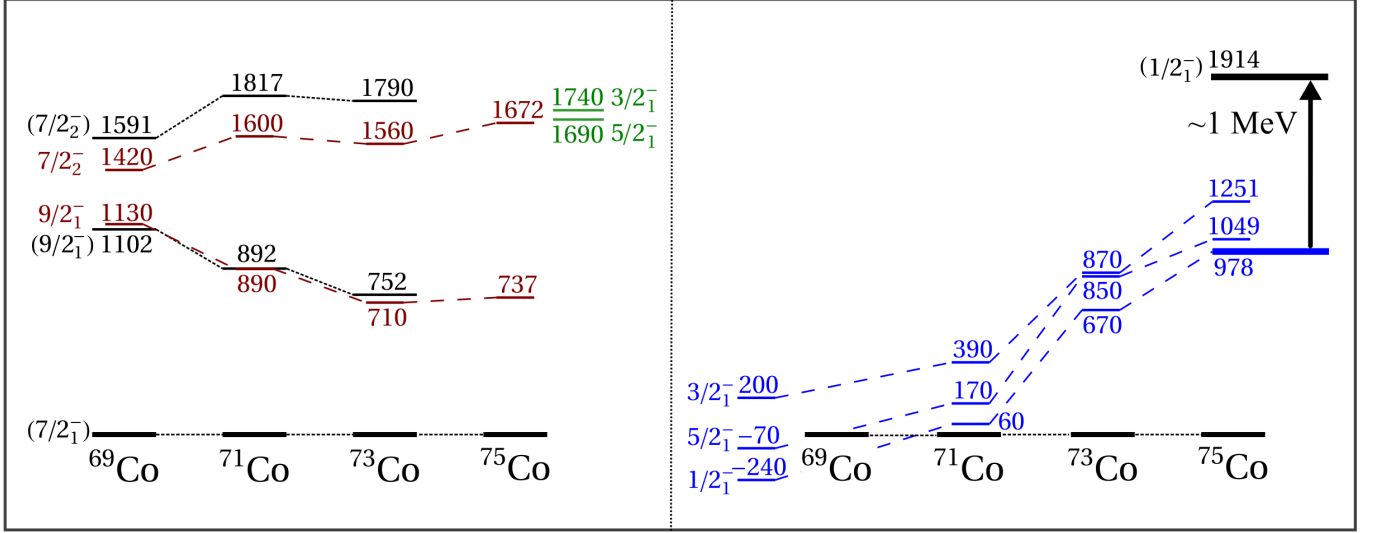


FIG. 6. (Color online) Evolution of the experimental (black) and theoretical (red, blue and green) levels discussed here for the odd-mass Co isotopes occupying the $\nu g_{9/2}$ shell. The experimental states are taken from [13] and the present work and the theoretical ones are calculated with the PFSDG-U interaction [15]. The left panel shows states with a spherical-like structure and the right one shows the lowest-lying levels of the $K = 1/2$ deformed band.

energy of the lowest-lying theoretical $5/2^-$ state lies at the excitation energy of the experimental ($5/2^-$) level, 1864 keV. This assumption is supported by the excellent agreement with the SM prediction discussed in Ref. [21], which places a $5/2^-$ level with a strong $\sim 76\%$ contribution from the $\nu f_{5/2}^{-1}$ configuration at 1821 keV.

The calculations clearly indicate an abundance of population to the lowest-lying $5/2^-$ state, with $B(GT) = 0.275$ and $\log ft = 4.35$, in good agreement with the experimentally measured $\log ft = 4.7(2)$. The slightly higher experimental value might be ascribed to the non-observation of some internal γ de-excitation branches from the ($5/2^-$) state, as the one connecting with the missing $1/2_1^-$ level. This state is expected to have $\nu p_{1/2}^{-1}$ as main configuration and is predicted about 700 keV below by the LNPS [71] and the calculations of Ref. [21]. Unfortunately, it has not been identified in the present work, maybe due to the limited statistics available or maybe because it lies at a higher excitation energy, as discussed for its lighter neighbour ^{73}Ni [21].

Theoretically, three negative-parity states are fed by β decay at excitation energies around that of the experimental 3057-keV level. Their energies and J^π values are 2.82 MeV and $7/2^-$, 2.96 MeV and $5/2^-$, and 3.15 MeV and $9/2^-$. The most strongly populated one is the $7/2^-$ level, with $B(GT) = 0.013$ and $\log ft = 5.31$. This result is close to the experimental $\log ft$ value of the 3057-keV state, $\log ft = 5.5(2)$. However, as the location of this level is only tentative due to the scarcity of $\beta(\gamma\gamma)$ coincidences, no spin-parity assignment is proposed. Another argument is that, given the large Pandemonium effect influencing our data, none of the other two spins and parities can be discarded, even if the calculated $\log ft$ values increase up to $\log ft = 6.79$.

In the case of ^{75}Co , the PFSDG-U levels are an extension of the SM calculations reported in Ref. [13] for ^{69}Co , ^{71}Co and ^{73}Co using the LNPS [71] and PFSDG-U [15] interactions. Similar to its lighter neighbours, two well defined structures associated to spherical (red and green) and deformed (blue) shapes are distinguished at low excitation energies in Fig. 6. The first is related to the coupling of the $\pi f_{7/2}^{-1}$ proton hole to the first 2^+ state in the ^{76}Ni core, which produces a multiplet of states with $J^\pi = 3/2^- - 11/2^-$. The second is attributed to proton and neutron excitations across the shell gaps $Z = 28$ and $N = 50$, and results in the development of a deformed $K = 1/2$ band with intrinsic quadrupole moment $Q_0 \approx 140 \text{ efm}^2$. On average, the deformed states have 1.5 protons and 1.5 neutrons above the closed shells, compared with 0.5 protons and 0.5 neutrons for the spherical-like states. According to the calculations, the E2 transitions within the deformed band are expected to be much stronger than those between the spherical-like states and between the two structures.

In the previous work of T. Lokotko *et al.* [13] on the lighter $\nu g_{9/2}$ odd-mass ^ACo isotopes, excited spherical ($7/2_2^-$) and ($9/2_1^-$) levels arising from the coupling of the $\pi f_{7/2}^{-1}$ hole to the (2_1^+) state in their ^{A+1}Ni cores were identified. These are shown in black on the left panel of Fig. 6, together with the $7/2_2^-$ and $9/2_1^-$ levels calculated with the PFSDG-U interaction, depicted in red for the sake of clarity. At first sight, one can notice that the excitation energies of the ($7/2_2^-$) candidates are on average 200 keV higher than their theoretical counterparts. If this systematic behaviour is extended to the lower-spin members of the multiplet in ^{75}Co (shown in green in the figure), their excitation energies could very likely be degenerated with or slightly above the observed

isomeric state. This shift would lend support to the interpretation of the isomer provided in Sec. V.2, which supports an M3 assignment for the 1914(2)-keV transition, corresponding to a $(1/2^-) \rightarrow (7/2^-)$ ground-state decay.

