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1 **Mercury anomalies across the end Permian mass extinction in South China from**
2 **shallow and deep water depositional environments**

3

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17 **Abstract**

18 Life on Earth suffered its greatest bio-crisis since multicellular organisms rose 600
19 million years ago during the end-Permian mass extinction. Coincidence of the mass
20 extinction with flood basalt eruptions in Siberia is well established, but the exact
21 causal connection between the eruptions and extinction processes in South China is

22 uncertain due to their wide spatial separation and the absence of direct geochemical
23 evidence linking the two. The concentration and stable isotope analysis of mercury
24 provides a way to test these links as its concentration is thought to be tied to igneous
25 activity. Mercury/total organic carbon ratios from three Permian-Triassic boundary
26 sections with a well-resolved extinction record in South China show elevated values
27 (up to 900 ppb/wt. % relative to a background of <100 ppb/wt. %) that exactly
28 coincides with the end-Permian mass extinction horizon. This enrichment does not
29 show any correlation with redox and sedimentation rate variations during that time.
30 Hg isotope mass-independent fractionation ($\Delta^{199}\text{Hg}$), with sustained positive values,
31 indicate a predominant atmospheric-derived signature of volcanic Hg in deep-shelf
32 settings of the Daxiakou and Shangsi sections. In contrast, the nearshore environment
33 of the Meishan section displays a negative $\Delta^{199}\text{Hg}$ signature, interpreted to be related
34 to terrestrial Hg sources. Such temporal differences in $\Delta^{199}\text{Hg}$ values shed new light
35 on Hg geochemical behavior in marine settings, and also on kill mechanisms
36 associated with volcanism that were responsible for biotic mortality at the end of the
37 Permian.

38

39 **1. Introduction**

40 The end-Permian mass extinction (EPME; 252 million years ago) was the most
41 severe biotic crisis of the Phanerozoic. Most workers agree that intense volcanic
42 activity of the Siberian Traps Large Igneous Province (STLIP) was a driver of

43 environmental change (Wignall, 2001; Svensen et al., 2009; Sun et al., 2012; Black et
44 al., 2014; Clarkson et al., 2015; Burgess et al., 2017). The STLIP has an estimated
45 volume of up to $3\text{--}4 \times 10^6 \text{ km}^3$, which is larger than any other continental basalt
46 province, including the Emeishan traps ($\sim 1 \times 10^6 \text{ km}^3$), Central Atlantic Magmatic
47 Province ($\sim 2 \times 10^6 \text{ km}^3$), Karoo and Ferrar traps ($\sim 2.5 \times 10^6 \text{ km}^3$) and Deccan traps
48 ($2\text{--}4 \times 10^6 \text{ km}^3$) (Courtilot and Renne, 2003, references therein). Its original volume
49 is likely considerably larger as most is buried and inaccessible beneath the younger
50 sediments of the West Siberian Basin (Reichow et al., 2009; Saunders, 2016). Recent
51 high-precision U-Pb dating of STLIP basalts have shown that the onset of eruptions
52 began shortly before the start of the mass extinction crisis (Burgess and Bowring,
53 2015). More recently, Burgess et al. (2017) further categorized the dated volcanic
54 rocks from STLIP into lava- and sill-originated rocks, and found that major transfer
55 from lava to sill eruptions coincided with the main episode of biotic extinction.
56 Although there is a clear temporal correlation between the extinction event and
57 volcanic activity, the precise mechanism that drove the environmental change is
58 unresolved. In recent years, stable isotope systems have been shown to provide
59 important insights into the response of environmental systems to key climatic changes
60 associated with extinction events (Payne et al., 2010; Clarkson et al., 2015; Song et al.,
61 2017; Liu et al., 2017). For example, calcium and boron isotopes were successfully
62 applied to demonstrate ocean acidification triggered by STLIP across the EPME
63 (Payne et al., 2010; Clarkson et al., 2015). But for any isotopic system there are often
64 multiple inputs and outputs that can control the isotopic fractionation process in the

65 environment. In this paper we attempt to evaluate environmental controls on Hg
66 associated with the EPME. Hg is a key gas associated with volcanic activity and has
67 been linked to the end-Permian environmental crisis (Sanei et al., 2012; Grasby et al.;
68 2013, 2017)—a link we aim to assess here using mercury concentrations and isotopes
69 measured in Permian-Triassic boundary (PTB) sections in South China.

70 The ratio of mercury concentrations over total organic carbon (Hg/TOC) in
71 sedimentary sections has been shown to provide a proxy for voluminous volcanism
72 (Sanei et al., 2012; Grasby et al., 2013, 2017). Mercury is a highly toxic heavy metal
73 and has a sufficiently long atmospheric residence time (>1.5 yr) for global distribution
74 (Blum et al., 2014). Explosive volcanic events inject abundant Hg into the atmosphere
75 ensuring its global reach (Pyle and Mather, 2003). Most volcanic mercury is released
76 as gaseous Hg⁰ and removed from the atmosphere mainly through oxidation to form
77 Hg²⁺, which then accumulates in oceans and on land through rainfall or adsorption
78 onto organic matter ensuring a strong association between Hg and TOC in sediments
79 (Gehrke et al., 2009; Ruiz and Tomiyasu, 2015). Hg isotopes can undergo both large
80 mass-dependent fractionations (MDF) and mass-independent fractionations (MIF) in
81 nature (Blum et al., 2014), and thus are capable of tracing Hg sources and cycling
82 (Grasby et al., 2017). Hg-MDF ($\delta^{202}\text{Hg}$) can result from many pathways, including
83 physical, chemical and biological reactions, whereas Hg-MIF ($\Delta^{199}\text{Hg}$) is controlled
84 by more limited pathways (mostly photochemical) and is unlikely to be altered in the
85 post-depositional processes (Blum et al., 2014; Thibodeau et al., 2016; Thibodeau and
86 Bergquist, 2017). Hence, Hg-MIF ($\Delta^{199}\text{Hg}$) is generally a more conservative tracer of

87 volcanic signature (Thibodeau and Bergquist, 2017).

88 In recent years, Hg concentrations and isotopes have been used to explore the
89 relationship between large igneous provinces and contemporary mass extinctions
90 (Sanei et al., 2012; Grasby et al.; 2013, 2017; Percival et al., 2015, 2017; Sial et al.,
91 2016; Thibodeau et al., 2016; Gong et al., 2017). Anomalous Hg deposition was
92 observed at the EPME crisis in the Sverdrup Basin, Canadian High Arctic that
93 occupied a paleogeographic position near the STLIP (Fig. 1A) (Sanei et al., 2012;
94 Grasby et al., 2013). More recently, Grasby et al. (2017) documented the difference in
95 Hg isotopes near the EPME between Sverdrup Basin at a deep water setting and the
96 shallower water Meishan section in South China. They attributed the negative $\Delta^{199}\text{Hg}$
97 values at the EPME in Meishan to terrestrial sources, and suggested that deeper water
98 sections that are isolated from terrestrial input provide better records of the volcanic
99 signature. Therefore, whether or not the signature of Hg enrichments associated with
100 STLIP is recorded in the South China sections is unresolved. We further this work by
101 examining a series of sections in South China covering a range of water depths. We
102 provide new data for two deeper-water sections (200–500m), at Daxiakou and
103 Shangsi in South China, and integrate this with the data from the shallow water
104 Meishan section (Grasby et al., 2017). We measured Hg concentrations and Hg
105 isotopic compositions through the *Clarkina changxingensis* (*C. changxingensis*) to
106 *Isarcicella isarcica* (*I. isarcica*) conodont zones, to clarify the timing and intensity of
107 the eruption across the PTB and link its relationship with EPME in South China. The
108 varied sedimentary environments in the three sections enable an assessment of the

109 geochemical behaviors of Hg in different water depths and indicated that the effects of
110 the STLIP extended to the Chinese sections.

111

112 **2. Geological background**

113 The Meishan, Daxiakou, and Shangsi sections are separated by over 1000 km and lie
114 along the northern part of South China Craton. During late Permian time, the craton
115 was situated at low latitudes in the eastern Paleo-Tethys (Fig. 1A). The craton was
116 characterized by marine facies in its interior, and bounded by lands to the east and
117 west (Fig. 1B). In the central part of the craton, the roughly east-west trending
118 Yangtze carbonate platform is flanked to the north and south by deeper water basins
119 (Feng et al., 1996; Fig. 1B). The three studied sections lie on the flanks of the
120 northern basin and the EPME has been calibrated at the bottom of the *C. meishanensis*
121 conodont zone, with the PTB was placed at the bottom of the *Hindeodus parvus* (*H.*
122 *parvus*) conodont zone (Jiang et al., 2011; Zhao et al., 2013; Chen et al., 2015).

