Rift Focussing and Magmatism During Late-Stage Rifting in Afar, Ethiopia

C. Moore¹, T. Wright¹, A. Hooper¹

¹COMET, School of Earth and Environment, University of Leeds, Leeds, LS2 9JT, UK

Key Points:

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6	•	From Sentinel-1 InSAR and GNSS observations, we resolve 3D average surface ve-
7		locities from 2014-19 across the whole Afar rift
8	•	Rift focusing increases with rift maturity, with distributed extension in southern
9		Afar, and localised extension in north and central Afar
10	•	We observe surface deformation related to magmatism at several volcanic centres
11		including Dallol, Erta 'Ale, Nabro, and Dabbahu-Manda-Hararo

Corresponding author: Chris Moore, ee12cm@leeds.ac.uk

12 Abstract

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Processes that facilitate the transition between continental rifting and sea-floor spread-13 ing remain unclear. Variations in the spatial distribution of extension through Afar and 14 into the Red Sea are indicative of temporal evolution of the rift. We develop a time se-15 ries of Sentinel-1 interferometric synthetic aperture radar (InSAR) observations of ground 16 deformation covering the whole Afar Rift from 2014-2019, to study the distribution of 17 extension across all magmatic segments. By incorporating GNSS observations, we re-18 solve 3D average velocities in the vertical, rift-perpendicular, and rift-parallel directions. 19 Results show the spatial distribution of long-term plate motions over the rift, as well as 20 deformation at individual volcanic centres, including Dallol, Nabro, and Erta 'Ale. We 21 find that in northern and central Afar, the majority of extension is accommodated within 22 \pm 15-30 km of magmatic spreading centres. In southern Afar, near the Nubia-Arabia-23 Somalia triple-junction, extension is distributed over 90-180 km, which may indicate an 24 increase in rift focussing with rift maturity. We also observe rapid surface uplift and rift-25 perpendicular extension at the Dabbahu-Manda-Hararo segment with velocities of 33 ± 4 mm/yr 26 and $37 \pm 4 \text{ mm/yr}$ respectively. These are higher than the background extension rate 27 of 18-20 mm/yr, but have decreased by 55-70 % since 2006-10. The data suggests that 28 this is due to an on-going long-lived response to the 2005-10 rifting episode, with poten-29 tial continued processes below the segment including a lower-crustal viscous response and 30 31 magma movement. Continued long-term observations of surface deformation provide key constraints on tectono-magmatic processes in Afar. 32

33 1 Introduction

Magma plays a significant role in accommodating the final stages of continental break-34 up and the transition into early sea-floor spreading, and the Afar region of Ethiopia is 35 uniquely placed to allow the observation of these processes sub-aerially (e.g. Ebinger, 36 2005; Kendall et al., 2005; Wright et al., 2012). Extensional strain in Afar is concentrated 37 onto elongate rifting segments (Ebinger & Casey, 2001) which are maintained by rift-38 ing episodes that include effusive eruptions and dyking (Wright et al., 2006). Magmatic 39 intrusions at these spreading segments produce time-dependent surface deformation, ob-40 servable using time series of interferometric synthetic aperture radar (InSAR). These long 41 time series can help separate surface deformation related to magma movement and vis-42 coelastic relaxation, and have been used to constrain, and highlight the importance of 43 viscous rheology in late-stage rifting (Pagli et al., 2014; J. Hamlyn et al., 2018). 44

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1.1 Regional Setting

Upwelling of a mantle plume initiated rifting in Afar around 30 Ma with abundant 46 flood-basalt volcanism, which has evolved into the ridge-ridge-ridge triple junction ob-47 served in Afar today (Wolfenden et al., 2004; Furman et al., 2006; Hammond et al., 2013). 48 Relative to the Nubian Plate, the Arabian Plate is moving at a rate of 18-20 mm/yr to 49 the NE (McClusky et al., 2010; ArRajehi et al., 2010), accommodated by the opening 50 of the Red Sea Rift (RSR); while the Somalian Plate is moving to the SE at $\sim 6 \text{ mm/yr}$ 51 (Saria et al., 2014), accommodated by the Main Ethiopian Rift (MER) which is the northern-52 most segment of the larger East African Rift. 53

The crust beneath Afar is significantly thinned in comparison to the surrounding Ethiopian highlands and MER. Crustal thicknesses range from 20-45 km in the Ethiopian plateau, 18-30 km in central Afar, and 15-20 km in northern Afar (Tiberi et al., 2005; Bastow & Keir, 2011; Hammond et al., 2011; Lavayssière et al., 2018). Low seismic velocities indicate partial melt within the crust (Gallacher et al., 2016), particularly below volcanic segments in Afar (Stork et al., 2013; Hammond, 2014; Hammond & Kendall, 2016). Seismicity in the upper and lower crust along the Erta 'Ale volcanic segment (EAVS) also indicates the presence of melt below the Erta 'Ale and Alu-Dalafilla volcanic cen tres (Illsley-Kemp et al., 2018).

