

This is a repository copy of *Kinetic Study of the Reactions PO* + *O2* and *PO2* + *O3* and *Spectroscopy of the PO Radical*.

White Rose Research Online URL for this paper: https://eprints.whiterose.ac.uk/165257/

Version: Accepted Version

Article:

Douglas, KM orcid.org/0000-0002-3281-3685, Blitz, MA, Mangan, TP orcid.org/0000-0001-7053-5594 et al. (2 more authors) (2020) Kinetic Study of the Reactions PO + O2 and PO2 + O3 and Spectroscopy of the PO Radical. The Journal of Physical Chemistry A. ISSN 1089-5639

https://doi.org/10.1021/acs.jpca.0c06106

© 2020 American Chemical Society. This is an author produced version of a journal article published in The Journal of Physical Chemistry A. Uploaded in accordance with the publisher's self-archiving policy.

Reuse

Items deposited in White Rose Research Online are protected by copyright, with all rights reserved unless indicated otherwise. They may be downloaded and/or printed for private study, or other acts as permitted by national copyright laws. The publisher or other rights holders may allow further reproduction and re-use of the full text version. This is indicated by the licence information on the White Rose Research Online record for the item.

Takedown

If you consider content in White Rose Research Online to be in breach of UK law, please notify us by emailing eprints@whiterose.ac.uk including the URL of the record and the reason for the withdrawal request.



eprints@whiterose.ac.uk https://eprints.whiterose.ac.uk/

Kinetic Study of the Reactions $PO + O_2$ and $PO_2 + O_3$ and Spectroscopy of the 1 2 **PO** Radical 3 Kevin M. Douglas^{a*}, Mark A. Blitz^{a,b}, Thomas P. Mangan^a, Colin M. Western^c, John M.C. 4 5 Plane^{a*} 6 ^aSchool of Chemistry, University of Leeds, Leeds, LS2 9JT, UK 7 ^bNational Centre for Atmospheric Science (NCAS), University of Leeds, Leeds, LS2 9JT, UK 8 ^cSchool of Chemistry, University of Bristol, BS8 1TS, UK 9 *Corresponding authors emails: K.M.Douglas@leeds.ac.uk; J.M.C.Plane@leeds.ac.uk 10 11 Abstract

12 The kinetics of the reactions of PO with O_2 and PO_2 with O_3 were studied at temperatures 13 ranging from ~ 190 to 340 K, using a pulsed laser photolysis-laser induced fluorescence 14 technique. For the reaction of $PO + O_2$ there is evidence of both a two- and three-body exit 15 channel, producing $PO_2 + O$ and PO_3 , respectively. Potential energy surfaces of both the PO + 16 O_2 and $PO_2 + O_3$ systems were calculated using electronic structure theory, and combined 17 with RRKM calculations to explain the observed pressure and temperature dependences. For 18 $PO + O_2$, at pressures typical of a planetary upper atmosphere where meteoric ablation of P will occur, the reaction is effectively pressure independent with a yield of $PO_2 + O$ of > 99%; 19 the rate coefficient can be expressed by: $\log_{10}(k, 120 - 500 \text{ K}, \text{ cm}^3 \text{ molecule}^{-1} \text{ s}^{-1}) = -13.915 + 1000 \text{ molecule}^{-1} \text{ s}^{-1}$ 20 21 2.470log₁₀(T) - 0.5020(log₁₀(T))², with an uncertainty of \pm 10 % over the experimental 22 temperature range (191 - 339 K). With increasing pressure, the yield of PO₃ increases, 23 reaching ~ 90% at a pressure of 1 atm and T = 300 K. For PO₂ + O₃, k(188 - 339 K) = $3.7 \times$ $10^{-11} \exp^{(-1131/T)}$ cm³ molecule⁻¹ s⁻¹, with an uncertainty of ± 26 % over the stated temperature 24 range. Laser induced fluorescence spectra of PO over the wavelength range of 245 – 248 nm 25 26 were collected and simulated using PGOPHER to obtain new spectroscopic constants for the 27 ground and v = 1 vibrational levels of the $X^2\Pi$ and $A^2\Sigma^+$ states of PO.

28

29 1. Introduction

30 Phosphorus, P, is one of main biogenic elements, a group which, along with H, C, O, 31 N, and S, are present in all known life forms. Compounds of phosphorus appear profusely in 32 biological systems, where they are involved in many fundamental biological functions, including replication, information transfer, and metabolism.¹ Despite its biological 33 34 importance, P is relatively scarce on a cosmic scale. With an elemental abundance of $3.4 \times$ 10⁻⁷ relative to H in the Sun's atmosphere,² it is the least abundant of the main biogenic 35 36 elements. Ortho-phosphate (oxidation state +5), is the dominant form of inorganic P at the 37 Earth's surface. However, P(V) salts typically exhibit poor bioavailability due to their low 38 water solubility, and low reactivity. Indeed, low concentrations of dissolved P make it a 39 limiting reaction in many ecosystems, as well as giving rise to the 'phosphorus problem' in 40 the origin or life.³⁻⁴ In contrast, the salts of more reduced forms of P (oxidation state $\leq +3$) are far more soluble and reactive, and thus have improved bioavailability. 41

42 One possible source of these reduced forms of P is from extra-terrestrial material that fell to Earth during the heavy bombardment period. A study by Pasek⁵ demonstrated how 43 44 phosphides, delivered directly to the surface of the Earth in iron nickel meteorites, can be 45 processed by aqueous phase chemistry to form several prebiotic P species. However, of the 46 total annual mass influx of exogenous material entering the Earth's atmosphere, iron nickel 47 meteorites only account for ~ 1%, with interplanetary dust particles (IDPs) accounting for the 48 other 99 %.⁶ An alternative entry route for reduced forms of exogenous P is the atmospheric processing of P that ablates in the upper atmosphere. In two previous publications we have 49 demonstrated that the meteoric ablation of IDPs is a significant source of atmospheric P, PO, 50 and PO₂ to the terrestrial planets,⁷ and how any P atoms will be rapidly oxidised to PO.⁸ 51 Continued atmospheric processing of PO and PO₂ will then result in a variety of compounds 52 53 in which P may exist in different oxidation states, due to the presence of both oxidizing and 54 reducing agents in these atmospheres. Figure 1 shows a schematic diagram of the likely 55 chemistry of meteor-ablated P species. This scheme was constructed by performing high-56 level electronic structure calculations (at the CBS-QB3 level of theory⁹) of P species reacting 57 with atmospherically relevant species to determine energetically viable reaction pathways. 58 Following initial oxidation of P and PO to PO₂ (via reactions R1 and R2a), there appear to be 59 two main channels: the formation of phosphoric acid (H₃PO₄) via the species HOPO₂ (reactions R6 and R7), or the formation of the bioavailable compound phosphonic acid 60 61 (H₃PO₃) via HPO₂ (reactions R8 and R9):

63		$\Delta H^{\circ}_{(0 \mathrm{K})} (\mathrm{kJ} \mathrm{mol}^{-1})$	
64	$P(^4S) + O_2 \rightarrow PO + O$	- 100	(R1)
65	$PO + O_2 \rightarrow PO_2 + O$	- 12	(R2a)
66	$PO + O_2 (+M) \rightarrow PO_3 (+M)$	- 416	(R2b)
67	$PO_2 + O_3 \rightarrow PO_3 + O_2$	- 305	(R3)
68	$PO_3 + H \rightarrow PO_2 + OH$	- 22	(R4)
69	$PO_3 + O \rightarrow PO_2 + O_2$	- 93	(R5)
70	$PO_2 + OH (+ M) \rightarrow HOPO_2 (+ M)$	- 441	(R6)
71	$HOPO_2 + H_2O (+ M) \rightarrow H_3PO_4 (+ M)$	- 188	(R7)
72	$PO_2 + H (+ M) \rightarrow HPO_2 (+ M)$	- 333	(R8)
73	$HPO_2 + H_2O (+ M) \rightarrow H_3PO_3 (+ M)$	- 173	(R9)
74			



Figure 1. Proposed reaction scheme for the neutral chemistry of P in the upper atmosphere ofa terrestrial planet.

75

79 At present, only two reactions from this scheme have been previously investigated 80 (R1 and R2). The temperature dependence of reaction R1 was reported in a previous 81 publication from this group, in which the reactions of both ground and excited state P atoms 82 with atmospherically relevant species were investigated.⁸ Prior to this, only room temperature determinations of the rate of R1 were available, and the literature rate coefficients reported 83 disagreed by over an order of magnitude ¹⁰⁻¹³. For reaction R2 there is similar disagreement in 84 85 the literature, with four previous room temperature rate coefficients reported. Two of the studies ¹⁴⁻¹⁵ put the rate at around 2×10^{-13} cm³ molecule⁻¹ s⁻¹, while the other two studies ¹⁶⁻¹⁷ 86 87 put the rate at around 60 times faster than this, with a value of ~ 1.3×10^{-11} cm³ molecule⁻¹ s⁻¹ ¹. This poor characterization of the gas-phase chemistry of phosphorus is not only of 88 consequence to planetary atmospheres, as P-bearing compounds, including PO¹⁸⁻²⁰, have 89 90 been detected in a range of astrochemical environments. Many current chemical models are 91 using the isovalence between nitrogen and phosphorus to derive chemical rates for P-bearing 92 species, the validity of which is questionable as these models tend to significantly 93 underpredict the observed abundances of PO. Indeed, modelled and observed abundances of 94 PO in a stellar outflow, in which PO is formed from the reaction of $P + OH \rightarrow PO + H$, disagree by greater than 3 orders-of-magnitude.²¹ Understanding the oxidation chemistry of 95 96 elemental phosphorus and its oxides is also of importance to combustion chemistry, as phosphorus-containing compounds may be useful as potential fire suppression agents.²²⁻²⁴ An 97 98 understanding of the combustion of organophosphate compounds is of particular importance 99 in the destruction of chemical warfare agents.²⁵⁻²⁶

100 The PO radical has been studied extensively by spectroscopic methods, with a brief 101 history of the experimental work given by Moussaoui *et al.*²⁷. Studies investigating the $A^2\Sigma^+$ 102 $\leftarrow X^2\Pi$ transition include those by Coquart *et al.*²⁸, Sausa *et al.*¹⁷, and Wong *et al.*²⁹. There 103 have also been numerous theoretical studies, with a good overview of these given by Liu *et* 104 $al.^{30}$. These theoretical studies have provided potential energy curves for the ground and 105 many excited states of PO and also transition dipole moments for electric dipole transitions. 106 However, as the $A^2\Sigma^+$ state is a low-lying Rydberg state, it is completely missed by the *ab* 107 *initio* calculations employed in the majority of these theoretical studies, making the available 108 literature on the $A^2\Sigma^+$ state rather sparse.³¹

109 In this paper, we present results from the second part of our investigation into the 110 reactions of meteor-ablated phosphorus, reporting temperature-dependent rate coefficients for 111 reactions R2 and R3. Rate coefficients were determined using a pulsed laser photolysis-laser 112 induced fluorescence (PLP-LIF) technique, which is described in Section 2. Using the same 113 experimental setup, LIF spectra of PO over the wavelength range of 245 – 248 nm were also 114 collected. These spectra are then assigned and new spectroscopic constants for the $X^2\Pi$ and 115 $A^2\Sigma^+$ states of PO reported.