Theoretically, the lowest-lying spherical-like state with $J^\pi = 1/2^-$ arises from the particle-core coupling configuration $\pi f_{7/2}^{-1} \otimes 4_1^+ (^{76}\text{Ni})$. The question comes naturally: Is this state expected at about the excitation energy of the observed isomer? Considering that the experimental (4_1^+) level in ^{76}Ni lies at 1920 keV [4] and that the energies of the members of the $\pi f_{7/2}^{-1} \otimes 4_1^+ (^{76}\text{Ni})$ multiplet are expected to increase at decreasing spin [72, 73], one can presume that the first spherical-like $1/2^-$ state lies above. Moreover, if the observed $(1/2^-)$ level is spherical, there still remains the question of why the deformed $1/2^-$ state, predicted at a much lower excitation energy, has not been observed as well. At least in the present data set, there are no indications for the existence of a low-spin β -decaying isomer. Based on these arguments, we presume that the $(1/2^-)$ isomeric level reported here for ^{75}Co is more likely the bandhead of the deformed $K = 1/2$ configuration. As it is shown in the right panel of Fig. 6, the deformed band would then be shifted up by ~ 1 MeV with respect to the predictions of the PFSDG-U calculations. This tentative conclusion implies that the increasing trend of the prolate-deformed band towards $N = 50$ in the odd-mass Co isotopes may be far more abrupt than expected by the PFSDG-U calculations (see the right panel of Fig. 6). This, in turn, points to a stronger dominance of spherical-like shapes at low excitation energies in the region immediately beneath ^{78}Ni , posing the question of how fast deformation develops in the $N = 50$ shell below ^{78}Ni . Further spectroscopic data on this and more exotic $N \leq 50$ nuclei will be necessary to provide answer to this question.

VI. SUMMARY AND CONCLUSIONS

The isomeric and β decays of ^{75}Co have been investigated at the RIBF facility at RIKEN (Japan) using the BigRIPS and EURICA setups. First spectroscopic infor-

mation is provided for ^{75}Co , for which a new isomeric transition at 1914(2) keV with a half-life of $t_{1/2} = 13(6)$ μs is reported. For the β -decay daughter ^{75}Ni , new levels extending beyond those recently reported in Refs. [30, 32] are provided. In the case of the β_n daughter ^{74}Ni , the population of the (8_1^+) candidate points to a similar feeding pattern as in the decay $^{73}\text{Co} \rightarrow ^{72}\text{Ni}$.

The nature of the observed states in ^{75}Co and ^{75}Ni has been discussed in terms of large-scale shell-model calculations using the PFSDG-U interaction in the $pf - sdg$ model space [15]. In general, a good agreement between experimental and calculated results is found in ^{75}Ni for excitation energies and $\log ft$ values. In the case of ^{75}Co , the observed isomeric state is proposed to have $J^\pi = (1/2^-)$, although a comparison with the PFSDG-U predictions reveals a 1-MeV discrepancy with the expected excitation energy of the prolate-deformed $J^\pi = 1/2^-$ bandhead, leaving no clear interpretation for the nature of the observed state.

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- [1] P. T. Hosmer, H. Schatz, A. Aprahamian, O. Arndt, R. R. C. Clement, A. Estrade, K.-L. Kratz, S. N. Liddick, P. F. Mantica, W. F. Mueller, F. Montes, A. C. Morton, M. Ouellette, E. Pellegrini, B. Pfeiffer, P. Reeder, P. Santi, M. Steiner, A. Stolz, B. E. Tomlin, W. B. Walters, and A. Wöhr, *Phys. Rev. Lett.* **94**, 112501 (2005).
 - [2] Z. Y. Xu, S. Nishimura, G. Lorusso, F. Browne, P. Doornenbal, G. Gey, H.-S. Jung, Z. Li, M. Niikura, P.-A. Söderström, T. Sumikama, J. Taprogge, Z. Vajta, H. Watanabe, J. Wu, A. Yagi, K. Yoshinaga, H. Baba, S. Franchoo, T. Isobe, P. R. John, I. Kojouharov, S. Kubono, N. Kurz, I. Matea, K. Matsui, D. Mengoni, P. Morfouace, D. R. Napoli, F. Naqvi, H. Nishibata, A. Odahara, E. Şahin, H. Sakurai, H. Schaffner, I. G. Stefan, D. Suzuki, R. Taniuchi, and V. Werner, *Phys. Rev. Lett.* **113**, 032505 (2014).
 - [3] T. Sumikama, S. Nishimura, H. Baba, F. Browne, P. Doornenbal, N. Fukuda, S. Franchoo, G. Gey, N. Inabe, T. Isobe, P. R. John, H. S. Jung, D. Kameda, T. Kubo, Z. Li, G. Lorusso, I. Matea, K. Matsui, P. Morfouace, D. Mengoni, D. R. Napoli, M. Niikura, H. Nishibata, A. Odahara, E. Sahin, H. Sakurai, P.-