Active volcansim in Afar is largely concentrated within discrete rifting segments. 63 The EAVS is the immediate on-land expression of the RSR. Erta 'Ale volcano on the EAVS 64 is host to an lava lake with recent overflows in 2010 (Field et al., 2012; Barnie, Oppen-65 heimer, & Pagli, 2016) and 2017, where a flank eruption indicated the presence of a shal-66 low magma body at ~ 1 km depth (Moore et al., 2019). At Gada 'Ale, magma withdrawal 67 and normal faulting caused subsidence from 1993-1996 (Amelung et al., 2000), and a dyke 68 intrusion fed from a magma chamber 2-3 km below Dallol was detected in 2004 (Nobile 69 et al., 2012). The 2008 eruption at Alu-Dalafilla was sourced from a ~ 1 km deep axis-70 aligned reservoir and a magma chamber at ~ 4 km depth (Pagli et al., 2012). 71

The largest recent volcano-tectonic rifting episode in Afar occurred from 2005-10
on the Dabbahu-Manda-Hararo volcanic segment (DMHVS) (Barnie, Keir, et al., 2016).
The initial dyke in September 2005 ruptured the whole 60 km long segment, and intruded
2.4-2.6 km³ of magma over ~ 2 weeks (Wright et al., 2006). Seismicity indicates that
this dyke initiated beneath the Dabbahu and Gabho volcanoes at the northern end of



Figure 1. The Afar Rift with white triangles indicating Holocene volcances and key volcances highlighted in yellow. Simplified volcanic segments (VS) are shown in red: AFVS - Afdera, ALVS - Alayta, AGVS - Asal-Ghoubbet, DMHVS - Dabbahu-Manda-Hararo, EAVS - Erta 'Ale, MI - Manda-Inakir, NMER - Northern Main Ethiopian Rift, TAVS - Tat 'Ale. A subset of GNSS velocity vectors with 95% confidence error ellipses (blue arrows) from (King et al., 2019) show the long-term plate motions. Grey box outlines show the Sentinel-1 coverage from 3 ascending (T014A, T087A, T116A) and 2 descending (T006D, T079D) tracks. Inset map shows the relative movement of the Arabian and Somalian plates to the Nubian plate, with plate boundaries from Bird (2003). RSR - Red Sea Rift, GAR - Gulf of Aden Rift, MER - Main Ethiopian Rift.

the DMHVS, before focussing on the segment centre (Ayele et al., 2009). 13 subsequent 77 dykes were emplaced between 2006-2010, drawing comparisons to the 1975-1984 Krafla 78 rifting episode where extension in both settings is accommodated by magmatic intru-79 sions (Hamling et al., 2010; Ferguson et al., 2010; Wright et al., 2012; Barnie, Keir, et 80 al., 2016). Throughout the rifting episode there was an ongoing post-rifting response to 81 the initial 2005 intrusion, which was modelled using magmatic sources in the upper crust, 82 and the inflation of a deeper source at the segment centre, as viscoelastic relaxation alone 83 was insufficient to reproduce geodetic observations (Grandin et al., 2010; Hamling et al., 84 2014). 85

The Manda-Inakir (MIVS) and Asal-Ghoubbet (AGVS) volcanic segments in southern Afar have also shown recent activity with the 1928-1929 eruption of Kammourta volcano in the MIVS (Audin et al., 1990), and the 1978 eruption at Ardoukoba volcano in the AGVS (Allard et al., 1979; Tarantola et al., 1979). Cattin et al. (2005) identified a post-rifting response in the AGVS to the 1978 eruption, with rift-perpendicular velocities decaying back to the long-term spreading rate 6-8 years after the eruption.

The Nabro Volcanic Range (NVR) is an off-axis volcanic-centre that sits within the Danakil Block, a rigid micro-plate which is moving away from Nubia with extension in Afar (Eagles et al., 2002). The NVR has hosted explosive eruptions at Dubbi volcano in 1861 (Wiart & Oppenheimer, 2000), and Nabro volcano in 2011 (J. E. Hamlyn et al., 2014; Goitom et al., 2015). Persistent subsidence was detected at Nabro for > 1 year following the 2011 eruption which was attributed to viscoelastic relaxation around a magma chamber at 6.4 ± 0.3 km depth (J. Hamlyn et al., 2018).

