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# HONO measurement by differential photolysis

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**Abstract.** Nitrous acid (HONO) has been quantitatively measured in situ by differential photolysis at 385 and 395 nm, and subsequent detection as nitric oxide (NO) by the chemiluminescence reaction with ozone (O<sub>3</sub>). The technique has been evaluated by Fourier transform infrared (FT-IR) spectroscopy to provide a direct HONO measurement in a simulation chamber and compared side by side with a long absorption path optical photometer (LOPAP) in the field. The NO–O<sub>3</sub> chemiluminescence technique is robust, well characterized, and capable of sampling at low pressure, whilst solid-state converter technology allows for unattended in situ HONO measurements in combination with fast time resolution and response.

### 1 Introduction

Nitrous acid (HONO) is a major source of hydroxyl (OH) radicals in the boundary layer (Elshorbany et al., 2008; Kim et al., 2014; Levy II, 1973). HONO can be formed homogeneously through reaction of nitric oxide (NO) with OH, heterogeneously through several pathways, or emitted directly (Kleffmann, 2007; Lammel and Cape, 1996; Spataro and Ianniello, 2014; Su et al., 2011). HONO is formed heterogeneously on surfaces through the reaction of NO<sub>2</sub> with  $H_2O$  (Bröske et al., 2003). This heterogeneous formation of HONO is a net source of OH radicals in the troposphere and is an important mediator of air quality, particularly in pol-

luted environments (Finlayson-Pitts et al., 2003; Gutzwiller et al., 2002; Lee et al., 2016). Combustion sources of HONO are direct emission from on-road vehicle exhausts (Rappengluck et al., 2013), aircraft and diesel emissions (Lee et al., 2011a), and biomass burning (Roberts et al., 2010). Emission from snowpack has also been documented (Beine et al., 2008; Zhou et al., 2001), and more recently biogenic sources of HONO have been identified from nitrite-producing bacteria (Oswald et al., 2013; Su et al., 2011) and soil crusts (Weber et al., 2015).

In urban areas HONO can be the major net source of OH (discounting radical cycling driven by e.g. NO), contributing up to 80% of daytime OH production in winter and 50% in summer (Elshorbany et al., 2008; Kleffmann, 2007; Villena et al., 2011b). However, the sources of HONO and the many processes by which it forms are not well understood (Kleffmann et al., 2006; Sörgel et al., 2011; Spataro and Ianniello, 2014; Villena et al., 2011a). There is a clear need for in situ measurement of HONO in order to better understand its chemistry and emissions.

Currently, methods of detecting HONO are either remotely through differential optical absorption spectroscopy (DOAS; Febo et al., 1996; Hendrick et al., 2014; Stutz et al., 2010) or by filter/denuder sampling (Acker et al., 2005, 2006; Febo et al., 1993, 1996; Ianniello et al., 2007). A variety of in situ techniques exist: namely, quantum cascade– tuneable infrared laser differential absorption spectrometry (QC-TILDAS) (Lee et al., 2011b); ion drift chemical ionization mass spectrometry (ID-CIMS) (Levy et al., 2014); ambient ion monitor-ion chromatography (AIM-IC) (Markovic et al., 2012; Vandenboer et al., 2014); stripping coil-visible absorption photometry (SC-AP) (Ren et al., 2011); negativeion proton-transfer chemical ionization mass spectrometry (NI-PT-CIMS) (Roberts et al., 2010); incoherent broadband cavity-enhanced absorption spectroscopy (IBBCEAS) (Pusede et al., 2014); and, as used in this study, long path absorption photometry (LOPAP) (Heland et al., 2001). LOPAP has been characterized quite extensively by other authors (e.g. Clemitshaw, 2004; Kleffmann and Wiesen, 2008; Kleffmann et al., 2006, 2013; Ródenas et al., 2013).

Here, we demonstrate the exploitation of a known HONO interference for photolytic  $NO_2$  conversion systems (Pollack et al., 2011; Ryerson et al., 2000; Sadanaga et al., 2010, 2014; Villena et al., 2012) to provide a simple photolytic technique for quantitative analysis of HONO.

#### 2 Experimental

The differential photolytic HONO technique, henceforth referred to as pHONO, was developed from an existing fast  $NO_x$  analyser, described in Sect. 2.1. The photolytic converter is described specifically in Sect. 2.2. Characterization and calibration are described in Sects. 2.3 and 2.4. The precision, limit of detection, artefacts, interferences, and uncertainties are described in Sects. 2.5 and 2.6.

#### 2.1 Differential photolysis instrument

Measurement were performed using a dual-channel Air Quality Design Inc. (Golden, Colorado, USA) instrument equipped with a UV-LED-based photolytic  $NO_2$  converter – commonly referred to as a blue-light converter (BLC) as described in Reed et al. (2016).

Briefly, a dual-channel NO chemiluminescence analyser operates in parallel with duplicated independent equipment. The two channels share a common inlet allowing for parallel calibration. One channel is equipped with a photolytic NO<sub>2</sub> converter so that NO<sub>x</sub> can be determined with that channel whilst also measuring NO concurrently. This allows for fast (1 Hz or greater) determination of NO and NO<sub>2</sub>. A chemiluminescent zero is taken every 5 min by increasing the NO + O<sub>3</sub> reaction time. Practically, the sample flow is diverted to a PFA volume so that >99% of NO is converted; any residual chemiluminescence signal arises therefore from slower (~ 2 orders of magnitude) O<sub>3</sub>+ alkene reactions (Drummond et al., 1985).

