

This is a repository copy of *Sub-barrier fusion cross section measurements with STELLA*.

White Rose Research Online URL for this paper: <a href="https://eprints.whiterose.ac.uk/id/eprint/131311/">https://eprints.whiterose.ac.uk/id/eprint/131311/</a>

Version: Published Version

# **Proceedings Paper:**

Heine, M., Courtin, S., Fruet, G. et al. (19 more authors) (2017) Sub-barrier fusion cross section measurements with STELLA. In: Nuclear Physics in Astrophysics VIII, NPA8 2017. 8th International Conference on Nuclear Physics in Astrophysics, NPA8 2017, 18-23 Jun 2017 EDP Sciences, ITA.

https://doi.org/10.1051/epjconf/201716501029

### Reuse

This article is distributed under the terms of the Creative Commons Attribution (CC BY) licence. This licence allows you to distribute, remix, tweak, and build upon the work, even commercially, as long as you credit the authors for the original work. More information and the full terms of the licence here: https://creativecommons.org/licenses/

#### Takedown

If you consider content in White Rose Research Online to be in breach of UK law, please notify us by emailing eprints@whiterose.ac.uk including the URL of the record and the reason for the withdrawal request.



## Sub-barrier fusion cross section measurements with STELLA

M. Heine<sup>1,\*</sup>, S. Courtin<sup>1,2</sup>, G. Fruet<sup>1</sup>, D.G. Jenkins<sup>3</sup>, D. Montanari<sup>1,2</sup>, P. Adsley<sup>3,4</sup>, C. Beck<sup>1</sup>, S. Della Negra<sup>4</sup>, P. Dené<sup>1</sup>, F. Haas<sup>1</sup>, F. Hammache<sup>4</sup>, G. Heitz<sup>1</sup>, O.S. Kirsebom<sup>5</sup>, M. Krauth<sup>1</sup>, J. Lesrel<sup>4</sup>, A. Meyer<sup>4</sup>, L. Morris<sup>3</sup>, P.H. Regan<sup>6,7</sup>, M. Richer<sup>1</sup>, M. Rudigier<sup>6</sup>, N. de Séréville<sup>4</sup>, and C. Stodel<sup>8</sup>

**Abstract.** The experimental setup STELLA (STELlar LAboratory) is designed for the measurement of deep sub-barrier light heavy ion fusion cross sections. For background suppression the  $\gamma$ -particle coincidence technique is used. In this project, LaBr<sub>3</sub> detectors from the UK FATIMA (FAst TIMing Array) collaboration are combined with annular silicon strip detectors customized at IPHC-CNRS, Strasbourg, and the setup is located at Andromède, IPN, Orsay. The commissioning of the experimental approach as well as a sub-barrier  $^{12}\text{C} + ^{12}\text{C} \rightarrow ^{24}\text{Mg}^*$  cross section measurement campaign are carried out.

## 1 Introduction

With the measurement of deep sub-barrier light heavy-ion fusion cross sections, both insights into nuclear cluster effects [1] as well as the S-factor at energies of astrophysical interest can be obtained. On the production route towards heavier elements during the carbon burning phase in massive stars,  $^{12}\mathrm{C}(^{12}\mathrm{C},^{24}\mathrm{Mg})$  was identified as a key reaction [2]. The relative energies of the fusion partners in the regime of astrophysical interest are as low as  $E_{rel}=2$  MeV. This is far below the Coulomb barrier of  $^{12}\mathrm{C}$  at  $E_{rel}=6.6$  MeV and the fusion cross sections for such measurements are expected to be in the order of nanobarn. Alongside to the use of high intensity beams up to 10  $\mu\mathrm{A}$  and compacted detector configurations  $^{12}\mathrm{C}$  fusion is detected *via* the coincident measurement of evaporation products and de-excitation gammas [3] from daughter nuclei of the  $^{24}\mathrm{Mg}$  compound.

In the  ${}^{12}C + {}^{12}C$  system the reaction channels, that include light charged particles, are

$$^{12}\text{C} + ^{12}\text{C} \rightarrow ^{24}\text{Mg}^* \rightarrow ^{23}\text{Na} + \text{p},$$
 (1)

$$^{24}\text{Mg}^* \rightarrow ^{20}\text{Ne} + \alpha,$$
 (2)

<sup>&</sup>lt;sup>1</sup> IPHC-CNRS/Université de Strasbourg, 67037 Strasbourg, (France)

<sup>&</sup>lt;sup>2</sup> USIAS/Université de Strasbourg, 67083 Strasbourg, (France)

<sup>&</sup>lt;sup>3</sup>University of York, YO10 5DD York, (UK)

<sup>&</sup>lt;sup>4</sup>IPN d'Orsay, UMR8608, CNRS/IN2P3, PSUD 11, F-91406 Orsay, (France)

<sup>&</sup>lt;sup>5</sup> Åarhus University, DK-8000 Åarhus, (Denmark)

<sup>&</sup>lt;sup>6</sup>University of Surrey, GU2 7XH Guildford, (UK)

<sup>&</sup>lt;sup>7</sup>National Physical Laboratory, Teddington, TW11 0LD, (UK)

<sup>&</sup>lt;sup>8</sup>GANIL, CEA/DSM-CNRS/IN2P3, 14076 Caen, (France)

<sup>\*</sup>e-mail: marcel.heine@iphc.cnrs.fr

with Q-values  $Q_p = 2.24$  MeV and  $Q_\alpha = 4.62$  MeV, respectively. Both reactions occur, in this energy domain, via the formation of an excited compound  $^{24}$ Mg nucleus. It was shown that the de-excitation of (1) and (2) predominantly cascades through the first excited states of the daughter nuclei [4]. Alongside with the measurement of the corresponding particle energies the detection of the characteristic gamma lines at 440 keV and 1634 keV, respectively, is targeted in the first instance.