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1.2 InSAR Velocity Methods & Applications in Afar

Methods for extracting a one-dimensional line-of-sight (LOS) displacement time 100 series from a sequence or network of interferograms are well established. These small-101 baseline algorithms utilise multiple interferogram connections between acquisition dates 102 to produce a more robust estimate of the incremental LOS ground displacement than 103 a simple stacking of interferograms (Berardino et al., 2002; Lanari et al., 2007; Biggs et 104 al., 2007). This methodology may be automated by software packages such as II-RATE 105 (Wang et al., 2012, and references therein), StaMPS (Hooper et al., 2012), GIAnT (Agram 106 et al., 2013), and LiCSBAS (Morishita et al., 2020) in order to obtain linear displace-107 ment rates and uncertainties at each pixel, while reducing the effect of common sources 108 of error such as atmospheric and orbital delays. These methods may be supplemented 109 by additional filtering to remove the atmospheric phase screen (APS) from the time se-110 ries, by firstly high-pass filtering in time, then low-pass filtering in space to calculate the 111 APS, which is then removed from the time series (e.g. Sousa et al., 2011). The conven-112 tional method for APS calculation relies on the assumption that the atmospheric delay 113 is not temporally correlated. With recent SAR missions providing shorter satellite re-114 visit times, this assumption may no longer be appropriate. Previous studies have pro-115 posed improvements to the APS correction, including applying a global weather model 116 (e.g. Jung et al., 2013), and accounting for the temporal variance of a pixel (e.g. Liu et 117 al., 2011; Refice et al., 2011). 118

The only previous InSAR derived velocity map covering the whole Afar region was 119 developed by Pagli et al. (2014), who used II-RATE to produce a displacement time se-120 ries between 2005-10. After removing large deformation steps associated with the DMH 121 dyke intrusion events, Pagli et al. (2014) smoothed the time series by removing the APS, 122 employing consistent Gaussian temporal and Butterworth spatial filters. Pagli et al. (2014) 123 extracted 3D (east, north, vertical) velocities from ascending and descending LOS and 124 Global Navigation Satellite Systems (GNSS) observations on a 10-20 km resolution mesh 125 following the method of Wang and Wright (2012). Surface velocities between 2005-10 126 from Pagli et al. (2014) showed a long-term plate spreading rate of 15-20 mm/yr in East-127

ern Afar (relative to a stable Nubian plate), and large extension rates of ~ 100 mm/yr
at Dabbahu associated with the background response during the 2005-10 DMH rifting
episode. Other InSAR velocity maps within Afar have focussed on individual rift segments, such as the DMHVS (Hamling et al., 2014) from 2006-10, the AGVS from 19972005 (Doubre & Peltzer, 2007), and around the Tendaho Graben from 2004-10 (Temtime
et al., 2018).

In this study, we make use of the high temporal and spatial resolution data from the Sentinel-1 satellite to extract surface velocities from 2014-19 across the whole of the Afar region. We develop and apply a small baseline methodology with spatial and temporal variance weighted filtering to improve the removal of the APS, reference the data to a stable Nubia GNSS reference frame, and calculate 3D (vertical, rift-perpendicular horizontal, and rift-parallel horizontal) velocities.

¹⁴⁰ 2 Data Processing and Time Series Methods

2.1 Sentinel-1 Data

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We use Sentinel-1A/B acquisitions from ascending tracks 14 (014A), 87 (087A), 142 and 116 (116A), and descending tracks 6 (006D) and 79 (079D) between October 2014 143 and August 2019. For processing efficiency, we divide each track into $12 (\sim 250 \times 250 \text{ km})$ 144 frames (Table S1). We produce a network of geocoded unwrapped interferograms for each 145 frame from single-look complex (SLC) images of each date using the LiCSAR software 146 (González et al., 2016; Lazecky et al., in review.), which automates the mass production 147 of interferograms using GAMMA (Werner et al., 2000). To reduce noise and data size, 148 we multi-look the SLCs at 20:4 range to azimuth looks, equating to $\sim 100 \times 100$ m pixel 149 size. We apply a topographic correction using an SRTM (Shuttle Radar Topography Mis-150 sion) 3-arc-second (~ 90 m resolution) DEM (Digital Elevation Model) (Farr & Kobrick, 151 2000), filter the interferograms using using a power spectrum filter (Goldstein et al., 1998), 152 and unwrap using SNAPHU (Chen & Zebker, 2002). We manually quality check the in-153 terferogram network for each frame to remove interferograms with decorrelation, co-registration, 154 or obvious unwrapping errors. We ensure that each epoch is connected to the network 155 by a minimum of 3 interferograms by creating new interferograms as needed. Each in-156 terferogram is referenced to the mean value, excluding areas of deformation around vol-157 canic centres. 158

