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A New Polarimetric Persistent Scatterer Interferometry Method using Temporal Coherence Optimization

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Abstract- While polarimetric Persistent Scatterer InSAR (PSI) is an effective technique for increasing the number and quality of selected persistent scatterer (PS) pixels, existing methods are suboptimal; a polarimetric channel combination is selected for each pixel based either on amplitude, which works well only for high amplitude scatterers such as man-made structures, or on the assumption that pixels in a surrounding window all have the same scattering mechanism. In this study, we present a new polarimetric PSI method in which we use a phase-based criterion to select the optimal channel for each pixel, which can work well even in non-urban environments. This algorithm is based on polarimetric optimisation of temporal coherence, as defined in the Stanford Method for Persistent Scatterers (StaMPS), to identify scatterers with stable phase characteristics. We form all possible co-polar and cross-polar interferograms from the available polarimetric channels and find the optimum coefficients for each pixel using defined search spaces to optimise the temporal coherence. We apply our algorithm, PolStaMPS, to an area in the Tehran basin that is covered primarily by vegetation. Our results confirm that the algorithm substantially improves on StaMPS performance, increasing the number of PS

pixels by 48%, 80% and 82% with respect to HH+VV, VV and HH channel, respectively, and increasing the signal-to-noise ratio of selected pixels.

Keywords: polarimetric Persistent Scatterer InSAR, StaMPS, temporal coherence.

1. INTRODUCTION

Persistent Scatterer InSAR (PSI) is a well-known technique to address decorrelation and atmospheric noise in conventional interferometry. This method identifies only those scatterers which display coherent scattering behaviour over time, known as persistent scatterers. A PSI algorithm was outlined first by Ferretti *et al.* [1], [2] with further algorithms quickly following [3], [4], [5] and [6]. In these algorithms, an initial set of PS pixels are identified by analysis of their amplitude scintillations in a series of co-registered SLC images and then refined based on the match of their phase with a pre-defined deformation model. Thus, in general, only bright scatterers with a deformation behaviour close to the assumed model are identified as PS pixels, and these algorithms work best where there are large numbers of man-made structures. Moreover, small baseline SAR differential interferometry approaches were presented by [7] and [8] based on appropriate combination of different interferograms produced by data pairs with small orbital separation (baseline) in order to limit the spatial decorrelation. In these methods, coherent pixels are selected through spatial coherence estimation.

An alternative PSI method, was put forward by [9] to identify large numbers of PS pixels in all terrains, including non-urban areas that lack man-made structures. This approach uses the spatial correlation of phase for identification of PS pixels. The parameter used to characterize phase stability in this approach is similar to a measure of coherence in time [10] and we refer to it as temporal coherence [9]. The ensemble phase coherence defined by [2], is not quite the same as the temporal coherence we refer to, as it requires a predefined deformation model.

Before the launch of radar sensors operating with a polarimetric configuration, SAR interferometry applications had been limited to a single polarimetric channel. Radar

polarimetry is a valuable technique for the extraction of geophysical parameters from SAR images [11] and [12]. Varying approaches to achieve this are based either on the statistical analysis of the polarimetric information [13] , [14] or on scattering models, which provide an understanding of the physics of the scattering process [15] , [16] and [17]. Therefore the introduction of polarimetric techniques in interferometric applications can improve performance of SAR interferometry. A general formulation for coherent conventional interferometry using polarimetry was introduced by Cloude and Papathanassiou [18]. This method sets up a spatial coherence optimisation problem using different polarimetric channels and then solves it to obtain the optimum linear combination of channels that leads to the best phase estimates. The decorrelation terms are decreased with the spatial coherence optimisation, and signal-to-noise ratio is therefore increased [19]. Another spatial coherence optimisation method was proposed by Colin *et al.,* [20]. This approach optimises the coherence using the same complex unitary vector for both antennae. This coherence is called single-mechanism coherence. Given a multi-baseline data set in this method, coherence can be optimised independently for every baseline. This can lead to identification of different dominant scattering centres depending on the chosen baseline. A more robust polarimetric optimisation approach to find the most coherent and dominant scatterer is a simultaneous optimisation of multi-baseline coherence, a technique first outlined by Neumann *et al.*, [19]. This approach generally leads to lower coherence magnitudes, but the corresponding linear combination of channels and their interferometric phases are estimated on the basis of all the available data and thus more accurately.

As density and quality of PS pixels are important factors in PSI algorithms, the concept of polarimetric optimisation in the PSI algorithms was proposed in [21] and [22] with zero-baseline ground based SAR (GB-SAR) data, to improve the number of reliable pixel candidates. In [21], the simplest coherence optimisation approach is performed based on

selection of the polarimetric channel with the highest average coherence value. A polarimetric PSI approach, known as ESPO (Exhaustive Search Polarimetric Optimisation), using spaceborne data set was presented first by Navarro-Sanchez *et al.*, [23]. This method finds the optimal weights for each available polarimetric channel to obtain an optimum combination of those channels that maximises the PS selection criterion. A study of the different polarimetric optimisation techniques using both zero-baseline and multi-baseline data was carried out by Iglesias *et al.*, [24]. The main goal was the exploitation of the available polarimetric optimisation methods, in the framework of differential interferometry, to improve the density and quality of PS pixels. Moreover, Sadeghi *et al*., compared the efficiency of different multi-baseline polarimetric optimisation techniques in terms of increasing the number of PS pixels and the signal-to-noise ratio, and also presented an enhanced multi-baseline coherence optimisation method [25]. It should be noted that the use of polarimetric SAR data entails two main drawbacks when compared to conventional single-polarimetric data: an increase in the amount of data to be processed (proportional to the number of polarimetric channels) and a reduction in the size of the images in the swath direction (hence the spatial coverage) due to the doubled pulse repetition frequency required to acquire fully polarimetric data.

Polarimetric PSI implementations, up to now, either optimise amplitude-based criteria for identification of PS pixels [23] , [24] , [26] and [27], or select the polarimetric channel combination that maximises the ensemble coherence of surrounding pixels [26] , [25] and [24]. The former approach can be quite successful for bright scatterers, such as buildings, but less for natural PS. A limitation of the latter approach is the common failure of the assumption that PS pixels are surrounded by scatterers with the same scattering properties, which leads to non-optimal weights for the polarimetric channels, and to a loss of spatial resolution. In this paper, we present a new method, PolStaMPS (Polarimetric StaMPS), which uses polarimetric optimisation of temporal coherence to increase the number of selected PS pixels in all terrains,

with or without buildings. We implement the temporal coherence optimisation after computing different interferogram channels for each master and slave image. The temporal coherence optimisation method was inspired by ESPO, as it finds the weights for each interferogram channels over search spaces. PolStaMPS codes will be included in the next release of StaMPS/MTI, with full instructions added to the manual.

This paper is organised as follows. Section 2 contains the basic principles of polarimetric interferometry and a brief review of ESPO, which is a polarimetric persistent scatterer interferometry method. The concept of temporal coherence in StaMPS is introduced in Section 3, followed by our new algorithm for optimisation of the temporal coherence, PolStaMPS, in Section 4. Section 5 describes the test site and the available dual polarimetric data set to evaluate the new algorithm. In Section 6, experimental results of PolStaMPS are shown and discussed. Finally, the main conclusions are summarized in Section 7.