In order to be able to also measure HONO, the NO<sub>x</sub> channel was redesigned so that the photolytic converter (Sect. 2.2) operates in a switching mode. That is, the two lamps of different wavelengths operate alternately on a 50 % duty cycle. The lamps switch every 30 s, allowing for 1 min time resolution data.

#### 2.2 NO<sub>2</sub>–HONO photolytic converter

Photolytic converters were based on those supplied by Air Quality Design and manufactured according to their proprietary standards (Buhr, 2004, 2007), and they are described in Reed et al. (2016). Two UV-LED arrays are positioned at opposing ends of a cavity which is highly reflective to UV. Sample gas is introduced at one end of the illuminated cavity, exiting at the other. NO in the sample exiting the converter is enhanced over the original by photolysis of NO<sub>2</sub> or HONO; thus by calibration of the conversion efficiency these can be quantified.

Modifications were made to the control of the UV-LED elements to allow independent switching of the lamps. The wavelength of one lamp was changed from standard (395 nm) to 385 nm in order to overlap better with the HONO absorption spectrum, whilst the actual UV-LEDs (3 W, LED Engin, Inc.) are more efficient and higher powered than those used in previous work (Reed et al., 2016).

The volume of the illuminated sample chamber is 16 mL, which, with a standard flow rate of 1 standard L min<sup>-1</sup>, gives a sample residence time of 0.96 s, assuming plug flow, at standard atmospheric temperature and pressure (SATP). The NO<sub>2</sub>-to-NO conversion efficiency of the standard BLC with the sample flow of 1 standard L min<sup>-1</sup> was ~ 89 % with both lamps illuminated. Individual lamp conversion efficiencies were 72.9 and  $81.3 \% \pm 0.1$  for the 385 and 395 nm lamps, respectively. Determination of the conversion efficiency is detailed in Sect. 2.4.

#### 2.3 Characterization

Spectral radiograms of the UV-LED output were obtained using the same procedure and equipment described in Reed et al. (2016) using an Ocean Optics QE65000 spectral radiometer coupled to a  $2\pi$  quartz collector within a light-sealed chamber.

Figure 1 shows the measured spectral emission of two UV-LED units of two different wavelengths; 385 and 395 nm. Also shown is the absorption cross section of HONO, BrONO<sub>2</sub>, and the NO<sub>2</sub> quantum yield (Burkholder et al., 2015). It is clear that there is greater overlap, calculated to be 30 %, of the HONO absorption features with the 385 nm LED than at 395 nm. In Reaction (R2) we see that NO is produced stoichiometrically through the photolysis of HONO. In this way, illuminating an air sample at either wavelength yields a signal, which we shall denote as NO<sup>†</sup><sub>2</sub> and which represents the sum of contributions from NO<sub>2</sub> and HONO (Reactions R1 + R2) in differing proportions depending upon wavelength.

$NO_2 + hv(\langle 410 \text{ nm}) \rightarrow NO + O(^3P)$	(R1)
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$$HONO + hv(<390\,\text{nm}) \rightarrow \text{NO} + \text{OH}$$
(R2)



Figure 1. The measured spectral output of two UV-LED elements, nominally 385 nm output in dark blue and 395 nm in red. The HONO absorption spectrum is shown in light blue, whilst the  $NO_2$  quantum yield is shown in dashed black. The absorption cross section of BrONO<sub>2</sub> is shown in green.

The difference in  $NO_2^{\dagger}$  signal measured at 385 and 395 nm corresponds to the difference in conversion efficiency of HONO and NO<sub>2</sub> between the two wavelengths. Differences in NO<sub>2</sub> conversion efficiency of each lamp may be readily calibrated for and so taken into account (see Sect. 2.4). The difference in  $NO_2^{\dagger}$  signal measured at 385 and 395 nm can therefore be used to calculate the HONO present in the sample Eq. (1):

$$\frac{NO_2^{\dagger}_{385} - NO_2^{\dagger}_{395}}{HONO \ CE_{385} - HONO \ CE_{395}} = [HONO].$$
(1)

Apparent HONO conversion efficiency (CE), HONO  $CE_{385}$ -HONO  $CE_{395}$ , is determined experimentally as described in Sect. 2.4.

#### 2.4 HONO and NO<sub>2</sub> conversion efficiencies

The NO<sub>2</sub>–HONO converter system was calibrated for both NO<sub>2</sub> and HONO conversion efficiency. NO<sub>2</sub> conversion efficiencies were determined following the procedure outlined by Lee et al. (2009). Conversion efficiency is the extent to which the equilibrium position of Reactions (R1) and (R2) are to the right. This is dependent on the jNO<sub>2</sub> and jHONO of the UV source, the residence time, and the concentration of oxidants of NO in the photolysis cell as described by Eq. (2) (Ryerson et al., 2000). Here *t* is the residence time within the photolysis cell, and k[Ox] is the concentration and rate constant of any oxidant that reacts with NO. Typically this would be ozone; however, OH formed from HONO photolysis must also be considered.

$$CE = \left[\frac{jt}{jt + k\left[Ox\right]t}\right] \left[1 - \exp^{(-jt - k\left[Ox\right]t)}\right]$$
(2)

The NO + OH back reaction occurs after an air sample has exited the photolytic converter but before entering the high vacuum of the analyser, causing a decrease in signal from HONO as discussed in Sect. 2.5. A similar reaction of  $NO + O_3$  also occurs; however this is ~ 3 orders of magnitude slower than NO + OH (Atkinson et al., 2004) and insignificant on the short (0.11 s) timescale in our instrument.