123 The Meishan section, which is located at Meishan Town, Changxing County,
124 Zhejiang Province, lay on the northeastern margin of the Yangtze Platform at a
125 shallow water depth of between 30–60 m based on sedimentary structures indicative
126 of fair-weather wave base and storm wave base (Yin et al., 2001; Chen et al., 2015;
127 Fig. 1B). This section is the Global Stratotype of Point and Section (GSSP) for the
128 PTB (Yin et al., 2001). Biostratigraphy, lithofacies, geochronology, and geochemistry
129 of the Meishan section have been extensively investigated since the 1980's, and have

130 been updated by Chen et al. (2015). Herein, the uppermost Permian succession (Beds
131 20–24) consists of bioclastic limestone of the Changxing Formation, whereas the
132 lowest Triassic strata consist of calcareous mudstone and marlstone of the Yinkeng
133 Formation (Beds 25–51) (Chen et al., 2015).

134 The Daxiakou section is situated near the town of Xiakou, Xingshan County,
135 Hubei province. This section was deposited on the northern margin of the South
136 China craton in a relatively deeper-water (~200–500 m) basin setting, as determined
137 by conodont biofacies and deeper-water foraminifera (Zhao et al., 2013; Zhang and
138 Gu, 2015; Fig. 1B). The uppermost Permian succession is dominated by black shale
139 with scattered interbeds of limestone and volcanic tuff of the Dalong Formation,
140 whereas the lowest Triassic sequence is composed of thin-bedded limestone and
141 calcareous mudstone interbeds of volcanic tuff of the Daye Formation. Zhao et al.
142 (2013) established the detailed conodont zonation at Daxiakou, which correlates well
143 with the Meishan section. Li et al. (2016) have also established high-resolution
144 astronomical cycles for the PTB beds in the same section.

145 The Shangsi section, one of the candidate sections for the GSSP of the PTB, is
146 located near Changjianggou village of Shangsi Town, Guangyuan City, Sichuan
147 Province. The Shangsi area lay on the northwestern margin of the South China craton
148 (Fig. 1B). The PTB succession is continuously exposed at this section and has been
149 well studied (Li et al., 1986; Wignall et al., 1995; Riccardi et al., 2007; Jiang et al.,
150 2011; Shen et al., 2011; Xiang et al., 2016). The latest Permian succession is mainly
151 composed of siliceous limestone interbedded with black shale of the Dalong

152 Formation, indicating that it represents a relatively deeper-water setting which is
153 supported by deep-water radiolarian assemblages (~200–500 m) (Jiang et al., 2011;
154 Xiao et al., 2017; Fig. 1B). However, the water-depth gradually started to shallow
155 from the latest Permian (Li et al., 1986). At the EPME, the lithology changed from
156 siliceous limestone to mudstone. The lowest Triassic strata mainly consist of interbeds
157 of marlstone and mudstone of the Feixianguan Formation.

158

159 **3. Methods**

160 3.1 Mercury concentrations

161 Mercury content was measured using a LECO AMA254 mercury analyzer at the
162 State Key Laboratory of Geological Processes and Mineral Resources (SKLGPMR),
163 China University of Geosciences (CUG-Wuhan). Prior to analysis, all samples were
164 freeze-dried to prevent decomposition of Hg. About 100 mg of mudstone or shale and
165 150–200 mg of limestone were analyzed. Data reliability was ensured by use of
166 international standard 502-685 (0.04 ± 0.008 ppm), which was analyzed after every
167 12 unknowns then followed by a repeat. The analytic precision is within 5 % and
168 reproducibility of sample concentrations was within 10 %.

169

170 3.2 Total organic carbon (TOC) measurement

171 TOC content measurement was conducted in the SKLGPMR (CUG, Wuhan) by
172 a vario macro cube elemental analyzer. About 10 g of powdered sample was put into

173 the 50 mL tube, then 50 % HCl was injected to dissolve carbonate minerals until there
174 was no further bubbling. After multiple centrifugal and lyophilization, the residue was
175 analyzed for total organic carbon (TOC). Data quality was assessed through multiple
176 analyses of standard sample DP-1 (65.44 ± 0.33 %). A standard sample and a repeat
177 were analyzed after every 12 unknowns, yielding an analytical accuracy of 2.5 % of
178 the reported values.

179

180 3.3 Hg isotopes

181 Hg isotopic compositions of samples from the three sections were analyzed at the
182 Institute of Geochemistry, Chinese Academy of Sciences, Guiyang, China. Detailed
183 analytical methods and processing procedure follow those described by Grasby et al.
184 (2017) and Gong et al. (2017). Approximately 0.5 g of ground sample and ~0.1 g of
185 standard reference materials (GSS-4, soil) were digested (95 °C for 90 min) using 5
186 mL of aqua regia (HCl/HNO₃ = 3, v/v). After digestion, the solution was centrifuged
187 (3000 rpm for 10 min) at room temperature and then decanted to obtain the
188 supernatant. Before conducting Hg isotope analysis, the digested sample solutions
189 were diluted to 2 and 1 ng mL⁻¹ using Milli-Q water based on total Hg values. All the
190 acid concentrations of the diluted solutions were < 20 %. Similarly, NIST SRM 3133
191 and 3177 Hg standard solutions with Hg concentrations of 2 and 1 ng mL⁻¹ were also
192 prepared.

193 Isotopic values were measured using a Neptune MC-ICP-MS with high
194 sensitivity X skimmer cone. NIST SRM 997 Tl was the internal standard used for

195 simultaneous instrumental mass bias correction of Hg. SnCl₂ (4 ng mL⁻¹) was used to
196 generate elemental Hg⁽⁰⁾ before being introduced into the plasma. The stability of the
197 instrument and measurement precision were monitored using NIST SRM 3133 and
198 3177 Hg standard solutions, respectively. Sample-standard bracketing was based on
199 NIST 3133. Digest recovery of the samples was monitored by MC-ICP-MS using
200 ²⁰²Hg signals. The sensitivity for ²⁰²Hg during Hg isotope analysis was ~1.2 V per ng
201 mL⁻¹ Hg.

202 Hg isotopic variations are reported in δ²⁰²Hg notation in units of per mil (‰)
203 referenced to the NIST-3133 Hg standard (once every 3 samples):

204

$$205 \delta^{202}\text{Hg} (\text{‰}) = [({}^{202}\text{Hg}/{}^{198}\text{Hg}_{\text{sample}})/({}^{202}\text{Hg}/{}^{198}\text{Hg}_{\text{standard}}) - 1] \times 1000 \quad (1)$$

206

207 Mass independent fractionation (MIF) of Hg isotopes is expressed in Δ notation
208 (Δ^{xxx}Hg), which describes the difference between the measured δ^{xxx}Hg and the
209 theoretically predicted δ^{xxx}Hg value, using the following equations:

210

$$211 \Delta^{199}\text{Hg} \approx \delta^{199}\text{Hg} - (\delta^{202}\text{Hg} \times 0.2520) \quad (2)$$

$$212 \Delta^{200}\text{Hg} \approx \delta^{200}\text{Hg} - (\delta^{202}\text{Hg} \times 0.5024) \quad (3)$$

$$213 \Delta^{201}\text{Hg} \approx \delta^{201}\text{Hg} - (\delta^{202}\text{Hg} \times 0.7520) \quad (4)$$

214

215 Replicate analysis of the NIST 3177 Hg intra lab isotope reference standard (n = 4
216 analytical sessions) were as follows: δ²⁰²Hg = -0.42 ± 0.06‰ (2 sd); Δ¹⁹⁹Hg = -0.04 ±

217 0.02‰ (2 sd); $\Delta^{200}\text{Hg} = 0.01 \pm 0.04\text{‰}$ (2 sd); $\Delta^{201}\text{Hg} = 0.01 \pm 0.04\text{‰}$ (2 sd).

218 The complete data set of Hg, TOC, Hg/TOC and Hg isotopes values can be
219 obtained in Online Supplementary Tables S1, S2, S3.

220

221 **4. Results**

222 4.1 TOC contents

223 In the Meishan section, TOC concentrations vary from 0.06 to 1.56 wt. %, and
224 have relatively higher values (~ 1 wt. %) at the bottom of Bed 22 and Beds 24–26,
225 44–49 (Fig. 2). Through the Daxiakou section, TOC concentrations range from 0.04 to
226 3.99 wt. %. The latest Permian rocks of Dalong Formation show very high TOC
227 concentrations (~ 4.0 wt. %; Fig. 3). While the earliest Triassic Daye Formation
228 display low TOC concentrations, almost all less than 1 wt. %. In the Shangsi section,
229 the TOC concentrations vary from 0.03 to 2.82 wt. %. The higher TOC values (~ 1
230 wt. %) occurs the bottom of Bed 22 of the Dalong Formation and Beds 30–33 in Daye
231 Formation (Fig. 4). In other horizons, TOC has lower values (mostly less than 0.4
232 wt. %).