2. POLARIMETRIC INTERFEROMETRY

Since there is a vector value for each pixel instead of a scalar one, polarimetric interferometry can be referred to as vector interferometry [18]. The general formulation is defined in Section 2.1. One of the most effective polarimetric PSI algorithms up to now, ESPO, was presented in [26]. This technique was formulated for two different criteria of PS selection to increase the number of PS pixels, which are amplitude dispersion index and average spatial coherence. A brief overview of this method is presented in Section 2.2.

2.1 GENERAL FORMULATION

A general formulation for polarimetric SAR interferometry, presented in full by Cloude and Papathanassiou [18], is reviewed in this section. Fully polarimetric radar systems measure a 2×2 complex scattering matrix [S] for each pixel in an image [28]. Through vectorization of the scattering matrix, a coherent scattering vector $\underline{k}$ can be extracted to generalise

interferometric phase and spatial coherence. Using Pauli basis matrices, the scattering vector for each pixel can be found as [18]

$$\underline{k} = \frac{1}{\sqrt{2}}[S^{HH+VV}, S^{HH-VV}, 2S^{HV}]^T, \quad (1)$$

where T indicates the matrix transposition operation, and S^{ij} (i , j = H or V) is the complex scattering coefficient for j transmitted and i received polarization in the HV polarisation basis. In the case of dual-polarisation interferometry, considering there is no data from the cross-polar channel, as provided by TerraSAR-X, the scattering vector changes to

$$\underline{k} = \frac{1}{\sqrt{2}}[S^{HH+VV}, S^{HH-VV}]^T. \quad (2)$$

Using the outer product formed from the scattering vectors $\underline{k}_m$ and $\underline{k}_s$ for master and slave images, a 4×4 matrix can be defined,

$$\mathbf{T}_4 = \begin{bmatrix} \mathbf{T}_{\mathrm{mm}} & \Omega_{ms} \\ \Omega_{ms}{}^H & \mathbf{T}_{ss} \end{bmatrix}, \quad (3)$$

where $^{\mathrm{H}}$ stands for conjugate transpose, and $\mathbf{T}_{\mathrm{mm}}$, $\mathbf{T}_{ss}$ and Ω_{ms} are 2×2 complex matrices given by

$$\mathbf{T}_{\mathrm{mm}} = \langle \underline{k}_m \underline{k}_m^H \rangle$$

$$\mathbf{T}_{ss} = \langle \underline{k}_s \underline{k}_s^H \rangle$$

$$\Omega_{ms} = \langle \underline{k}_m \underline{k}_s^H \rangle. \quad (4)$$

In order to extend standard SAR interferometry, which uses a scalar formulation, into a vector formulation, two normalised complex vectors $\underline{\omega}_m$ and $\underline{\omega}_s$ for master and slave images, are introduced. These vectors can be called projection vectors and interpreted as linear combination of channels. The scalar complex value for each pixel can be defined as $\mu = \underline{\omega}^H \underline{k}$, which is a linear combination of the elements of $\underline{k}$. The vector interferogram is obtained as

147 $$\mu_m \mu_s^* = \left(\underline{\omega}_m^H \underline{k}_m\right)\left(\underline{\omega}_s^H \underline{k}_s\right)^H = \underline{\omega}_m^H \Omega_{\mathrm{ms}} \underline{\omega}_s,$$

148 (5)

149 where * is the conjugate operation. The interferometric phase can be extracted using

150 $$\varphi_{int} = \arg\left(\underline{\omega}_m^H \Omega_{\mathrm{ms}} \underline{\omega}_s\right). \quad (6)$$

151 Optimum values of the projection vectors can be found through polarimetric optimisation of
152 spatial coherence. The generalised vector expression for the spatial coherence ρ is given by

153 $$\rho = \frac{\left|E(\underline{\omega}_m^H \Omega_{\mathrm{ms}} \underline{\omega}_s)\right|}{\sqrt{E(\underline{\omega}_m^H \mathbf{T}_{\mathrm{mm}}\, \underline{\omega}_m) E(\underline{\omega}_s^H \mathbf{T}_{\mathrm{ss}} \underline{\omega}_s)}},$$

154 (7)

155 where $E(\ldots)$ indicates the expectation operator. In order to estimate the spatial coherence, a
156 window is required and it is assumed that the surrounding pixels in the window have similar
157 scattering properties. Therefore, in addition to the loss of the spatial details, the optimisation
158 process will not work properly in the common case where this is not true.
159 The $\underline{\omega}$ vector can be constrained to be the same all along the whole stack of images. This is
160 referred to as Equal Scattering Mechanisms (ESM), which selects the most stable scattering
161 mechanism over time for each pixel of an image set covering a case study [26]. Moreover, in
162 the case of multi-baseline spatial coherence optimisation, the averaged spatial coherence, $\overline{\rho}$, is
163 optimised according to (8).

164 $$|\overline{\rho}| = \frac{1}{K} \sum_{k=1}^{K} |\rho_k|,$$

165 (8)

166 where K is the number of interferometric pairs.

167 2.2 **ESPO**

168 Polarimetric PSI was first introduced by Navarro-Sanchez *et al.* in [23] through ESPO, which
169 is a multi-baseline ESM optimisation method. This optimisation approach consists of searching
170 for the unitary vector $\underline{\omega}$ that maximises the PS selection criteria, which can be either average

spatial coherence or amplitude dispersion index. The optimum interferogram can be found with a parametrisation of $\omega(\alpha, \psi)$, in the case of dual-polarimetry, as

$$\underline{\omega} = [\cos\alpha \ , \sin\alpha \, e^{j\psi}]^T, \begin{cases} 0 \le \alpha \le \pi/2 \\ -\pi \le \psi < \pi \end{cases} \quad (9)$$

This parametrisation of the projection vector assumes that it is unitary, $|\underline{\omega}| = 1$, and rotated such that the phase of the first element is zero. Through an exhaustive search, optimum values are found for α and ψ for each pixel. The α parameter we define here should not be confused with the α angle widely used in polarimetry after its definition in [17].

After optimisation of the quality criteria, PS pixels are selected based on a threshold average spatial coherence in multi-looked data, or a threshold amplitude dispersion index in single-looked data. More recently, the amplitude dispersion index was optimised through ESPO to improve the PS analysis in [27]. Moreover, an alternative way to optimise the coherence was proposed to decrease the computation time [29].

3. TEMPORAL COHERENCE IN StaMPS

StaMPS is a PSI technique designed to work in non-urban environments, with deformation that may be highly non-linear in time. The PS identification step in this method is based primarily on phase characteristics and can identify low-amplitude pixels more effectively than traditional amplitude-based algorithms [9].