To mitigate for atmospheric phase delay for each interferogram we compare the ef-159 fectiveness of corrections from a linear trend of phase with elevation over the whole frame 160 (e.g. Elliott et al., 2008), and the GACOS atmospheric model (e.g. Yu et al., 2017, 2018). 161 For a linear phase-elevation trend correction, the mean root-mean-square (RMS) mis-162 fit for all 12 frames is reduced by 2.9 mm in comparison to the mean RMS misfit for all 163 frames with no atmospheric correction. The GACOS correction gives a reduction in mean 164 RMS misfit of 4.2 mm, but increases the RMS misfit in 29 % of interferograms. To re-165 duce this we follow an approach similar to Shen et al. (2019), scaling the GACOS cor-166 rection for each interferogram in order to minimise the resulting RMS misfit. This im-167 proves the atmospheric correction further, producing a reduction in mean RMS misfit 168 of 5.7 mm (see Figure S1). In order to account for any residual topographic atmospheric 169 signal in each frame, we remove a linear trend of phase with elevation from each epoch, 170 after time series filtering. 171

172 2.2 Time Series

We apply an SBAS style methodology to invert for the displacement time series at each pixel in the frame, using all interferograms where a pixel is coherent. We assess the spatial and temporal variance of the time series by firstly, filtering using a Laplacian filter with a temporal width of 3 epochs and scale factor of 3, then calculating the RMS misfit from this trend at each pixel for every epoch. We calculate the spatial distribution of RMS misfit from the time series misfits at each point, and the temporal distribution of RMS misfit from the misfits of all pixels at each epoch (see Figure S2). To resolve the RMS misfit value of each pixel at every epoch, we scale the spatial RMS misfit map to the temporal RMS misfit value at each epoch (see Supplementary Materials).
We use these error estimates to provide weights during time series filtering, and in the
inversion of filtered displacement time series for average velocities.

In order to reduce the remaining APS in the displacement time series, we filter the 184 time series using a high-pass temporal and a low-pass spatial filter to produce the APS, 185 which we then remove from the time series. To calculate a low-pass temporal filter, we 186 apply a weighted linear trend with a fixed temporal width of 1 year centred on each point. 187 To calculate the weighting for the local trend, we use the RMS misfit as a proxy for stan-188 dard error, and convert the RMS misfit values into weights using the Bi-Square function 189 where no weight is given to RMS values that exceed 6 standard deviations of the local 190 misfits (e.g. Cleveland & Devlin, 1988). We also scale these weights by their temporal 191 distance from the target epoch of the local time series (see Supplementary Materials). 192 Having calculated the low-pass temporal filter, we remove it from the time series to cre-193 ate a high-pass temporal filter. We then apply a Gaussian spatial filter with a half-width 194 of ~ 2 km in order to resolve the APS for each epoch. 195

After we remove the APS, we remove a planar ramp in space and a linear trend 196 with height to correct for any remnant long-wavelength and elevation-correlated atmo-197 spheric delay. We later correct for long-wavelength deformation using GNSS observa-198 tions. For each frame, we compute the average velocity at each pixel by inverting for a 199 single linear trend through time, allowing for a constant offset. We produce a variance-200 covariance matrix (VCM) for each pixel, treating the temporal variation of the scaled 201 RMS misfit as independent errors. By including the VCM in the inversion, we can quan-202 tify the uncertainty of the resulting velocities. 203

De Zan et al. (2015) demonstrate how a potential systematic phase-bias in inter-204 ferograms with decreasing temporal baseline can influence the resulting time series. We 205 test the magnitude of this bias by selecting consecutive 12, 24, and 36-day interferograms 206 from frame $079D_07694_{-}131313$ covering ~ 1 year (see Figure S3). We use a 'daisy-chain' 207 stack approach to resolve the cumulative displacements from the 12, 24, and 36 day un-208 wrapped interferograms between December 2017 and February 2019. Any differences between these stacks indicates the presence of phase-bias and/or unwrapping errors. We 210 find residual differences between the 12 and 24-day, and 12 and 36-day stacks of up to 211 50 mm, and residuals of up to 10 mm between the 24 and 36-day stacks. While this in-212 dicates that the 12-day interferograms are susceptible to a phase-bias, we find that re-213 moving the 12-day interferograms from the network effects our displacement time series 214 by < 5 mm per epoch, and our average velocities by < 1 mm/yr. While we are not able 215 to account for any bias in the 36-day interferograms, Ansari et al. (2020) indicate that 216 the velocity bias is small in comparison to 12-day interferograms. 217