233

234 4.2 Hg and Hg/TOC variations

235 The Meishan, Daxiakou and Shangsi sections all record high Hg concentrations
236 and a short-lived positive excursion in Hg/TOC values coincident with the EPME. Hg
237 concentrations are directly tied to TOC contents (Grasby et al., 2013), and this is

238 demonstrated by the covariant relationship between TOC and Hg concentrations of
239 the studied sections where Daxiakou and Shangsi have the highest Hg values in the
240 late Permian along with anomalously high TOC contents (Figs. 2, 3, 4).

241 In the Meishan section, the highest Hg values (up to 176 ppb) occur at the top of
242 Bed 24 and in Beds 25–26. High values also occur at the bottom of Bed 22 and Beds
243 44–49. TOC values are high in these layers, indicating the Hg concentrations are
244 probable driven by lithologically controlled changes in organic matter content.

245 Compared to Hg concentrations, positive excursions in Hg/TOC are only observed in
246 the uppermost parts of Bed 24 and Bed 25, corresponding to the EPME (Fig. 2). In the
247 Daxiakou section, the uppermost Dalong Formation (Bed 9), in the *C. meishanensis*
248 conodont zone, corresponds to Bed 25 in the Meishan section and records high values
249 in Hg concentrations (~300 ppb) and a positive excursion in Hg/TOC ratios (Fig. 3).

250 High Hg concentrations are also present through the upper Permian of the Dalong
251 Formation, although in these layers there is no Hg/TOC excursion (Fig. 3). This may
252 imply that the high Hg contents are due to high levels of TOC in the shale. Similar to
253 TOC, high Hg concentrations also occur at the bottom of Bed 22 and Beds 30–33 in
254 the Shangsi section (Fig. 4), except near the EPME where high Hg values (~100 ppb)
255 were recorded, whereas TOC contents are relatively low. A Hg/TOC excursion was
256 only recorded near the EPME.

257

258 4.3 Hg isotope compositions of the study sections

259 The Hg isotope values of the Meishan, Daxiakou and Shangsi sections are

260 shown in figures 2, 3, and 4, respectively, and analytical results are also provided in
261 Online Supplementary Tables S1, S2, and S3.

262 MDF ($\delta^{202}\text{Hg}$) at Meishan ranges from -2.3 ‰ to -0.58 ‰, and MIF ($\Delta^{199}\text{Hg}$)
263 varies from -0.12 ‰ to 2.0 ‰. They both show negative shifts at the EPME (to -1.74 ‰
264 in $\delta^{202}\text{Hg}$ and -0.12 ‰ in $\Delta^{199}\text{Hg}$; Fig. 2). MDF ($\delta^{202}\text{Hg}$) also shows similar negative
265 values above and below the EPME. MDF ($\delta^{202}\text{Hg}$) values at Daxiakou are negative
266 and vary from -1.78 ‰ to -0.49 ‰. The values rise below the EPME (to -0.49 ‰) and
267 decline above it (Fig. 3). MIF ($\Delta^{199}\text{Hg}$) values vary through the section, from -0.02 ‰
268 to 0.27 ‰. For uppermost Permian samples, $\Delta^{199}\text{Hg}$ values are relatively high and
269 have an average of 0.18 ‰. The high values continue through the EPME and drop
270 below 1.0 ‰ in the earliest Triassic. Overall, $\Delta^{199}\text{Hg}$ values at Daxiakou display
271 positive values with no significant change across the EPME (Fig. 3). At the Shangsi
272 section, MDF ($\delta^{202}\text{Hg}$) and MIF ($\Delta^{199}\text{Hg}$) values range from -2.26 ‰ to -0.58 ‰, and
273 -0.01 ‰ to 0.21 ‰, respectively. The $\Delta^{199}\text{Hg}$ values are similar to those at Daxiakou,
274 with high values in uppermost Permian samples and lower values in the earliest
275 Triassic, and also show sustained positive values across the EPME (Fig. 4). The main
276 difference between these two sections is that greater MDF ($\delta^{202}\text{Hg}$) and MIF ($\Delta^{199}\text{Hg}$)
277 levels are reached at the EPME in Daxiakou. The maximum $\delta^{202}\text{Hg}$ and $\Delta^{199}\text{Hg}$ at the
278 EPME in the Shangsi section are -0.75 ‰ and 0.11 ‰, respectively, whereas the
279 maximum values at Daxiakou are -0.49 ‰ and 0.15 ‰.

280

281 5. Discussion

282 5.1 Cause of the Hg and Hg/TOC anomalies

283 High Hg concentrations and elevated Hg/TOC values are all recorded across the
284 EPME in the studies of the Meishan, Daxiakou and Shangsi sections. Background Hg
285 concentrations are directly tied to marine organic matter (OM) deposition rates. For
286 example, high Hg values are observed at the bottom of Bed 22, and Beds 44–49 in
287 Meishan, through the upper Permian Dalong Formation at Daxiakou, and at the
288 bottom of Bed 22 and Beds 30–33 at Shangsi, and all are associated with high TOC
289 contents in shale or siliceous limestone. A constant and low Hg/TOC ratio is a
290 common feature of background marine conditions (Grasby et al., 2013, 2017). Our
291 results also show a pronounced increase in Hg/TOC values coincident with the
292 extinction interval (Figs. 2–5).

293 The role of lithology, sedimentation rate, and marine anoxia on Hg/TOC values
294 must all be considered to explain these shifts, because all are known to vary in the
295 studied sections. Lithologic variations are pronounced near the EPME and include
296 bioclastic limestone, mudstone, shale and volcanic ash beds, but there is no
297 relationship to the Hg/TOC changes. Within the *C. meishanensis* conodont zone at
298 Daxiakou for example, constant high Hg/TOC values were recorded in different
299 lithologic samples, including volcanic ash, shale, mudstone and limestone. In addition
300 to lithology, if the atmospheric mercury level is constant, a reduced sedimentation rate
301 may result in elevated Hg concentrations (Percival et al, 2015). This affect can be
302 accounted for by comparison with TOC concentrations, as these would also covary

303 with sedimentation rates (e.g. dilution effects). Using a recently-devised
304 cyclostratigraphic scheme (Li et al., 2016), sedimentation rates in the Meishan and
305 Daxiakou sections do indeed vary substantially (Fig. 2, 3). This reveals though that
306 there is no clear correlation between Hg/TOC values and sedimentation rate. Marine
307 anoxia was also widespread during the PTB transition, and its intensity varied
308 considerably. The PTB transition saw the temporary establishment of oxic/weakly
309 dysoxic conditions in Meishan (Chen et al., 2015) whereas the deeper-water Shangsi
310 and Daxiakou sections saw more oxygen-restricted conditions (Xiang et al., 2016;
311 Shen et al., 2016). The Hg/TOC fluctuations are not influenced by these oxygenation
312 trends, the highest values only occur at the EPME (Beds 27) in Shangsi (Fig. 4),
313 indicating that redox conditions have no significant control.

314 The Hg isotopic compositions can help to trace the origin and the pathways of
315 the observed Hg and Hg/TOC excursions (Thibodeau et al., 2016; Grasby et al., 2017;
316 Gong et al., 2017). Volcanic Hg has $\delta^{202}\text{Hg}$ values between -2.0 ‰ and 0 ‰, and has
317 insignificant MIF ($\Delta^{199}\text{Hg} \approx 0$ ‰) (Zambardi et al., 2009; Yin et al., 2016; Grasby et
318 al., 2017; Thibodeau and Bergquist, 2017). Once emitted to the environment, MDF
319 ($\delta^{202}\text{Hg}$) can result in indistinguishable signatures from many pathways, including
320 hydrothermal reactions, sorption/deposition, organic matter burial and diagenesis
321 (Blum et al., 2014; Thibodeau et al., 2016; Grasby et al., 2017). In contrast, MIF
322 ($\Delta^{199}\text{Hg}$) mostly occurs through photochemical reduction, such as in aqueous droplets
323 and surface waters, resulting in positive $\Delta^{199}\text{Hg}$ values (Bergquist and Blum, 2007;
324 Blum et al., 2014). Aquatic environments can receive Hg through atmospheric Hg^{2+}

325 deposition and direct uptake of $\text{Hg}^0_{(g)}$ by terrestrial sources (Thibodeau and Bergquist,
326 2017). Thus, marine sediments dominated by atmospheric Hg^{2+} deposition tend to
327 have positive $\Delta^{199}\text{Hg}$ because of the photo reduction in the surface environments
328 (Grasby et al., 2017; Thibodeau and Bergquist, 2017). However, sediments that
329 received $\text{Hg}^0_{(g)}$ through terrestrial runoff tend to have greater negative $\delta^{202}\text{Hg}$ and
330 $\Delta^{199}\text{Hg}$ than atmospheric $\text{Hg}^0_{(g)}$, and terrestrial resources acquire negative $\delta^{202}\text{Hg}$ and
331 $\Delta^{199}\text{Hg}$ values when plants and soils sequester $\text{Hg}^0_{(g)}$ (Thibodeau and Bergquist,
332 2017). In our study, the two relatively deeper water sections, Daxiakou and Shangsi,
333 both recorded higher Hg/TOC values and sustained positive $\Delta^{199}\text{Hg}$ values at the
334 EPME (Fig. 3, 4). This finding is consistent with the data of coeval Buchanan lake
335 sediments (deep water facies) that show positive $\Delta^{199}\text{Hg}$ values with no significant
336 change across the EPME despite a large Hg spike (Grasby et al., 2017). The
337 excursions in Hg/TOC indicate that increased Hg flux was imported to the oceans at
338 that time. Meanwhile, the positive $\Delta^{199}\text{Hg}$ values indicate sediments in these deep
339 water sections received Hg primarily through atmospheric Hg deposition, instead of
340 terrestrial runoff. Thus, the observed Hg peaks are due to enhanced flux of Hg from
341 atmosphere consistent with enhanced atmospheric content. The best way to explain
342 this is that enhanced volcanic activity injected massive quantities of Hg into the
343 atmosphere and increased the global atmospheric Hg level.