- A. Söderström, G. I. Stefan, D. Suzuki, H. Suzuki, H. Takeda, R. Taniuchi, J. Taprogge, Z. Vajta, H. Watanabe, V. Werner, J. Wu, Z. Y. Xu, A. Yagi, and K. Yoshinaga, *Phys. Rev. C* **95**, 051601 (2017).
- [4] P.-A. Söderström, S. Nishimura, Z. Y. Xu, K. Sieja, V. Werner, P. Doornenbal, G. Lorusso, F. Browne, G. Gey, H. S. Jung, T. Sumikama, J. Taprogge, Z. Vajta, H. Watanabe, J. Wu, H. Baba, Z. Dombradi, S. Franchoo, T. Isobe, P. R. John, Y.-K. Kim, I. Kojouharov, N. Kurz, Y. K. Kwon, Z. Li, I. Matea, K. Matsui, G. Martínez-Pinedo, D. Mengoni, P. Morfouace, D. R. Napoli, M. Niikura, H. Nishibata, A. Odahara, K. Ogawa, N. Pietralla, E. Şahin, H. Sakurai, H. Schaffner, D. Sohler, I. G. Stefan, D. Suzuki, R. Taniuchi, A. Yagi, and K. Yoshinaga, *Phys. Rev. C* **92**, 051305 (2015).
- [5] G. Benzoni, A. Morales, H. Watanabe, S. Nishimura, L. Coraggio, N. Itaco, A. Gargano, F. Browne, R. Daido, P. Doornenbal, Y. Fang, G. Lorusso, Z. Patel, S. Rice, L. Sinclair, P.-A. Söderström, T. Sumikama, J. Wu, Z. Xu, R. Yokoyama, H. Baba, R. Avigo, F. Bello Garrote, N. Blasi, A. Bracco, F. Camera, S. Ceruti, F. Crespi, G. de Angelis, M.-C. Delattre, Z. Dombradi, A. Gottardo, T. Isobe, I. Kuti, K. Matsui, B. Melon, D. Mengoni, T. Miyazaki, V. Modamio-Hoybjor, S. Momiyama, D. Napoli, M. Niikura, R. Orlandi, H. Sakurai, E. Sahin, D. Sohler, R. Taniuchi, J. Taprogge, Z. Vajta, J. Valiente-Dobón, O. Wieland, and M. Yalcinkaya, *Phys. Lett.* **751**, 107 (2015).
- [6] C. Santamaria, C. Louchart, A. Obertelli, V. Werner, P. Doornenbal, F. Nowacki, G. Authelet, H. Baba, D. Calvet, F. Château, A. Corsi, A. Delbart, J.-M. Gheller, A. Gillibert, T. Isobe, V. Lapoux, M. Matsushita, S. Momiyama, T. Motobayashi, M. Niikura, H. Otsu, C. Péron, A. Peyaud, E. C. Pollacco, J.-Y. Roussé, H. Sakurai, M. Sasano, Y. Shiga, S. Takeuchi, R. Taniuchi, T. Uesaka, H. Wang, K. Yoneda, F. Browne, L. X. Chung, Z. Dombradi, S. Franchoo, F. Giacoppo, A. Gottardo, K. Hadynska-Klek, Z. Korkulu, S. Koyama, Y. Kubota, J. Lee, M. Lettmann, R. Lozeva, K. Matsui, T. Miyazaki, S. Nishimura, L. Olivier, S. Ota, Z. Patel, N. Pietralla, E. Sahin, C. Shand, P.-A. Söderström, I. Stefan, D. Steppenbeck, T. Sumikama, D. Suzuki, Z. Vajta, J. Wu, and Z. Xu, *Phys. Rev. Lett.* **115**, 192501 (2015).
- [7] L. Olivier, S. Franchoo, M. Niikura, Z. Vajta, D. Sohler, P. Doornenbal, A. Obertelli, Y. Tsunoda, T. Otsuka, G. Authelet, H. Baba, D. Calvet, F. Château, A. Corsi, A. Delbart, J.-M. Gheller, A. Gillibert, T. Isobe, V. Lapoux, M. Matsushita, S. Momiyama, T. Motobayashi, H. Otsu, C. Péron, A. Peyaud, E. C. Pollacco, J.-Y. Roussé, H. Sakurai, C. Santamaria, M. Sasano, Y. Shiga, S. Takeuchi, R. Taniuchi, T. Uesaka, H. Wang, K. Yoneda, F. Browne, L. X. Chung, Z. Dombradi, F. Flavigny, F. Giacoppo, A. Gottardo, K. Hadynska-Klek, Z. Korkulu, S. Koyama, Y. Kubota, J. Lee, M. Lettmann, C. Louchart, R. Lozeva, K. Matsui, T. Miyazaki, S. Nishimura, K. Ogata, S. Ota, Z. Patel, E. Sahin, C. Shand, P.-A. Söderström, I. Stefan, D. Steppenbeck, T. Sumikama, D. Suzuki, V. Werner, J. Wu, and Z. Xu, *Phys. Rev. Lett.* **119**, 192501 (2017).
- [8] A. Welker, N. A. S. Althubiti, D. Atanasov, K. Blaum, T. E. Cocolios, F. Herfurth, S. Kreim, D. Lunney, V. Manea, M. Mougeot, D. Neidherr, F. Nowacki, A. Poves, M. Rosenbusch, L. Schweikhard, F. Wienholtz, R. N. Wolf, and K. Zuber, *Phys. Rev. Lett.* **119**, 192502 (2017).