The main criterion of PS identification, temporal coherence, is estimated using phase analysis. After forming interferograms and removing most of topographic phase, the residual phase of the *x*th pixel in the *k*th interferogram, $\varphi_{int,x,k}$, contains a contribution from several sources as

$$\varphi_{int,x,k} = \varphi_{def,x,k} + \varphi_{\alpha,x,k} + \varphi_{orb,x,k} + \varphi_{\varepsilon,x,k} + \varphi_{n,x,k}, \quad (10)$$

where $\varphi_{def,x,k}$ is the phase change due to deformation in the satellite line-of-sight (LOS) direction, $\varphi_{\alpha,x,k}$ is the phase due to difference in atmospheric delay between passes, $\varphi_{orb,x,k}$ is

the phase due to orbit inaccuracies, $\varphi_{\varepsilon,x,k}$ is the residual topographic phase due to error in the DEM, and $\varphi_{n,x,k}$ is the decorrelation noise term.

Quantification of the noise term is used to identify which scatterers are persistent [30]. Assuming spatial correlation of most of phase contributions over a specified distance, the spatial average of residual phase, $\bar{\varphi}_{int,x,k}$, is estimated using a spatial filtering as

$$\bar{\varphi}_{int,x,k} = \overline{\varphi}_{def,x,k} + \overline{\varphi}_{\alpha,x,k} + \overline{\varphi}_{orb,x,k} + \overline{\varphi}_{\varepsilon,x,k}, \quad (11)$$

where the bar denotes the spatially filtered phase, and $\overline{\varphi}_{\varepsilon,x,k}$ is the spatially filtered sum of $\varphi_{\varepsilon,x,k}$ and $\varphi_{n,x,k}$. Subtracting the spatially correlated phase, equation (11), from residual phase, equation (10), yields

$$\varphi_{int,x,k} - \bar{\varphi}_{int,x,k} = \varphi_{\varepsilon,x,k} + \varphi_{n,x,k} - \overline{\varphi}'_{\varepsilon,x,k}, \quad (12)$$

where $\overline{\varphi}'_{\varepsilon,x,k} = \overline{\varphi}_{\varepsilon,x,k} - \left(\varphi_{def,x,k} - \overline{\varphi}_{def,x,k}\right) - \left(\varphi_{\alpha,x,k} - \overline{\varphi}_{\alpha,x,k}\right) - \left(\varphi_{orb,x,k} - \overline{\varphi}_{orb,x,k}\right)$, and is assumed to be insignificant. The residual topography phase is proportional to the perpendicular component of the baseline, $B_{\perp,x,k}$, so $\varphi_{\varepsilon,x,k} = B_{\perp,x,k} G_{\varepsilon,x}$ where $G_{\varepsilon,x}$ is a proportionality constant that can be estimated. Temporal coherence, which is a measure of phase noise level and indicator of whether the pixel is a PS [30] and [31], is defined as follows

$$\gamma_x = \frac{1}{K}\left|\sum_{k=1}^{K} exp\left\{\sqrt{-1}\,(\varphi_{int,x,k} - \bar{\varphi}_{int,x,k} - \hat{\varphi}_{\varepsilon,x,k})\right\}\right|, \quad (13)$$

where K is the number of available interferograms and $\hat{\varphi}_{\varepsilon,x,k}$ is the estimate of residual topographic phase. For each PS candidate, $\bar{\varphi}_{int,x,k}$, $\hat{\varphi}_{\varepsilon,x,k}$ and relevant γ_x are estimated in an iterative process until temporal coherence convergence is achieved. Finally, PS pixels are selected based on the probability that their phase time series is not just noise, by comparing the joint probability density function (PDF) of coherence and amplitude dispersion index to that for simulated pixels with random phase.

4. TEMPORAL COHERENCE OPTIMISATION IN PolStaMPS

All polarimetric PSI algorithms to date have utilised spatial coherence or the amplitude dispersion index to optimise the weights for the different polarimetric channels. Amplitude-based polarimetric PSI is only useful for high amplitude PS. On the other hand, using spatial coherence to select PS pixels relies on surrounding pixels having the same mechanism, which is often not the case for PS pixels.

In our new algorithm we extend the approach of StaMPS, which uses temporal coherence to select PS with high-density in non-urban areas. The main goal of the algorithm is to find the weights for the polarimetric channels that optimise the temporal coherence for each pixel. In addition to optimising the phase-based criterion, implementing the optimisation process after forming interferograms and removing the topographic contribution is a difference of PolStaMPS compared to other polarimetric PSI algorithms.

The optimum interferogram phase, $\varphi_{opt-int,x,k}$, obtained from substituting equation (2) in equation (5) is

$$\varphi_{opt-int,x,k} = \arg\left(\underline{\omega}_m^H \Omega_{\mathrm{ms}} \underline{\omega}_s\right)$$

$$= \arg\left(\frac{1}{2}\begin{bmatrix}\omega_m^{i*} & \omega_m^{ii*}\end{bmatrix}\begin{bmatrix}S_m^{HH+VV}.S_s^{HH+VV*} & S_m^{HH+VV}.S_s^{HH-VV*} \\ S_m^{HH-VV}.S_s^{HH+VV*} & S_m^{HH-VV}.S_s^{HH-VV*}\end{bmatrix}\begin{bmatrix}\omega_s^{i} \\ \omega_s^{ii}\end{bmatrix}\right)$$

$$= \arg\left(f_1.\frac{1}{2}(S_m^{HH+VV}.S_s^{HH+VV*}) + f_2.\frac{1}{2}(S_m^{HH+VV}.S_s^{HH-VV*}) + f_3.\frac{1}{2}(S_m^{HH-VV}.S_s^{HH+VV*}) + f_4.\frac{1}{2}(S_m^{HH-VV}.S_s^{HH-VV*})\right)$$

$$= \arg\left(f_1.\Phi_{int-1,x,k} + f_2.\Phi_{int-2,x,k} + f_3.\Phi_{int-3,x,k} + f_4.\Phi_{int-4,x,k}\right), \quad (14)$$

where $\Phi_{int-1,x,k} \ldots \Phi_{int-4,x,k}$, elements of $[\Omega_{ms}]$, are 4 different types of interferogram, whose linear combination forms the optimum kth interferogram for the xth pixel. ω^i and ω^{ii} are the first and second element of $\underline{\omega}$. $f_1 \ldots f_4$ are coefficients for the 4 types of interferogram as

$$f_1 = \omega_m^{i*}.\omega_s^{i}$$
$$f_2 = \omega_m^{ii*}.\omega_s^{i}$$

$$f_3 = \omega_m^{i*}.\,\omega_s^{ii}$$
$$f_4 = \omega_m^{ii*}.\,\omega_s^{ii}.$$
(15)

The polarimetric expression of temporal coherence is introduced in (16). Similar to standard StaMPS, there is an iterative process to estimate $\bar{\varphi}_{opt-int,x,k}$, which is substituted by the spatially correlated phase of $\Phi_{int-1,x,k}$ in the first iteration. In every iteration, after applying a spatial filtering to calculate $\bar{\varphi}_{opt-int,x,k}$, the optimum values for f_1 … f_4 and $\hat{\varphi}_{\varepsilon,x,k}$ are found in the defined search spaces to optimise $\gamma_{pol,x}$ and then the final value of the $\hat{\varphi}_{\varepsilon,x,k}$ is estimated through the obtained optimum phase. In the final iteration, polarimetric temporal coherence converges, and the coefficients and the optimum interferograms, according to (14), are obtained.