116

117 2. Experimental Procedure

118 2.1 Reaction kinetics

119 The experimental apparatus employed in this study has been discussed in detail in recent publications.^{8, 32-34} so only a brief synopsis is given here. All experiments were carried 120 121 out in a slow flow reaction cell, using a PLP-LIF technique, with detection of either PO or 122 PO₂ radicals. The reaction cell consists of a cylindrical stainless steel chamber with four orthogonal horizontal side arms, and a fifth vertical side arm. The chamber was enclosed in a 123 124 thermally insulated container, which can be operated as a furnace or filled with dry ice, 125 providing a temperature range from 190 - 800 K. Temperatures inside the reactor were 126 monitored by a shielded K-type thermocouple inserted directly into the centre of the chamber. 127 Radical precursors, reagents, and bath gases were introduced to the chamber via the 5 side 128 arms, after being combined in a mixing manifold to ensure homogeneous mixing. Radical 129 precursors were prepared as dilute mixtures of between 0.5 and 5 % in N₂ (total precursor 130 concentrations in the cell were typically around 0.1 %). Flow rates were controlled using 131 calibrated mass flow controllers (MKS instruments), with total flow rates ranging from 100 -132 400 standard cm⁻³ min⁻¹. These flow rates were sufficient to ensure a fresh flow of gas 133 through the interaction region for each photolysis laser pulse. The pressure inside the reaction 134 chamber was measured by a calibrated capacitance manometer (Baratron MKS PR 4000), 135 and controlled by a needle valve on the exit line to the pump. The photolysis and probe laser 136 beams were introduced collinearly on opposite sides of the cell, and the fluorescence signal 137 collected using a photomultiplier tube (PMT) (Electron Tubes, model 9816QB) mounted 138 orthogonally to the laser beams. To increase the solid angle of collected fluorescence, a glass 139 tube of ~1.5 cm diameter was positioned ~1 cm above the interaction region to act as a 140 waveguide to transport the fluorescence light along the vertical side arm to the PMT.

For reaction R2, PO radicals were produced from the multiphoton dissociation of POCl₃ at 248 nm (R10). For reaction R3, PO₂ radicals were produced by the photolysis of either POCl₃ or PCl₃ at 248 nm, in the presence of O₃. O₃ was produced by passing a flow of 144 O₂ though a corona discharge in an ozonizer (Edwards, E28), with a typical conversion 145 efficiency of ~ 2 %. O_3 concentrations were monitored downstream of the reactor, by either 146 UV or green light absorption spectroscopy. UV absorption spectroscopy was carried out at 147 253.7 nm from a Hg Pen-ray lamp, using a 1 m absorption cell with a PMT (Hamamatsu 148 Type H9306-13) fitted with a monochromator set at 257.3 nm (Minichrom, 300 µm slits). 149 Green light absorption spectroscopy was carried out using a multipass Herriott cell.³⁵ with a total pass length of 1.38 m. Green light at 532 nm was produced using a solid state laser, and 150 151 monitored using a photodiode detector (Thorlabs SM05PD3A) connected to digital 152 oscilloscope (LeCroy, LT262). Ozone concentrations were determined using absorption cross sections taken from the MPI-Mainz Spectral Atlas³⁶ (1.16×10^{-17} cm² at 253.7 nm and $3.0 \times$ 153 10⁻²¹ cm² at 532 nm). Corrections were made to the measured ozone concentrations to 154 155 account for photolysis by the 248 nm photolysis laser in the reaction chamber. This was done 156 by first determining the intensity of 248 nm light entering the reaction cell (as measured by a 157 Gentec UP19-VR power meter) and reaching the interaction region, after accounting for loses 158 due to windows and attenuation by O_3 in the 15 cm side-arm. Then, using the intensity of 159 light reaching the interaction region, the number of photons of 248 nm light absorbed by O_3 160 over the 1 cm interaction region was calculated. Assuming that each photon of 248 nm light absorbed led to the photo-dissociation of one O_3 molecule,³⁷ the remaining concentration of 161 162 O₃ in the interaction region could then be determined. In a typical experiment, at high ozone concentrations of ~ 6 × 10¹⁵ molecule cm⁻³, around 10 % of the O₃ within the 1 cm³ 163 interaction region was removed, while at lower ozone concentrations of ~ 6×10^{14} molecule 164 165 cm⁻³ around 20 % of the ozone was removed. To validate this method of correcting the ozone 166 concentration, the reduction in the intensity of the 248 nm light due to absorption by O₃ as it 167 passed through the entire cell was also calculated, and shown to be in good agreement with 168 measured reductions in the 248 nm light intensity exiting the cell.

169 In all experiments, the 248 nm photolysis light was generated from a KrF excimer 170 laser (Lambda Physik COMPEX 102). In experiments requiring multiphoton dissociation of 171 the precursor, the excimer beam was loosely focused using a 50 cm focal length lens, with the 172 focal point positioned approximately 10 cm beyond the centre of the reaction chamber, giving 173 a beam cross-section in the interaction region of ~ 8 mm². The KrF laser fluence at the interaction region ranged from between 10 - 20 mJ cm⁻² for experiments in which no 174 focusing was required, and from between 270 - 650 mJ cm⁻² for experiments in which 175 176 focusing was required. PO and PO₂ radicals were observed by time-resolved LIF 177 spectroscopy, using the frequency doubled output of a Nd:YAG pumped dye laser (a Quantel Q-smart 850 pumping a Sirah Cobra-Stretch with a BBO doubling crystal). PO($X^2\Pi$) was 178 probed using the $A^2\Sigma^+$ (v'=0) $\leftarrow X^2\Pi$ (v''=0) transition at ~ 246.3 nm.¹⁷ The non-resonant 179 180 (v',v'', 0,1) fluorescence at ~ 255.4 nm was collected using a PMT fitted with an appropriate filter (Laser2000 Brightline Bandpass filter, $\lambda_{max} - 255$ nm, fwhm - 8 nm), and recorded 181 182 using a digital oscilloscope (LeCroy, LT262). PO₂ was probed using either the $B^{2}B_{1}$ (2,7,0) $\leftarrow X^{2}A_{1}$ (0,0,0) transition at ~ 286.1 nm or the $B^{2}B_{1}$ (4,2,0) $\leftarrow X^{2}A_{1}$ (0,0,0) at 287.0 nm, 183 184 and the non-resonant fluorescence at $\lambda > 350$ nm discriminated using a 350 nm cut on filter. 185 The temporal evolution of the LIF signal was recorded by varying the time delay between the 186 photolysis and probe laser. A typical time-revolved LIF profile (Figures 2 and 4) consisted of 187 between 150 and 250 delay steps, and resulted from the average of between 5 and 10188 individual delay scans.

189 In the experiments monitoring LIF from PO₂, the raw PMT signal needed to be 190 corrected for chemiluminescence produced following the photolysis of POCl₃ and PCl₃ in the 191 presence of O₃. Although this chemiluminescence has been observed in other studies involving oxygen-phosphorus systems,³⁸⁻⁴¹ the precise reactions producing it are unknown, 192 although two studies^{38, 41} do suggest it arises from PO₂ itself. Indeed the observed loss of the 193 194 chemiluminescence signal with $[O_3]$ was consistent with the observed loss of the PO₂ LIF signal with $[O_3]$, suggesting PO_2^* as a possible source. However, of the possible reactions 195 196 forming PO₂ explored in section 3.2 below, only the reaction between PO and O₃ is 197 sufficiently exothermic to populate the first excited state of PO₂, which lies ~ 250 kJ mol⁻¹ above the ground state.⁴² As PO may only be produced in experiments using PCl₃ as a 198 199 precursor (see scheme 3 in section 3.2), this reaction cannot explain the chemiluminescence 200 observed when using POCl₃ as a precursor. Furthermore, the observed growth of the 201 chemiluminescence signal with [O₃] was typically only around half the observed growth of 202 the PO₂ LIF signal with [O₃], suggesting a species other than PO₂ as the source of the 203 chemiluminescence. As the exact nature of the chemiluminescence signal is uncertain, its 204 contribution to the raw PMT signal was removed. To do this, back-to-back experiments were 205 conducted, one with the probe laser on, and one with the probe laser off (Figure 4a). In the 206 experiment with the probe laser off, all signal measured by the PMT results from 207 chemiluminescence initiated by the photolysis laser. In the experiment with the probe laser 208 on, the measured signal contained both chemiluminescence initiated by the probe laser, and 209 signal from the LIF of PO₂. The PO₂ LIF signal (Figure 4) is then obtained by subtraction. 210 Note that although chemiluminescence would also have been produced in the experiments 211 monitoring PO LIF signal, no correction to the raw PMT signal was required as the 212 chemiluminescence was effectively excluded by the interference filter used in those 213 experiments.

214

215 2.2 PO LIF spectrum

216 PO LIF spectra were collected using the same experimental setup as the kinetics 217 experiments. PO radicals were generated either by the reaction with O₂ of ground- and 218 excited-state P atoms produced from the multiphoton dissociation of PCl₃ or PBr₃ at 248 nm, 219 or directly from the multiphoton dissociation of POCl₃ at 248 nm. These three regimes were 220 used to ensure a positive identification of PO. In the experiments in which LIF spectra were 221 collected, the delay between the photolysis and probe lasers was fixed, while the probe laser 222 wavelength was linearly scanned between 245 and 248 nm. In this manner, each new laser 223 pulse produced the PO LIF signal at a new wavelength, rather than at a later delay time. The pulse width of the probe laser employed is 6 ns, with a linewidth of 0.003 nm. The output of 224 225 the probe laser was calibrated using a wavemeter (Bristol Wavemeter 871).

226

227 Materials

He (99.999 %, BOC), N₂ (99.9995 %, BOC), O₂ (99.999 %, BOC), CO₂ (99.999 %, BOC) were used without further purification. PCl₃ (\geq 99.0 %, VWR), POCl₃ (99 %, Sigma Aldrich), dimethyl methylphosphonate (DMMP) (97 %, Sigma Aldrich) were initially degassed by freeze-pump-thawing to remove volatile contaminants, and then made up as dilute vapours in N₂ or He.

233

3. Results

235 3.1 $PO + O_2$

We initially thought that the rate of the reaction between PO and O₂ (R2) could be determined from the PO LIF profiles produced following PLP of PCl₃ in the presence of O₂, via the following reaction scheme:

 $P(^{4}S) + O_{2} \rightarrow PO + O \tag{R1}$

$$PO + O_2 \rightarrow PO_2 + O \tag{R2a}$$

However, as discussed previously,⁸ the multiphoton dissociation of PCl₃ produces significant 241 amounts of the first two excited states of P (the ²D and ²P states, hereafter collectively 242 referred to as P^*). The formation of P^* together with ground state $P(^4S)$ complicated the 243 244 reaction scheme so that information regarding the removal of PO could not be extracted. 245 Therefore, in order to measure the rate of reaction R2, a photolytic source of PO was used. Previous studies into the rate of reaction R2 employed DMMP as a PO precursor.¹⁵⁻¹⁷ 246 247 However, we observed that photolysis of DMMP at 248 nm produced substantial amounts of 248 P*. Indeed, photolysis of DMMP produced so much $P(^{2}P)$ that we were able to use it as precursor when measuring the rate of $P(^{2}P)$ with O_{2} .⁸ The formation of P^{*} from the 249 250 multiphoton dissociation of DMMP resulted in the reaction scheme for PO formation and 251 removal becoming over complicated, making information regarding the removal of PO 252 difficult to extract. This led us to investigate a cleaner source of photolytic PO. Tests on the 253 PLP of phosphoryl chloride, POCl₃, indicated that it was a good source of PO, and did not 254 produce any evidence of P* formation, even when using concentrations of POCl₃ and 255 photolysis energies substantially higher than required for the kinetics experiments. We cannot rule out the significant formation of ground state $P(^4S)$, as we are unable to observe this 256 species directly. However, as the formation of PO from the reaction of $P(^4S)$ and O_2 (R1) 257 258 would occur on a much longer timescale than the removal of the photolytically produced PO with O₂,¹⁶ this would not affect the measurement of the rate of R2. Numerical simulations of 259 260 the system (using the chemical kinetics software Kintecus⁴³), involving reactions R1 and R2, 261 indicate that 5 times more P(⁴S) than PO would need to be produced from the photolysis of 262 POCl₃ in order to reduce the measured rate of R2 by only 6 %, a scenario that is unlikely to occur due to the large amount of energy required to produce $P(^{4}S)$ from POCl₃. 263