344

345 5.2 Variations of Hg/TOC values and Hg isotopic compositions in different water

346 depths

347 The Meishan, Daxiakou and Shangsi sections all showed short-duration, albeit
348 variable, positive excursions, in Hg/TOC across the EPME (Figs. 2–5). The maximum
349 Hg/TOC at Meishan, Shangsi and Daxiakou are 926 ppb/wt. %, 553 ppb/wt. %, and
350 254 ppb/wt. %, respectively. The prominent negative shifts of $\delta^{202}\text{Hg}$ and $\Delta^{199}\text{Hg}$
351 values at the EPME horizon in the Meishan section (Fig. 5), which is a clastic
352 sediment-starved carbonate platform setting, indicate that enhanced Hg influx from
353 terrestrial sources could be imported to oceans and overwhelmed the volcanic
354 signature from the Siberian Traps (Grasby et al., 2017). It is possible that most of Hg
355 coming from the terrestrial source might have also been derived from the STLIP and
356 this is indicated by the similar Hg/TOC baselines of the studied sections. The larger
357 spike at Meishan suggests a focusing of the volcanic Hg via the terrestrial ecosystems
358 to marine deposition above that of direct atmospheric inputs. Similarly, highest
359 Hg/TOC excursions were also noted in terrestrial and coastal sections by Percival et al.
360 (2015, 2017) who have reported Hg anomalies in several different sedimentary
361 records that span Pliensbachian-Toarcian transition and end-Triassic mass extinction.
362 For example, the Pliensbachian-Toarcian transition at Bornholm, Danish Basin, which
363 preserves a fluvial to marginal marine setting, shows greatest its Hg/TOC excursion
364 during the Toarcian Oceanic anoxic events. The end-Triassic record from a
365 fluvial-deltaic section of Astartekløft, Greenland, also shows its greatest Hg/TOC
366 excursion during the end-Triassic mass extinction interval. These suggest that the
367 different values in Hg/TOC at the EPME recorded in our studied sections are not due
368 to slightly different sampling resolution at the sites (e.g., lower resolution at

369 Daxiakou), rather they show that coastal/terrestrial sediments include additional
370 terrestrial Hg relative to deeper sections.

371 The Daxiakou and Shangsi sections both represent deeper water depositional
372 environments (~200–500 m) (Jiang et al., 2011; Zhao et al., 2013; Zhang and Gu,
373 2015; Xiao et al., 2017). At the EPME, the Daxiakou section is inferred to have been
374 at a greater water depth than the Shangsi section, based on its overall finer grain-size
375 of the sediment. The $\Delta^{199}\text{Hg}$ values of Daxiakou at the EPME have greater positive
376 values than those from Shangsi (Fig. 5), consistent with the Daxiakou section being
377 least affected by terrestrial Hg. The positive $\Delta^{199}\text{Hg}$ values at the EPME in the
378 Daxiakou and Shangsi sections are consistent with those observed in modern
379 sediments that receive insignificant terrestrial runoff (Thibodeau and Bergquist, 2017,
380 references therein), as well as those obtained from deeper water section across the
381 Cretaceous-Paleogene transition (Bajada del Jagüel section, Sial et al., 2016).
382 Conversely, no significant Hg-MIF ($\Delta^{199}\text{Hg} \approx 0 \text{ ‰}$) signature with Hg enrichments
383 that was associated with an overloading of volcanic Hg from the Central Atlantic
384 Magmatic Province, was observed during end-Triassic mass extinction in a shallow
385 shelf sediments at Muller Canyon, Nevada (USA), indicating that the sediments
386 received volcanic Hg through both atmosphere Hg^{2+} deposition and terrestrial runoff,
387 and the Hg-MIF signatures were overwhelmed (Thibodeau et al., 2016; Thibodeau
388 and Bergquist, 2017).

389 In summary, sediments receive variable proportions of their Hg from direct
390 atmospheric and terrestrial inputs. It is likely that Hg, including Hg/TOC and Hg

391 isotopes in shallow water sections are more sensitive to terrestrial signals, such as the
392 portion of Hg entering from the atmosphere to land systems and ancient Hg
393 originating from eroding rocks. In contrast, deep water sections better record the
394 timing and strength of eruptions in the geological record because of their more precise
395 sensitivity to Hg atmospheric perturbations unmediated by the terrestrial realm
396 (Grasby et al., 2017; Thibodeau and Bergquist, 2017).

397

398 5.3 Causal links between biotic extinction and STLIP

399 The STLIP is recognized as the largest continental basalt province (Dobretsov et
400 al., 2008). The preserved area of the province is of the order of 2.5 million km² and a
401 working estimate for the total volume of eruptives and shallow-level intrusions is 3-4
402 million km³ (Fedorenko et al., 2000). The coincidence of the STLIP with the EPME
403 (Wignall, 2001; Courtillot and Renne, 2003; Burgess and Bowing, 2015), suggests a
404 link between the two events. Release of volcanic CO₂ and thermogenic methane
405 through magma emplacement into the West Siberian Tunguska basin, which is
406 composed of a thick succession of evaporite, clastic, carbonate, and
407 hydrocarbon-bearing rocks, is thought to have triggered strong climatic warming and
408 ocean acidification (Svensen et al., 2009; Sun et al., 2012; Black et al., 2014;
409 Clarkson et al., 2015), as evidenced 3 to 4 ‰ negative shift in C-isotopic records
410 coincident with the EPME horizon (Fig. 5).

411 Siberian Traps magmatic activity is divided into three stages based on
412 high-precision U-Pb dates: early extrusion (252.24–251.907 Ma; ~C.

413 changxingenss–C. yini conodont zones), characterized by initial pyroclastic eruptions
414 followed by lava effusion, then medium-term sill intrusion (251.907–251.483 Ma; ~ C.
415 meishanensis–lower I. isarcica conodont zones), and terminal stage (lasting until at
416 least 251.354 Ma; above the lower I. isarcica conodont zone), marked by initial lava
417 extrusion after a ~420 ka hiatus and then sill intrusion (Burgess and Bowring, 2015;
418 Burgess et al., 2017). This implies that STLIP involved three or more original magma
419 pulses. Burgess et al. (2017) showed that the initial pulse of Siberian Traps sills
420 following the initial extrusive pyroclastics and lava, coincided with the onset of
421 EPME and an abrupt negative C isotope shift. Hg enrichment also occurs at that
422 horizon (this study; Sanei et al., 2012; Grasby et al., 2013; 2017). This suggests that a
423 large volume of Hg was released into the atmosphere at the time of initial sill
424 intrusion. Most of the sills are exposed at the periphery of exposures of lavas and
425 volcanoclastic rocks, and are up to 1200 m thick (Svensen et al., 2009; Burgess and
426 Bowring, 2015). Compared to the pyroclastic lava, the sills take up a much larger area
427 and volume (Vasiliev et al., 2000; Burgess and Bowring, 2015). Total sill volume in
428 the STLIP is difficult to constrain but has been conservatively estimated in excess of 2
429 $\times 10^6$ km³ (Vasiliev et al., 2000), which maybe exceed 50 % of the STLIP's total
430 volume. Such massive sills provided extensive horizontal and vertical magmatic
431 channels and diatremes, by which volatiles are delivered to the atmosphere (Svensen
432 et al., 2009; Burgess et al., 2017). As a heat sensitive element, the major proportion of
433 Hg in sedimentary rocks is released if the temperature exceeds 340 °C (Sial et al.,
434 2013). Therefore, during the early intrusion stage, massive gas Hg⁰ would be

435 produced by interaction between the magma and country rocks and easily released to
436 the atmosphere through magmatic vents (Burgess et al., 2017). The Hg content of
437 mantle material is 0.01 $\mu\text{g/g}$ (McDonough, 2003). Using a density of basic rock of
438 $2.8 \times 10^3 \text{ kg/m}^3$, total Hg released from $2 \times 10^6 \text{ km}^3$ sills would be $5.6 \times 10^9 \text{ Mg}$. A
439 similar estimated value of $3.8 \times 10^9 \text{ Mg}$ was given by Sanei et al. (2012). In addition,
440 heating of sediments over the large area encompassed by the sill complex could
441 provide additional Hg release, the estimated quantity of which is about $1.75 \times 10^5 \text{ Mg}$
442 (Sanei et al., 2012). If most of the $2 \times 10^6 \text{ km}^3$ sills intruded during 251.907–251.483
443 Ma, then 13000 Mg yr^{-1} volcanic Hg would be released to the atmosphere during that
444 time, which is ~ 5.6 times current anthropogenic emissions of $\sim 2320 \text{ Mg yr}^{-1}$ and is
445 ~ 144.4 times modern volcanic emissions of $\sim 90 \text{ Mg yr}^{-1}$ (Pirrone et al., 2010; Sial et
446 al., 2013). The general elevation of Hg/TOC values in the C. meishanensis–lower I.
447 isarcica conodont zones, marked by green on figure 5, corresponds with the high rate
448 of Hg emission during the Siberian sill emplacement. Furthermore, the highest values
449 at the EPME correspond to the strongest release. Depending on the high rate of
450 emission, the enhanced Hg release at the EPME horizon may have temporarily
451 exceeded the capacity of organic matter to fix Hg resulting in its buildup in ocean
452 water to toxic levels (Sanei et al., 2015).