$$\gamma_{pol,x} = \frac{1}{K}\left|\sum_{k=1}^{K} exp\{\sqrt{-1}(\varphi_{opt-int,x,k} - \bar{\varphi}_{opt-int,x,k} - \hat{\varphi}_{\varepsilon,x,k})\}\right|.$$
(16)

In order to optimise $\gamma_{pol,x}$, the coefficients are parametrised based on the definition of $\underline{\omega}$ in ESPO as

$$f_1 = \cos\alpha\ .\cos\alpha\ = \cos\alpha^{\,2}$$
$$f_2 = \sin\alpha\, e^{-j\psi}.\cos\alpha = \sin\alpha.\cos\alpha.\, e^{-j\psi}$$
$$f_3 = \cos\alpha\ .\sin\alpha\, e^{j\psi} = \cos\alpha\ .\sin\alpha\, e^{j\psi}$$
$$f_4 = \sin\alpha\, e^{-j\psi}.\sin\alpha\, e^{j\psi} = \sin\alpha^2.$$
(17)

Therefore, only a two-dimensional search space is defined by α and ψ in each iteration, and the best values are extracted for each one. In order to define coefficients and then optimise the temporal coherence, we specified a grid for the search space of α and ψ values, with 10 degrees steps. Steps larger than 10 degrees would yield a shorter computing time, but due to the relatively complex pattern of the temporal coherence function, may cause convergence on a local maximum rather than the absolute one.

5. CASE STUDY AND DATA SET

Since the main priority of this research is increasing PS density in non-urban areas, we selected Tehran basin, which contains areas primarily covered by vegetation, as a test case. The Tehran basin suffers from a high-rate of land subsidence and is located in the north of Iran, between the Alborz Mountains to the north and the Arad and Fashapouye mountains to the south. This subsidence was first revealed by geodetic observations from precise levelling surveys carried out across the area between 1995 and 2002 [32]. Due to poor coherence, conventional interferometry has generally not been successful in measuring deformation. Therefore, a number of enhanced algorithms based on PSI have been applied to this region [33], [34] and [35]. We applied our new PolStaMPS method to a 2.6 × 1.2 km portion of the Tehran basin containing pixels with the highest rate of deformation and covered mostly by agricultural fields (Figure 1).

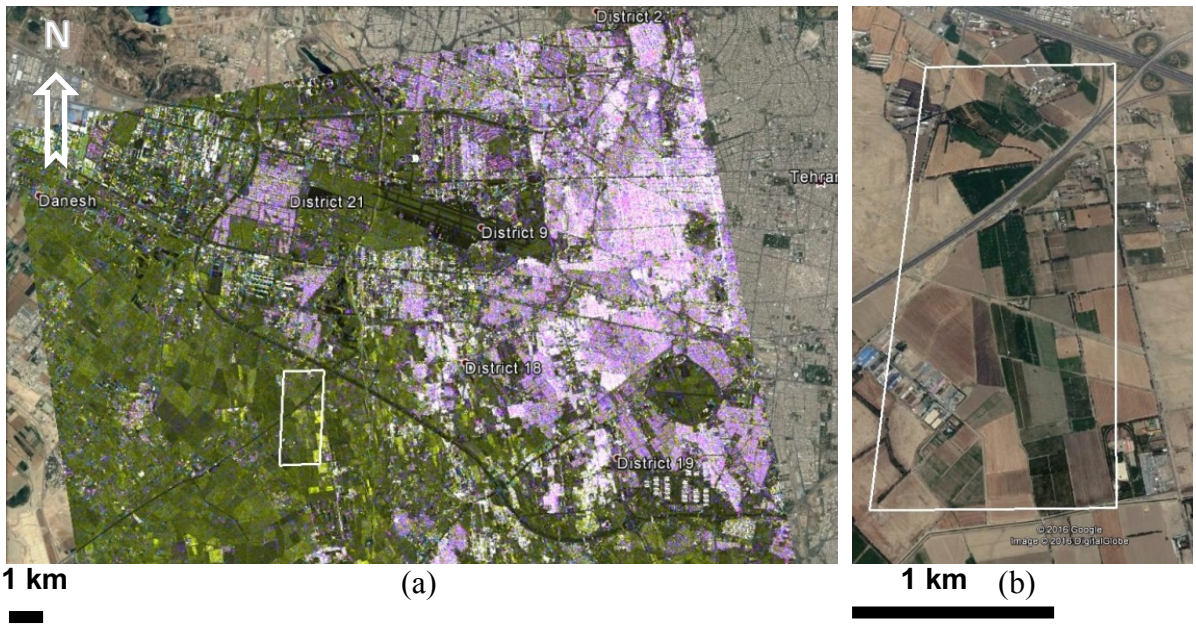


Fig. 1 (a) Spatial location of the case study (outlined polygon) over the composite RGB of master image (20131211), Channels: R=HH, G=VV, B: Absolute value of the difference between channels. (b) The case study (outlined rectangle) with detailed features.

In order to optimise the temporal coherence using polarimetric data, we tasked TerraSAR-X to acquire dual-polarisation (HH/VV) images. A set of 22 dual-polarisation Strip-map images from 21 July 2013 to 22 April 2014 were obtained. Azimuth and slant-range resolution are 6.6 and 1.17 m, respectively, whereas pixel dimensions are 2.4 and 0.91 m, respectively. Fig. 2 illustrates the spatial and temporal baselines of all slave images with respect to the master one.

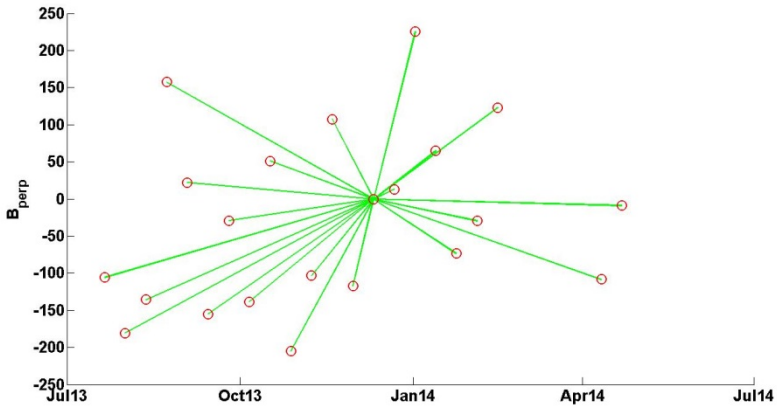


Fig. 2. Spatial baselines vs. temporal baselines of slave images with respect to the master (20131211).

6. PolStaMPS RESULTS AND DISCUSSION

In addition to the linear channels (HH and VV), we also ran StaMPS on the HH+VV channel, which forms the initial co-polar interferogram in PolStaMPS, $\Phi_{int-1,x,k}$, as its phase values are expected to be more stable over surface scattering areas, e.g. rural ones, than the linear channels.