A typical PO LIF profile produced following PLP of POCl₃ in the presence of O_2 is shown in Figure 2, which demonstrates a small growth of the PO LIF signal at short times.. We attribute this small growth to the formation of rotationally excited PO^{*} during photolysis, which is then rapidly relaxed to the chamber temperature by the N₂ bath gas. No consideration of vibrational relaxation is required, firstly as vibrational relaxation has been shown to be much slower than reactive removal of PO (v = 0),¹⁶ and secondly as PO LIF spectra collected when using POCl₃ as a precursor indicate no vibrationally excited (v = 1)PO is formed (Figure 6). Thus, the reaction scheme for the formation and removal of PO (v = 0) is given by:

$$POCl_3 + hv \rightarrow PO, PO^* + co-products$$
 (R10)

274
$$\operatorname{PO}^*(+\operatorname{N}_2) \xrightarrow{\kappa_{\operatorname{rel}}} \operatorname{PO}$$
 (R11)

275
$$PO + O_2 \xrightarrow{k_2} loss$$
 (R2)

where PO^{*} is an initially rotationally excited PO molecule formed from the photolysis of POCl₃, k_{rel} is the combined rate coefficient for relaxation of any rotationally excited PO, and k_2 is the bimolecular rate coefficient for the reaction of ground state PO with O₂. As experiments were carried out under pseudo-first order conditions ([PO] « [O₂]), the temporal evolution of the PO is given by:

281
$$[PO]_{t} = \left(\frac{k_{\text{rel}}}{k' - k_{\text{rel}}}\right) [PO^{*}]_{0} \left(e^{-k_{\text{rel}} \cdot t} - e^{-k' \cdot t}\right) + [PO]_{0} e^{-k' \cdot t}$$
(E1)

282 and

283
$$k' = k_2[R] + k'_{diff}$$
 (E2)

284 where [PO]₀ and [PO^{*}]₀ are the initial concentrations of PO and PO^{*} produced from reaction 285 R10 respectively, t is the time delay between the photolysis and probe lasers, k' is the pseudo-286 first order rate constant for the removal of PO (i.e., $k_2[O_2]$), and k'_{diff} is the total rate 287 coefficient for other minor loss processes of PO (such as diffusion out of the probe laser 288 beam and reaction with the POCl₃ precursor). Equation E1 was fitted to the PO profiles to extract the parameters k_{rel} , k', [PO]₀, and [PO^{*}]₀. A plot of k' vs [O₂] then yields a straight line 289 290 of gradient k_2 and intercept k'_{diff} . The inset in Figure 3 shows an example of such a plot, and 291 the small intercept (relative to the total removal rate) demonstrates that reaction with O₂ 292 dominates PO removal. For a given experimental run (in which the temperature and total 293 pressure are kept constant), the parameter $k_{\rm rel}$ was found to be effectively independent of $[O_2]$, over the range employed (typically $0 - 2 \times 10^{15}$ molecule cm⁻³). This allowed us to do a 294 global fit of the PO traces from an experimental run, with a single fitted value for k_{rel} . For 295 296 different experimental runs, k_{rel} varied linearly with pressure (see Table 1), as would be 297 expected of rotational relaxation processes. When fitting a biexponential growth and loss, if 298 the growth and loss rates are approaching one another, the fitting procedure is sometimes 299 unable to find a unique solution and returns a result in which the growth and loss rates are 300 equal, meaning one rate has been artificially inflated, and the other decreased. This can be a 301 source of error when analysing biexponential profiles. To minimise the possibility of this 302 affecting our results, for a given pressure (and thus k_{rel}), the [O₂] range was selected to ensure 303 that, even at the highest $[O_2]$, the removal rate was around 4 times lower than the growth rate.



Figure 2. PO LIF signal following PLP of POCl₃ at a total pressure of 19.8 Torr and $[O_2] = 307$ 8.35 × 10¹⁴ molecule cm⁻³, at T = 335 K. Inset: a bimolecular plot for reaction R3 at T = 3358.35 K, giving $k_2 = (1.44 \pm 0.03) \times 10^{-11}$ cm³ molecule⁻¹ s⁻¹.

309

310 The bi-molecular rate coefficients for the reaction of PO with O₂ (R2) are presented as 311 a function of temperature in Table 1 and Figure 3a, and compared with available literature. 312 No effects were observed on the bimolecular rate coefficients determined in this study as the 313 radical concentration and photolysis energy were varied by around a factor of 2. The temperature range over which R2 could be studied was limited by thermal decomposition of 314 315 the POCl₃ precursor at temperatures higher than ~ 350 K, which resulted in a large 316 background signal of PO, and a PO removal rate with O_2 that appeared to be ~ 1000 times 317 slower than that measured at ~ 340 K. At the lowest temperature that we investigated (~ 191 318 K), loss of PO signal indicated significant freezing out of POCl₃ on the walls of the reaction 319 chamber. As PO fluorescence is also quenched by the N₂ bath gas, we were unable to go 320 above pressures of ~ 6 Torr at this temperature, due to loss of the PO signal. At temperatures 321 of ~ 294 and 338 K, rate coefficients were measured over the pressure range of ~ 10 - 40322 Torr. At both these temperatures, plotting the biexponential rate coefficients vs the total bath 323 gas concentration (Figures 3b and 3c), suggests evidence for a small positive pressure 324 dependence, which would be due to a three-body removal channel:

325
$$PO + O_2 (+M) \rightarrow PO_3$$
 $\Delta H^{\circ}_{(0 \text{ K})} (\text{kJ mol}^{-1}) = -416$ (R2b)

326 An Arrhenius fit to the experimental data (dotted line Figure 3a) yields k_2 (PO + O₂, 191 – 327 339 K) = (1.91 ± 0.33) × 10⁻¹¹ × exp^{(-84 ± 52)/T}. This fit was determined after increasing the 328 errors at temperatures 294 and 338 K to 10 % of their mean value to avoid the fit being 329 overly weighted to some of the rate coefficients with smaller statistical errors.

In addition to monitoring the loss of PO, experiments were also carried out in which the growth of the PO₂ product from R2 was monitored. These experiments confirmed PO₂ as a product of reaction R2, and the observed rise of PO₂ was consistent with the measured rate of removal of PO with O₂. However, due to the poor signal to noise of these experiments, no rate coefficients for reaction R2 were determined monitoring the growth of the PO₂ product.

335



Figure 3. Experimentally determined rate coefficients for the reaction of PO + O₂ vs temperature (a). Pressure dependence of the PO + O₂ reaction at: b) T = 294 K and c) T = 337K. Open symbol are experimental data from: this study (black downward facing triangles); Long *et al.*¹⁶ (dark blue pentagon); Sausa *et al.*¹⁷ (dark green star). Solid symbols and lines are rates calculated by MESMER⁴⁴ for: PO + O₂ \rightarrow PO₂ + O (channel R2a, red squares); PO + O₂ (+N₂) \rightarrow PO₃ (channel R2b, blue diamonds); total reaction PO + O₂ (R2, green upward

facing triangles). The dotted line is an Arrhenius fit to the experimental data. At temperatures
of 294 and 337 K, experimental errors have been increased to 10 % of their mean value to
avoid any one point being overly weighted.

346

347 Ta	ble 1. Rate coeffi	eients for the reaction	n of PO + O_2 and z	relevant experimental conditions
--------	--------------------	-------------------------	-------------------------	----------------------------------

						-	
348	(errors are s	tatistical a	t the 1σ level).	*Literature values	taken from:	^{<i>a</i>} Aleksandrov <i>et al</i> . ¹⁴	, full
040	1 11 0	•	· · · · · br · ·	115 60	117 dr	1 16	

detail of experiment unavailable; ^bWong et al.¹⁵; ^cSausa et al.¹⁷; ^dLong et al.¹⁶.

Reaction	Precursor	Excitation λ	<i>T</i> (K)	Bath gas	Pressure (Torr)	$k_2 (\times 10^{-11} \text{ cm}^3 \text{ molecule}^{-1} \text{ s}^{-1})$	$k'_{\rm diff}~({ m s}^{-1})$	$k_{\rm rel} (\times 10^3 {\rm s}^{-1})$
$PO + O_2$	POCl ₃	246.3	191	N ₂	5.8	1.03 ± 0.26	1279 ± 891	57 ± 11
	POCl ₃	246.3	192	N ₂	6.1	1.29 ± 0.28	791 ± 246	44 ± 3
	POCl ₃	246.3	295	N ₂	10.2	1.45 ± 0.07	1943 ± 103	62 ± 5
	POCl ₃	246.3	295	N ₂	18.9	1.34 ± 0.06	2640 ± 169	89 ± 10
	POCl ₃	246.3	293	N ₂	19.2	1.47 ± 0.09	2692 ± 196	81 ± 2
	POCl ₃	246.3	295	N ₂	20.7	1.46 ± 0.06	2509 ± 122	88 ± 8
	POCl ₃	246.3	295	N_2	40.2	1.44 ± 0.07	3989 ± 265	183 ± 30
	POCl ₃	246.3	293	N_2	42.2	1.55 ± 0.10	2161 ± 2285	187 ± 26
*a			300			0.022 ± 0.010		
*b	DMMP	~ 325	303	Ar	~ 2.4	$0.02\substack{+0.02\\-0.01}$		
*c	DMMP	246.3	298	O ₂	~ 10	1.2 ± 0.2		
*d	DMMP	246.3	298	O ₂	~ 1.5	1.4 ± 0.2		
	POCl ₃	246.3	339	N_2	10.5	1.35 ± 0.05	1935 ± 110	53 ± 4
	POCl ₃	246.3	335	N_2	19.8	1.44 ± 0.03	2296 ± 91	77 ± 8
	POCl ₃	246.3	337	N_2	22.8	1.33 ± 0.07	1814 ± 134	99 ± 17
	POCl ₃	246.3	337	N_2	38.4	1.65 ± 0.10	3224 ±281	132 ± 20

350

351 3.2 $PO_2 + O_3$

352 A typical PO₂ LIF profile produced following the photolysis of POCl₃ in the presence 353 of O₃ can be seen in Figure 4b. Unlike the experiments monitoring PO fluorescence, the 248 354 nm photolysis laser did not need to be focused in order to observe PO₂, suggesting only a single photon is required to initiate the reaction leading to PO₂. Energetically, a single 248 355 356 nm photon has enough energy to remove either 1 Cl atom (R12), or 2 atoms as Cl₂, from POCl₃. However, using the time-dependent density functional (TD-DFT) excited states 357 method,⁴⁵ within the Gaussian 16 suite of programs,⁴⁶ shows that at the TD//B3LYP/6-358 359 311+g(2d,p) level of theory, vertical excitation of POCl₃ to its first excited state (relevant for 360 a 248 nm photon), and then allowing the molecule to relax, there is clear dissociation to 361 $POCl_2 + Cl$. Thus, we propose the following reaction scheme for the formation of PO_2 :

362
$$POCl_3 + hv \rightarrow POCl_2 + Cl$$
 $\Delta H^{\circ}_{(0 \text{ K})} (\text{kJ mol}^{-1}) = +365$ R12

363 POCl₂ + O₃
$$\xrightarrow{k_{13}}$$
 PO₂ + Cl₂ + O₂ $\Delta H^{\circ}_{(0 \text{ K})}$ (kJ mol⁻¹) = -127 R13

where k_{13} is the bimolecular rate coefficient for the reaction of POCl₂ with O₃. Together with reaction for the removal of PO₂ (R3):

and as experiments were carried out under pseudo-first order conditions ([PO₂] « [O₃]), the
 temporal evolution of PO₂ is given by:

369
$$[PO_2]_t = \left(\frac{k'_{growth}}{k'_{loss} - k'_{growth}}\right) [POCl_2]_0 \left(e^{-k'_{growth} \cdot t} - e^{-k'_{loss} \cdot t}\right)$$
(E3)

370 where k'_{growth} and k'_{loss} are the pseudo-first order rate coefficients for the reactions producing 371 and removing PO₂, and [POCl₂]₀ is the initial amount of POCl₂ formed following photolysis 372 of POCl₃. Equation E3 was fitted to the PO₂ profiles (Figure 4b) and the parameters k'_{growth} 373 and k'_{loss} , and [POCl₂]₀ extracted. Plots of the parameters k'_{growth} and k'_{loss} should then yielded 374 straight lines, with gradients equal to the bimolecular rate coefficients for reactions R3 and R13 (Equation E2). As we have discussed previously,⁸ when analysing biexponential traces 375 376 some prior knowledge about the system is required to determine whether k'_{growth} , and thus the 377 plot of k'_{growth} vs [O₃], relates to the reaction producing or removing PO₂, and vice versa for 378 $k'_{\rm loss}$ i.e., we need to know which reaction is faster than the other one. As there have been no 379 previous studies of either reaction R3 or R13, this problem was solved by carrying out some 380 additional experiments using PCl₃ as a precursor. This second set of experiments would again 381 yield two bimolecular rate coefficients, one for the production and one for the removal of 382 PO₂. Comparing the two pairs of rate coefficients, one rate coefficient (for the removal of 383 PO_2 by O_3 , R3) should be common to both and thus identifiable.



Figure 4. a) PMT signal detected following the photolysis of POCl₃ in the presence of O₃, both with (black circles), and without (red squares) the probe laser. The PO₂ LIF signal (green triangles) is then the difference of the two traces. b) PO₂ LIF signal following PLP of POCl₃ precursor at a total pressure of 10.6 Torr and $[O_3] = 1.6 \times 10^{15}$ molecule cm⁻³, at T =

390 337 K. Inset: bimolecular plot for reaction R3 at T = 337 K, giving $k_3 = (9.32 \pm 1.37) \times 10^{-13}$ 391 cm³ molecule⁻¹ s⁻¹. c) PO₂ LIF signal following PLP of PCl₃ precursor at a total pressure of 392 9.36 Torr and $[O_3] = 3.8 \times 10^{15}$ molecule cm⁻³, at T = 292 K. Inset: bimolecular plot for 393 reaction R3 at T = 291 K, giving $k_3 = (9.51 \pm 2.79) \times 10^{-13}$ cm³ molecule⁻¹ s⁻¹.

394

395 A typical PO₂ LIF profile produced following the photolysis of PCl₃ in the presence 396 of O₃ can be seen in Figure 4c. As with the experiments conducted using POCl₃ as a 397 precursor, the 248 nm photolysis laser did not need to be focused in order to observe PO₂, 398 suggesting that only a single photon is required to initiate the reaction leading to PO₂. 399 Energetically, a single 248 nm photon has enough energy to remove either 1 Cl atom, or 2 Cl 400 atoms as Cl₂, from PCl₃. The first excited state of PCl₃ is 480 kJ mol⁻¹ vertically above the ground state (at the TD//B3LYP/6-311+g(2d,p) level of theory,⁹) which is very close to the 401 402 energy of a single 248 nm photon. A relaxed scan of the potential surface of this excited state 403 shows that it should dissociate to $PCl_2 + Cl$. We therefore propose one of the two following 404 reaction schemes for the formation of PO₂ when using PCl₃ as a precursor:

405

Scheme 1		Scheme 2	
$PCl_3 + hv \rightarrow PCl_2 + Cl$	+ 325 kJ mol ⁻¹	$PCl_3 + hv \rightarrow PCl_2 + Cl$	+ 325 kJ mol ⁻¹
$PCl_2 + O_3 \rightarrow PO + Cl_2 + O_2$	- 96 kJ mol ⁻¹	$PCl_2 + O_3 \rightarrow POCl + Cl + O_2$	- 216 kJ mol ⁻¹
$PO + O_3 \rightarrow PO_2 + O_2$	- 414 kJ mol ⁻¹	$POCl + O_3 \rightarrow PO_2 + Cl + O_2$	- 43 kJ mol ⁻¹

406

407 From our experiments we are unable to tell definitively whether one or the other (or both) of 408 these schemes forms the observed PO₂. Nevertheless, the PO₂ traces can be satisfactorily fit 409 using a biexponential (Figure 4c), where the formation of PO₂ is treated as a single process, 410 described by the term k_{PO_2} . With the substitution of the [POCl₂]₀ term with the term [PCl₂]₀ 411 (the initial amount of PCl₂ formed following photolysis of PCl₃), we were able to fit the PO₂ 412 traces using equation E3 and extract the parameters, k'_{growth} , k'_{loss} , and [PCl₂]₀. Plots of k'_{growth} 413 and k'_{loss} vs [O₃] then yielded straight lines with gradients equal to the bimolecular rate 414 coefficients k_{PO_2} and k_3 , and intercepts k'_{diff} (equation E2).

415 Figure 5 shows the rate coefficients obtained when using both POCl₃ and PCl₃ as a precursor. At $T \sim 293$ K the rate coefficients obtained by plotting $k'_{\text{growth}} vs$ [O₃] for the 416 417 different precursors differ significantly (Figure 5a), ranging from $(7.11 \pm 0.74) \times 10^{-12} \text{ cm}^3$ 418 molecule⁻¹ s⁻¹ for POCl₃ to only $(1.54 \pm 0.37) \times 10^{-12}$ cm³ molecule⁻¹ s⁻¹ for PCl₃. In contrast, there is quite good agreement between the rate coefficients obtained by plotting $k'_{loss} vs$ [O₃] 419 for the different precursors (Figure 5b): $(8.93 \pm 0.19) \times 10^{-13} \text{ cm}^3 \text{ molecule}^{-1} \text{ s}^{-1}$ for POCl₃ vs 420 $(7.58 \pm 0.50) \times 10^{-13}$ cm³ molecule⁻¹ s⁻¹ for PCl₃. This implies that k'_{growth} describes the 421 422 reactions forming PO₂, while k'_{loss} describes the reaction removing PO₂, R3. With this 423 assignment, k_3 was determined over a range of temperatures and pressures (Figure 5 and 424 Table 2). k_3 did not vary significantly as radical concentrations and photolysis energies were 425 varied by a factor of ~2, or pressures varied by a factor of ~6. The upper temperature limit at 426 which k_3 could be studied was set by the thermal decomposition of O₃ on the reactor walls at 427 temperatures higher than 380 K. An Arrhenius fit (dotted line in Figure 5b) to the experimental data yields $k_3(PO_2 + O_3, 186 - 339 \text{ K}) = (2.2 \pm 0.7) \times 10^{-11} \exp^{(-990 \pm 81)/T} \text{ cm}^3$ 428 429 molecule⁻¹ s⁻¹.

430 A possible source of error in the determination of k_3 is interference from O atoms 431 produced from the photolysis of O₃. These O atoms may react with either the precursors or 432 other photolysis products to form PO₂. As discussed above, between 10 and 20 % of the O₃ at 433 the interaction region is photolysed by the 248 nm photolysis laser. Photolysis of O₃ at 248 434 nm produces either excited state $O(^{1}D) + O_{2}$ with a quantum yield of 0.9, or ground state $O(^{3}P) + O_{2}$ with a quantum yield of $0.1.^{37}$ The excited $O(^{1}D)$ atoms will be rapidly relaxed 435 436 down to ground state $O({}^{3}P)$ by collisions with the N₂ bath gas. This relaxation occurs on a 437 significantly faster timescale than the observed growth of PO₂ when using either precursor (~ 438 200 times faster at a pressure of 5 Torr), suggesting that reaction of $O(^{1}D)$, either with the 439 precursors or other photolysis products, is not a significant source of PO₂. With regards to 440 reaction of the relaxed $O({}^{3}P)$ with the precursors, with POCl₃ the reaction to PO₂ + Cl₂ + Cl is 441 thermodynamically unfavourable, being endothermic by ~ 120 kJ mol^{-1} . In contrast, for the 442 reaction of $O({}^{3}P)$ with POCl₂ there are two energetically viable reactions:

443
$$POCl_2 + O({}^{3}P) \rightarrow PO_2 + Cl_2$$
 $\Delta H^{\circ}_{(0 \text{ K})} (\text{kJ mol}^{-1}) = -226$ R14a

$$POC_{12} + O(P) \rightarrow PO_2 + C_{12} \qquad \Delta H (0 \text{ K}) (\text{KJ IIIOI}) = -220 \qquad \text{K14a}$$

 $POCl_2 + O(^{3}P) \rightarrow PO_2Cl + Cl$ $\Delta H^{\circ}_{(0 \text{ K})}$ (kJ mol⁻¹) = - 317 R14b

445 The more exothermic channel (R14b), which also involves a single bond cleavage, should be 446 favoured, suggesting the contribution of reaction R14a to the PO₂ observed in our experiments is likely to minor. However, as we are not able to unequivocally rule out 447 contribution from reaction R14a, numerical simulations using Kintecus⁴³ have been 448 449 conducted, in which for a range of experimentally applicable [O₃] and [O], the rates of R13 450 and R14a were varied and their effect on the determined rate of R3 evaluated. In these simulations, k_{13} was varied between 0 and 8 × 10⁻¹² cm³ molecule⁻¹ s⁻¹, while k_{14a} was varied 451 from 0 up to 1×10^{-10} cm³ molecule⁻¹ s⁻¹ (using values higher than these upper limits results 452 453 in the gradient of a plot of $k'_{\text{growth}} vs$ [O₃] being larger than that observed experimentally). The 454 numerical simulations showed that the rate coefficient determined for R3 was unaffected by 455 varying either k_{13} or k_{14a} between the ranges given. Thus, the occurrence of the side reaction 456 R14a producing PO₂ in our experiments would not have affected the value determined for k_3 . 457 The reason for this is due to the approximately linear relationship between [O₃] and [O] in 458 our experiments (at high $[O_3]$ the percentage of O_3 photolysis at the interaction region is 459 smaller than at low [O₃], so the relationship is not precisely linear). Whether PO₂ is produced 460 from the reaction of POCl₂ with O_3 (R13) or from the reaction of POCl₂ with O(³P) (R14a), 461 or from a combination of both, as a doubling of the concentration of O₃ will be mirrored by 462 an approximately doubling of the concentration of $O({}^{3}P)$, the rate of PO₂ formation will also 463 be doubled. As such, we are able to fit a biexponential (rather than a tri- or higher 464 exponential) to our PO₂ LIF traces to extract the pseudo-first order rate coefficients, and the 465 value determined for k_3 is unaffected.