- [9] C. Shand, Z. Podolyák, M. Górski, P. Doornenbal, A. Obertelli, F. Nowacki, T. Otsuka, K. Sieja, J. Tostevin, Y. Tsunoda, G. Authelet, H. Baba, D. Calvet, A. Château, S. Chen, A. Corsi, A. Delbart, J. Gheller, A. Giganon, A. Gillibert, T. Isobe, V. Lapoux, M. Matsushita, S. Momiyama, T. Motobayashi, M. Niikura, H. Otsu, N. Paul, C. Péron, A. Peyaud, E. Pollacco, J.-Y. Roussé, H. Sakurai, C. Santamaria, M. Sasano, Y. Shiga, D. Steppenbeck, S. Takeuchi, R. Taniuchi, T. Uesaka, H. Wang, K. Yoneda, T. Ando, T. Arici, A. Blazhev, F. Browne, A. Bruce, R. Carroll, L. Chung, M. Cortés, M. Dewald, B. Ding, Z. Dombradi, F. Flavigny, S. Franchoo, F. Giacoppo, A. Gottardo, K. Hadynska-Klek, A. Jungclaus, Z. Korkulu, S. Koyama, Y. Kubota, J. Lee, M. Lettmann, B. Linh, J. Liu, Z. Liu, C. Lizarazo, C. Louchart, R. Lozeva, K. Matsui, T. Miyazaki, K. Moschner, M. Nagamine, N. Nakatsuka, S. Nishimura, C. Nita, C. Nobs, L. Olivier, S. Ota, R. Orlandi, Z. Patel, P. Regan, M. Rudigier, E. Şahin, T. Saito, P.-A. Söderström, I. Stefan, T. Sumikama, D. Suzuki, Z. Vajta, V. Vaquero, V. Werner, K. Wimmer, J. Wu, and Z. Xu, *Phys. Lett.* **773**, 492 (2017).
- [10] E. Sahin, F. L. Bello Garrote, Y. Tsunoda, T. Otsuka, G. de Angelis, A. Görgen, M. Niikura, S. Nishimura, Z. Y. Xu, H. Baba, F. Browne, M.-C. Delattre, P. Doornenbal, S. Franchoo, G. Gey, K. Hadynska-Klek, T. Isobe, P. R. John, H. S. Jung, I. Kojouharov, T. Kubo, N. Kurz, Z. Li, G. Lorusso, I. Matea, K. Matsui, D. Mengoni, P. Morfouace, D. R. Napoli, F. Naqvi, H. Nishibata, A. Odahara, H. Sakurai, H. Schaffner, P.-A. Söderström, D. Sohler, I. G. Stefan, T. Sumikama, D. Suzuki, R. Taniuchi, J. Taprogge, Z. Vajta, H. Watanabe, V. Werner, J. Wu, A. Yagi, M. Yalcinkaya, and K. Yoshinaga, *Phys. Rev. Lett.* **118**, 242502 (2017).
- [11] A. Morales, G. Benzoni, H. Watanabe, G. de Angelis, S. Nishimura, L. Coraggio, A. Gargano, N. Itaco, T. Otsuka, Y. Tsunoda, P. Van Isacker, F. Browne, R. Daido, P. Doornenbal, Y. Fang, G. Lorusso, Z. Patel, S. Rice, L. Sinclair, P.-A. Söderström, T. Sumikama, J. Valiente-Dobón, J. Wu, Z. Xu, A. Yagi, R. Yokoyama, H. Baba, R. Avigo, F. Bello Garrote, N. Blasi, A. Bracco, A. Bruce, F. Camera, S. Ceruti, F. Crespi, M.-C. Delattre, Z. Dombradi, A. Gottardo, T. Isobe, I. Kojouharov, N. Kurz, I. Kuti, S. Lalkovski, K. Matsui, B. Melon, D. Mengoni, T. Miyazaki, V. Modamio-Hoybjor, S. Momiyama, D. Napoli, M. Niikura, R. Orlandi, Z. Podolyák, P. Regan, H. Sakurai, E. Sahin, D. Sohler, H. Schaffner, R. Taniuchi, J. Taprogge, Z. Vajta, O. Wieland, and M. Yalcinkaya, *Phys. Lett.* **781**, 706 (2018).
- [12] L. Olivier, S. Franchoo, M. Niikura, Z. Vajta, D. Sohler, P. Doornenbal, A. Obertelli, Y. Tsunoda, T. Otsuka, G. Authelet, H. Baba, D. Calvet, F. Château, A. Corsi, A. Delbart, J.-M. Gheller, A. Gillibert, T. Isobe, V. Lapoux, M. Matsushita, S. Momiyama, T. Motobayashi, H. Otsu, C. Péron, A. Peyaud, E. C. Pollacco, J.-Y. Roussé, H. Sakurai, C. Santamaria, M. Sasano, Y. Shiga, S. Takeuchi, R. Taniuchi, T. Uesaka, H. Wang, K. Yoneda, F. Browne, L. X. Chung, Z. Dombradi, F. Flavigny, F. Giacoppo, A. Gottardo, K. Hadynska-Klek, Z. Korkulu, S. Koyama, Y. Kubota, J. Lee, M. Lettmann, C. Louchart, R. Lozeva, K. Matsui, T. Miyazaki, S. Nishimura, K. Ogata, S. Ota, Z. Patel, E. Sahin, C. Shand, P.-A. Söderström, I. Stefan, D. Steppenbeck, T. Sumikama, D. Suzuki, V. Werner, J. Wu, and