2.3 3D Velocities

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We tie frames together within their respective tracks by sub-sampling the InSAR 219 data points to a 5×5 km spacing in the overlap between frames, and 10×10 km spac-220 ing elsewhere, then solving for and removing a planar ramp for each frame that minimises 221 residuals in the along-track frame overlap regions. Removing these ramps does not bias 222 the results as long-wavelength signals are later corrected using GNSS. In the frame over-223 lap region, we use the mean value of LOS velocity for each point. We find that using lin-224 ear ramps to combine frames within tracks produces the fewest boundary artefacts when 225 compared to using a single offset value calculated from the median value in the frame 226 overlap region, or solving for a 2D quadratic function for each frame. Boundary arte-227

facts within tracks can occur due to differences between frames in time series length, the variation in acquisition dates used, relative weighting during time-series filtering, and orbital ramp removal. Although in principle it would be possible to only process and use interferograms that cover the whole along-track extent of the study region, this would require excluding several epochs where data were not acquired over the whole area, resulting in truncated time series.

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To reference the LOS velocity in each track to a stable Nubian plate, we use a net-234 work of 105 GNSS stations in the Afar region to characterise long-wavelength plate mo-235 tions. The data are a subset of the GeoPRISMS community velocity field for East Africa 236 in a Nubia-fixed International Terrestrial Reference Frame (ITRF2014) (King et al., 2019) 237 We remove 32 stations in central Afar where the velocities are dominated by the ground 238 motions associated with the 2005-10 DMH rifting episode. As the resulting network is 239 sparse, with the majority of stations concentrated in Eastern Afar and few points on the 240 Nubian and Somalian plates; we add 17 additional fabricated GNSS stations on the sta-241 ble Nubian plate, with an assumed zero velocity (with uncertainties of $\pm 1 \text{ mm/yr}$ and 242 $\pm 2 \text{ mm/yr}$ in the horizontal and vertical components), to help constrain the velocity 243 field where data are sparse. We project East and North GNSS horizontal velocities into 244 the rift-perpendicular (e.g. Hamling et al., 2014), and rift-parallel directions, oriented 245 at 61°N and -29°N respectively. 246

From this network, we interpolate a smooth GNSS velocity field in the rift-perpendicular 247 and rift-parallel directions over the whole Afar region at 100×100 m grid spacing (Fig-248 ure S4) using the natural neighbour algorithm (e.g. Boissonnat & Cazals, 2002). As the 249 additional pseudo-observations define where this interpolated field reaches zero veloc-250 ity, we selected these points such that they are on the Nubian plate, away from the rift 251 border faults. We are not concerned with the precise locations, as where the interpolated 252 velocity field reaches zero does not significantly influence the data within the Afar from 253 the Rift. We estimate the error in the interpolated velocities by systematically remov-254 ing each GNSS station from the network, interpolating new velocity fields in the rift-perpendicular 255 and rift-parallel directions from the reduced network, then calculating the residual be-256 tween the interpolated fields and the GNSS observation. We take the standard devia-257 tion of these residuals as the error in the rift-perpendicular and rift-parallel GNSS ve-258 locity field. 259

We sub-sample the InSAR LOS track velocities as previously, then extract points 260 where there are ascending, descending, and interpolated GNSS data. We also mask points 261 around the active rift segments so that volcanic ground deformation does not interfere 262 with the referencing to the long-term plate motions. Using these points, we solve for the 263 3D velocity (rift-perpendicular, rift-parallel, vertical) at each point and a residual 2D (East, 264 North) quadratic function for each track. We remove the respective quadratic from each 265 InSAR track to resolve LOS velocity in a stable Nubia reference frame. The resulting 266 LOS velocities and standard deviations are shown in Figures 2 and S5 respectively. 267

In order to resolve a full 3D velocity field (vertical, rift-perpendicular horizontal, 268 rift-parallel horizontal) at 100×100 m resolution, we use the smooth rift-parallel GNSS 269 field to provide a constraint on the rift-parallel velocity at each point, as the rift-parallel 270 velocities are small in comparison to the rift-perpendicular and vertical velocities. We 271 include this constraint with the ascending and descending LOS InSAR observations to 272 calculate 3D velocities at each point using a least-squares inversion (Wright et al., 2004; 273 Hussain et al., 2016; Weiss et al., 2020). The decomposition of InSAR LOS velocities (L)274 into rift-perpendicular (H1), rift-parallel (H2), and vertical (Z) velocities is shown in 275 276 Equation 1 for the incidence angle (θ) , satellite heading (α) , rift angle from North (ϕ) . We weight the inversion and resolve uncertainties by including a diagonal VCM using 277 the previously calculated variance at each point. 278



Figure 2. LOS average velocities over the Afar region between November 2014 and August 2019 from Sentinel-1 tracks T087A, T014A, T116A, T079D, and T006D. LOS velocities are referenced to a stable Nubia reference frame using long-term plate motions from the regional GNSS network (King et al., 2019). Arrows indicate the track look directions, and the volcanic segments are shown as black outlines.