#### 2 Measurement Station

In the first campaign at Andromède [5], a  $^{12}C^{2,3+}$  beam with energies ranging from 11.1 MeV to 5.5 MeV is directed on (thin) fixed targets of natural  $^{12}C$  with densities of a few tens of  $\mu g/cm^2$ . In the course of the measurements the beam intensity was gradually increased from 40 pnA to 800 pnA. The accuracy of the detection systems as well as the time alignment of the data acquisition are guaranteed and cross section measurements are carried out [6].

For the detection of de-excitation gammas, an array of 36 (1.5" x 2") LaBr<sub>3</sub> detectors from the UK FATIMA collaboration is developed with the focus on the maximization of the detection efficiency. By means of Geant4 simulations the geometrical acceptance was optimized to 21% of the 4II solid angle with the full-energy peak detection efficiency of 8% (440 keV) and 2% (1634 keV) [7]. The simulations are furthermore utilized to reconstruct the decay pattern of the natural contaminant <sup>138</sup>La. In the comparison to experimental data a correction of an electronic walk of 10 keV per day is carried out. This ensures that the QDC channels of the targeted gamma energies don't drift and a similar correction is foreseen for the charged particle detectors.

The nominal positions for annular in-beam double sided silicon strip detectors are 3 cm upstream as well as 6 cm up- and downstream with respect to the target are equipped with annular detectors based on S1, S3 Micron @ chips. They have a strip pitch of 1505  $\mu$ m (S1) and 886  $\mu$ m (S3) and are protected by absorber foils against scattered beam as well as electrons produced in the target. Due to the compact placement the geometrical acceptance is 25% of the  $4\Pi$  solid angle.

To withstand the heat input during bombardment with up to  $10 \,\mu\text{A}$  beam, a rotating target mechanism is developed and commissioned. Amongst fixed target positions the target wheel hosts three slots for frames with a diameter of 4.6 cm that spin with around 1000 rpm.

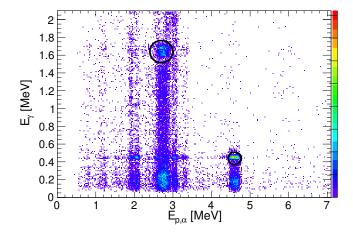
For the absolute normalization, two monitor detectors are placed in extensions around 25 cm from the beam spot position. These barrier surface silicon detectors measure the well know Mott scattering rates from the  $^{12}\text{C}+^{12}\text{C}$  reaction at 45°. In addition, the beam current is recorded by an integrator behind the experiment. From a combination of both measurements, the target thickness can be recognized as well directly from the experimental data.

#### 3 First Results

During the commissioning, the <sup>12</sup>C+<sup>12</sup>C reaction is measured to compare with sub-barrier data from the literature and to guarantee the accuracy of the analysis. The time synchronization of the time stamped trigger less data acquisition system is tested extracting time correlations between gammas (Compton events) as well as the gamma- and particle detectors (fusion reactions). In this first attempt the time coincidence window is set to 400 ns.

In the course of the experiment several calibration measurements with <sup>152</sup>Eu and <sup>241</sup>Am sources are carried out to guarantee the reliability of the energy gates for the gamma- and particle detectors, respectively. It turns out that a crude calibration at the beginning is sufficient for the gammas, while afterward data are calibrated in blocks of one hour based on the comparison to the simulated self activity pattern mentioned earlier.

The correlation of simultaneous gamma- and charged particle detector entries is shown in figure 1 for a beam energy of around  $E_{beam} = 10.4$  MeV for a subset of data. The coincidence window of



**Figure 1.** [Color online] Correlation of simultaneous gamma- and charged particle detector entries with a coincidence window of 400 ns. Fusion events from the  $^{12}\text{C}+^{12}\text{C}$  reaction, populating the first excited state of the  $^{23}\text{Na}$  or the  $^{20}\text{Ne}$  residual nucleus, can be selected like indicated for the characteristic gammas at  $E_{\gamma}=0.440$  MeV or  $E_{\gamma}=1.634$  MeV resulting from their subsequent de-excitation.

400 ns used in this first analysis can still be minimized to reduce random coincidence background. For the calculation of fusion cross sections, the regions of interest around the full-energy depositions of the characteristic gammas from the  $^{23}$ Na as well as  $^{20}$ Ne daughter nuclei are highlighted with ellipses at the gamma energies of  $E_{\gamma} = 0.440$  MeV or  $E_{\gamma} = 1.634$  MeV, respectively.

It is also noteworthy to mention that the timing resolution of the detection system allows for a discrimination of protons and alpha particles based on the rise time of the electronic pulse. In this way, a background free identification of the respective reaction is realized and a comprehensive analysis for further measurements at different beam energy is on the way.

### References

- [1] D. Jenkins and S. Courtin, Jour. Phys. G 42, 034010 (2015)
- [2] A. Chieffi et al., Astr. Phys. Jour. **502**, 737 (1998)
- [3] C.L. Jiang et al., Nucl. Instr. and Methods A 682, 12 (2012)
- [4] K.A. Erb et al., Phys. Rev. C 22, 22 (1980)
- [5] S. Della Negra, Inn. Rev. p. 38 (2016)
- [6] G. Fruet et al., submitted to EPJ Web of Conferences (2017)
- [7] M. Heine et al., J. Phys. Conf. Ser **763**, 012005 (2016)