- Z. Xu, Phys. Rev. Lett. **119**, 192501 (2017).
- [13] T. Lokotko, S. Leblond, J. Lee, P. Doornenbal, A. Obertelli, A. Poves, F. Nowacki, K. Ogata, K. Yoshida, G. Authelet, H. Baba, D. Calvet, F. Château, S. Chen, A. Corsi, A. Delbart, J.-M. Gheller, A. Gillibert, T. Isobe, V. Lapoux, M. Matsushita, S. Momiyama, T. Motobayashi, M. Niikura, H. Otsu, C. Péron, A. Peyaud, E. C. Pollacco, J.-Y. Roussé, H. Sakurai, C. Santamaria, Z. Y. Xu, M. Sasano, Y. Shiga, S. Takeuchi, R. Taniuchi, T. Uesaka, H. Wang, V. Werner, F. Browne, L. X. Chung, Z. Dombradi, S. Franchoo, F. Giacoppo, A. Gottardo, K. Hadynska-Klek, Z. Korkulu, S. Koyama, Y. Kubota, M. Lettmann, C. Louchart, R. Lozeva, K. Matsui, T. Miyazaki, S. Nishimura, L. Olivier, S. Ota, Z. Patel, E. Sahin, C. Shand, P.-A. Söderström, I. Stefan, D. Steppenbeck, T. Sumikama, D. Suzuki, Z. Vajta, and J. Wu, Phys. Rev. C **101**, 034314 (2020).
- [14] F. H. Garcia, C. Andreoiu, G. C. Ball, A. Bell, A. B. Garnsworthy, F. Nowacki, C. M. Petrache, A. Poves, K. Whitmore, F. A. Ali, N. Bernier, S. S. Bhattacharjee, M. Bowry, R. J. Coleman, I. Dillmann, I. Djianto, A. M. Forney, M. Gascoine, G. Hackman, K. G. Leach, A. N. Murphy, C. R. Natzke, B. Olaizola, K. Ortner, E. E. Peters, M. M. Rajabali, K. Raymond, C. E. Svensson, R. Umashankar, J. Williams, and D. Yates, Phys. Rev. Lett. **125**, 172501 (2020).
- [15] F. Nowacki, A. Poves, E. Caurier, and B. Bounthong, Phys. Rev. Lett. **117**, 272501 (2016).
- [16] T. Otsuka, T. Suzuki, R. Fujimoto, H. Grawe, and Y. Akaishi, Phys. Rev. Lett. **95**, 232502 (2005).
- [17] R. Taniuchi, C. Santamaria, P. Doornenbal, A. Obertelli, K. Yoneda, G. Authelet, H. Baba, D. Calvet, F. Château, A. Corsi, A. Delbart, J.-M. Gheller, A. Gillibert, J. D. Holt, T. Isobe, V. Lapoux, M. Matsushita, J. Menéndez, S. Momiyama, T. Motobayashi, M. Niikura, F. Nowacki, K. Ogata, H. Otsu, T. Otsuka, C. Péron, S. Péru, A. Peyaud, E. C. Pollacco, A. Poves, J.-Y. Roussé, H. Sakurai, A. Schwenk, Y. Shiga, J. Simonis, S. R. Stroberg, S. Takeuchi, Y. Tsunoda, T. Uesaka, H. Wang, F. Browne, L. X. Chung, Z. Dombradi, S. Franchoo, F. Giacoppo, A. Gottardo, K. Hadynska-Klek, Z. Korkulu, S. Koyama, Y. Kubota, J. Lee, M. Lettmann, C. Louchart, R. Lozeva, K. Matsui, T. Miyazaki, S. Nishimura, L. Olivier, S. Ota, Z. Patel, E. Sahin, C. Shand, P.-A. Söderström, I. Stefan, D. Steppenbeck, T. Sumikama, D. Suzuki, Z. Vajta, V. Werner, J. Wu, and Z. Y. Xu, Nature (London) **569**, 53 (2019).
- [18] N. Tsunoda, T. Otsuka, K. Takayanagi, N. Shimizu, T. Suzuki, Y. Utsuno, S. Yoshida, and H. Ueno, Nature (London) **587**, 66 (2020).
- [19] O. Perru, O. Sorlin, S. Franchoo, F. Azaiez, E. Bouchez, C. Bourgeois, A. Chatillon, J. M. Daugas, Z. Dlouhy, Z. Dombradi, C. Donzaud, L. Gaudefroy, H. Grawe, S. Grévy, D. Guillemaud-Mueller, F. Hammache, F. Ibrahim, Y. Le Coz, S. M. Lukyanov, I. Matea, J. Mrazek, F. Nowacki, Y.-E. Penionzhkevich, F. de Oliveira Santos, F. Pougheon, M. G. Saint-Laurent, G. Sletten, M. Stanoiu, C. Stodel, C. Theisen, and D. Verney, Phys. Rev. Lett. **96**, 232501 (2006).
- [20] N. Aoi, S. Kanno, S. Takeuchi, H. Suzuki, D. Bazin, M. Bowen, C. Campbell, J. Cook, D.-C. Dinca, A. Gade, T. Glasmacher, H. Iwasaki, T. Kubo, K. Kurita, T. Motobayashi, W. Mueller, T. Nakamura, H. Sakurai, M. Takashina, J. Terry, K. Yoneda, and H. Zwahlen, Phys. Lett. **692**, 302 (2010).
- [21] M. M. Rajabali, R. Grzywacz, S. N. Liddick, C. Mazzocchi, J. C. Batchelder, T. Baumann, C. R. Bingham, I. G. Darby, T. N. Ginter, S. V. Ilyushkin, M. Karny, W. Królas, P. F. Mantica, K. Miernik, M. Pfützner, K. P. Rykaczewski, D. Weisshaar, and J. A. Winger, Phys. Rev. C **85**, 034326 (2012).
- [22] Y. Tsunoda, T. Otsuka, N. Shimizu, M. Honma, and Y. Utsuno, Phys. Rev. C **89**, 031301 (2014).
- [23] T. Marchi, G. de Angelis, J. J. Valiente-Dobón, V. M. Bader, T. Baugher, D. Bazin, J. Berryman, A. Bonaccorso, R. Clark, L. Coraggio, H. L. Crawford, M. Doncel, E. Farnea, A. Gade, A. Gadea, A. Gargano, T. Glasmacher, A. Gottardo, F. Gramegna, N. Itaco, P. R. John, R. Kumar, S. M. Lenzi, S. Lunardi, S. McDaniel, C. Michelagnoli, D. Mengoni, V. Modamio, D. R. Napoli, B. Quintana, A. Ratkiewicz, F. Recchia, E. Sahin, R. Stroberg, D. Weisshaar, K. Wimmer, and R. Winkler, Phys. Rev. Lett. **113**, 182501 (2014).
- [24] K. Kolos, D. Miller, R. Grzywacz, H. Iwasaki, M. Al-Shudifat, D. Bazin, C. R. Bingham, T. Braunroth, G. Cerizza, A. Gade, A. Lemasson, S. N. Liddick, M. Madurga, C. Morse, M. Portillo, M. M. Rajabali, F. Recchia, L. L. Riedinger, P. Voss, W. B. Walters, D. Weisshaar, K. Whitmore, K. Wimmer, and J. A. Tostevin, Phys. Rev. Lett. **116**, 122502 (2016).
- [25] A. I. Morales, G. Benzon, H. Watanabe, S. Nishimura, F. Browne, R. Daido, P. Doornenbal, Y. Fang, G. Lorusso, Z. Patel, S. Rice, L. Sinclair, P.-A. Söderström, T. Sumikama, J. Wu, Z. Y. Xu, A. Yagi, R. Yokoyama, H. Baba, R. Avigo, F. L. Bello Garrote, N. Blasi, A. Bracco, F. Camera, S. Ceruti, F. C. L. Crespi, G. de Angelis, M.-C. Delattre, Z. Dombradi, A. Gottardo, T. Isobe, I. Kojouharov, N. Kurz, I. Kuti, K. Matsui, B. Melon, D. Mengoni, T. Miyazaki, V. Modamio-Hoyborg, S. Momiyama, D. R. Napoli, M. Niikura, R. Orlandi, H. Sakurai, E. Sahin, D. Sohler, H. Shaffner, R. Taniuchi, J. Taprogge, Z. Vajta, J. J. Valiente-Dobón, O. Wieland, and M. Yalcinkaya, Phys. Rev. C **93**, 034328 (2016).
- [26] A. Morales, G. Benzon, H. Watanabe, Y. Tsunoda, T. Otsuka, S. Nishimura, F. Browne, R. Daido, P. Doornenbal, Y. Fang, G. Lorusso, Z. Patel, S. Rice, L. Sinclair, P.-A. Söderström, T. Sumikama, J. Wu, Z. Xu, A. Yagi, R. Yokoyama, H. Baba, R. Avigo, F. Bello Garrote, N. Blasi, A. Bracco, F. Camera, S. Ceruti, F. Crespi, G. de Angelis, M.-C. Delattre, Z. Dombradi, A. Gottardo, T. Isobe, I. Kojouharov, N. Kurz, I. Kuti, K. Matsui, B. Melon, D. Mengoni, T. Miyazaki, V. Modamio-Hoybjor, S. Momiyama, D. Napoli, M. Niikura, R. Orlandi, H. Sakurai, E. Sahin, D. Sohler, H. Schaffner, R. Taniuchi, J. Taprogge, Z. Vajta, J. Valiente-Dobón, O. Wieland, and M. Yalcinkaya, Phys. Lett. **765**, 328 (2017).
- [27] M. L. Cortés, P. Doornenbal, M. Dupuis, S. M. Lenzi, F. Nowacki, A. Obertelli, S. Péru, N. Pietralla, V. Werner, K. Wimmer, G. Authelet, H. Baba, D. Calvet, F. Château, A. Corsi, A. Delbart, J.-M. Gheller, A. Gillibert, T. Isobe, V. Lapoux, C. Louchart, M. Matsushita, S. Momiyama, T. Motobayashi, M. Niikura, H. Otsu, C. Péron, A. Peyaud, E. C. Pollacco, J.-Y. Roussé, H. Sakurai, C. Santamaria, M. Sasano, Y. Shiga, S. Takeuchi, R. Taniuchi, T. Uesaka, H. Wang, K. Yoneda, F. Browne, L. X. Chung, Z. Dombradi, S. Franchoo, F. Giacoppo, A. Gottardo, K. Hadynska-