Results (Figure 3), discussed in Section 3, show both the long-term plate motion and surface deformation associated with magmatism. As we use the interpolated GNSS velocity field as an additional constraint in the rift-parallel direction, the resulting riftparallel error estimates are lower than the rift-perpendicular errors, which we calculate using only the InSAR observations to retain high spatial resolution (see Figure 3).

$$L = \begin{bmatrix} -\sin(\theta)\cos(\alpha + \phi) + \sin(\theta)\sin(\alpha + \phi) - \cos(\theta) \end{bmatrix} \begin{bmatrix} H1\\H2\\Z \end{bmatrix}$$
(1)

²⁷⁹ **3** Key Findings & Discussion

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3.1 Plate Motions & Rift Focussing

Our 2014-2019 horizontal velocity maps (Figure 3) show the rift-perpendicular extension over the Afar rift at rates of up to $25 \pm 5 \text{ mm/yr}$, with negligible motions in the rift-parallel direction. We also observe the 'trapdoor' motion of the Danakil micro-plate relative to the stable Nubian plate (Kidane, 2016), where rift-perpendicular extension in the RSR is gradually transferred into Afar between 13-16°N. Areas of noise up to $\pm 10 \text{ mm/yr}$ over the Ethiopian highlands region on the Nubain plate, are highlighted in the standard deviation maps shown in Figure 3. The regions of high error in T087A, and the northernmost portion of T014A (see Figure S5), are a result of the shorter time series length in these regions producing more uncertainty in the long-term velocity estimates. Elsewhere, errors of up to \pm 5 mm/yr are likely due to uncorrected atmospheric delays, and artefacts over track boundaries, where we are unable to account for small LOS velocity variations between overlapping tracks.

Profiles taken across the rift highlight the focussing of extension in Afar. Profiles 293 covering the Alid graben, at northern-most tip of the Afar rift (Figures 4B and 5B), show that a broad uplift and extensional signal of up to $\sim 20 \text{ mm/yr}$ is centred within ± 10 -295 15 km of the rift axis. This may be indicative of deep magmatic intrusion in an area with 296 the smallest background extension rates throughout Afar, but comparable to extension 297 rates at the active volcanic islands at the southern end of the oceanic RSR (Eyles et al., 298 2018). Profiles traversing the EAVS in the Danakil Depression (Figures 4C and 5C) high-299 light that the majority of extension here is focussed into a region within \pm 15-20 km of 300 the rift axis. We also observe subsidence on the EAVS between the Erta 'Ale and Alu 301



Figure 3. Vertical, rift-perpendicular, and rift-parallel average velocities and standard deviation over the Afar region between November 2014 and August 2019. All velocities are referenced to a stable Nubian plate. Vertical velocities are positive upwards, rift-perpendicular velocities are positive to the NE (61°N), and rift-parallel velocities are positive to the NW (-29°N). Standard deviation colour-scale is limited to 5 mm/yr to highlight variation in regions of low variance. Maximum standard deviations are (2 s.f.): 7.2 mm/yr (vertical), 9.5 mm/yr (rift-perpendicular), and 1.7 mm/yr (rift-parallel). GNSS stations are shown as circles with GNSS velocities on the same colour-scale as the InSAR velocities, and volcanic segments as black outlines.

³⁰² Dalafilla volcances (Figure 4C), which could be linked to magma withdrawal associated ³⁰³ with the 2017-19 eruption at Erta 'Ale (Moore et al., 2019). Profiles between the EAVS ³⁰⁴ and the DMHVS (Figure 5C) show that extension in this region is shared between the ³⁰⁵ ALVS and the TAVS, and focussed to within \pm 10-20 km of the rift segments.

Profiles covering the DMHVS (Figures 4D and 5D) also show that the long-term extension is concentrated near the rift-axis, with only small variations in rift-perpendicular velocity away from \pm 20-30 km of the segment centre. Elevated velocities close to the segment centre are associated with the 2005-10 DMHVS rifting episode, and are discussed



Figure 4. (B-E) Vertical (blue, positive up) and rift-perpendicular (red, positive towards 61°N) velocities over 4 10 km wide cross-rift profiles marked on insert map, covering (B) the northern tip of the Afar triangle, (C) the Erta 'Ale volcanic segment, (D) the Dabbahu-Manda-Hararo volcanic segment (DMHVS), and (E) the Asal-Ghoubbet volcanic segment (AGVS). Black lines show surface elevation along the profiles, with vertical dashed lines indicating the location of the major rift axis on the profile. (F-I) Vertical and rift-perpendicular velocity maps at (F) Dallol volcano, (G) Nabro volcano, (H) the AGVS, and (I) the DMHVS. Velocities in each subset are referenced to the local background mean value, with contours indicating elevation. The location of modelled deformation sources for a 0.9-1.3 km deep sill (Okada, 1985) at Dallol (F, Figure S6) and a 5.5-6.8 km deep point source (Mogi, 1958) at Nabro (G, Figure S7) are shown as black outlines.