453 Other than a direct link by poisoning, most extinction scenarios invoke a
454 volcanism-extinction nexus caused by the subsequent effects of CO_2 release and rapid
455 global warming (Sun et al., 2012), which include ocean anoxia, acidification, and the
456 direct consequences of warming. Such effects are longer term than direct Hg

457 poisoning which, if significant, is likely to have occurred on a timescale of decades to
458 centuries because of the short residence time of this element in the ocean-atmosphere
459 system. Carbon dioxide-driven global warming is a longer term effect (centuries,
460 millennia) because of the longer residence time in the Earth surface system (Wignall,
461 2001). Nonetheless, our data show a tight coincidence between the onset of extinction
462 in South China and Hg enrichment. Any lead in time between the onset of STLIP
463 volcanism and the environmental consequences was below the temporal resolution of
464 our study. Phrased another way, the devastation caused by the onset of Siberian Trap
465 sill emplacement was geologically instantaneous.

466 **6. Conclusions**

467 Three sections across the end-Permian mass extinction at Meishan, Daxiakou,
468 and Shangsi in South China all record high Hg concentrations and a short-lived
469 positive excursion in Hg/TOC values. Hg isotopes with positive values ($\Delta^{199}\text{Hg}$) and
470 limited variability in the Daxiakou and Shangsi sections indicate a predominant
471 atmospheric-derived Hg signature from STLIP, whereas in the Meishan section,
472 $\delta^{202}\text{Hg}$ and $\Delta^{199}\text{Hg}$ both display negative shifts, indicating a significant proportion of
473 terrestrial sources (Grasby et al., 2017). Our data indicate that Hg, including Hg/TOC
474 and Hg isotopes, show different geochemical behavior at different water depths. For
475 Meishan, a clastic sediment-starved carbonate platform setting, enhanced Hg influx
476 from terrestrial sources could be imported to the marine setting and overwhelmed the
477 volcanic signature from Siberian Traps. At the deeper water marine sections, at

478 Daxiakou and Shangsi, a lower terrestrial signature was recorded. Although more
479 studies are undoubtedly needed on the geochemical behavior of Hg in different
480 sedimentary facies, our finding suggests that deep water marine facies can more
481 directly record the atmospheric-derived Hg signature from volcanism than near shore
482 shallow marine environments where signals are mediated to a greater extent by
483 terrestrial ecosystems via continental run off. Furthermore, our study demonstrates
484 that the effects of the STLIP extended to the Chinese sections, and clarifies the
485 relationship between Siberian eruptions and the EPME in South China.

486

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501 **References**

502 Bergquist, B.A., Blum, J.D., 2007. Mass-dependent and -independent fractionation of
503 Hg isotopes by photoreduction in aquatic systems. *Science* 318, 417–420.

504 Black, B.A., Lamarque, A.F., Shields, C.A., Elkins-Tanton, L.T., Kiehl, J.T., 2014.

505 Acid rain and ozone depletion from pulsed Siberian Traps magmatism. *Geology*
506 42, 67–70.

507 Blum, J.D., Sherman, L.S., Johnson, M.W., 2014. Mercury isotopes in earth and
508 environmental sciences. *Annu. Rev. Earth Planet. Sci. Lett.* 42, 249–269.

509 Burgess, S.D., Bowring, S.A., Shen, S.Z., 2014. High-precision timeline for Earth's
510 most severe extinction. *Proc. Natl. Acad. Sci. U. S. A.* 111, 3316–3321.

511 Burgess, S.D., and Bowring, S.A., 2015. High-precision geochronology confirms
512 voluminous magmatism before, during, and after Earth's most severe extinction.
513 *Sci. Adv.* 1: 1500470e.

514 Burgess, S.D., Muirhead, J.D., Bowring, S.A., 2017. Initial pulse of Siberian Traps
515 sills as the trigger of the end-Permian mass extinction. *Nat. Commun.* 8, 164.

516 Chen, Z.Q., Yang, H., Luo, M., Benton, M.J., Kaiho, K., Zhao, L.S., Huang, Y.G.,
517 Zhang, K.X., Fang, Y.H., Jiang, H.S., Qiu, H., Li, Y., Tu, C.Y., Shi, L., Zhang, L.,

518 Feng, X.Q., Chen, L., 2015. Complete biotic and sedimentary records of the
519 Permian–Triassic transition from Meishan section, South China: ecologically
520 assessing mass extinction and its aftermath. *Earth-Sci. Rev.* 149, 67–107.

521 Clarkson, M.O., Kasemann, S.A., Wood, R.A., Lenton, T.M., Daines, S.J., Richoz, S.,

522 OhnemueLLer, F., Meixner, A., Poulton, S.W., Tipper, E.T., 2015. Ocean
523 acidification and the Permo-Triassic mass extinction. *Science* 348, 229–232.

524 Courtillot, V.E., Renne, P.R., 2003. On the ages of flood basalt events. *C. R.
525 Geoscience* 335, 113–140.

526 Dobretsov, N.L., Kirdyashkin, A.A., Kirdyashkin, A.G., Vernikovskiy, V.A., Gladkov,
527 I.N., 2008. Modelling of thermochemical plumes and implications for the origin
528 of the Siberian traps. *Lithos* 100, 66–92.

529 Feng, Z., Yang, Y., Jin, Z., He, Y., Wu, S., Xin, W., Bao, Z., Tan, J., 1996.
530 Lithofacies paleogeography of the Permian of South China. *Acta
531 Sedimentologica Sinica* 14, 1–11.

532 Fedorenko, V., Czamanske, G., Zen'ko, T., Budahn, J., Siems, D., 2000. Field and
533 geochemical studies of the melilite-bearing Arydzhangsky Suite, and an overall
534 perspective on the Siberian alkaline-ultramafic flood volcanic rocks. *Int. Geol.
535 Rev.* 42, 769–804.

536 Grasby, S.E., Sanei, H., Beauchamp, B., Chen, Z.H., 2013. Mercury deposition
537 through the Permo–Triassic biotic crisis. *Chem. Geol.* 351, 209–216.

538 Grasby, S.E., Shen, W.J., Yin, R.S., Gleason, G.D., Blum, J.D., Lepak, R.F., Hurley,
539 J.P., Beauchamp, B., 2017. Isotopic signatures of mercury contamination in latest
540 Permian Oceans. *Geology* 45, 55–58.

541 Gehrke, G.E., Blum, J.D., Meyers, P.A., 2009. The geochemical behaviour and
542 isotopic composition of Hg in a mid-Pleistocene western Mediterranean sapropel.
543 *Geochim. Cosmochim. Acta* 73, 1651–1665.

544 Jiang, H.S., Lai, X.L., Yan, C.B., Aldridge, R.J., Wignall, P., Sun, Y.D., 2011. Revised
545 conodont zonation and conodont evolution across the Permian–Triassic boundary
546 at the Shangsi section, Guangyuan, Sichuan, South China. *Global and Planetary*
547 *Change* 77, 103–115.

548 Li, Z.S., Zhan, L.P., Zhu, X.F., Zhang, J.H., Jin, R.G., Liu, G.F., Sheng, H.B., Shen,
549 G.M., Dai, J.Y., Huang, H.Q., Xie, L.C., Yan, Z., 1986. Mass extinction and
550 geological events between Palaeozoic and Mesozoic era. *Acta geological sinica*
551 60, 1–15 (in Chinese with English abstract).

552 Li, M.S., Ogg, J., Zhang, Z., Huang, C.J., Hinnov, L., Chen, Z.Q., Zou, Z.Y., 2016.
553 Astronomical tuning of the end-Permian extinction and the Early Triassic Epoch
554 of South China and Germany. *Earth Planet. Sci. Lett.* 441, 10–25.