- Klek, Z. Korkulu, S. Koyama, Y. Kubota, J. Lee, M. Lettmann, R. Lozeva, K. Matsui, T. Miyazaki, S. Nishimura, L. Olivier, S. Ota, Z. Patel, E. Sahin, C. M. Shand, P.-A. Söderström, I. Stefan, D. Steppenbeck, T. Sumikama, D. Suzuki, Z. Vajta, J. Wu, and Z. Xu, *Phys. Rev. C* **97**, 044315 (2018).
- [28] B. Elman, A. Gade, R. V. F. Janssens, A. D. Ayangeakaa, D. Bazin, J. Belarge, P. C. Bender, B. A. Brown, C. M. Campbell, M. P. Carpenter, H. L. Crawford, B. P. Crider, P. Fallon, A. M. Forney, J. Harker, S. N. Liddick, B. Longfellow, E. Lunderberg, C. J. Prokop, J. Sethi, R. Taniuchi, W. B. Walters, D. Weisshaar, and S. Zhu, *Phys. Rev. C* **100**, 034317 (2019).
- [29] Z. Elekes, A. Kripkó, D. Sohler, K. Sieja, K. Ogata, K. Yoshida, P. Doornenbal, A. Obertelli, G. Authalet, H. Baba, D. Calvet, F. Château, A. Corsi, A. Delbart, J.-M. Gheller, A. Gillibert, T. Isobe, V. Lapoux, M. Matsushita, S. Momiyama, T. Motobayashi, H. Otsu, C. Péron, A. Peyaud, E. C. Pollacco, J.-Y. Roussé, H. Sakurai, C. Santamaria, Y. Shiga, S. Takeuchi, R. Taniuchi, T. Uesaka, H. Wang, K. Yoneda, F. Browne, L. X. Chung, Z. Dombrádi, F. Flavigny, S. Franchoo, F. Giacoppo, A. Gottardo, K. Hadyńska-Klęk, Z. Korkulu, S. Koyama, Y. Kubota, J. Lee, M. Lettmann, C. Louchart, R. Lozeva, K. Matsui, T. Miyazaki, M. Niikura, S. Nishimura, L. Olivier, S. Ota, Z. Patel, E. Sahin, C. Shand, P.-A. Söderström, I. Stefan, D. Steppenbeck, T. Sumikama, D. Suzuki, Z. Vajta, V. Werner, J. Wu, and Z. Xu (SunFlower Collaboration), *Phys. Rev. C* **99**, 014312 (2019).
- [30] A. Gottardo, G. de Angelis, P. Doornenbal, L. Coraggio, A. Gargano, N. Itaco, K. Kaneko, P. Van Isacker, T. Furumoto, G. Benzoni, J. Lee, H. Liu, M. Matsushita, D. Mengoni, V. Modamio-Hoybjør, S. Momiyama, T. Motobayashi, D. R. Napoli, M. Niikura, E. Sahin, Y. Shiga, H. Sakurai, R. Taniuchi, S. Takeuchi, H. Wang, J. J. Valiente-Dobón, R. Avigo, H. Baba, N. Blasi, F. L. Bello Garrote, F. Browne, F. C. L. Crespi, S. Ceruti, R. Daido, M.-C. Delattre, D. Fang, Z. Dombradi, T. Isobe, I. Kuti, G. Lorusso, K. Matsui, B. Melon, T. Miyazaki, S. Nishimura, R. Orlandi, Z. Patel, S. Rice, L. Sinclair, P. A. Söderström, D. Sohler, T. Sumikama, J. Taprogge, Z. Vajta, H. Watanabe, O. Wieland, J. Wu, Z. Y. Xu, M. Yalcinkaya, and R. Yokoyama, *Phys. Rev. C* **102**, 014323 (2020).
- [31] F. L. Bello Garrote, E. Sahin, Y. Tsunoda, T. Otsuka, A. Görgen, M. Niikura, S. Nishimura, G. de Angelis, G. Benzoni, A. I. Morales, V. Modamio, Z. Y. Xu, H. Baba, F. Browne, A. M. Bruce, S. Ceruti, F. C. L. Crespi, R. Daido, M.-C. Delattre, P. Doornenbal, Z. Dombradi, Y. Fang, S. Franchoo, G. Gey, A. Gottardo, K. Hadyńska-Klęk, T. Isobe, P. R. John, H. S. Jung, I. Kojouharov, T. Kubo, N. Kurz, I. Kuti, Z. Li, G. Lorusso, I. Matea, K. Matsui, D. Mengoni, T. Miyazaki, S. Momiyama, P. Morfouace, D. R. Napoli, F. Naqvi, H. Nishibata, A. Odahara, R. Orlandi, Z. Patel, S. Rice, H. Sakurai, H. Schaffner, L. Sinclair, P.-A. Söderström, D. Sohler, I. G. Stefan, T. Sumikama, D. Suzuki, R. Taniuchi, J. Taprogge, Z. Vajta, J. J. Valiente-Dobón, H. Watanabe, V. Werner, J. Wu, A. Yagi, M. Yalcinkaya, R. Yokoyama, and K. Yoshinaga, *Phys. Rev. C* **102**, 034314 (2020).
- [32] S. Go, R. Grzywacz, C. Mazzocchi, S. N. Liddick, M. Alshudifat, J. C. Batchelder, T. Baumann, A. A. Ciemny, T. N. Ginter, C. J. Gross, K. Kolos, A. Korgul, S. V. Paulauskas, C. J. Prokop, M. M. Rajabali, K. P. Rykaczewski, S. Taylor, and Y. Xiao, *Phys. Rev. C* **102**, 044331 (2020).
- [33] I. Talmi, *Simple Models of Complex Nuclei: The Shell Model and Interacting Boson Model* (Chur, Switzerland: Harwood Academic Publishers, 1993).
- [34] R. Casten, *Nuclear Structure from a Simple Perspective* (Oxford University press, New York, 2000).
- [35] A. Escuderos and L. Zamick, *Phys. Rev. C* **73**, 044302 (2006).
- [36] C. Qi, *Phys. Rev. C* **83**, 014307 (2011).
- [37] L. Zamick, *Phys. Rev. C* **75**, 064305 (2007).
- [38] P. Van Isacker and S. Heinze, *Phys. Rev. Lett.* **100**, 052501 (2008).
- [39] C. Qi, *Phys. Rev. C* **81**, 034318 (2010).
- [40] P. Van Isacker and I. Celikovic, in *Proc. 14th International Symposium on Capture Gamma-Ray Spectroscopy and Related Topics*, edited by P. Garret and B. Hadinia (World Scientific, 2013) p. 44.
- [41] H. Mach, A. Korgul, M. Górski, H. Grawe, I. Matea, M. Stănoiu, L. M. Fraile, Y. E. Penionzkevich, F. D. O. Santos, D. Verney, S. Lukyanov, B. Cederwall, A. Covello, Z. Dlouhý, B. Fogelberg, G. De France, A. Gargano, G. Georgiev, R. Grzywacz, A. F. Lisetskiy, J. Mrazek, F. Nowacki, W. A. Plóciennik, Z. Podolyák, S. Ray, E. Ruchowska, M.-G. Saint-Laurent, M. Sawicka, C. Stodel, and O. Tarasov, *Phys. Rev. C* **95**, 014313 (2017).
- [42] C. Qi, *Phys. Lett.* **773**, 616 (2017).
- [43] H. L. Crawford, R. M. Clark, P. Fallon, A. O. Macchiavelli, T. Baugher, D. Bazin, C. W. Beausang, J. S. Berryman, D. L. Bleuel, C. M. Campbell, M. Cromaz, G. de Angelis, A. Gade, R. O. Hughes, I. Y. Lee, S. M. Lenzi, F. Nowacki, S. Paschalis, M. Petri, A. Poves, A. Ratkiewicz, T. J. Ross, E. Sahin, D. Weisshaar, K. Wimmer, and R. Winkler, *Phys. Rev. Lett.* **110**, 242701 (2013).
- [44] S. N. Liddick, W. B. Walters, C. J. Chiara, R. V. F. Janssens, B. Abromeit, A. Ayres, A. Bey, C. R. Bingham, M. P. Carpenter, L. Cartegni, J. Chen, H. L. Crawford, I. G. Darby, R. Grzywacz, J. Harker, C. R. Hoffman, S. Ilyushkin, F. G. Kondev, N. Larson, M. Madurga, D. Miller, S. Padgett, S. V. Paulauskas, M. M. Rajabali, K. Rykaczewski, D. Seweryniak, S. Suchyta, and S. Zhu, *Phys. Rev. C* **92**, 024319 (2015).
- [45] M. Cortés, W. Rodriguez, P. Doornenbal, A. Obertelli, J. Holt, S. Lenzi, J. Menéndez, F. Nowacki, K. Ogata, A. Poves, T. Rodríguez, A. Schwenk, J. Simonis, S. Stroberg, K. Yoshida, L. Achouri, H. Baba, F. Browne, D. Calvet, F. Château, S. Chen, N. Chiga, A. Corsi, A. Delbart, J.-M. Gheller, A. Giganon, A. Gillibert, C. Hilaire, T. Isobe, T. Kobayashi, Y. Kubota, V. Lapoux, H. Liu, T. Motobayashi, I. Murray, H. Otsu, V. Panin, N. Paul, H. Sakurai, M. Sasano, D. Steppenbeck, L. Stuhl, Y. Sun, Y. Togano, T. Uesaka, K. Wimmer, K. Yoneda, O. Aktas, T. Aumann, L. Chung, F. Flavigny, S. Franchoo, I. Gašparić, R.-B. Gerst, J. Gibelin, K. Hahn, D. Kim, T. Koizumi, Y. Kondo, P. Koseoglou, J. Lee, C. Lehr, B. Linh, T. Lokotko, M. MacCormick, K. Moschner, T. Nakamura, S. Park, D. Rossi, E. Sahin, D. Sohler, P.-A. Söderström, S. Takeuchi, H. Toernqvist, V. Vaquero, V. Wagner, S. Wang, V. Werner, X. Xu, H. Yamada, D. Yan, Z. Yang, M. Yasuda, and L. Zanetti, *Phys. Lett.* **800**, 135071 (2020).