The sensitivity of a detector in counts per second per part per trillion (cps/ppt) is determined by adding a 7.5 mL min<sup>-1</sup> mass-flow-controlled flow (MFC) of NO calibration gas (4.78 ppm NO in N<sub>2</sub>, BOC) to the inlet of the analyser whilst sampling an overflow of zero air free from  $NO_x$ , VOC, and ozone. This equates to a calibration concentration of 12.5 ppbV NO per channel. Zero air was generated by scrubbing dried  $(-40 T_d)$  compressed air using Sofnofil (Molecular Products) and activated charcoal (Sigma Aldrich) traps. As described by Reed et al. (2016), this combination results in the lowest NO<sub>2</sub> signal. The sensitivity was ~ 6.8 and ~ 6.4 ( $\pm$ 5%) cps/ppt for the NO and NO<sub>x</sub> channels, respectively. In order to determine the NO<sub>2</sub> converter efficiency, a portion of the NO added to the inlet is first titrated to NO<sub>2</sub> by reaction with ozone, typically generating 10.0 ppbV NO<sub>2</sub>. Ozone is generated by illuminated a small flow (~  $10 \text{ mL min}^{-1}$ ) of O<sub>2</sub> with a broad-output lowpressure mercury UV lamp (BHK Inc.) The analyser signal (photomultiplier counts in Hz) is then recorded with neither UV-LED illuminated and then with each illuminated in turn to determine the increase in signal arising for each lamp. The conversion efficiency (CE) is then determined as in Eq. (3), where, when calibrating in zero air or at high mixing ratios relative to the measurement, ambient (Amb) values may be neglected. Ambient signals must be considered when calibration and measurement levels are comparable, and in a changing background.

$$CE = 1 -$$
(3)  
(NO.Hz<sub>Untitrated</sub> - NO.Hz<sub>Amb.1</sub>) - (NO<sub>x</sub>.Hz<sub>Illuminated</sub> - NO<sub>x</sub>.Hz<sub>Amb</sub>)

(NO.Hz<sub>Untitrated</sub> – NO.Hz<sub>Amb.2</sub>) – (NO.Hz<sub>Titrated</sub> – NO.Hz<sub>Amb.3</sub>)

The NO<sub>2</sub> conversion efficiency was determined to be 72.9  $(j = 1.3 \text{ s}^{-1})$  and  $81.2 \% (j = 1.7 \text{ s}^{-1}) \pm 0.1$  for the 385 and 395 nm lamps, respectively.

Calibration for HONO was achieved by sampling a permeation source over a range of dilutions using methods modified from Taira and Kanda (1990) and Febo et al. (1995). Nitrous acid was generated by the reaction of hydrochloric acid with sodium nitrite salt as described by Febo et al. (1995), shown in Reaction (R3).

$$HCl + NaNO_2 \rightarrow HONO + NaCl$$
 (R3)

In order to achieve a continuous source of HONO, a permeation tube (Kin-Tek, HRT-010.00-BLANK/U) was filled with HCl (37 %, Fluka, AR grade) and placed in a thermostated (30 to 55 °C) permeation oven (Kin-Tek, 585) with NaNO<sub>2</sub> salt (Fluka, AR grade). The permeation oven was flushed with 1.5 standard L min<sup>-1</sup> zero air. The reaction is limited by HCl which permeates at a low rate thus allowing low concentrations (< 50 ppb) of HONO to be generated continuously.

As side products of Reaction (R3) can also be produced, the output of the permeation source was continuously analysed for impurities. In Reaction (R4) NO and NO<sub>2</sub> can be formed by the gas phase self-reaction of HONO. In Reaction (R5), HNO<sub>3</sub> can be formed by reaction between adsorbed and gas phase HONO.

$$2\text{HONO}_{(g)} \rightarrow \text{NO} + \text{NO}_2 + \text{H}_2\text{O}$$
 (R4)

$$HONO_{(ads)} + NO_2 \rightarrow HNO_3 + NO \tag{R5}$$

To quantify HONO without any direct measurement and close the nitrogen balance, NO, NO<sub>2</sub>, and total NO<sub> $\nu$ </sub> (NO+NO<sub>2</sub>+ other reactive oxidized nitrogen species such as HNO<sub>3</sub>, HONO, PAN) were measured continuously. The differential photolysis instrument itself was used to quantify the NO. NO2 was measured directly using an EPA certified Teledyne API T500U which uses Cavity Attenuated Phase Shift (CAPS) spectroscopy from Aerodyne Research Inc. (Kebabian and Freedman, 2007; Kebabian et al., 2005, 2008), to avoid any HONO interference which would have been present in a photolytic measurement. This is because the T500U uses a 450 nm band-pass for NO<sub>2</sub>, whereas HONO absorbs only below 390 nm as shown in Fig. 1. Glyoxal is a possible interference in the CAPS technique as it also absorbs in the 440-460 nm range (Volkamer et al., 2005), however all VOC are scrubbed from the zero air during calibration. The nitrate radical (from particulate nitrate) is also an interference in the CAPS technique as it absorbs above  $\sim$  400 nm (Stark et al., 2007), however the T500U is fitted with a high-efficiency particulate air (HEPA) filter to mitigate this. Nitric acid which absorbs below 340 nm (Rattigan et al., 1992) is also not an interference. Total  $NO_v$  was quantified using a Thermo Environmental 42c TL  $NO_x$  analyser equipped with a molybdenum catalytic converter which has been shown to quantify  $NO_{\nu}$  species such as HONO and HNO<sub>3</sub> (Clemitshaw, 2004; Fehsenfeld et al., 1987; Villena et al., 2012; Williams et al., 1998). The TEI 42c TL and Teledyne API T500U were calibrated either directly with an NO standard or by gas phase titration of NO to NO<sub>2</sub> using a Monitor Europe S6100 Multi Gas Calibrator. Production of HNO<sub>3</sub> (Reaction R5) would be indicated by an enhancement in NO over NO<sub>2</sub>, as NO and NO<sub>2</sub> are produced stoichiometrically through the self-reaction of HONO (Reaction R4), whereas HNO<sub>3</sub> production consumes NO<sub>2</sub> and produces NO. Thus, HNO<sub>3</sub> can be indirectly quantified by the NO: NO<sub>2</sub> ratio, and was found to be a minimal contribution to total  $NO_{y}$ . As such, HONO can reasonably be presumed to be equivalent to  $[NO_{\nu}] - ([NO] + [NO_2] + [HNO_3])$ . Measured quantities are shown in Table 1.