555 Liu, S.A., Wu, H.C., Shen, S.Z., Jiang, G.Q., Zhang, S.H., Lv, Y.W., Zhang, H., Li,
556 S.G., 2017. Zinc isotope evidence for intensive magmatism immediately before
557 the end-Permian mass extinction. *Geology* 45, 343–346.

558 McDonough, W. F., 2003. Compositional model for the Earth's core. *Treatise on*
559 *geochemistry* 2, 568.

560 Payne, J.L., Turchyn, A.V., Paytan, A., DePaolo, D.J., Lehrmann, D.J., Yu, M.Y., Wei,
561 J.Y., 2010. Calcium isotope constraints on the end-Permian mass extinctions.
562 *PNAS* 107, 8543–8548.

563 Percival, L.M.E., Witt, M.L.I., Mather, T.A., Hermoso, M., Jenkyns, H.C., Hesselbo,
564 S.P., Al-Suwaidi, A.H., Storm, M.S., Xu, W., Ruhl, M., 2015. Globally enhanced
565 mercury deposition during the end-Pliensbachian extinction and Toarcian OAE:

566 A link to the Karoo–Ferrar Large Igneous Province. *Earth Planet. Sci. Lett.* 428,
567 267–280.

568 Percival, L.M.E., Ruhl, M., Hesselbo, S.P., Jenkyns, H.C., Mather, T.A., Whiteside,
569 J.H., 2017. Mercury evidence for pulsed volcanism during the end-Triassic mass
570 extinction. *Proc. Natl. Acad. Sci. U. S. A.* 114, 7929–7934.

571 Pirrone, N., Cinnirella, S., Feng, X., Finkelman, R.B., Friedli, H.R., Leaner, J., Mason,
572 R., Mukherjee, A.B., Stracher, G.B., Streets, D.G., Telmer, K., 2010. Global
573 mercury emissions to the atmosphere from anthropogenic and natural sources.
574 *Atmos. Chem. Phys. Discuss.* 10, 4719–4752.

575 Pyle, D.M., Mather, T.A., 2003. The importance of volcanic emissions for the global
576 atmospheric mercury cycle. *Atmos. Environ.* 37, 5115–5124

577 Gong, Q., Wang, X.D., Zhao, L.S., Grasby, S.E., Chen, Z.Q., Zhang, L., Li, Y., Cao,
578 L., Li, Z.H., 2017. Mercury spikes suggest volcanic driver of the
579 Ordovician-Silurian mass extinction. *Sci.Rep.* 7, 5304.

580 Reichow, M.K., Pringle, M.S., Al'Mukhamedov, A.I., Allen, M.B., Andreichev, V.L.,
581 Buslov, M.M., Davies, C.E., Fedoseev, G.S., Fitton, J.G., Inger, S., Medvedev,
582 A.Y., Mitchell, C., Puchkov, V.N., Safonova, I.Y., Scott, I.A., Saunders, A.D.,
583 2009. The timing and extent of the eruption of the Siberian Traps large igneous
584 province: Implications for the end-Permian environmental crisis. *Earth Planet.*
585 *Sci. Lett.* 277, 9–20.

586 Riccardi, A., L Kump, R., Arthur, M.A., D'Hondt, S., 2007. Carbon isotopic evidence
587 for chemocline upward excursions during the end-Permian event. *Palaeogeogr.*

588 Palaeoclimatol. Palaeoecol. 248, 73–81.

589 Ruiz, W.L.G., and Tomiyasu, T., 2015. Distribution of mercury in sediments from
590 Kagoshima Bay, Japan, and its relationship with physical and chemical factors.
591 Environ. Earth Sci. 74, 1175–1188.

592 Sanei, H., Grasby, S.E., Beauchamp, B., 2012. Latest Permian mercury anomalies.
593 Geology 40, 63–66.

594 Sanei, H., Grasby, S.E., Beauchamp, B., 2015. Contaminants in marine sedimentary
595 deposits from coal fly ash during the latest Permian extinction. Develop.
596 Paleoenviron. Res. 18, 89–99.

597 Saunders, A.D., 2016. Two LIPs and two Earth-system crises: the impact of the North
598 Atlantic Igneous Province and the Siberian Traps on the Earth-surface carbon
599 cycle. Geol. Mag. 153, 201–222.

600 Sial, A.N., Lacerda, L.D., Ferreira, V.P., Frei, R., Marquillas, R.A., Barbosa, J.A.,
601 Gaucher, C., Windmüller, C.C., Pereira, N.S., 2013. Mercury as a proxy for
602 volcanic activity during extreme environmental turnover: The
603 Cretaceous–Paleogene transition. Palaeogeogr. Palaeoclimatol. Palaeoecol. 387,
604 153–164.

605 Sial, A.N., Chen, J.B., Lacerda, L.D., Frei, R., Tewari, V.C., Pandit, M.K., Gaucher,
606 C., Ferreira, V.P., Cirilli, S., Peralta, S., Korte, C., Barbosa, J.A., Pereira, N.S.,
607 2016. Mercury enrichment and Hg isotopes in Cretaceous–Paleogene boundary
608 successions: Links to volcanism and palaeoenvironmental impacts. Cretaceous
609 Res. 66, 60–81.

610 Shen, S.Z., Crowley, J.L., Wang, Y., Bowring, S.A., Erwin, D.H., Sadler, P.M., Cao,
611 C.Q., Rothman, D.H., Henderson, C.M., Ramezani, J., Zhang, H., Shen, Y.N.,
612 Wang, X.D., Wang, W., Wu, L., Li, W.Z., Tang, Y.G., Liu, X.L., Liu, L.J., Zeng,
613 Y., Jin, Y.G., 2011, Calibrating the end-Permian mass extinction. *Science* 334,
614 1367–1372.

615 Shen, J., Feng, Q.L., Algeo, T.J., Li, C., Planavsky, N.J., Zhou, L., Zhang, M.L., 2016.
616 Two pulses of oceanic environmental disturbance during the Permian–Triassic
617 boundary crisis. *Earth Planet. Sci. Lett.* 443, 139–152.

618 Song, H.Y., Song, H.J., Algeo, T.J., Tong, J.N., Romaniello, S.R., Zhu, Y.Y., Chu,
619 D.L., Gong, Y.M., Anbar, A.D., 2017. Uranium and carbon isotopes document
620 global-ocean redoxproductivity relationships linked to cooling during the
621 Frasnian- Famennian mass extinction. *Geology* 45, 887–890.

622 Sun, Y.D., Joachimski, M.M., Wignall, P.B., Yan, C.B., Chen, Y.L., Jiang, H.S.,
623 Wang, L.N., Lai, X.L., 2012. Lethally Hot Temperatures during the Early
624 Triassic Greenhouse. *Science* 338, 366-370.

625 Svensen, H., Planke, S., Polozov, A.G., Schmidbauer, N., Corfu, F., Podladchikov,
626 H.H., Jamtveit, B., 2009. Siberian gas venting and the end-Permian
627 environmental crisis. *Earth Planet. Sci. Lett.* 277, 490–500.

628 Thibodeau, A.M., Ritterbush, K., Yager, J.A., West, A.J., Ibarra, Y., Bottjer, D.J.,
629 Berelson, W.M., Bergquist, B.A., Corsetti, F.A., 2016. Mercury anomalies and
630 the timing of biotic recovery following the end-Triassic mass extinction. *Nat.*
631 *Commun.* 7, 11147.

632 Thibodeau, A.M., Bergquist, B.A., 2017. Do mercury isotopes record the signature of
633 massive volcanism in marine sedimentary records. *Geology* 45, 96–96.

634 Vasiliev, Y.R., Zolotukhin, V.V., Feoktistov, G.D., Pruskaya, S. N., 2000.
635 Evaluation of the volumes and genesis of the Permo–Triassic trap magmatism on
636 the Siberian platform. *Russ. Geol. Geophys* 41, 1696–1705.

637 Wignall, P.B., Hallam, A., Lai, X.L., Yang, F.Q., 1995. Paleoenvironmental changes
638 across the Permian/Triassic boundary at Shangsi (N. Sichuan, China). *Hist.*
639 *Biol.* 10, 175–189.

640 Wignall, P.B., 2001. Large provinces and mass extinctions. *Earth Sci. Rev.* 53, 1-33.

641 Xiang, L., Schoepfer, S.D., Zhang, H., Yuan, D.X., Cao, C.Q., Zheng, Q.F.,
642 Henderson, C.M., Shen, SZ., 2016. Oceanic redox evolution across the
643 end-Permian mass extinction at Shangsi, South China. *Palaeogeogr.*
644 *Palaeoclimatol. Palaeoecol.* 448, 59–71.

645 Xiao, Y.F., Suzukib, N., He, W.H., 2017. Water depths of the latest Permian
646 (Changhsingian) radiolarians estimated from correspondence analysis. *Earth-Sci.*
647 *Rev.* 173, 141–158.