- [46] A. Gade, R. V. F. Janssens, J. A. Tostevin, D. Bazin, J. Belarge, P. C. Bender, S. Bottoni, M. P. Carpenter, B. Elman, S. J. Freeman, T. Lauritsen, S. M. Lenzi, B. Longfellow, E. Lunderberg, A. Poves, L. A. Riley, D. K. Sharp, D. Weisshaar, and S. Zhu, *Phys. Rev. C* **99**, 011301 (2019).
- [47] L. Canete, S. Giraud, A. Kankainen, B. Bastin, F. Nowacki, A. Poves, P. Ascher, T. Eronen, V. Alcindor, A. Jokinen, A. Khanam, I. D. Moore, D. A. Nesterenko, F. De Oliveira Santos, H. Penttilä, C. Petrone, I. Pohjalainen, A. de Roubin, V. A. Rubchenya, M. Vilen, and J. Äystö, *Phys. Rev. C* **101**, 041304 (2020).
- [48] T. Otsuka and Y. Tsunoda, *J. Phys. G* **43**, 024009 (2016).
- [49] T. Togashi, Y. Tsunoda, T. Otsuka, and N. Shimizu, *Phys. Rev. Lett.* **117**, 172502 (2016).
- [50] T. Otsuka, Y. Tsunoda, T. Abe, N. Shimizu, and P. Van Duppen, *Phys. Rev. Lett.* **123**, 222502 (2019).
- [51] T. Otsuka, A. Gade, O. Sorlin, T. Suzuki, and Y. Utsuno, *Rev. Mod. Phys.* **92**, 015002 (2020).
- [52] T. Otsuka, *Physica Scripta* **T152**, 014007 (2013).
- [53] T. Shafer, J. Engel, C. Fröhlich, G. C. McLaughlin, M. Mumpower, and R. Surman, *Phys. Rev. C* **94**, 055802 (2016).
- [54] P. Hosmer, H. Schatz, A. Aprahamian, O. Arndt, R. R. C. Clement, A. Estrade, K. Farouqi, K.-L. Kratz, S. N. Liddick, A. F. Lisetskiy, P. F. Mantica, P. Möller, W. F. Mueller, F. Montes, A. C. Morton, M. Ouellette, E. Pellegrini, J. Pereira, B. Pfeiffer, P. Reeder, P. Santi, M. Steiner, A. Stolz, B. E. Tomlin, W. B. Walters, and A. Wöhr, *Phys. Rev. C* **82**, 025806 (2010).
- [55] T. Ohnishi, T. Kubo, K. Kusaka, A. Yoshida, K. Yoshida, M. Ohtake, N. Fukuda, H. Takeda, D. Kameda, K. Tanaka, N. Inabe, Y. Yanagisawa, Y. Gono, H. Watanabe, H. Otsu, H. Baba, T. Ichihara, Y. Yamaguchi, M. Takechi, S. Nishimura, H. Ueno, A. Yoshimi, H. Sakurai, T. Motobayashi, T. Nakao, Y. Mizoi, M. Matsushita, K. Ieki, N. Kobayashi, K. Tanaka, Y. Kawada, N. Tanaka, S. Deguchi, Y. Satou, Y. Kondo, T. Nakamura, K. Yoshinaga, C. Ishii, H. Yoshii, Y. Miyashita, N. Uematsu, Y. Shiraki, T. Sumikama, J. Chiba, E. Ideguchi, A. Saito, T. Yamaguchi, I. Hachiuma, T. Suzuki, T. Moriguchi, A. Ozawa, T. Ohtsubo, M. A. Famiano, H. Geissel, A. S. Nettleton, O. B. Tarasov, D. P. Bazin, B. M. Sherrill, S. L. Manikonda, and J. A. Nolen, *J. Phys. Soc. Jpn* **79**, 073201 (2010).
- [56] N. Fukuda, T. Kubo, T. Ohnishi, N. Inabe, H. Takeda, D. Kameda, and H. Suzuki, *Nucl. Instrum. Methods* **317**, 323 (2013).
- [57] Y. Mizoi, T. Kubo, H. Sakurai, K. Kusaka, K. Yoshida, and A. Yoshida, *RIKEN Accel. Prog. Rep.* **38**, 297 (2005).
- [58] S. Nishimura, *Prog. Theor. Exp. Phys.* **2012** (2012).
- [59] P.-A. Söderström, S. Nishimura, P. Doornenbal, G. Lorusso, T. Sumikama, H. Watanabe, Z. Xu, H. Baba, F. Browne, S. Go, G. Gey, T. Isobe, H.-S. Jung, G. Kim, Y.-K. Kim, I. Kojouharov, N. Kurz, Y. Kwon, Z. Li, K. Moschner, T. Nakao, H. Nishibata, M. Nishimura, A. Odahara, H. Sakurai, H. Schaffner, T. Shimoda, J. Taprogge, Z. Vajta, V. Werner, J. Wu, A. Yagi, and K. Yoshinaga, *Nucl. Instrum. Methods* **317**, 649 (2013).
- [60] S. Nishimura, G. Lorusso, Z. Xu, J. Wu, R. Gernhauser, H. Jung, Y. Kwon, Z. Li, K. Steiger, and H. Sakurai, *RIKEN Accel. Prog. Rep.* **46**, 182 (2013).
- [61] S. E. A. Orrigo, B. Rubio, W. Gelletly, P. Aguilera, A. Algora, A. I. Morales, J. Agramunt, D. S. Ahn, P. Ascher, B. Blank, C. Borcea, A. Boso, R. B. Cakirli, J. Chiba, G. de Angelis, G. de France, F. Diel, P. Doornenbal, Y. Fujita, N. Fukuda, E. Ganioglu, M. Gerbaux, J. Giovinnazzo, S. Go, T. Goigoux, S. Grévy, V. Guadilla, N. Inabe, G. Kiss, T. Kubo, S. Kubono, T. Kurtukian-Nieto, D. Lubos, C. Magron, F. Molina, A. Montaner-Pizá, D. Napoli, D. Nishimura, S. Nishimura, H. Oikawa, Y. Shimizu, C. Sidong, P. A. Söderström, T. Sumikama, H. Suzuki, H. Takeda, Y. Takei, M. Tanaka, P. Vi, J. Wu, and S. Yagi, (2020), arXiv:2008.10576 [nucl-ex].
- [62] A. I. Morales, J. Benlliure, M. Górski, H. Grawe, S. Verma, P. H. Regan, Z. Podolyák, S. Pietri, R. Kumar, E. Casarejos, A. Algora, N. Alkhomashi, H. Álvarez-Pol, G. Benzoni, A. Blazhev, P. Boutachkov, A. M. Bruce, L. S. Cáceres, I. J. Cullen, A. M. Denis Bacelar, P. Doornenbal, M. E. Estévez-Aguado, G. Farrelly, Y. Fujita, A. B. Garnsworthy, W. Gelletly, J. Gerl, J. Grebosz, R. Hoischen, I. Kojouharov, N. Kurz, S. Lalkowski, Z. Liu, C. Mihai, F. Molina, D. Mücher, W. Prokopowicz, B. Rubio, H. Schaffner, S. J. Steer, A. Tamii, S. Tashenov, J. J. Valiente-Dobón, P. M. Walker, H. J. Wollersheim, and P. J. Woods, *Phys. Rev. C* **88**, 014319 (2013).
- [63] W. Verkerke and D. P. Kirkby, *eConf C0303241*, MOLT007 (2003), arXiv:physics/0306116.
- [64] F. James, *MINUIT Function Minimisation and Error Analysis: Reference Manual, Version 94.1* (CERN, 1994).
- [65] A. I. Morales, A. Algora, B. Rubio, K. Kaneko, S. Nishimura, P. Aguilera, S. E. A. Orrigo, F. Molina, G. de Angelis, F. Recchia, G. Kiss, V. H. Phong, J. Wu, D. Nishimura, H. Oikawa, T. Goigoux, J. Giovinnazzo, P. Ascher, J. Agramunt, D. S. Ahn, H. Baba, B. Blank, C. Borcea, A. Boso, P. Davies, F. Diel, Z. Dombrádi, P. Doornenbal, J. Eberth, G. de France, Y. Fujita, N. Fukuda, E. Ganioglu, W. Gelletly, M. Gerbaux, S. Grévy, V. Guadilla, N. Inabe, T. Isobe, I. Kojouharov, W. Korten, T. Kubo, S. Kubono, T. Kurtukian-Nieto, N. Kurz, J. Lee, S. Lenzi, J. Liu, T. Lokotko, D. Lubos, C. Magron, A. Montaner-Pizá, D. R. Napoli, H. Sakurai, H. Schaffner, Y. Shimizu, C. Sidong, P.-A. Söderström, T. Sumikama, H. Suzuki, H. Takeda, Y. Takei, M. Tanaka, and S. Yagi, *Phys. Rev. C* **95**, 064327 (2017).
- [66] J. Hardy, L. Carraz, B. Jonson, and P. Hansen, *Phys. Lett.* **71**, 307 (1977).
- [67] M. Wang, G. Audi, F. G. Kondev, W. Huang, S. Naimi, and X. Xu, *Chinese Phys. C* **41**, 030003 (2017).
- [68] W. F. Mueller, B. Bruyneel, S. Franchoo, H. Grawe, M. Huyse, U. Köster, K.-L. Kratz, K. Kruglov, Y. Kudryavtsev, B. Pfeiffer, R. Raabe, I. Reusen, P. Thirolf, P. Van Duppen, J. Van Roosbroeck, L. Vermeeren, W. B. Walters, and L. Weissman, *Phys. Rev. Lett.* **83**, 3613 (1999).
- [69] N. Al-Dahan, P. H. Regan, Z. Podolyák, P. M. Walker, N. Alkhomashi, G. D. Dracoulis, G. Farrelly, J. Benlliure, S. B. Pietri, R. F. Casten, P. D. Stevenson, W. Gelletly, S. J. Steer, A. B. Garnsworthy, E. Casarejos, J. Gerl, H. J. Wollersheim, J. Grebosz, M. Górski, I. Kojouharov, H. Schaffner, A. Algora, G. Benzoni, A. Blazhev, P. Boutachkov, A. M. Bruce, I. J. Cullen, A. M. Denis Bacelar, A. Y. Deo, M. E. Estevez, Y. Fujita,