Figure 3 displays the polarimetric temporal coherence values as a function of (α, ψ) for four representative pixels with different values of optimum temporal coherence. The shape of the temporal coherence function is smooth enough to allow numerical methods to approximate the maximum value. For this reason, a point close to the absolute maximum of the temporal coherence is first found using a grid search, and then a gradient-based method is used to find the maximum, hence reducing the computational cost.

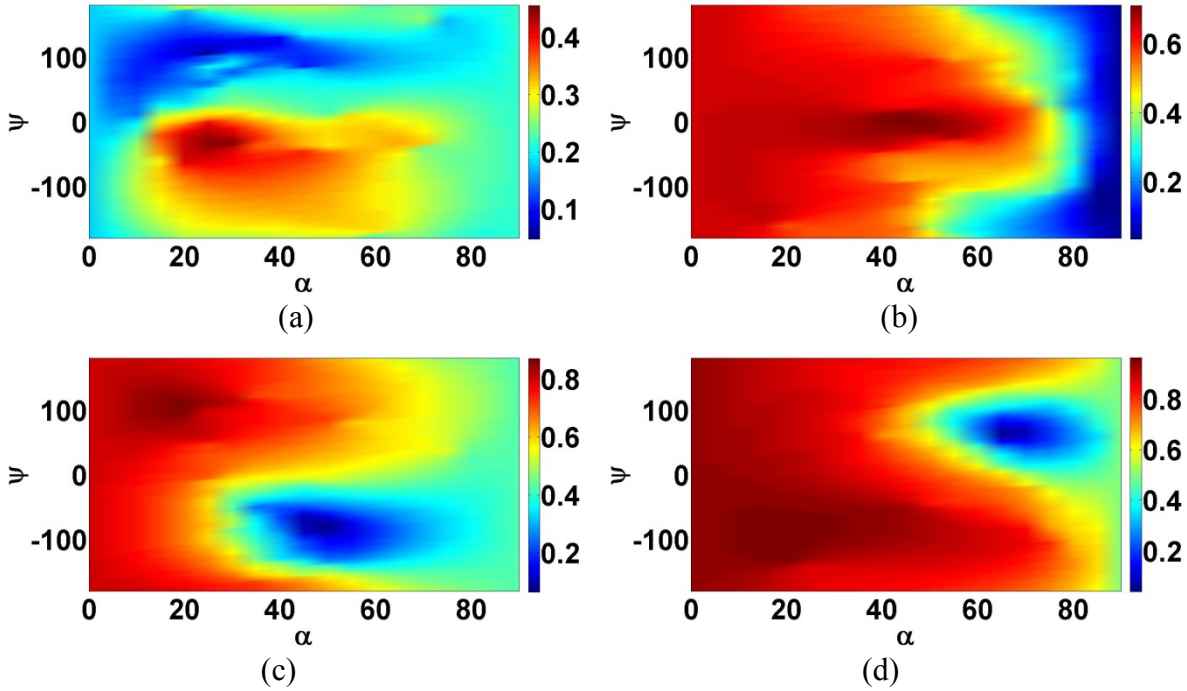


Fig .3. Temporal coherence values as a function of (α, ψ) for four representative pixels with different values of $\gamma_{pol,x}$. (a) $\gamma_{pol,x}$=0.456, (b) $\gamma_{pol,x}$=0.711, (c) $\gamma_{pol,x}$ =0.871, (d) $\gamma_{pol,x}$=0.962.

Histograms of the estimated $\gamma_{pol,x}$ in PolStaMPS and the estimated γ_x in standard StaMPS for initial selected pixels are compared in Figure 4. This comparison shows a significant increase in the number of pixels with high temporal coherence for the optimum channel, compared to the HH, VV and HH+VV channels. The increase in coherence will be, in part, due to an increase in the bias. For instance, coherence estimated on the sea is not zero (as it should be theoretically) due to the estimation bias in any single channel and, moreover, increases in the optimum polarimetric combination. To test whether the entire coherence increase can be explained by an increase in the bias, we check (below) the spatial distribution of the optimum coefficients, and compare the noise levels of selected points in the original channels to those in the optimum channel. We note, however, that in any case, the increase in bias should not lead to more pixels being selected, due to the StaMPS mechanism for pixel selection, which depends on a comparison of the coherence distribution to that for simulated pixels, rather than simple thresholding.

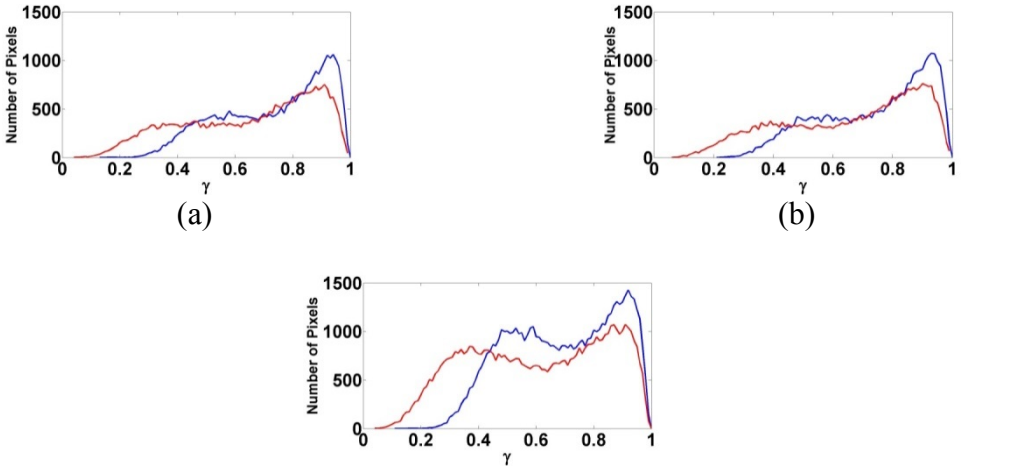


Fig. 4. Histogram of the γ_x and the $\gamma_{pol,x}$ for initial selected pixels related to (a) HH and Optimum channel, (b) VV and Optimum channel, (c) HH+VV and optimum channel. Blue and red line indicate the optimum and single-pol channel behaviour, respectively.

In homogeneous areas, the scattering properties of neighbouring pixels are expected to be spatially similar. Therefore, if the projection vectors and the optimum coefficients reflect the actual scattering properties, rather than taking values that just increase the coherence bias of each pixel, they will generally be spatially smooth.

As can be seen in Figure 5., the estimated coefficients are not randomly distributed and there is spatial consistency for the distribution of all coefficients, especially f_1 and f_4, which are real numbers and correspond to the two co-polar interferograms. The coefficient of the first co-polar interferogram, f_1, which enhances surface scattering behaviour, has large values in most of the areas. Moreover, a clear complementarity between f_1 and f_4 is observed since where f_1 is small, f_4 is large. f_2 and f_3 are complex coefficients for the two cross-polar interferograms and their maps are similar for amplitude and phase.