648 Yin, H.F., Zhang, K.X., Tong, J.N., Yang, Z.Y., Wu, S.B., 2001. The global stratotype
649 section and point (GSSP) of the Permian-Triassic boundary. *China basic science*
650 24, 102–114.

651 Yin, R., Feng, X., Hurley, J.P., Krabbenhoft, D.P., Lepak, R.F., Hu, R., Zhang, Q., Li,
652 Z., Bi, X., 2016. Mercury isotopes as proxies to identify sources and
653 environmental impacts of mercury in sphalerites. *Sci. Rep.* 6, 18686.

654 Zambardi, T., Sonke, J.E., Toutain, J.P., Sortino, F., Shinohara, H., 2009. Mercury
655 emissions and stable isotopic compositions at Vulcano Island (Italy). *Earth*
656 *Planet. Sci. Lett.* 277, 236–243.

657 Zhao, L.S., Chen, Y.L., Chen, Z.Q., Cao, L., 2013. Uppermost Permian to Lower
658 Triassic conodont zonation from three gorges area, South China. *Palaios* 28,
659 523–540.

660 Zhang, M.H., Gu, S.Z., 2015. Latest Permian deep-water foraminifers from Daxiakou,
661 Hubei, South China. *J. Paleontol.* 89, 448–464.

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666 **Figure captions:**

667

668 **Fig. 1 Global (A) and South China (SC) (B) paleogeographic configurations over**
669 **the Permian-Triassic transition showing the locations of the Meishan (MS),**
670 **Daxiakou (DXK) and Shangsi (SS) sections, and the Siberia Traps large igneous**
671 **province (STLIP). Global paleographic map is available online**
672 **<http://jan.ucc.nau.edu/~rcb7/>, and map B follows Feng et al. (1996).**

673

674 **Fig. 2 TOC, Hg, Hg/TOC and Hg isotopic compositions over the**
675 **Permian–Triassic transition at the Meishan section, South China. Hg, Hg/TOC**

676 ratios and Hg isotopes through bed 24 to bed 29 are from Grasby et al. (2017), and
677 shown by red circles. Sedimentation rates (SR) are calculated by Li et al. (2016)'s age
678 model. Conodont zones follow Chen et al. (2015).

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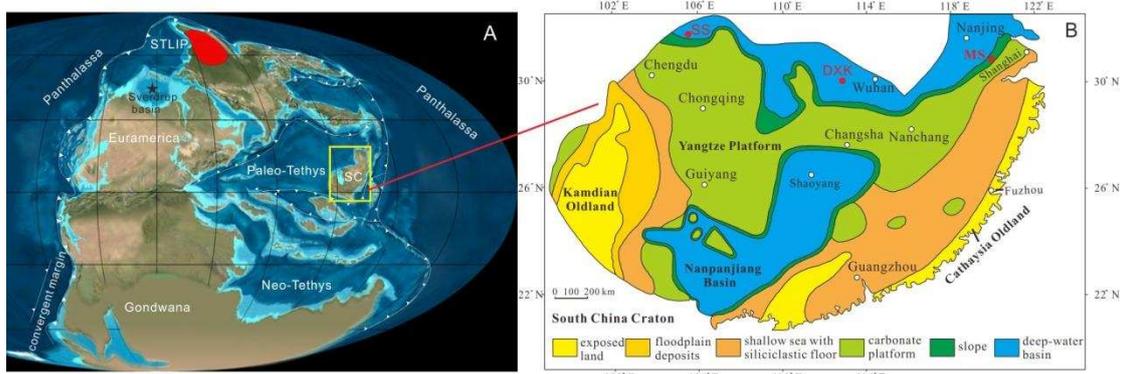
681 **Fig. 3 TOC, Hg, Hg/TOC and Hg isotopic compositions over the**
682 **Permian–Triassic transition at the Daxiakou section, South China.** Sedimentation
683 rates (SR) are calculated by Li et al. (2016)'s age model. Conodont zones follow Zhao
684 et al. (2013).

685

686 **Fig. 4 TOC, Hg, Hg/TOC and Hg isotopic compositions over the**
687 **Permian–Triassic transition at the Shangsi section, South China.** Conodont zones
688 follow Jiang et al. (2011).

689

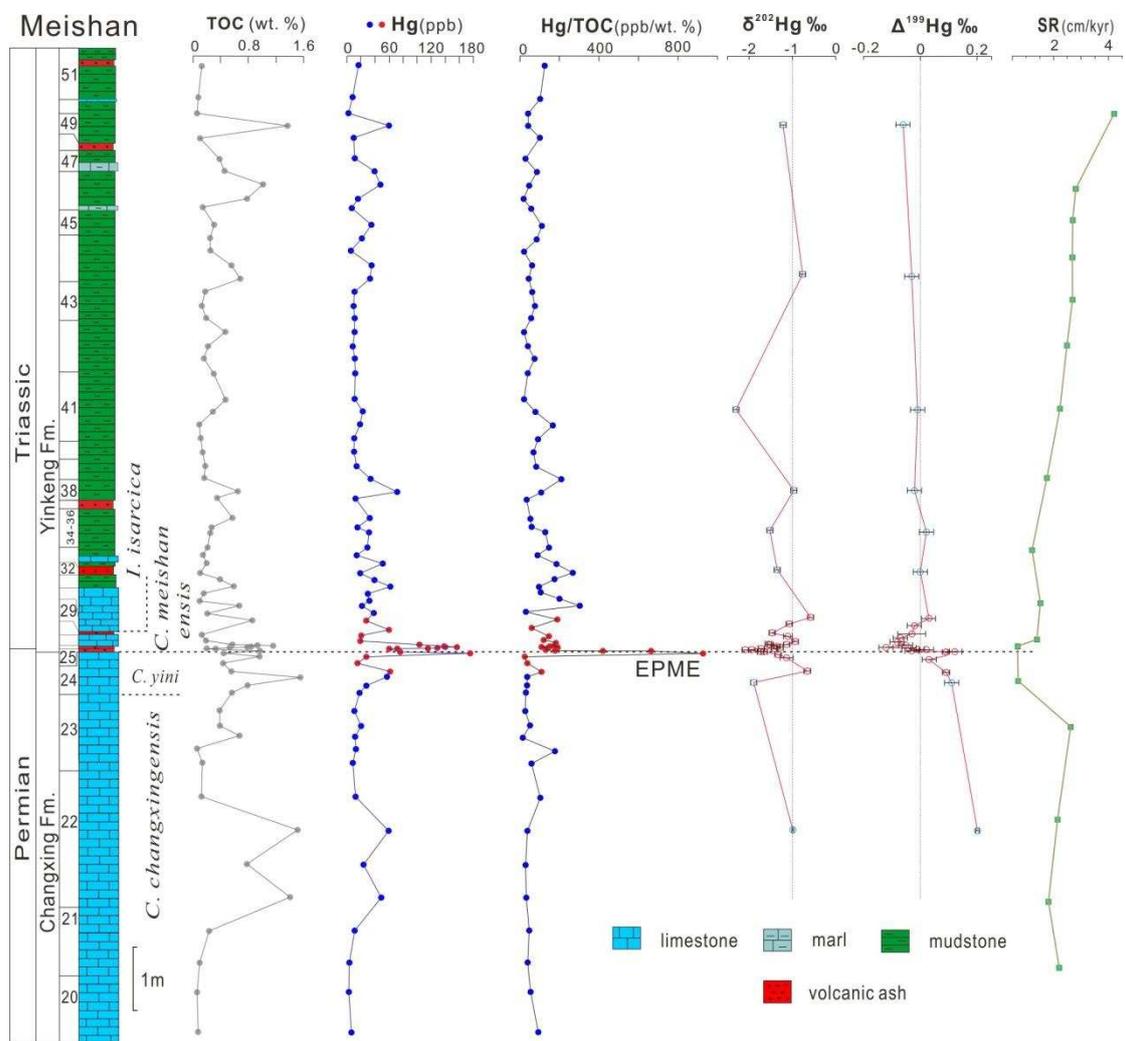
690 **Fig. 5 Integrated excursions of Hg/TOC ratios, Hg isotopes, $\delta^{13}\text{C}$ values (Burgess**
691 **et al., 2014), species richness variations (Song et al., 2013) and STLIP stages**
692 **(Burgess et al., 2017) over the Permian–Triassic transition.** Ages are from
693 Meishan section reported by Burgess et al. (2014). C = Clarkina, H = Hindeodus, I =
694 Isarcicella