The stability of the HONO permeation source was recorded over a 12 h period using NO<sub>x</sub> measured by the differential photolysis analyser (the most sensitive measurement available) as a proxy for NO, NO<sub>2</sub>, and HONO. The stability was found to be  $\pm 0.01$  ppb h<sup>-1</sup>, with a standard deviation of 0.4 ppb. The uncertainty in the HONO source is determined by a combination of the accuracy of the NO, NO<sub>2</sub>, and NO<sub>y</sub> measurements and their respective calibra-



**Figure 2.** Difference in HONO conversion between 385 and 395 nm UV-LEDs over a range of dilutions. Median values are in red, whilst all data are shown in grey. Linear fit is in black; error bars show standard deviation of the mean values.

tions. The NO calibration uncertainty, due to MFC flows and standard gas accuracy is 5%, similarly for the CAPS  $NO_2$  and Thermo 42i TL  $NO_y$ . This results in an overall uncertainty in [HONO] of 8.7%.

In Fig. 2 the observed conversion of HONO, that is the difference between HONO conversion by the 385 and 395 nm lamps, is shown. As can be seen HONO conversion is consistently  $6.54 \pm 0.21$  % more at 385 nm than 395 nm. The fact that the "apparent HONO conversion" (HONO CE<sub>385</sub> –HONO CE<sub>395</sub> in Eq. 1) is constant as a function of HONO means that the determination of [HONO] should be a linear function of the difference in NO<sup>†</sup><sub>7</sub> signal at 385 and 395 nm.

#### 2.5 Measurement precision and limit of detection

The precision of NO chemiluminescence detection is the ability to discriminate signal from photon counting noise. The photon counting noise is taken as the standard deviation of the pre-chamber zero signal (determined every 5 min as described in Sect. 2.1) in Hz (counts per second) divided by the sensitivity in cps pptV<sup>-1</sup>. The average zero count rate was ~ 3000 Hz on both channels with a 1  $\sigma$  standard deviation of ~ 75 Hz, giving a precision for NO of 11.0 and 11.7 pptV for 1 s data with sensitivities of 6.8 and 6.4 cps pptV<sup>-1</sup> for channels 1 and 2, respectively. The precision for NO<sub>2</sub> and HONO (channel 2) must be divided by their conversion efficiencies. This results in a precision for NO<sub>2</sub> of 16.0 and 14.4 pptV for 1 s data at 385 and 395 nm, respectively. The HONO precision is calculated from the (constant) differential conversion efficiency as 179.9 pptV for 1 s data.

The limit of detection is defined as the precision multiplied by  $1/\sqrt{n}$ , where *n* is the number of data points (Lee et al., 2009). For 1 min data the LOD for NO is 1.4 pptV. As the

No.	NO <sub>y</sub> ppb measured	NO ppb measured	NO <sub>2</sub> ppb measured	HNO <sub>3</sub> ppb calculated	HONO ppb calculated
1	20.40	3.34	2.64	0.35	14.08
2	19.29	2.96	2.35	0.30	13.68
3	16.82	2.59	2.10	0.26	11.89
4	14.95	2.27	1.87	0.20	10.62
5	13.40	2.05	1.73	0.16	9.45
6	12.15	1.86	1.58	0.14	8.57
7	11.09	1.70	1.46	0.12	7.81
8	10.17	1.60	1.35	0.11	7.14
Percent	_	15.5	12.8	1.3	70.4

Table 1. Showing the distribution of  $NO_y$  species NO,  $NO_2$ , HNO<sub>3</sub>, and HONO produced from the HONO permeation source.

 $NO_x$  channel operates on a 30 s switching cycle, the 1 min data are actually only averaged over 30 s. Thus, the 1 min  $NO_2$  limit of detection is 2.6 pptV (calculated from the CE of the 395 nm lamp), and the limit of detection of HONO is 32.8 pptV averaged over 1 min. However the HONO limit of detection must also account for the  $NO_2$  limit of detection at both 385 and 395 nm, resulting in an overall HONO LOD of 38.3 pptV averaged over 1 min.

# 2.6 Measurement artefacts, interferences, and uncertainties

It is noted that at both 385 and 395 nm there is potential photolytic interference from  $BrONO_2$  (or in fact any other compounds which photolyse to give NO at either wavelength), with similar spectral overlap (Fig. 1). Assuming a quantum yield of 1 integrated over all wavelengths for  $BrONO_2$ , 21.5 ppt of  $BrONO_2$  at 385 nm and 18.1 ppt at 395 nm would be required to produce a 1 ppt error in the  $NO_2$ –HONO signal. Due to the low abundance (< 10 pptV) of  $BrONO_2$  in the lower atmosphere (Yang et al., 2005), interference is therefore likely to be minimal (Pollack et al., 2011). The difference in conversion for the different lamps equates to a maximum error in HONO determination of 3.4 % [ $BrONO_2$ ]: typically much less than 1 pptV.