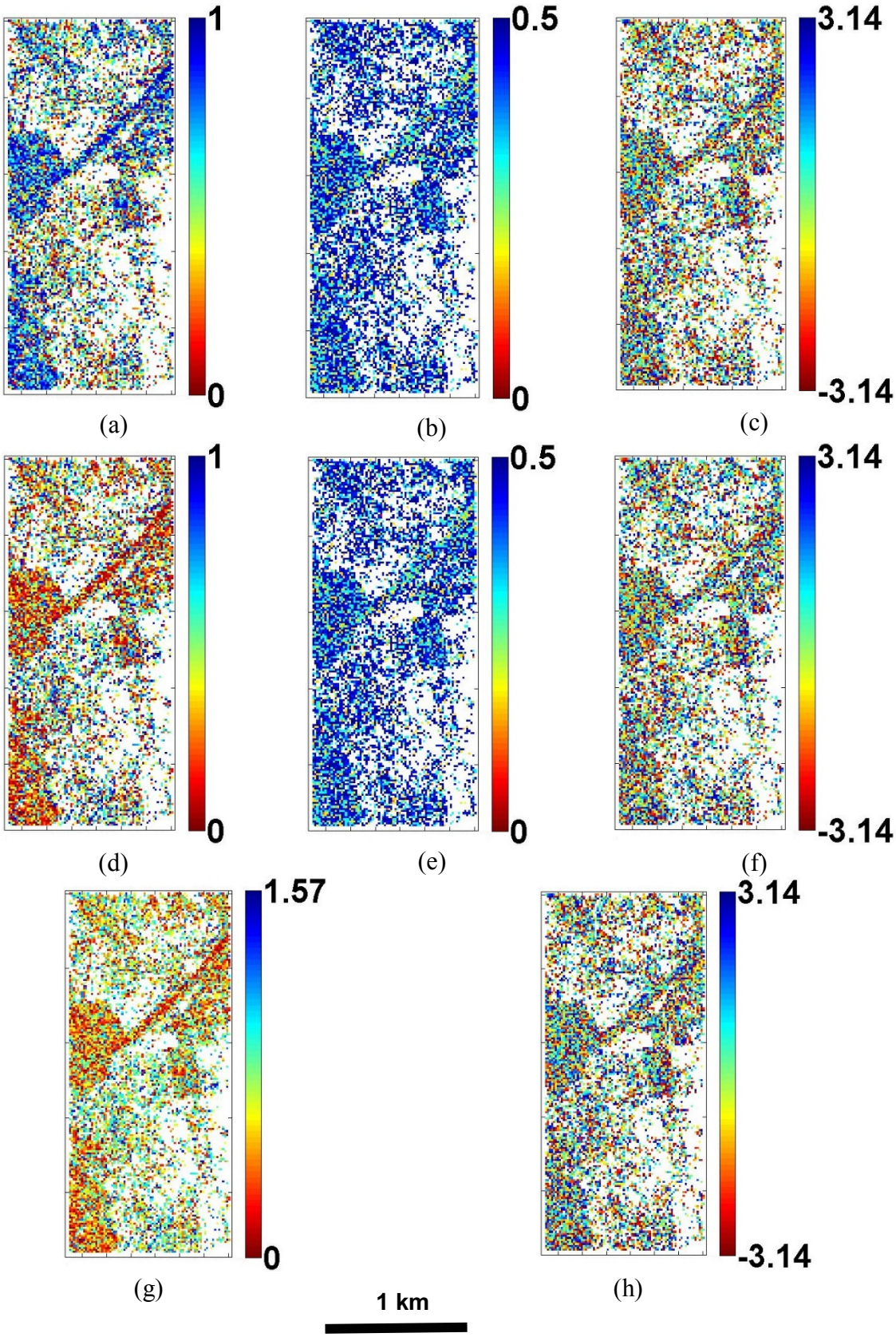


Fig. 5. Maps of optimum coefficients and parameters for an interferogram. (a) f_1, (b) amplitude of f_2, (c) phase of f_2 (d) f_4, (e) amplitude of f_3, (f) phase of f_3, (g) α, (h) ψ.

The number of final selected PS pixels over the case study using standard StaMPS for different channels (HH, VV, HH+VV) and PolStaMPS is presented in Table. 1. It is clear that the increase in the number of PS pixels using the HH+VV channel in standard StaMPS compared to the linear channels is trivial. However, using PolStaMPS the number increases by 48%, 80% and 82% with respect to HH+VV, VV and HH channel, respectively. There are some PS pixels which are not identified by StaMPS with linear channels, but they are selected by both PolStaMPS and StaMPS with HH+VV. In fact, approximately 40% of the additional PS pixels that are selected by PolStaMPS with respect to StaMPS with linear channels are also selected by StaMPS with HH+VV channel.