Figure 5. Map of 21 10 km wide rift-perpendicular velocity (relative to stable Nubia, positive towards 61°N) profiles over the Afar rift (A), with rift segments outlined in white. To help distinguish between profiles, profiles alternate between solid and dashed lines. The velocity profiles cover northern Afar (B), the Erta 'Ale and Tat 'Ale segments (C), the Dabbahu-Manda-Hararo segment (D), and southern Afar (E). The standard deviation of rift-perpendicular velocities varies from \pm 2-7 mm/yr (see Figure 3). Profile distances are relative to the rift axis, with positive towards the NE. The location of the rift axis is marked by circles on the map (A), and dashed black lines on profiles (B-E). Grey shading on the profiles indicates the region where the majority of rift extension is accommodated.

in Section 3.2. Profiles over southern Afar and the AGVS (Figures 4E and 5E) show a
more distributed pattern of extension with an increase in rift-perpendicular velocities
80-160 km to the SW of the rift axis, before velocities stabilise at ~ 20 mm/yr on the
Danakil micro-plate within 10-20 km to the NE of the rift axis. In southern Afar, extension between 2014-19 may be largely accommodated by tectonic rather than magmatic
mechanisms, with strain being distributed across a sequence of horst and graben structures.

Our results indicate that at the more mature segments with active magmatism in central and northern Afar, extension is largely focussed to within \pm 15-30 km of the rift axis; while at less mature segments without active magmatism in southern Afar, extension may be distributed of a over 90-180 km. The broad distribution of strain in southern Afar is in general agreement with 1992-2010 GNSS observations from Kogan et al. (2012), who suggest that extension along a profile in southern Afar occurs over ~ 175 km.
Kogan et al. (2012) also suggest that extension becomes more distributed with rift development. In contrast, our results suggest an increase in focussing with rift maturity
during late-stage continental break-up, in keeping with strain localisation assisting the
transition into oceanic spreading centres.

3.2 Magmatic Deformation

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Figure 4 highlights the localised surface deformation at Dallol (4F) and Nabro (4G) 328 volcanoes, and at the DMHVS (4I), where localised deformation, likely associated with 329 magma migration, is visible. Figure 4H over the AGVS indicates the lack of magma re-330 lated deformation at this segment between 2014-19. As magmatic deformation may not 331 be steady in time, we look at time series for points located in the middle of these cen-332 tres. Time series of vertical displacements at the Dallol, Nabro, Dabbahu, and Manda-333 Hararo volcanic centres show that the deformation is linear through time, indicating that 334 the velocities are representative (Figure 6). For Erta 'Ale volcano, we select a point $\sim 2 \text{ km}$ 335 to the north of the summit caldera in order to avoid the step surface deformation asso-336 ciated with a dyke intrusion in January 2017 (Moore et al., 2019). Following this intru-337 sion the Erta 'Ale edifice shows linear subsidence at a rate of $15 \pm 4 \text{ mm/yr}$ (Figure 6). 338

At Dallol volcano, at the northern end on the EAVS, we observe a high rate of subsidence of up to 45 ± 4 mm/yr, with negligible horizontal movement. The subsidence signal is focussed on the central cone at Dallol. We model this signal using the Markov-Chain Monte-Carlo Geodetic Bayesian Inversion Software (GBIS) (Bagnardi & Hooper, 2018). For the T014A and T079D LOS deformation between 2014-19, we test source ge-



Figure 6. 2014-19 time series of vertical displacements (uplift positive) with 1 standard deviation estimates, at 5 deforming volcanic centres in Afar. The points used for the time series are representative of the whole edifice deformation and are positioned at the centre of the volcanic edifice; with the exception of Erta 'Ale where the point selected is ~ 2 km North of the summit lava lake in order to avoid surface deformation associated with the Jan 2017 dyke intrusion event (Moore et al., 2019). Coordinates of points used for each volcanic centre: Dallol - 40.3°E/14.243°N, Nabro - 41.685°E/13.38°N, Erta 'Ale - 40.65°E/13.62°N, Dabbahu - 40.48°E/12.58°N, Manda-Hararo - 40.88°E/12.13°N.

ometries including a point pressure source (Mogi, 1958), a planar dislocation (Okada, 1985), and a penny-shaped crack (Fialko et al., 2001). We find that a $\sim 1 \times 2$ km horizontal sill at 0.9-1.3 km depth with ~ 0.27 m of contraction gives the lowest residual RMS misfit (Figure S6). This is in agreement with Nobile et al. (2012) who inferred the presence of a deflating magma body at 1.5-3.3 km depth beneath Dallol between 2004-06.