695

696 Fig. 1

697



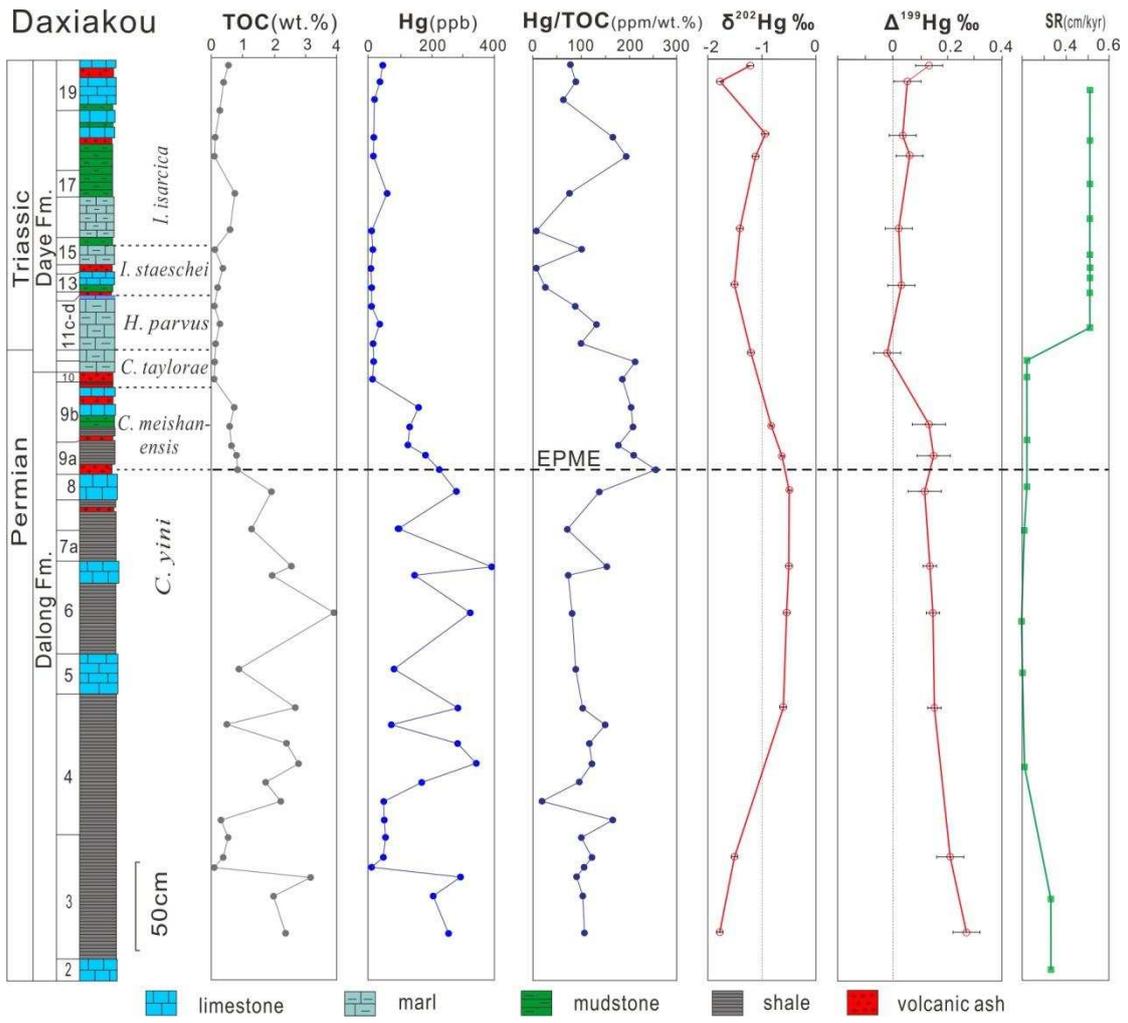
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699

700 Fig. 2

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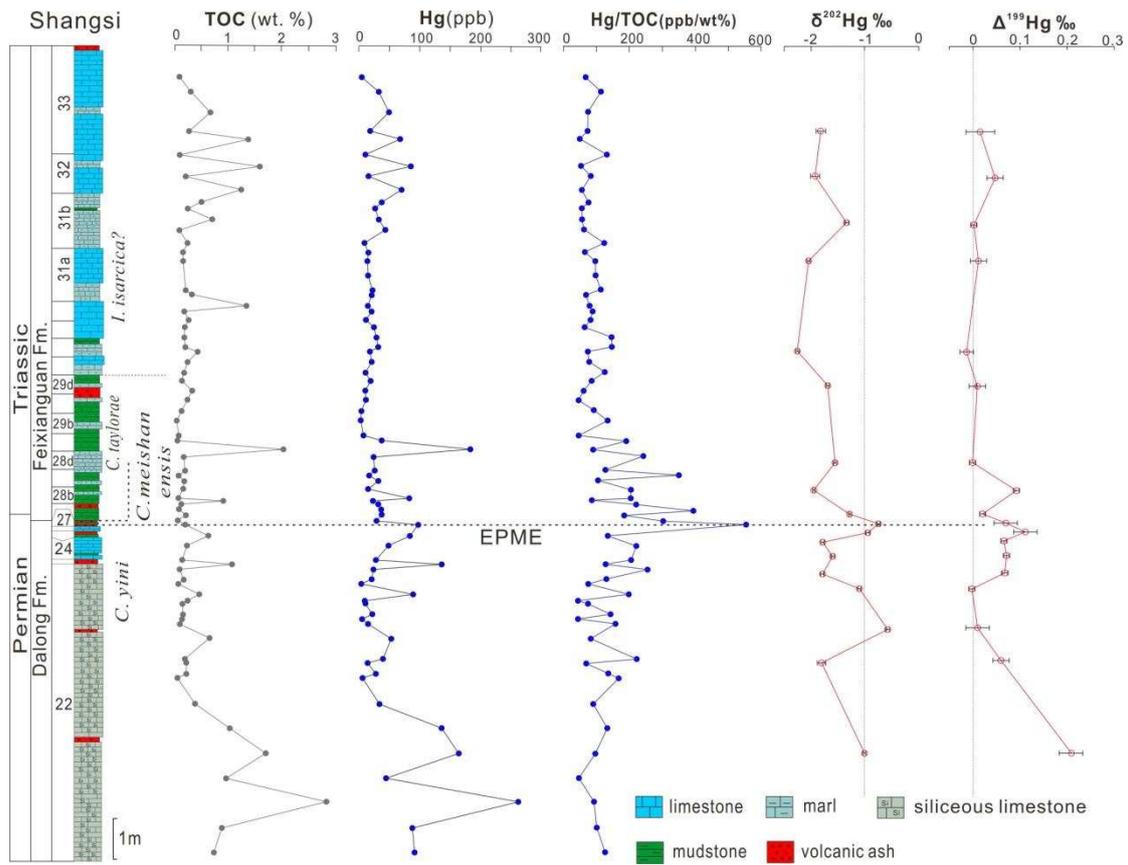
702



703

704 Fig. 3

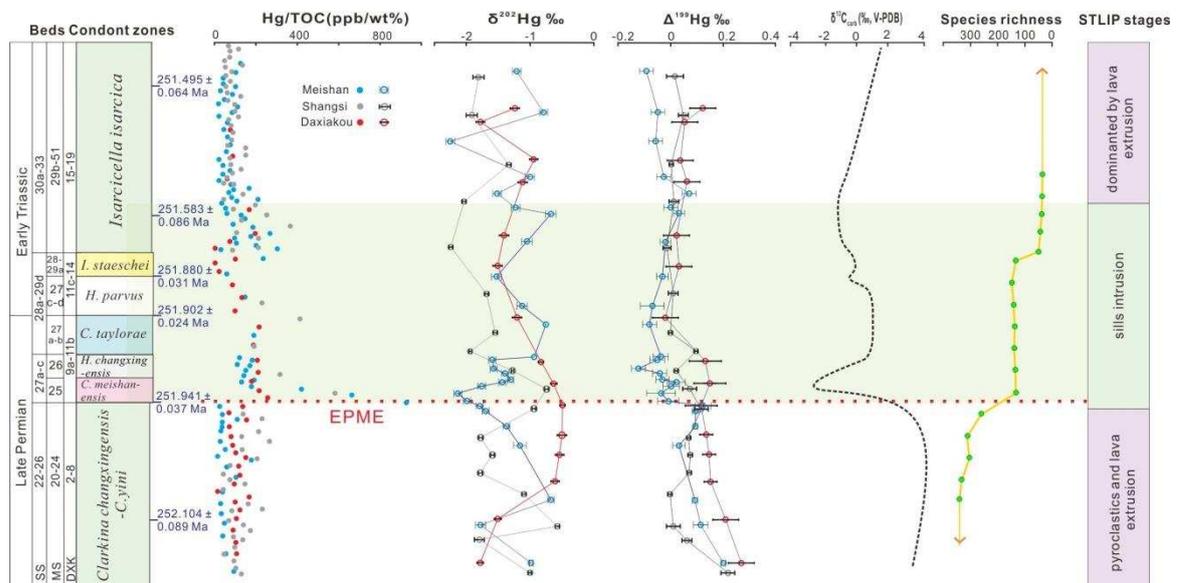
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707 Fig. 4

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710 Fig. 5

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