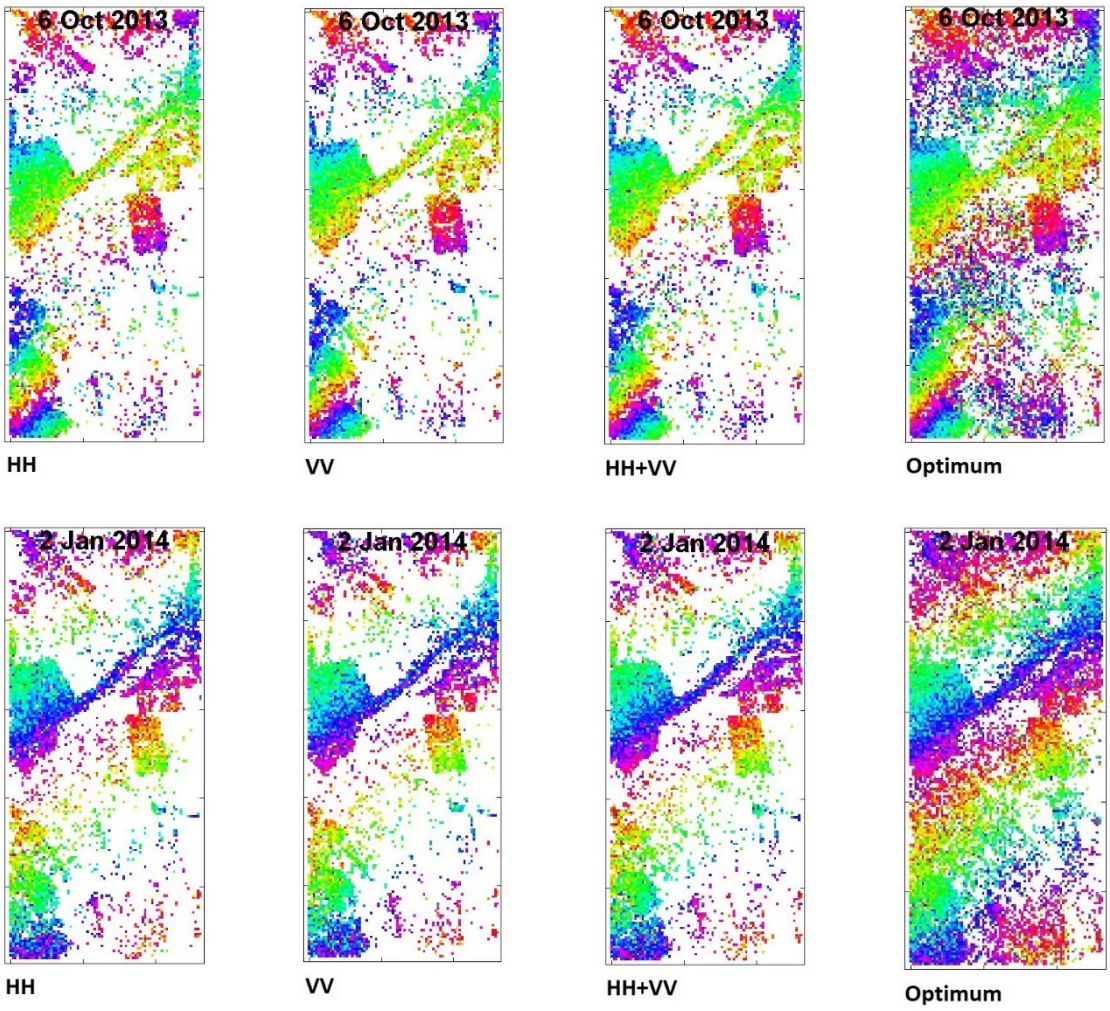
Table 1. Number of identified PS pixels

HH	VV	HH+VV	Optimum
26322	26694	32374	47997

Figure 6 shows the wrapped phase of selected pixels for optimum interferograms and HH, VV and HH+VV interferograms. As can be seen, the additional PS pixels in the optimum channel look clearly coherent. Furthermore, there are some common PS pixels whose phases are less noisy in the optimum interferogram. In order to assess the phase quality for the interferograms obtained by PolStaMPS in comparison to the original StaMPS, phase noise is estimated according to [9]. First, the PS pixels are connected to form a network using Delaunay triangulation. Then for each arc connecting two PS pixels, a weighted-average phase is calculated from the entire time series, and removed from the original phase of the arc, which is then low-pass filtered in time. The resulting phase, with the weighted-average phase added back in, provides an estimate for the smooth underlying signal. Phase noise is estimated by subtracting the smooth phase from the original phase of the arc. Finally, the phase noise of each PS pixel is obtained from the phase noise of its corresponding arcs. Figure 7 shows a comparison of histograms of phase noise standard deviation for commonly identified PS pixels in single-polar and optimum channels. The optimum channel shows a 7%, 16% and 17%

reduction in the number of PS pixels with standard deviation above 0.5 radians with respect to HH+VV, VV and HH channel. This confirms that, in addition to increasing PS density, the proposed algorithm is also successful in reducing the noise level of those PS pixels selected by standard StaMPS, although the reduction in the noise level is less pronounced than the increase in the number of selected PS pixels.

The resulting velocity maps of PolStaMPS and standard StaMPS are plotted in figure 8. The pattern of deformation rate is very similar, as expected, but the density of measurements is greater in the PolStaMPS case. The maximum velocity for this case study is -139.7 mm/year for the optimum channel.



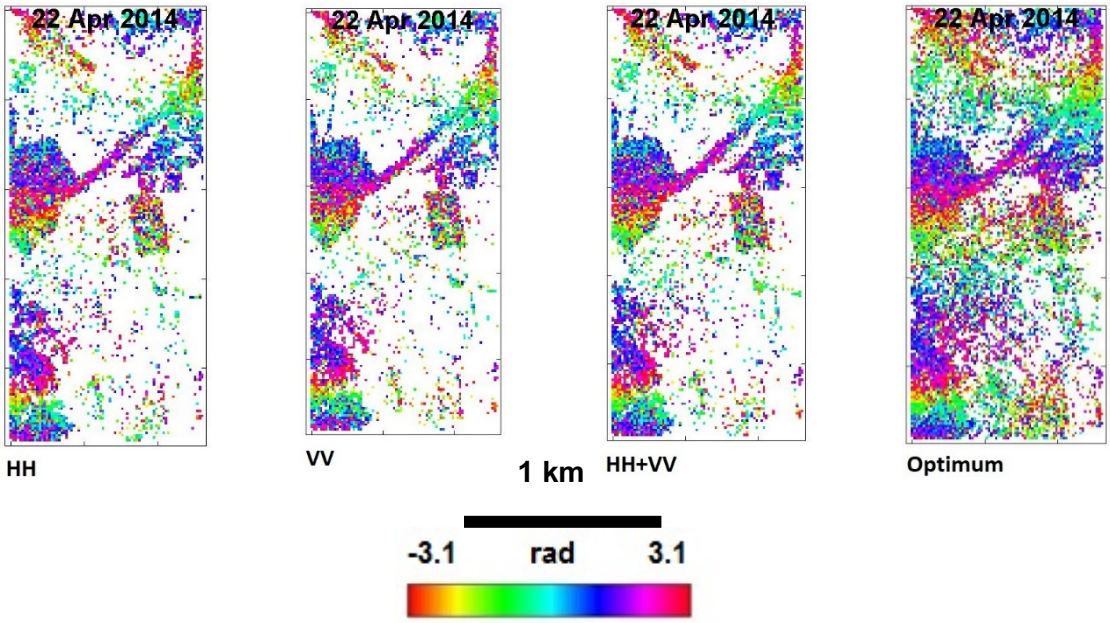


Fig. 6. A selection of wrapped interferograms formed from available data set acquired using HH, VV, HH+VV and Optimum channel over the case study. The master acquisitions date is 11 Dec 2013. Each colour fringe represents 1.55 cm of displacement in the LOS.

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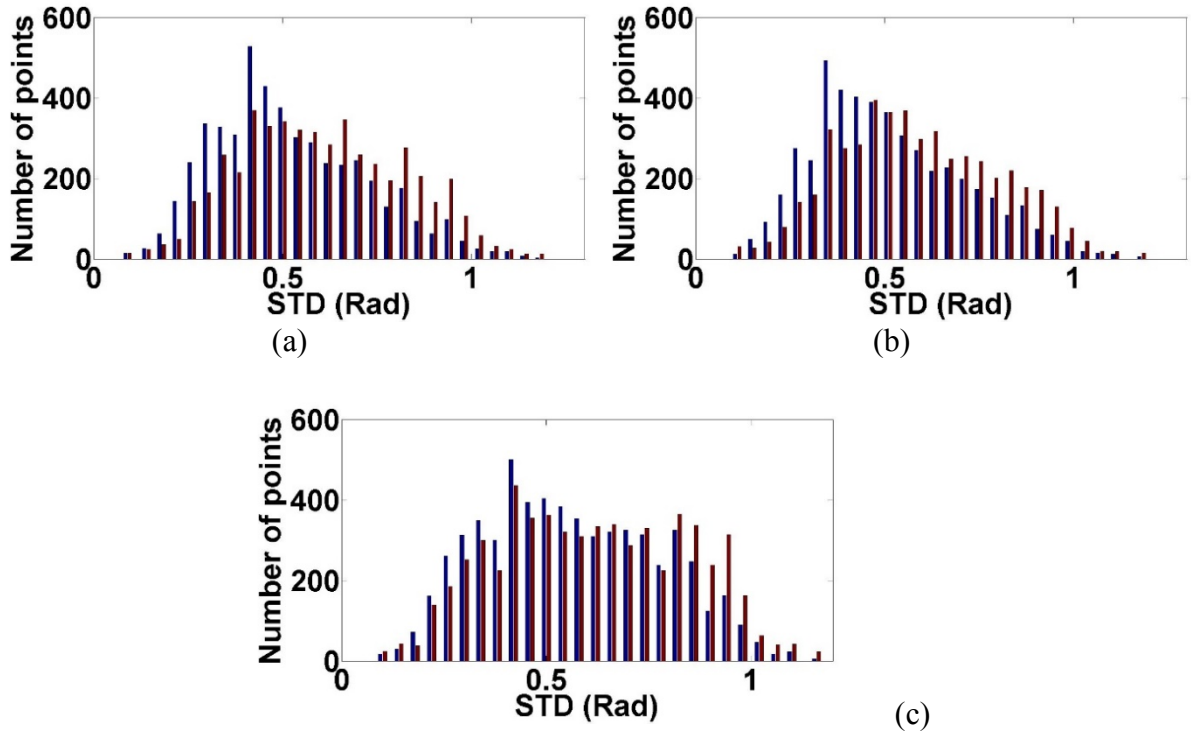