- R. Hoischen, R. Kumar, S. Lalkovski, Z. Liu, P. J. Mason, C. Mihai, F. Molina, D. Mücher, B. Rubio, A. Tamii, S. Tashenov, J. J. Valiente-Dobón, and P. J. Woods, *Phys. Rev. C* **85**, 034301 (2012).
- [70] A. Gottardo, J. J. Valiente-Dobón, G. Benzoni, R. Nicolini, A. Gadea, S. Lunardi, P. Boutachkov, A. M. Bruce, M. Górska, J. Grebosz, S. Pietri, Z. Podolyák, M. Pfützner, P. H. Regan, H. Weick, J. Alcántara Núñez, A. Algora, N. Al-Dahan, G. de Angelis, Y. Ayyad, N. Alkhomashi, P. R. P. Allegro, D. Bazzacco, J. Benlliure, M. Bowry, A. Bracco, M. Bunce, F. Camera, E. Casarejos, M. L. Cortes, F. C. L. Crespi, A. Corsi, A. M. Denis Bacelar, A. Y. Deo, C. Domingo-Pardo, M. Doncel, Z. Dombradi, T. Engert, K. Eppinger, G. F. Farrelly, F. Farinon, E. Farnea, H. Geissel, J. Gerl, N. Goel, E. Gregor, T. Habermann, R. Hoischen, R. Janik, S. Klupp, I. Kojouharov, N. Kurz, S. M. Lenzi, S. Leoni, S. Mandal, R. Menegazzo, D. Mengoni, B. Million, A. I. Morales, D. R. Napoli, F. Naqvi, C. Nociforo, A. Prochazka, W. Prokopowicz, F. Recchia, R. V. Ribas, M. W. Reed, D. Rudolph, E. Sahin, H. Schaffner, A. Sharma, B. Sitar, D. Siwal, K. Steiger, P. Strmen, T. P. D. Swan, I. Szarka, C. A. Ur, P. M. Walker, O. Wieland, H.-J. Wollersheim, F. Nowacki, E. Maglione, and A. P. Zuker, *Phys. Rev. Lett.* **109**, 162502 (2012).
- [71] S. M. Lenzi, F. Nowacki, A. Poves, and K. Sieja, *Phys. Rev. C* **82**, 054301 (2010).
- [72] L. Satpathy and S. Gujrathi, *Nucl. Phys.* **110**, 400 (1968).
- [73] V. Paar, *Nucl. Phys.* **211**, 29 (1973).