Thermal interferences can also become apparent if the photolytic converter raises the temperature of the sample gas above ambient (Reed et al., 2016). Thermally labile NO<sub>y</sub> compounds, e.g. peroxyacetyl nitrate, may decompose within the converter, resulting in a positive signal when measuring in NO<sub>x</sub> mode. This will be true at either 385 or 395 nm wavelengths. Under the reasonable assumption that the temperature of the photolysis cell is the same at both wavelengths, then, as shown in our previous work (Reed et al., 2016), the interference scales with NO<sub>2</sub> conversion efficiency. In this case the difference in thermal interference at 385 nm is 10.2 % greater than at 395 nm owing to having a lower NO<sub>2</sub> conversion efficiency. This represents a small bias in the HONO measurement, which can be eliminated by hav-

ing NO<sub>2</sub> conversion efficiencies equal at both wavelengths as was the case during field measurements in Sect. 3.2. When NO<sub>2</sub> conversion efficiencies are equal, thermal interferences affect only the NO<sub>2</sub> measurement; when they are unequal, some uncertainty proportional to the difference in NO<sub>2</sub> CE is introduced.

Zero offset artefacts have been shown to manifest in photolytic converters (Gao et al., 1994; Del Negro et al., 1999; Pollack et al., 2011; Ryerson et al., 2000). That is, when sampling  $NO_x$ -free synthetic air, a non-zero signal is observed when the photolytic converter is illuminated. This is attributed to  $NO_x$  production from species adsorbed on the walls of the photolytic converter once illuminated by UV (Pollack et al., 2011). These artefacts were accounted for during calibration by sampling an overflow of  $NO_x$ -free zero air whilst recording the analyser signal when the photolysis cell was illuminated at 385 and 395 nm. The artefacts were found to be 368 and 319 ( $\pm 5\%$ ) pptV, respectively. There was 0 pptV artefact in NO. UV-induced artefacts vary with time (Ryerson et al., 2000) and have intrinsic uncertainty in their determination; they thus contribute to overall measurement uncertainty.

This spurious artefact signal can be minimized by periodic cleaning of the inside of the photolysis cell (Pollack et al., 2011; Reed et al., 2016; Ryerson et al., 2000), by choice of cell material (Reed et al., 2016), and by lensing the UV light so as not to illuminate the walls of the cell – sacrificing some conversion efficiency (Ridley et al., 1988).

The effect of the back reaction of OH + NO to reform HONO before detection of NO, thus reducing the NO signal in the NO<sub>x</sub>-HONO measurement in the presence of HONO, is another possible source of uncertainty. As a greater proportion of HONO is photolysed at one wavelength respective to another, the sample inside and exiting the converter necessarily has differing OH concentrations. The effect on the NO signal detected was calculated using a box model in FACSIMILE kinetic modelling software (MCPA Software Ltd.). Kinetic data for O<sub>x</sub>, HO<sub>x</sub>, and NO<sub>x</sub> reactions are taken from IUPAC Evaluated Kinetic Data (Atkinson et al., 2004). The box model was initiated with the NO, NO<sub>2</sub> HONO, and HNO<sub>3</sub> concentrations shown in Table 1. NO, NO<sub>2</sub>, O<sub>3</sub>, and OH concentrations at the outlet of the photolysis cell after 1 s illumination at each wavelength were calculated. The residence time between an air sample exiting the photolysis cell and entering the high vacuum of the NO analyser through the  $\sim$  25 cm of 1/4 in. PFA tubing is 0.11 s. The air sample is a mixture of mostly NO, O<sub>3</sub>, OH, and unconverted NO<sub>2</sub>. The absence of UV irradiation results in chemistry analogous to night-time  $NO_x$  chemistry with the addition of a significant OH source. The interference from the OH + NO reaction was determined as the decrease in [NO] during the 0.11 s residence time as a percentage of measured [HONO]. The discrepancy was calculated to vary linearly with [HONO] from -0.97 to -2.10 % over the range of calibrations, with differences between lamps well within the accuracy of the calibration. The degree of interference from OH in NO2 and HONO determination was found to be a function of k([OH] + [NO])on the timescale here (0.11 s). Reducing the residence time after the photolysis cell would reduce the error in HONO and NO<sub>2</sub> (in the presence of HONO). Conversely, a system with a suitably long residence time between the photolysis cell and detector may experience little-to-no HONO interference as the OH + NO back reaction begins to dominate. There is of course a trade-off in that the data must be corrected for ambient ozone affecting the  $NO: NO_2$  ratio. It is important to note that there can never be any negative interference in NO2 caused by the presence of HONO, only positive or none.

Outside of calibration the effect of the OH back reaction with NO is likely to be less significant due to the presence of volatile organic compounds (VOCs) which also react with OH with comparable rates to NO. It is therefore difficult to know the absolute HONO conversion of each UV-LED without very accurate OH reactivity/VOC concentration measurements. Due to these unknowns, it would not be possible to correct the NO<sub>2</sub> signal for HONO interference as might be hoped.