At Nabro volcano, we observe edifice uplift and extension at rates of up to $12 \pm 3 \text{ mm/yr}$, 350 combined with the subsidence of lava flows from the 2011 eruption (J. E. Hamlyn et al., 351 2014), and a highly localised subsidence and contraction signal of up to $14 \pm 3 \text{ mm/yr}$ 352 at the centre of the Nabro caldera. This uplift of the Nabro edifice shows a change from 353 InSAR observed subsidence of 150-200 mm/yr from 2011-12, after the 2011 eruption (J. Ham-354 lyn et al., 2018). As our vertical displacement time series (Figure 6) indicates that the 355 uplift at Nabro is linear between 2014-19, we suggest that this post-eruption edifice sub-356 sidence must have stopped between 2012-14. We model T006D, T014A, and T079D LOS 357 observations at Nabro volcano between 2014-19 using a point pressure source (Mogi, 1958), 358 and penny-shaped crack (Fialko et al., 2001) using GBIS (Bagnardi & Hooper, 2018). 359 Figure S7 shows the optimal Mogi source at 5.5-6.8 km depth with a volume increase 360 of $7-11 \times 10^6$ m³. The location of this source is in agreement with the magma chamber 361 at 6.4 ± 0.3 km depth inferred by (J. Hamlyn et al., 2018) between 2011-12, indicating 362 re-charge of the melt storage below Nabro volcano. 363

After the initial 2005 intrusion at the DMHVS, background uplift and rift-perpendicular 364 extension continued throughout the 2005-10 rifting episode at rates of up to 80-240 mm/yr365 and 110-180 mm/yr respectively, from 2006-10 at the segment centre (Pagli et al., 2014; 366 Hamling et al., 2014). We show that this uplift and extension is ongoing between 2014-367 19 at average rates of $33 \pm 4 \text{ mm/yr}$ and $37 \pm 4 \text{ mm/yr}$ respectively (Figure 4c,h). Our 368 vertical velocities also show that an area of $\sim 20 \text{ mm/yr}$ subsidence from 2006-10 at the 369 southern end of the DMHVS (40.9°E/12.1°N) uplifts at an average rate of $18 \pm 4 \text{ mm/yr}$ 370 from 2014-19. This decaying post-rifting response to the initial 2005 intrusion may be 371 indicative of continued magma movement beneath the DMHVS and/or a time-dependent 372 viscous response due to a more ductile rheology generated from repeated intrusions at 373 the rift segment. 374

4 Conclusions

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We develop Sentinel-1 displacement time series at 100×100 m resolution between 2014-19 over 3 ascending and 2 descending tracks, covering the whole Afar rift. We implement a RMS misfit weighted APS correction to clean the time series, and produce average velocity maps for each frame. Using GNSS observations of long-term plate motions, we reference the InSAR velocities to the stable Nubian plate, and convert LOS to 3D velocities (vertical, rift-perpendicular, rift-parallel).

We are able to resolve deformation at individual volcanic centres, with subsidence of $45 \pm 4 \text{ mm/yr}$ at Dallol volcano, consistent with the deflation of a shallow sill at 0.9-1.3 km depth. We also show that edifice uplift at Nabro volcano of $12 \pm 3 \text{ mm/yr}$ is sourced from a magma chamber at 5.5-6.8 km depth, consistent with the source of post-eruption subsidence observed between 2011-12 (J. Hamlyn et al., 2018).

Pagli et al. (2014) and Hamling et al. (2014) identify vertical and rift-perpendicular horizontal surface velocities between 2006-10 of 80-240 mm/yr and 110-180 mm/yr, respectively associated with a background post-rift response to the initial 2005 dyking episode at the DMHVS. We show that this response is ongoing between 2014-19, but at lower rates of 33 ± 4 mm/yr and 37 ± 4 mm/yr respectively. We suggest that this ~ 15 year response to the 2005 dyke intrusion is indicative of continued magma movement and/or time-dependant viscous processes within the crust below the rift segment.

We resolve the long-term motion of the Danakil micro-plate with rift-perpendicular 394 velocities of up to 25 ± 5 mm/yr, with negligible motions in the rift-parallel direction. 395 From cross-rift profiles, we find that extension is largely focussed to within \pm 15-30 km 396 of the rift-axis on the active magmatic segments in northern Afar, while strain in south-397 ern Afar is distributed across 90-180 km of the rift. This trend of increased focussing of 398 extension with rift maturity is in contrast to the trend suggested by Kogan et al. (2012), 399 but consistent with strain localisation assisting the transition into oceanic spreading cen-400 tres. 401

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