466 With regards to the reaction of $O({}^{3}P)$ with the PCl₃ precursor, there are energetically 467 viable pathways to POCl₂ and POCl, both of which could go on to react with O₃ or another 468 $O(^{3}P)$ to produce PO₂. However, there are two reasons we can rule out $O(^{3}P)$ reacting with 469 the PCl₃ precursor as a source of PO₂. First, if this were the case, as PCl₃, $O(^{3}P)$, and O₃ are 470 all in excess in our experiments, we would expect to see vastly more PO₂ LIF signal in 471 experiments using PCl_3 as a precursor than in experiments using $POCl_3$, in which PO_2 is only 472 produced from reaction of the deficient POCl₂ photolysis product. Second, if a reaction 473 between $O({}^{3}P)$ and PCl₃ was generating PO₂, we would expect the PO₂ LIF signal to be 474 directly proportional to $[O(^{3}P)]$ (and thus $[O_{3}]$). As neither of these cases are true, the reaction 475 between $O({}^{3}P)$ and PCl_{3} can be ruled out as a source of PO₂. For the reaction of $O({}^{3}P)$ with 476 PCl₂, there are energetically viable product channels, producing either PO or POCl. This 477 leads us to suggest a further two reaction schemes for the formation of PO₂ when using PCl₃ 478 as a precursor:

479

Scheme 3		Scheme 4	
$PCl_3 + h\nu \rightarrow PCl_2 + Cl$	+ 325 kJ mol ⁻¹	$PCl_3 + hv \rightarrow PCl_2 + Cl$	+ 325 kJ mol ⁻¹
$PCl_2 + O(^{3}P) \rightarrow PO + Cl$	- 195 kJ mol ⁻¹	$PCl_2 + O(^{3}P) \rightarrow POCl + Cl$	- 315 kJ mol ⁻¹
$PO + O_3 \rightarrow PO_2 + O_2$	- 414 kJ mol ⁻¹	$POCl + O(^{3}P) \rightarrow PO_{2} + Cl$	- 142 kJ mol ⁻¹

480

481 From our experiments we are unable to tell which of the four reaction schemes involving 482 PCl₂ reacting with O_3 or $O({}^{3}P)$, or combination thereof, are generating PO₂ in our 483 experiments. Nevertheless, as stated above, the PO₂ traces were satisfactorily fit using the 484 biexponential given in Equation E3 (Figure 4c). Furthermore, the numerical simulations on the POCl₂ with O_3 and $O({}^{3}P)$ system indicated that whether PO₂ is generated from reactions 485 486 of O_3 or $O({}^{3}P)$, the rate coefficient determined for R3 was unaffected due to the 487 approximately linear relationship between $[O_3]$ and $[O(^{3}P)]$. As such, we expect any 488 contribution from schemes 3 and 4 to the production of PO_2 to have little effect on k_3 .



491 Figure 5. Bimolecular rate coefficients determined by plotting: a) $k'_{\text{growth}} vs$ [O₃] and 492 determined to relate to reactions R13 and/or R14b (for POCl₃ as a precursor), and the 493 summed processes forming PO₂ (for PCl₃ as a precursor); and b) $k'_{\text{loss}} vs$ [O₃] and determined 494 to relate to reaction R3. Black circles are experiments using POCl₃ as a precursor, while red 495 triangles are those using PCl₃ as a precursor. The dotted line is an Arrhenius fit.

Table 2. Rate coefficients for reaction of $PO_2 + O_3$ and relevant experimental conditions (uncertainties reported at the 1 σ level for the linear least-squares fitting of the pseudo firstorder coefficients as a function of $[O_3]$)

Reaction	Precursor	Excitation λ	<i>T</i> (K)	Pressure (Torr)	$k_3 (\times 10^{-13} \text{ cm}^3 \text{ molecule}^{-1} \text{ s}^{-1})$
$PO_2 + O_3$	POCl ₃	286.1	188	4.49	0.82 ± 0.15
	POC1 ₃	286.1	191	5.48	0.47 ± 0.15
	POC1 ₃	286.1	186	8.22	0.96 ± 0.17
	POC1 ₃	286.1	195	10.48	1.90 ± 0.28
	POC1 ₃	286.1	193	19.91	0.85 ± 0.44
	POCl ₃	287.0	295	5.26	9.32 ± 1.37
	POC1 ₃	286.1	294	10.18	8.50 ± 1.05
	POC1 ₃	286.1	295	13.1	9.05 ± 1.01
	POC1 ₃	287.0	293	30.0	8.26 ± 1.14
	POCl ₃	286.1	293	30.2	9.10 ± 1.10
	POCl ₃	287.0	295	32.4	9.48 ± 1.09
	PCl ₃	287.0	291	6.69	6.07 ± 1.85
	PCl ₃	287.0	292	9.39	9.85 ± 3.86
	PCl ₃	287.0	291	10.37	9.51 ± 2.79
	PCl ₃	287.0	292	15.5	6.73 ± 7.57
	POCl ₃	286.1	337	5.14	9.07 ± 2.17
	POCl ₃	287.0	339	9.45	9.62 ± 4.22
	POC1 ₃	286.1	337	10.6	9.04 ± 1.16
	POCl ₃	287.0	337	16.27	8.39 ± 8.95

501 3.2 PO LIF Spectra

502 LIF spectra of PO, collected using different PO production regimes, are shown in 503 Figure 6. The good agreement between spectra producing using different precursors allows 504 for a positive identification of PO, while comparison of line positions to earlier emission and fluorescence studies allows for positive identification of the A-X system.^{17, 28-29, 47} The spectra 505 506 have been baseline corrected to remove any contribution from scattered probe laser light. No 507 corrections were made to account for laser dye efficiency, or the spectral efficiency of the 508 PMT, as neither of these change significantly over the narrow spectral range of the 509 measurements.

510 Figure 6 illustrates that there is a significant population of vibrationally excited PO (v 511 = 1) when using PCl₃ or PBr₃ as a precursor, indicating that the reaction of $P^* + O_2$ (and 512 possibly $P(^{4}S) + O_{2}$) produces both ground and vibrationally excited PO (v = 1). As with the 513 detection of PO (v = 0), PO (v = 1) is also detected by non-resonant fluorescence via the (v', v''514 1,2) transition at ~ 253.6 nm. This is due to the resonant fluorescence being excluded by our 515 interference filter. When using POCl₃ as a precursor, we see no evidence of vibrationally 516 excited PO, even at very short delay times (30 µs), indicating that the small growth in PO 517 signal observed in the kinetic traces of PO (Figure 2) is due to rotational relaxation, rather 518 than relaxation of higher vibrational states.



Figure 6. PO LIF spectra taken using three different precursors, together with labels indicating the $X^2\Pi_{3/2}$ and $X^2\Pi_{1/2}$ sub-bands, and the upper and lower vibrational levels: Blue upper trace – PCl₃; red middle trace – PBr₃; grey lower trace – POCl₃.

524

525 4. Discussion

526 4.1 Comparison with previous work

527 The bimolecular rate coefficients for the removal of PO with O₂ (R2) determined in this study are compared with the literature values in Table 1 and Figure 3. As discussed 528 529 above, there is significant discrepancy in the literature over the room temperature rate of reaction R2, with two studies reporting a rate constant of ~ 2×10^{-13} cm³ molecule⁻¹ s⁻¹, ¹⁴⁻¹⁵ 530 and two others a rate constant 60 times faster, of ~ 1.3×10^{-11} cm³ molecule⁻¹ s⁻¹.¹⁶⁻¹⁷ Taking 531 532 an average of our room temperature measurements, we determine a value of $k_2 = (1.44 \pm$ $(0.11) \times 10^{-11} \text{ cm}^3 \text{ molecule}^{-1} \text{ s}^{-1}$ at T = 294 K, putting the rate towards the top end of the 533 available literature, in good agreement with the value reported by Long *et al.*¹⁶ and lying 534 535 slightly above (just outside mutual error limits) the value reported by Sausa et al.¹⁷. The 536 probable reason for the large discrepancy between the measured values of k_2 is the different 537 experimental techniques employed. The two lower values were measured in fast flow tubes, 538 using a microwave discharge of DMMP to produce PO, the loss of which was then monitored using LIF. The two larger rate coefficients were measured by the PLP-LIF technique, similar 539 540 to that employed in this study, with the exception that DMMP rather than POCl₃ was used as the PO precursor. As in the present work, Sausa et al.¹⁷ used a KrF excimer laser (248 nm) as 541 542 the photolysis source. They discuss in their paper that the focused KrF radiation may result in 543 the dissociation of O_2 , so that they observed the reaction of PO with O rather than $PO + O_2$. 544 Indeed they do note a significant deviation from exponential decay of PO when the photolysis source is a shorter wavelength ArF laser (193 nm), which would have generated a much 545 546 larger fraction of O atoms. However, interference from O atoms can in fact be ruled out,

because Long *et al.*¹⁶ produced PO from the infrared multiphoton dissociation of DMMP, a method which cannot generate O atoms from O₂, and they obtained a value for k_2 in good agreement with that obtained with the KrF lasers employed in the present study and that by Sausa *et al.*¹⁷.

551 Another possible source of error in measurements of k_2 is interference from P atoms, 552 which has not been discussed in the previous studies. As we have demonstrated previously,⁸ multiphoton dissociation of DMMP by focused 248 nm light produces significant amounts of 553 554 P*. Although not directly detectable in our experiments, we also presume the presence of ground state $P(^{4}S)$, formed either directly from the multiphoton dissociation of DMMP, or 555 from the relaxation of P*. These excited state P* atoms react with O₂ to form PO on a 556 timescale similar to that at which PO is removed by O₂.⁸ Thus, if significant amounts of these 557 558 excited states are present, the observed removal of PO would be slower, resulting in a smaller 559 rate coefficient for the reaction. This may be less of a problem in the study by Long et al.¹⁶ in which the photolysis source is 10.6 μ m light from an infrared CO₂ laser. At this wavelength, 560 561 many more photons would need to be absorbed to excite PO to a high enough level in which 562 it may dissociate to ground or excited state P atoms. Indeed, the rate constant measured by Long *et al.*¹⁶ agrees well with that determined in our study where we used a PO precursor that 563 does not produce P* when photolysed. $P({}^{4}S)$ atoms also react with O₂ to produce PO (R1), 564 565 however the timescale of this reaction is around 50 times slower than the removal of PO by 566 O₂, and as such the formation of any ground state P atoms should only have a minor effect on 567 the measurement of k_2 (see Section 3.1).

Figure 3 indicates that there is a small pressure dependence for reaction R2, which is evidence for both a two- and three-body channel for this reaction (R2a and R2b, respectively), which are explored using theoretical calculations in Section 4.2. The presence of the three-body channel forming PO₃ has not previously reported in the literature, presumably as the previous studies did not measure the removal of PO₂ over a range of pressures.

574 The have been no previous investigation into the kinetics of the reaction between PO₂ 575 and O_3 . Taking the average of our room temperature measurements, we determine that $k_3 =$ $(8.47 \pm 0.30) \times 10^{-13}$ cm³ molecule⁻¹ s⁻¹ at T = 293 K. This is over 4 orders of magnitude 576 faster than the reaction between the isovalent NO₂ with O₃ (~ 3.5×10^{-17} cm³ molecule⁻¹ s⁻¹ at 577 578 T = 298 K).⁴⁸ Although this reaction is unlikely be important astrochemically, the fact that 579 many current chemical models are using the isovalence between nitrogen and phosphorus to 580 derive rate coefficients for P-bearing species is called into question, and highlights the 581 importance of experimentally determined rate coefficients.

582

583 4.2 Theoretical Calculations

Electronic structure calculations were carried out using the Gaussian 16 suite of the programs.⁹ Vibrational frequencies, rotational constants and energies were calculated with the Complete Basis Set (CBS-QB3) method of Montgomery et al.⁴⁹ The Cartesian coordinates, molecular parameters and heats of formation of the phosphorus oxides are listed in Table 3.

Table 3. Molecular properties of the stationary points on the potential energy surfaces for PO $+ O_2$ and PO₂ + O₃. The molecular geometries are illustrated in Figure 7.