Additional uncertainties arise in the HONO measurement from sampling a changing  $NO_2^{\dagger}$  background. Changes in signal when illuminating a sample at 385 nm cannot be attributed to NO<sub>2</sub> or HONO without also knowing the corresponding signal at 395 nm which necessarily is measured at a different time. This uncertainty can be reduced by increasing the differential conversion (which was 6.54 % in our system). We make the first order assumption that the NO<sub>2</sub> or HONO background changes on slower timescales than the instrument response, and that it does not change over the course of two 30 s cycles.

Thus, the uncertainty in the differential conversion is a combination of the uncertainty in the HONO calibration source (8.7%), the uncertainty of the sensitivities (5%), NO<sub>2</sub>, and HONO conversion efficiencies (5%), and the uncertainty in the artefact (5%). This results in an overall root sum of squares uncertainty of 15.8% in [HONO].

#### 3 Results and discussion

The pHONO instrument was evaluated in an atmospheric simulation chamber (Sect. 3.1) and compared in the field side by side with LOPAP (Sect. 3.2).

#### 3.1 Chamber measurements

The Highly Instrumented Reactor for Atmospheric Chemistry (HIRAC) is a simulation chamber facility based at the School of Chemistry, University of Leeds (Glowacki et al., 2007a). HIRAC is a cylindrical stainless-steel chamber with a total volume of  $\sim 2.25 \text{ m}^3$ , containing four fans for mixing throughout the chamber, and with a total mixing time of  $\sim 60 \text{ s}$ . The stainless-steel structure of HIRAC allows for pressure-dependent experiments to be carried out, over the range of  $\sim 10\text{--}1000 \text{ mbar}$ . Numerous sample ports are located around the chamber allowing the attaching of instruments or introduction of gas. A multi-pass Fourier transform infrared (FT-IR) instrument (Bruker IFS/66, 128.52 m path length) is present to allow spectra of the gas within the chamber to be taken (Glowacki et al., 2007b). HIRAC is also capable of operating over a range of temperatures ( $-40 \text{ to } 70 \,^\circ\text{C}$ ).

Experiments were carried out at ambient temperature (20 °C) and pressure (1000 mbar), whilst the chamber was kept dark. HIRAC was filled with 80 % N<sub>2</sub> (BOC, UHP, 99.998 %) and 20 % O<sub>2</sub> (BOC) before HONO was synthesized external to the chamber following a modified procedure described previously by Taira and Kanda (1990). A 1 % aqueous sodium nitrite solution was added dropwise to a 30 % aqueous solution of sulfuric acid. The resulting Reaction (R6) produces HONO, which was added directly to the chamber via a continuous flow of N<sub>2</sub> over the reaction mixture. This is analogous to the permeation source; however, side products need not be considered due to the direct HONO measurement afforded by FT-IR.

$$2NaNO_2 + H_2SO_4 \rightarrow 2HONO + Na_2HSO_4$$
(R6)

FT-IR spectra were taken at 60s intervals with a spectral resolution of  $1 \text{ cm}^{-1}$ , whilst the differential photolysis analyser sampled from the chamber. Dilution of the HONO-NO-NO<sub>2</sub> mixture was achieved by partial evacuation of the chamber and subsequent refilling with synthetic air  $(N_2/O_2)$ . The average HONO concentration determined from the average of two distinct absorbance lines at  $1264 \text{ cm}^{-1}$  (trans-HONO, Q-branch) and  $853 \text{ cm}^{-1}$ (cis-HONO, Q-branch) in the FT-IR using absolute line strength data from Barney et al. (2000). The absorptivity data were  $5.22 \pm 0.52 \times 10^{-19}$  cm<sup>2</sup> molecule<sup>-1</sup> (1264 cm<sup>-1</sup>) *trans*-HONO) and  $9.00 \pm 0.90 \times 10^{-19} \text{ cm}^2 \text{ molecule}^{-1}$  $(853 \text{ cm}^{-1}, \text{ cis-HONO})$ . Barney et al. (2000) showed there to be discrepancies between various published line strengths, as did Lee et al. (2012); thus there is some inherent uncertainty in the absolute [HONO] determined by FT-IR. Some of the spectra used in quantification are shown in Fig. 3.



Figure 3. FT-IR spectra of dominant HONO absorbance lines at 1264 and  $853 \text{ cm}^{-1}$  over a range of concentrations.



**Figure 4.** HONO determined by FT-IR (*y* axis) versus HONO measured by the photolytic/chemiluminescence differential photolysis instrument (*x* axis). Median values at each dilution are in red; all values are shown in grey. The 1 : 1 line is shown for reference as well as an exponential fit above 140 ppbV HONO. Error bars show the 15.8 % calibration uncertainty of the pHONO instrument and the 10.0 % uncertainty in the FT-IR cross section (Barney et al., 2000).

Figure 4 shows the strong positive correlation between the HONO measured by differential photolysis and by FT-IR within the HIRAC chamber up to ~140 ppbV, deviating at higher mixing ratios. Systematic offset between the pHONO and FT-IR can be attributed to uncertainties in the FT-IR line strength data for HONO and the uncertainty in the pHONO calibration as indicated by the *x* (15.8 %) and *y* (10.0 %) error bars. However, the data lie close to or on the 1 : 1 correlation line when these uncertainties are included. Studies have also shown there to be large discrepancies between different published line strengths for HONO (Barney et al., 2000; Lee et al., 2012), which constitutes a greater uncertainty in [HONO] derived by FT-IR of up to +33 % and -25 to -30 % (Barney et al., 2000) depending on the particular cross section used.