Fig. 7. Histogram of phase noise standard deviation for commonly identified PS pixels between optimum channel and (a) HH channel, (b) VV channel and (c) HH+VV channel. Blue and red bar indicate the optimum and single-polar channel behaviour, respectively.

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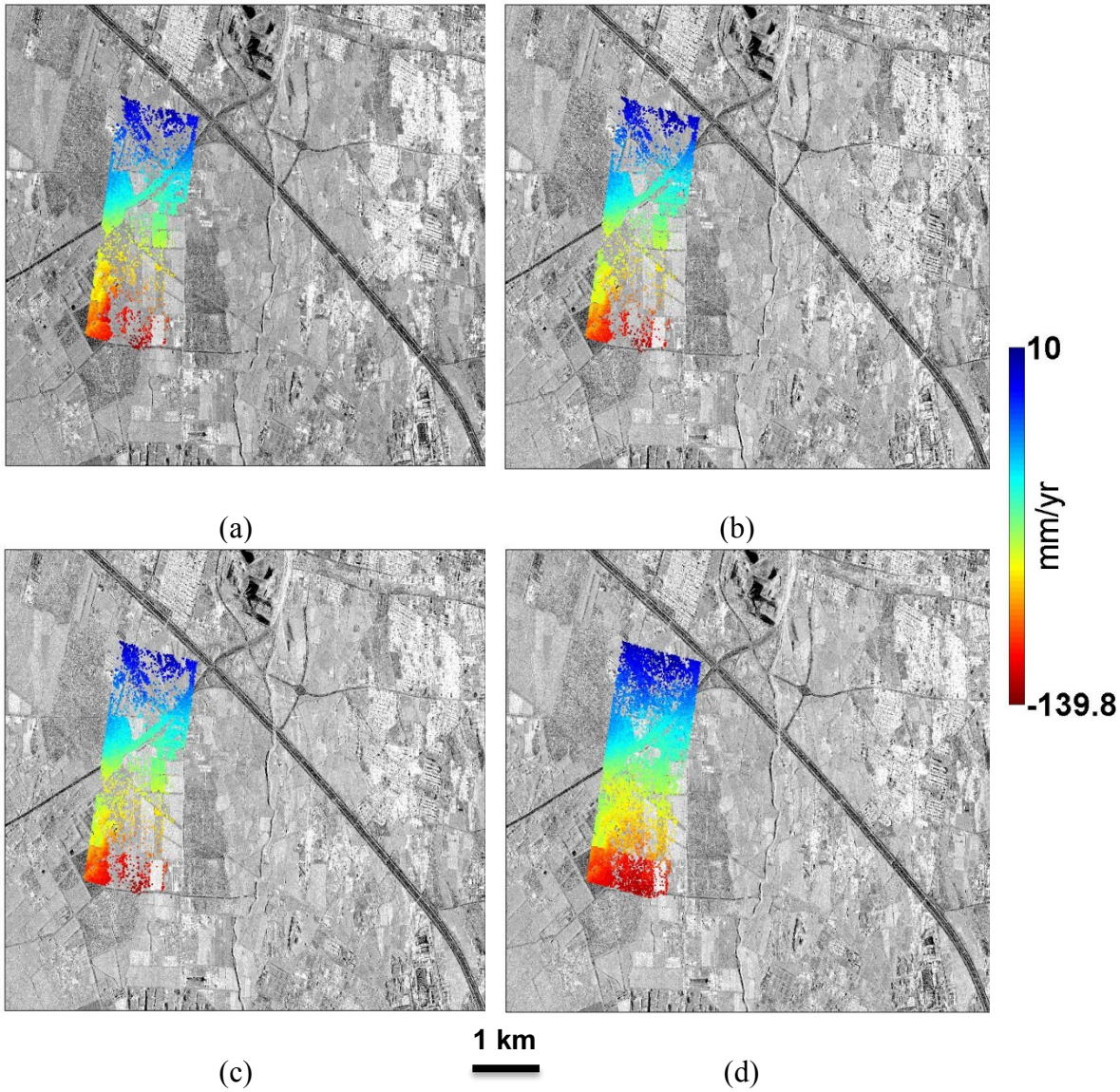

Fig. 8. Mean LOS velocities on the case study between 21 July 2013 and 22 April 2014 plotted on interferogram amplitude, (a) HH channel, (b) VV channel, (c) HH+VV channel, d) optimum channel.

370 The polarimetric PSI method leads to an increase in the number of selected PS pixels when
371 compared to standard PSI, although this comes with a computational cost. PolStaMPS is
372 inspired by ESPO and consequently finds the coefficients in the defined search spaces to
373 optimise the temporal coherence. This leads to an increase in the computation time of ~80 times
374 with respect to standard StaMPS. The computation time depends on the defined step in the
375 search spaces; larger steps decrease the computation time, although they could lead to
376 convergence on local optima instead of global ones. Optimising the temporal coherence using
377 other existing optimisation methods, e.g. Union, in which the optimum channel is selected from

a polarimetric channel with limited availability [21], may work with a lower computational cost, but the solutions are suboptimal. It should be mentioned that PolStaMPS can be applied over areas larger than the case study in this research, and the computation cost increases approximately linearly with the number of pixels of the scene.

7. CONCLUSIONS

In this study, we present a new polarimetric PSI approach that i) is applicable in areas lacking man-made structures and ii) retains the full spatial resolution of the input images. Using this technique we are able to identify natural targets that the standard PSI approach fails to select: the number of PS is improved by 48%, 80% and 82% with respect to the HH+VV, VV and HH channels, respectively. Moreover, the phase quality of the selected PS pixels is also improved. We have successfully applied this new algorithm to a rural part of the Tehran basin to monitor high-rate land subsidence and envisage that it can be used to estimate crustal deformation in most terrains. Future work should include a comparison of the results and performance of PolStaMPS with respect to other polarimetric PSI methods.

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