Molecule	Geometry (Cartesian co-ordinates in Å) ^a	Rotational constants (GHz) ^a	Vibrational frequencies (cm ⁻¹) ^a	Δ _f H ^e (0 K) (kJ mol ⁻¹) ^b
	$PO + O_2$	surface (Figure 7a)		
РО	P, -0.319, 0.057, 0. O, 0.176, -1.341, 0.	21.78779	1247	-31
cis-OPO ₂	P, 0.760, 0.487, 0.000 O, 1.208, -0.902, 0.000 O, -0.977, 0.570, 0.000 O, -1.654, -0.583, 0.000	16.31182 5.04050 3.85062	164, 217, 500, 540, 1071, 1300	-150
TS1 (from <i>cis</i> - OPO ₂ to <i>trans</i> -OPO ₂	P, 0.896, 0.377, -0.134 O, 1.582, 1.646, 0.097 O, -0.756, 0.712, 0.037 O, -1.976, 0.225, 0.003	35.26271 3.53849 3.21579	209 <i>i</i> , 94, 342, 547, 1139, 1292	-139
trans-OPO ₂	P, -0.549, -0.457, 0.005 O, -1.730, 0.403, -0.001 O, 0.830, 0.613, 0.001 O, 1.934, -0.149, -0.001	32.75631 3.94718 3.52270	105, 246, 370, 638, 1105, 1300	-161
TS2 (from <i>trans</i> -OPO ₂ to PO ₃)	P, -0.332, -0.114, 0.460 O, -1.536, -0.050, 0.389 O, 1.002, 0.809, -0.052 O, 1.161, -0.605, 0.366	18.38211 5.63914 5.25830	1506 <i>i</i> , 237, 354, 479, 930, 1175	-115
TS3 (from <i>cis</i> - OPO ₂ to PO ₂ + O	P, -0.136, 0.898, -0.071 O, 0.163, 0.149, 1.155 O, -0.244, 0.372, -1.488 O, -0.036, -1.398, -2.161	17.01510 3.67813 3.02436	391 <i>i</i> , 103, 223, 405, 889, 1264	-29
PO ₃	P, 0.000, 0.000, 0.002 O, 0.000, 0.000, 1.479 O, 1.278, 0.000, -0.742 O, -1.278, 0.000, -0.742	9.64387 9.64387 4.82194	192, 192, 434, 1015, 1123, 1123	-447
PO ₂ (or OPO)	P, 0.001, 0., 0.000 O, 0.0006, 0., 1.474 O, 1.056, 0., -1.028	98.10129 8.57057 7.88196	388, 1068, 1314	-290
PO ₂ (or OPO)	P, 0., 0., 0.276 O, 0., 1.376, -0.269 O, 0., -1.376, -0.269	108.11428 8.33894 7.74181	398, 1077, 1478	-260 ^d
	$PO_2 + O_3$	surface (Figure 7b)		
TS1 (from PO ₂ + O ₃)	O, 2.091, 0.500, 0.770 O, 2.073, -0.246, -0.284 O, 1.174, -1.168, -0.259 P, -1.332, 0.070, -0.277 O, -2.321, -0.589, 0.605 O, -0.703, 1.407, -0.375 °	5.87904 1.52137 1.35133 °	36 <i>i</i> , 38, 63, 66, 99, 176, 400, 699, 1035, 1077, 1101, 1454	-135 ^d
OPO-O ₃	P, 0.622, 0.000, 0.000 O, 1.170, -1.344, 0.000 O, 1.169, 1.344, -0.001 O, -1.721, 0.000, 0.000 O, -0.886, 0.000, 1.071 O, -0.887, 0.000, -1.070	5.35021 3.06448 2.71613	201, 215, 227, 346, 374, 411, 447, 819, 924, 1036, 1161, 1458	-380
TS2 (from OPO-O ₃ to $PO_3 + O_3$)	P, -0.676, 0.000, 0.022 O, -1.179, 1.348, -0.166 O, -1.179, -1.348, -0.165 O, 1.749, 0.000, 0.165	5.28567 2.86956 2.56924	144 <i>i</i> , 123, 234, 314, 383, 403, 552, 734, 828, 1155, 1180,	-370

O, 0.666, 0.000, 1.113 O, 1.223, 0.000, -1.013	1	1476	

^a Calculated at the CBS-QB3 level of theory. ^b Heat of formation at 0 K, using reference values for $P_{(g)}$, O and 592 O_3 of $\Delta_f H^{e}$ (0 K) = 315.7, 246.8 and 144.5 kJ mol⁻¹, respectively.⁵⁰ ^c Calculated at the MP2/6-311+g(2d,p) level 593 of theory. ^dEnergy at the CCSD(T)//MP2/6-311+g(2d,p) level of theory.

594

595 The resulting potential energy surfaces for PO + O_2 (R2) and PO₂ + O_3 (R3) are 596 shown in Figures 7a and 7b, respectively, which also illustrate the geometries of the stationary points. RRKM calculations were then performed for these two reactions using the 597 Master Equation Solver for Multi-Energy well Reactions (MESMER) program.⁴⁴ R2 initially 598 599 involves addition of O₂ to PO, forming the *cis*-OPO₂ complex. This can either dissociate 600 directly to $PO_2 + O$ via TS3, or first rearrange to trans-OPO₂ via TS1 followed by 601 rearrangement to PO₃ via TS2. PO₃ will then either be quenched by collision with the third 602 body (N_2) , or dissociate to $PO_2 + O_2$.



Figure 7. Potential energy surfaces for (a) $PO + O_2$ (R2), determined at the CBS-QB3 level of theory, and (b) $PO_2 + O_3$ (R3), determined at the CBS-QB3 level of theory apart from barrier TS1 which is a CCSD(T)//MP2/6-311+g(2d,p) calculation with respect to the reactants.

609

610 The internal energy of each species on the potential energy surface was divided into a 611 contiguous set of grains (width 200 cm⁻¹), each containing a bundle of rovibrational states. 612 Each grain was then assigned a set of microcanonical rate coefficients for dissociation back to 613 the reactants (PO + O_2), and forward to the products (PO₂ + O), using an inverse Laplace 614 transformation to link them directly to the high-pressure limiting recombination coefficients 615 (k_{∞}) . An Arrhenius expression was optimised by floating both the pre-exponential factor and 616 the activation energy to give the best fit of the RRKM model to the experimental data, yielding $k_{\infty}(PO + O_2) = 3.5 \times 10^{-11} \exp^{(-168/T)} \text{ cm}^3$ molecule⁻¹ s⁻¹. For the reaction of PO₂ + O, 617 the rate coefficient was set to be reasonably fast with a small T-dependence, $k_{\infty}(PO_2 + O) =$ 618

619 $2.4 \times 10^{-11} \exp^{(-25/T)} \text{ cm}^3$ molecule⁻¹ s⁻¹. The exponential down model was used to estimate the 620 probability of collisional transfer between grains. For N₂ as the third body, the average 621 energy for downward transitions $\langle \Delta E \rangle_{\text{down}}$ was set to 300 cm⁻¹ at 300 K, with a temperature 622 dependence of $T^{0.25}$.⁵¹

623 Figure 3 shows a satisfactory fit to the experimental data at 294 and 337 K. Figure 8 624 illustrates the RRKM-predicted overall rate coefficient k_2 (Figure 8a) and the branching ratio 625 to produce PO₃, as a function of $[N_2]$ and T (Figure 8b). Figure 8 shows that the variation of 626 the total rate coefficient and the branching ratio to form the two products OPO and PO₃ have relatively complex dependences on P and T. This arises from the nature of the PES for the 627 reaction (Figure 7a). As a result, a Troe-type fall-off expression⁵² does not give a good fit to 628 the rate coefficient for the recombination reaction forming PO₃. Nevertheless, the data plotted 629 630 in Figure 8 can be used to estimate the rate coefficients for both reaction channels at a 631 selected P and T. As shown in Figure 8a, k_2 is predicted to have a rather small T and P 632 dependence. At $[N_2] < 10^{17}$ cm⁻³, over the temperature range 120 - 500 K, the fraction of PO₃ 633 formed is below 1%. At pressures typical of planetary upper atmospheres where meteoric ablation of P will occur,⁷ the reaction should essentially be pressure-independent and can be 634 fitted to the following expression: $\log_{10}(k_2, 120 - 500 \text{ K}, \text{ cm}^3 \text{ molecule}^{-1} \text{ s}^{-1}) = -13.915 + 1000 \text{ molecule}^{-1} \text{ s}^{-1}$ 635 2.470log₁₀(T) - 0.5020(log₁₀(T))², with an uncertainty of \pm 10 % within the experimental 636 637 temperature range (191 – 339 K).



Figure 8. RRKM calculations for the reaction PO + O_2 on the surface in Figure 7a: a) the overall rate coefficient, k_2 ; b) the branching ratio to PO₃; both plotted as a function of [N₂] and *T*.

643

For the reaction between PO_2 and O_3 , the RRKM fit using the potential energy surface in Figure 7b was used. The measured positive temperature dependence of the reaction (Figure 5) implies that there is a barrier in the entrance channel. Initial exploration of the reaction entrance channel using the B3LYP/6-311+g(2d,p) level of theory did not indicate the 648 presence of a barrier. However, this DFT functional is not appropriate for long-range 649 interactions; guided by previous work on the analogous reaction between NO₂ and O₃ (which has a significantly larger barrier in the entrance channel),⁵³ we used the MP2 level of theory 650 to determine that a small barrier exists at a $PO_2 - O_3$ separation of 2.7 Å. An accurate coupled 651 cluster calculation at the CCSD(T)//MP2/6-311+g(2d,p) level of theory⁵⁴ gave a barrier 652 653 height of 9.4 kJ mol⁻¹ (Figure 7b and Table 3), in good agreement with the 8.2 ± 0.7 kJ mol⁻¹ 654 activation energy determined from the Arrhenius fit to the experimental data. A satisfactory 655 fit to the experimental data was then obtained by optimising the pre-exponential factor in an Arrhenius expression for the inverse Laplace transformation, yielding $k_{\infty}(PO_2 + O_3, 180 -$ 656 370 K = $3.7 \times 10^{-11} \exp^{(-1131/T)} \text{ cm}^3$ molecule⁻¹ s⁻¹ ($\sigma = \pm 26$ % over the experimental 657 658 temperature range, 188 - 339 K), where the activation energy is set to 9.4 kJ mol⁻¹.

659

660 4.3 Fitting the PO spectra

661 To simulate the PO spectra obtained in this study, we have used the PGOPHER spectral simulator,⁵⁵ together with the molecular constants provided by Verma and Singhal⁴⁷ for the X 662 ² Π state and by Coquart *et al.*²⁸ for the $A^2\Sigma^+$ state. The simulated PO spectrum was then 663 664 compared to the experimental spectrum obtained using PCl₃ as a precursor, which was chosen 665 as it has the best signal-to-noise (although it should be noted that for the lines that are visible 666 in the spectra produced using PBr₃ and POCl₃ as precursors, there is excellent agreement 667 between the line positions – Figure 6). Next, the simulated spectrum was refined by first 668 assigning simulated lines to the corresponding peaks in the experimental spectra, then 669 refitting the simulated spectrum to the experimental allowing the molecular constants to float. 670 This process was iterative, with more peaks being able to be assigned and fitted as agreement 671 between the simulated and experimental spectra improved. Both the (0,0) and (1,1) bands 672 were fit simultaneously.