Figure 4 shows that at lower HONO mixing ratios, <140 ppb, there is better agreement between the pHONO and FT-IR measurements, whereas the response of the differential photolysis technique appears to be suppressed at high



**Figure 5.** Simulated conversion (open circles) and difference in conversion (closed circles) for photolytic converters with different j in the presence of OH (red) and O<sub>3</sub> (grey) oxidants.

[HONO]. This is a result of how a photolytic converter operates as expressed by Eq. (2) in Sect. 2.4

Having two LEDs with different HONO absorption overlap results in two values for j(HONO). Using the j(NO<sub>2</sub>) values already found (1.3 and  $1.7 \text{ s}^{-1}$ ) as an easily determined proxy for j(HONO), the change in conversion with oxidant concentration can be approximated.

Figure 5 shows how the percentage conversion of any precursor that dissociates to NO, in this case HONO and NO<sub>2</sub>, changes with increasing oxidant concentration. In the case of O<sub>3</sub> the total conversion decreases linearly with increasing [Ox], whilst the difference between the two remains constant (9%). Conversely, with OH, conversion decays exponentially in total, and as a difference between two LEDs of different *j*. This effect can be seen clearly above 150 ppbV HONO in Fig. 4. Below 140 ppbV a constant difference in conversion of 6.54% is a reasonable approximation.

The high HONO mixing ratios within HIRAC, necessary to be detected by FT-IR (LOD  $\sim$  40 ppb), were several orders of magnitude higher than would be expected in the atmosphere where ppt (Beine et al., 2006; Ren et al., 2010; Zhang et al., 2009, 2012) to low ppb (Acker et al., 2006; Febo et al., 1996; Hendrick et al., 2014; Kanaya et al., 2007; Stutz, 2004) are typical. Thus, this non-linearity at high [HONO] is unlikely to pose a serious limitation of the differential photolysis method, with the possible exception of areas with very high  $NO_x$  backgrounds. This could be partially mitigated by having greater photolysis power at 385 nm, in combination with moving to shorter wavelengths with better overlap with the HONO absorption cross section. It is clear in Fig. 1 that the 385 nm UV-LED has significantly lower light output than at 395 nm; this is reflected in their respective NO<sub>2</sub> conversion efficiencies (72.9 and 81.3%). Alternatively, separate 385 and 395 nm converters can be employed working in parallel, thus doubling the number of UV-LEDs and doubling the photolysis power at each respective wavelength. This would also allow for fast measurement simultaneously, i.e. 1 Hz or faster. Alternatively, the lower conversion efficiency at high [HONO] could be calibrated for, though, as shown in the following section, in typical atmospheric conditions no calibration or correction was required.

# 3.2 Field measurements

The Weybourne Atmospheric Observatory (WAO; Penkett et al., 1999) is a regional Global Atmosphere Watch (GAW) station located on the North Norfolk coast, UK (52°57′01.5″ N 1°07′19″ E). The WAO has a long history of atmospheric measurements stretching back to its inception in 1994. During summer 2015, the WAO was host to the Integrated Chemistry of Ozone in the Atmosphere (ICOZA) campaign, ostensibly measuring ozone production rates. As part of the campaign a long path absorption photometer (LOPAP-03, QUMA Elektronik & Analytik GmbH) (Heland et al., 2001) was deployed in order to measure HONO. Alongside the LOPAP, the NO, NO<sub>2</sub>, HONO (differential photolysis) instrument described in Sect. 2.1 measured at a 1 min time resolution.

During the ICOZA campaign, a high variation of HONO concentrations (up to  $\sim 500 \,\text{pptV}$ ) were observed by the LOPAP between 29 June and 7 July, providing an ideal opportunity for comparison between the two methods. The pHONO was deployed with replacement UV-LEDs with greater output. Both the 385 and 395 nm lamps had the same photon flux, indicated by identical individual NO<sub>2</sub> conversion efficiencies ( $\sim 89\%$ ), in the expectation that better HONO conversion, and therefore sensitivity, would be achieved. The estimated increase in overlap with the HONO adsorption spectrum of the new 385 nm LED was 45 %, compared to 30% calculated for the original LED. Thus lamps were installed as is without calibration to mitigate the fall in output over time that affects the LEDs, particularly the 385 nm LED. The decreasing output is believed to be a result of the power control circuitry of the LEDs which does not limit the current draw immediately after power is supplied, only after a few seconds. This means every time the lamp is switched on it outputs its maximum (with corresponding heat), which, when used in a  $30 \, \text{s}^{-1}$  switching mode as here, shortens the life considerably.

The pHONO instrument sampled from an inlet box (also housing a NO<sub>y</sub> converter) located ~4 m from ground level on the sampling tower at Weybourne. The sample point was connected to the instrument by a 12 m 1/4 in. PFA line (Swagelok) which was shared by the CAPS NO<sub>2</sub> instrument; thus the flow rate was ~3 standard L min<sup>-1</sup>, resulting in a residence time of ~3 s. The LOPAP instrument, which has its own inlet, sampled from the roof of a specially converted van located 20 m away upslope. Consequently, both instruments sampled at a similar height and there was clear, unob-



**Figure 6.** HONO time series during July 2015 at the Weybourne Atmospheric Observatory (WAO) measured by LOPAP (grey) and pHONO (red).

structed line of sight between them. The pHONO inlet was only  $\sim 1 \,\mathrm{m}$  above the Weybourne observatory roof, which may have contributed to the turbulent dynamics observed in the data especially during daytime. The pHONO instrument was calibrated for NO sensitivity in ambient air twice nightly at 00:00 and 04:00 LT; NO offset was taken between these times by assuming it is equivalent to a stable night-time NO value in remote regions away from any source, where NO should be zero in the presence of ozone (Lee et al., 2009). NO<sub>2</sub> conversion efficiencies were determined in zero air once per week. Limits of detection were 1.5 and 1.9 ppt averaged over 1 min for NO and NO<sub>2</sub>, respectively. The LOPAP was operated and calibrated according to the standard procedures described in Kleffmann and Wiesen (2008), with a detection limit of 3 pptV and time resolution of 5 min. Zero measurements using high-purity N2 (N5 grade, BOC) were performed every 12 h on the LOPAP.