673 When fitting the spectra, we encountered some difficulty in obtaining a good fit to 674 sections of the fine structure within the (0,0) bands. Verma and Singhal⁴⁷ reported that the $A^{2}\Sigma^{+}$ (v = 0) level of PO is perturbed by the ${}^{4}\Sigma^{-}$ (v= 4) level. By including this perturbing state 675 676 in our simulation, we were able to obtain good a fit to the data. Figure 9 shows the simulated 677 spectrum overlaid on the experimental spectrum, as well as the residual for the contour fit, 678 while Tables 4 - 6 give the molecular constants employed in the simulation. The best match 679 was obtained using a rotational temperature of 350 K, and by convolving the simulated spectra with a Gaussian with 0.6 cm⁻¹ FWHM, which accounts for Doppler broadening and 680 681 spectral resolution. In addition to allowing the molecular constants to float, the relative 682 intensities of the $A^2\Sigma^+$ (v = 0) $\leftarrow X^2\Pi$ (v = 0), the $A^2\Sigma^+$ (v = 1) $\leftarrow X^2\Pi$ (v = 1) transitions were 683 also floated. The ${}^{4}\Sigma^{-}(v=4) \leftarrow A^{2}\Sigma^{+}(v=0)$ matrix element was 4.593 ± 0.011 .

As can be seen from Figure 9, there is good agreement between the line positions of the simulated and experimental spectra, for both the (0,0) and (1,1) bands. In general there is also satisfactory agreement between the simulated and experimental line intensities. There does, however seem to be some systematic differences between branches, implying some additional Ω (but not *J*) dependent mixing between branches. Comparing the molecular 689 constants obtained in this study with the literature values (Tables 4 - 6) shows that all the 690 values are essentially consistent, with the differences most likely arising from minor 691 differences in the definitions of the constants.

692



693

Figure 9: Measured and calculated $A^2\Sigma^+ \leftarrow X^2\Pi$ band of PO, together with the residual (experimental – simulated). Plot a) shows the $X^2\Pi_{3/2}$ sub-band, while plot b) shows the $X^2\Pi_{1/2}$ sub-band.

698 Table 4. Molecular constants of the $X^{2}\Pi$ state employed in the simulation of the PO spectrum. Units of cm⁻¹.

	$X^2\Pi$, v =	$= 0 (cm^{-1})$
	This study	Verma and Singhal ⁴⁷
В	0.73063 ± 0.00056	0.73106 ± 0.00003

А	223.908 ± 0.023	224.01 ± 0.01
γ	0.0909 ± 0.0085	
р	0.0201 ± 0.0015	-0.007
q	-0.00118 ± 0.00023	
D	$(1.34 \pm 0.36) \times 10^{-6}$	$(1.31 \pm 0.05) \times 10^{-6}$
	$X^2\Pi, v =$	1, (cm ⁻¹)
	This study	Verma and Singhal ⁴⁷
Т	1220.173 ± 0.013	1220.22 ± 0.02
В	0.72547 ± 0.00085	0.72551 ± 0.00003
А	223.323 ± 0.026	224.18 ± 0.01
γ	0.119 ± 0.017	
р	0.0215 ± 0.0020	-0.007
q	-0.00397 ± 0.00059	
D	$-(1.29 \pm 0.91) \times 10^{-7}$	$(1.41 \pm 0.05) \times 10^{-6}$

Table 5. Molecular constants of the $A^2\Sigma^+$ state employed in the simulation of the PO spectrum. Units of cm⁻¹.

	$A^2\Sigma^+, v = 0$	$0 (cm^{-1})$
	This study	Coquart et al.28
Т	40486.613 ± 0.021	40485.61
В	0.77740 ± 0.00055	0.77725
D	$(1.23 \pm 0.34) \times 10^{-6}$	1×10^{-6}
	$A^2\Sigma^+, v =$	$1, (cm^{-1})$
	This study	Coquart et al.28
Т	41863.138 ± 0.013	41862.76
В	0.77251 ± 0.00089	0.77195
D	$(1.3 \pm 1.0) \times 10^{-6}$	1×10^{-6}

703

Table 6. Molecular constants of the ${}^{4}\Sigma^{-}$ (*v*= 4) perturbing level employed in the simulation of

705 the PO spectrum. Units of cm^{-1} .

	${}^{4}\Sigma^{-}, v = 4 \text{ (cm}^{-1})$	
	This study	Verma and Singhal ⁴⁷
Т	40518.77 ± 0.24	40517.69 ± 0.23
В	0.5982 ± 0.0012	0.5984 ± 0.0012
λ_{SS}	0.935 ± 0.066	0.33 ± 0.11
γ	0.0297 ± 0.0069	0.020 ± 0.008
D	0	0.6×10^{-6}
$6^{\frac{1}{2}} < {}^{4}\Sigma^{-} H A^{2}\Sigma^{+}, v = 0 >$	4.593 ± 0.011	-7.235 ± 0.093

707 4.4 Atmospheric Implications

708 Interplanetary dust particles entering the upper atmosphere of a terrestrial planet will undergo meteoric ablation at around the 1 μ bar pressure region.⁵⁶ Phosphorus will ablate 709 from these particles as either PO or PO₂,⁷ and these molecules will mostly dissociate to form 710 711 ground and excited state P atoms following hyperthermal collisions with atmospheric molecules. Peak ablation in the Earth's atmosphere occurs around 85 km,⁵⁶ and the P atoms 712 713 formed in this region will be quickly oxidised by O₂ to PO.⁸ These PO molecules will then 714 react with O₂, with an e-folding lifetime of around 2.5 ms. At the low pressure at this altitude 715 (< 10^{-5} bar), the reaction will proceed *via* the 2-body bimolecular channel to form PO₂ and O. As the concentration of O_3 at this altitude is around 10^5 times lower than O_2 ,⁵⁷ removal of 716 PO_2 by O_3 to PO_3 will be a very slow process: if this were the only process removing PO_2 , it 717 718 would have a lifetime of ~ 7.5 days. This suggests PO_3 will not be a major reservoir of 719 meteor ablated P. As there are no exothermic processes to convert PO₂ back to PO, the 720 speciation of phosphorus should essentially depend on the reactions converting PO₂ into 721 HPO₂, HOPO, and HOPO₂, as illustrated in Figure 1.

722

723 5. Conclusions

724 The reactions of PO with O_2 (R2) and PO₂ with O_3 (R3) have been studied both 725 experimentally and theoretically, with this being the first study to our knowledge of reaction 726 R3. For reaction R2, there are significant discrepancies in the literature values reported for k_2 727 at room temperature, with values varying by around a factor of 60. We have determined with 728 the PLP-LIF technique a rate coefficient at the top end of the literature values, with $k_2(294 \text{ K})$ = $(1.44 \pm 0.11) \times 10^{-11}$ cm³ molecule⁻¹ s⁻¹ at $P \sim 10$ Torr. Rate coefficients for R2 determined 729 730 over a range of pressures also indicate the presence of a three-body channel, in which PO + 731 O_2 combine to form PO₃ (R2b). The potential energy surfaces of both the PO + O_2 and PO₂ + 732 O₃ systems were determined using electronic structure theory, and these calculations 733 combined with RRKM theory to explain the observed pressure and temperature dependences. 734 For PO + O_2 , at pressures typical of a planetary upper atmosphere where meteoric ablation of 735 P will occur, the reaction is effectively pressure independent with a yield of $PO_2 + O$ of > 736 99%, and can be expressed by: $\log_{10}(k_2, 120 - 500 \text{ K}, \text{ cm}^3 \text{ molecule}^{-1} \text{ s}^{-1}) = -13.915 + 1000 \text{ molecule}^{-1} \text{ s}^{-1}$ 737 2.470log₁₀(T) - 0.5020(log₁₀(T))², with an uncertainty of \pm 10 % over the experimental 738 temperature range (191 - 339 K). With increasing pressure, the yield of PO₃ increases, 739 reaching ~ 90% at a pressure of 1 atm and T = 300 K. For PO₂ + O₃, $k_3(PO_2 + O_3) = 3.7 \times 10^{-10}$ 740 ¹¹ exp^(-1131/T) cm³ molecule⁻¹ s⁻¹, with an uncertainty of \pm 26 % over the experimental 741 temperature range (188 - 339 K). Laser induced fluorescence spectra of PO over the 742 wavelength range of 245 – 248 nm were collected and fitted using the PGOPHER program, 743 which yields new spectroscopic constants for the ground and v = 1 vibrational levels of the X 744 $^{2}\Pi$ and $A^{2}\Sigma^{+}$ states of PO.

745 Acknowledgements

- 746 This study was supported by funding from the UK Science and Technology Facilities Council
- 747 (grant ST/P000517/1). The raw and processed data produced by this study is archived at
- T48 Leeds University and is available upon request to JMCP.
- 749
- 750 References
- 751
- 752 1. Maciá, E. The role of phosphorus in chemical evolution. *Chem. Soc. Rev.* 2005, *34*, 691-701.
- 2. Lodders, K. Solar system abundances and condensation temperatures of the elements. *Astrophys. J.* 2003, *591*, 1220-1247.
- 756 3. Redfield, A. C. The biological control of chemical factors in the environment. *Am.*757 *Sci.* 1958, *46*, 205-221.
- 758 4. Schwartz, A. W. Phosphorus in prebiotic chemistry. *Philos. T. Roy. Soc. B* 2006, *361*,
 759 1743-1749.
- 760 5. Pasek, M. A. Rethinking early Earth phosphorus geochemistry. P. Natl. A. Sci. 2008,
 761 105, 853-858.
- 762 6. Plane, J. M. C.; Flynn, G. J.; Määttänen, A.; Moores, J. E.; Poppe, A. R.; Carrillo763 Sanchez, J. D.; Listowski, C. Impacts of cosmic dust on planetary atmospheres and surfaces.
 764 Space Sci. Rev. 2017, 214, 23.
- 765 7. Carrillo-Sánchez, J. D.; Bones, D. L.; Douglas, K. M.; Flynn, G. J.; Wirick, S.;
 766 Fegley, B.; Araki, T.; Kaulich, B.; Plane, J. M. C. Injection of meteoric phosphorus into
 767 planetary atmospheres. *Planet. Space. Sci.* 2020, *187*, 104926.
- 8. Douglas, K. M.; Blitz, M. A.; Mangan, T. P.; Plane, J. M. C. Experimental Study of
 the Removal of Ground- and Excited-State Phosphorus Atoms by Atmospherically Relevant
 Species. J. Phys. Chem. A 2019, 123, 9469-9478.
- Frisch, M. J.; Trucks, G. W.; Schlegel, H. B.; Scuseria, G. E.; Robb, M. A.;
 Cheeseman, J. R.; Scalmani, G.; Barone, V.; Petersson, G. A.; Nakatsuji, H., et al. *Gaussian 16 Rev. B.01*, Wallingford, CT, 2016.
- 10. Henshaw, T. L.; MacDonald, M. A.; Stedman, D. H.; Coombe, R. D. The $P({}^{4}S_{u}) + N_{3}({}^{2}\Pi_{g})$ reaction: chemical generation of a new metastable state of PN. *J. Phys. Chem.* **1987**, 91, 2838-2842.
- 777 11. Clyne, M. A. A.;Ono, Y. Kinetic studies of ground-state phosphorus atoms. *J. Chem.*778 *Soc. Farad. T.* 2 1982, 78, 1149-1164.
- 12. Husain, D.;Slater, N. K. H. Time-resolved resonance fluorescence studies of ground state phosphorus atoms, $P[3p^{3}({}^{4}S_{3/2})]$. *J. Chem. Soc. Farad. T.* 2 **1978**, *74*, 1627-1643.
- 781 13. Husain, D.;Norris, P. E. Reactions of phosphorus atoms, $P(3^4S_{3/2})$, studied by 782 attenuation of atomic resonance radiation in vacuum ultraviolet. *J. Chem. Soc. Farad. T.* 2 783 **1977**, 73, 1107-1115.
- 14. Aleksandrov, E. N.; Arutyunov, V. S.; Dubrovina, I. V.; Kozlov, S. N. On the role of PO radicals in the reaction of phosphorus oxidation. *Dokl. Akad. Nauk. SSSR* **1982**, *267*, 110-
- PO radicals in the reaction of phosphorus oxidation. *Dokl. Akad. Nauk. SSSR* 1982, 267, 110
 113.
 - 30

- 787 15. Wong, K. N.; Anderson, W. R.; Kotlar, A. J.; Dewilde, M. A.; Decker, L. J. Lifetimes 788 and quenching of the B $^{2}\Sigma^{+}$ PO by atmospheric gases. *J. Chem. Phys.* **1986**, *84*, 81-90.
- 16. Long, S. R.; Christesen, S. D.; Force, A. P.; Bernstein, J. S. Rate constant for the reaction of PO radical with oxygen. *J. Chem. Phys.* **1986**, *84*, 5965-5966.

791 17. Sausa, R. C.; Miziolek, A. W.; Long, S. R. State distributions, quenching, and
792 reaction of the PO radical generated in Excimer laser photofragmentation of dimethyl
793 methylphosphonate. *J. Phys. Chem.* **1986**, *90*, 3994-3998.

- De Beck, E.; Kaminski, T.; Patel, N. A.; Young, K. H.; Gottlieb, C. A.; Menten, K.
 M.; Decin, L. PO and PN in the wind of the oxygen-rich AGB star IK Tauri. *Astron. Astrophys.* 2013, 558, 9.
- Tenenbaum, E. D.; Dodd, J. L.; Milam, S. N.; Woolf, N. J.; Ziurys, L. M. Comparatie
 spectra of oxygen-rich versus carbon-rich circumstellar shells: VY Cannis Majoris and IRC +
 10216 at 215-285 GHz. *Astrophys. J. Lett* 2010, 720, L102-L107.
- Ziurys, L. M.; Milam, S. N.; Apponi, A. J.; Woolf, N. J. Chemical complexity in the
 winds of the oxygen-rich supergiant star VY Canis Majoris. *Nature* 2007, 447, 1094-1097.
- 802 21. Gobrecht, D.; Cherchneff, I.; Sarangi, A.; Plane, J. M. C.; Bromley, S. T. Dust
 803 formation in the oxygen-rich AGB star IK Tauri. *Astron. Astrophys.* 2016, 585, A6.
- Jayaweera, T. M.; Melius, C. F.; Pitz, W. J.; Westbrook, C. K.; Korobeinichev, O. P.;
 Shvartsberg, V. M.; Shmakov, A. G.; Rybitskaya, I. V.; Curran, H. J. Flame inhibition by
 phosphorus-containing compounds over a range of equivalence ratios. *Combust. Flame* 2005, *140*, 103-115.
- 808 23. Korobeinichev, O. P.; Shvartsberg, V. M.; Bol'shova, T. A.; Shmakov, A. G.;
 809 Knyaz'kov, D. A. Inhibition of methane-oxygen flames by organophosphorus compounds.
 810 *Cobust, Explo. Shock Waves* 2002, *38*, 127-133.
- 811 24. MacDonald, M. A.; Jayaweera, T. M.; Fisher, E. M.; Gouldin, F. C. Inhibition of
 812 nonpremixed flames by phosphorus-containing compounds. *Combust. Flame* 1999, *116*, 166813 176.
- 814 25. Korobeinichev, O. P.; Shvartsberg, V. M.; Shmakov, A. G. The chemistry of combustion of organophosphorus compounds. *Russ. Chem. Rev.* 2007, *76*, 1021-1047.
- 816 26. Zegers, E. J. P.;Fisher, E. M. Gas-phase pyrolysis of diisopropyl methylphosphonate.
 817 *Combust. Flame* 1998, *115*, 230-240.
- 818 27. Moussaoui, Y.; Ouamerali, O.; De Mare, G. R. Properties of the phosphorus oxide
 819 radical, PO, its cation and anion in their ground electronic states: comparison of theoretical
 820 and experimental data. *Int. Rev. Phys. Chem.* 2003, 22, 641-675.
- 821 28. Coquart, B.; Paz, M. D.; Prudhomme, J. C. Transition $A^2\Sigma^+ X^2 \Pi$ des molécules P¹⁶O 822 et P¹⁸O. Perturbations de l'état $A^2\Sigma^+$. *Can. J. Phys.* **1975**, *53*, 377-384.
- 823 29. Wong, K. N.; Anderson, W. R.; Kotlar, A. J. Radiative processes following laser 824 excitation of the A ${}^{2}\Sigma^{+}$ state of PO. *J. Chem. Phys.* **1986**, *85*, 2406-2413.
- 825 30. Liu, H.; Shi, D. H.; Sun, J. F.; Zhu, Z. L. Accurate potential energy curves and
 826 spectroscopic properties of the 27 Lambda-S states and 73 Omega states of the PO radical.
 827 *Mol. Phys.* 2017, *115*, 714-730.

- 828 31. Izzaouihda, S.; El Makarim, H. A.; Komiha, N.; Lahmar, S.; Ghalila, H. Ab-initio
 829 potential energy curves of valence and Rydberg electronic states of the PO radical. *Comput.*830 *Theor. Chem.* 2014, *1049*, 102-108.
- 32. Gómez Martín, J. C.; Blitz, M. A.; Plane, J. M. C. Kinetic studies of atmospherically
 relevant silicon chemistry Part I: Silicon atom reactions. *Phys. Chem. Chem. Phys.* 2009, *11*,
 671-678.
- Mangan, T. P.; McAdam, N.; Daly, S. M.; Plane, J. M. C. Kinetic study of Ni and
 NiO reactions pertinent to the Earth's upper atmosphere. *J. Phys. Chem. A* 2019, *123*, 601610.
- 837 34. Totterdill, A.; Gómez Martín, J. C.; Kovács, T.; Feng, W.; Plane, J. M. C.
 838 Experimental Study of the Mesospheric Removal of NF3 by Neutral Meteoric Metals and
 839 Lyman-α Radiation. J. Phys. Chem. A 2014, 118, 4120-4129.
- 840 35. Lewis, T.; Heard, D. E.; Blitz, M. A. A novel multiplex absorption spectrometer for
 841 time-resolved studies. *Rev. Sci. Instrum.* 2018, 89, 024101.
- 842 36. Keller-Rudek, H.; Moortgat, G. K.; Sander, R.; Sörensen, R. The MPI-Mainz UV/VIS
 843 Spectral Atlas of Gaseous Molecules of Atmospheric Interest. *Earth Syst. Sci. Data* 2013, *5*, 365-373.
- 845 37. Matsumi, Y.;Kawasaki, M. Photolysis of Atmospheric Ozone in the Ultraviolet
 846 Region. *Chem. Rev.* 2003, 103, 4767-4782.
- 847 38. Hamilton, P. A. The laser-induced fluorescence-spectrum and radiative lifetime of
 848 PO₂. J. Chem. Phys. 1987, 86, 33-41.
- 849 39. Hamilton, P. A.;Murrells, T. P. Kinetics and mechanism of the reactions of PH_3 with 850 $O({}^{3}P)$ and $N({}^{4}S)$ atoms. *J. Chem. Soc. Farad. T.* 2 **1985**, *81*, 1531-1541.
- 40. Hamilton, P. A.; Murrells, T. P. Mechanism for the chemiluminescence in oxygenphosphorus systems. *J. Phys. Chem.* **1986**, *90*, 182-185.
- 853 41. Kampf, R. P.;Parson, J. M. Chemiluminescent pathways in reactions of phosphorus,
 854 antimony, and bismuth with ozone to form dioxides and monoxides. *J. Chem. Phys.* 1998,
 855 108, 7595-7606.
- 42. Cai, Z. L.; Hirsch, G.; Buenker, R. J. Ab initio study of the electronic spectrum of the
 PO2 radical. *Chem. Phys. Lett.* 1996, 255, 350-356.
- 858 43. Ianni, J. C. Kintecus. <u>www.kintecus.com</u>.
- 859 44. Glowacki, D. R.; Liang, C. H.; Morley, C.; Pilling, M. J.; Robertson, S. H. MESMER:
 860 An open-source master equation solver for multi-energy well reactions. *J. Phys. Chem. A*861 2012, *116*, 9545-9560.
- 862 45. Stratmann, R. E.; Scuseria, G. E.; Frisch, M. J. An efficient implementation of time863 dependent density-functional theory for the calculation of excitation energies of large
 864 molecules. J. Chem. Phys. 1998, 109, 8218-8224.
- Frisch, M. J.; Trucks, G. W.; Schlegel, H. B.; Scuseria, G. E.; Robb, M. A.;
 Cheeseman, J. R.; Scalmani, G.; Barone, V.; Petersson, G. A.; Nakatsuji, H., et al. *Gaussian 16, Revision B.01*, Gaussian, Inc.: Wallingford, CT, USA, 2016.
- 868 47. Verma, R. D.;Singhal, S. R. New Results on the $B^2\Sigma^+$, $b^4\Sigma^-$, and $X^2\Pi$ States of PO. 869 *Can. J. Phys.* **1975**, *53*, 411-419.

48. Atkinson, R.; Baulch, D. L.; Cox, R. A.; Crowley, J. N.; Hampson, R. F.; Hynes, R.
63.; Jenkin, M. E.; Rossi, M. J.; Troe, J. Evaluated kinetic and photochemical data for atmospheric chemistry: Volume I - gas phase reactions of O-x, HOx, NOx and SOx species.
Atmos. Chem. Phys. 2004, 4, 1461-1738.

49. Montgomery, J. A., Jr.; Frisch, M. J.; Ochterski, J. W.; Petersson, G. A. A complete
basis set model chemistry. VI. Use of density functional geometries and frequencies. J. *Chem. Phys.* 1999, 110, 2822-2827.

- 50. Chase, M. W., Jr.; Davies, C. A.; Downey, J. R., Jr.; Frurip, D. J.; McDonald, R. A.;
 Syverud, A. N., NIST-JANAF Thermochemical Tables 1985. In *NIST Standard Reference Database 13*, 1998 ed.; Standard Reference Data Program, National Institute of Standards
 and Technology Gaithersburg, MD, 1998.
- 881 51. Gilbert, R. G.;Smith, S. C. *Theory of Unimolecular and Recombination Reactions*.
 882 Blackwell: Oxford, 1990.
- 883 52. Burkholder, J. B.; Sander, S. P.; Abbatt, J.; Barker, J. R.; Huie, R. E.; Kolb, C. E.;
 884 Kurylo, M. J.; Orkin, V. L.; Wilmouth, D. M.; Wine, P. H. Chemical Kinetics and
 885 Photochemical Data for Use in Atmospheric Studies, Evaluation No. 18.
 886 <u>http://jpldataeval.jpl.nasa.gov</u>.
- 887 53. Peiró-García, J.;Nebot-Gil, I. Ab initio study of the mechanism of the atmospheric 888 reaction: $NO_2+O_3 \rightarrow NO_3 + O_2$. J. Comput. Chem. 2003, 24, 1657-1663.
- Scuseria, G. E.; Janssen, C. L.; III, H. F. S. An efficient reformulation of the closedshell coupled cluster single and double excitation (CCSD) equations. *J. Chem. Phys.* 1988,
 89, 7382-7387.
- 892 55. Western, C. M. PGOPHER: A program for simulating rotational, vibrational and
 893 electronic spectra. J. Quant. Spectrosc. Radiat. Transfer 2017, 186, 221-242.
- 56. Plane, J. M. C.; Feng, W. H.; Martin, J. C. G.; Gerding, M.; Raizada, S. A new model
 of meteoric calcium in the mesosphere and lower thermosphere. *Atmos. Chem. Phys.* 2018,
 18, 14799-14811.
- 897 57. Plane, J. M. C. Atmospheric chemistry of meteoric metals. *Chem. Rev.* 2003, 103, 4963-4984.
- 899
- 900
- 901
- 902
- ---
- 903
- 904
- 905

906

907

908

911 TOC Graphic



Reaction coordinate