Figure 6 shows the HONO time series from both the LOPAP and pHONO instruments during the 8 days of concurrent HONO measurements.

There is reasonable agreement ( $R^2 = 0.58$ , v = 0.82x) between the established LOPAP method of HONO measurement and that provided by the pHONO instrument without correction or calibration (Fig. 6). During the high-ozone and high-HONO events observed on the first and second especially there is very good agreement ( $R^2 = 0.70$ , y = 0.83x) between the two. Gaps in the data represent times where the pHONO limit of determination was reached - where there are too few points in the averaging window after statistical analysis of the data to be meaningful, e.g. <2 points in a 5 min averaging window. This is because in real atmospheric conditions the pHONO instrument is hampered by the time resolution at which data are collected; i.e. if there is strong turbulence, meaning the NO2 or HONO concentration varies rapidly on a timescale shorter than that at which data are collected, then wide scatter is observed as was the case at Weybourne. Strong boundary layer transport meant that the NO2 measurement varied by up to 1.5 ppb in a minute; this was most evident during the mornings. The limitation is because

of the way the data must be processed by interpolating between measurements and subtraction of the 395 nm signal form the 385 nm signal. Decreasing the time between photolysis switching (from 30 s) would obviously decrease this effect, but, ultimately, separate 385 nm (or lower) and 395 nm analyser channels are required. Consequently the data analysis routine for the pHONO data includes tests for the variability of the data, discarding points which show >5 % variation from the subsequent point. Data failing this test are discarded and result in gaps; this is the effective limit of determination. The data are then treated with a robust locally weighted scatter plot smoothing (LOESS; Cleveland, 1979) algorithm to remove extreme values. The gaps in the time series of LOPAP (Fig. 5) were due to the removal of zero measurements and false spikes due to bubbles passing the detector.

Figure 7 demonstrates the level of agreement in the measured HONO concentration by the LOPAP and pHONO methods from 29 June and 7 July. In Fig. 7, the data have been fitted by orthogonal distance regression (Boggs et al., 1987), which allows for errors in both x and y data but which has been shown to possibly overestimate the slope (Carroll and Ruppert, 1996), and by simple linear regression. In either case a strong positive correlation is exhibited, suggesting the replacement UV-LEDs had the desired effect without the application of corrections for the HONO conversion efficiency. The slopes of 0.82 and 1.02 suggest that the new 385 nm lamp was able to convert the majority of HONO. The discrepancy suggests that up to  $\sim$  18 % of HONO was converted by the 395 nm lamp. The scatter evident in Fig. 7 is attributed to atmospheric dynamic effects resulting in a rapidly changing NO<sub>2</sub> background on timescales faster than the response of the instrument  $(30 \text{ s}^{-1})$ , leading to a coefficient of determination of 0.58. Non-zero intercepts indicate a small systematic offset may be present in the pHONO instrument, though more likely is an effect of scatter on the regressions. Disagreement between the two methods (e.g. 29 June) is likely due to local contamination from diesel-powered agricultural and construction equipment operating nearby, and by the exhaust of the fluorescence assay by gas expansion (FAGE) instrument (Whalley et al., 2013). The FAGE can discriminate HO<sub>2</sub> and RO<sub>2</sub> by complete reaction of HO<sub>2</sub>, with NO necessitating high NO concentrations. The exhaust is vented through a vat of Sofnofil sorbent which removes most NO<sub>x</sub> but may also oxidize NO to HONO, which then may be sampled differently by the pHONO instrument and LOPAP due to their relative locations to the FAGE exhaust.

Accuracy and uncertainty in unstable conditions could be improved by measuring at the two different wavelengths concurrently, rather than consecutively. In the same way photolytic  $NO_2$  measurement is improved by measuring concurrently with NO, rather than consecutively. This would require three chemiluminescent analysers in parallel, with two photolytic converters. However, in ambient indoor air quality monitoring, where HONO is seen as increasingly im-



**Figure 7.** Correlation between HONO measured by LOPAP (x axis) and pHONO (y axis). The simple linear regression is shown in blue, and the orthogonal distance regression in green. The 1 : 1 line is shown for reference.

portant (Gligorovski, 2016), a simple single-channel, dualwavelength design might be appropriate and useful.

# 4 Conclusions

An instrument for in situ determination of HONO photolytically has been developed, characterized, and deployed in the field as a proof of concept. During an atmospheric simulation chamber comparison, the HONO measured corresponded well with FT-IR measurement. During field tests the photolytic HONO instrument agreed reasonably ( $R^2 = 0.58$ , y = 0.82x) well with the established LOPAP instrument, though the limitations of having a two-channel sequential measurement were apparent at times; this would be easily overcome in a three-channel concurrent system. Calibration would gain from a pure HONO source; currently the pHONO calibration requires an independent, direct NO<sub>2</sub> measurement and NO<sub>y</sub> measurement.

#### Data availability

We thank the NASA Jet Propulsion Laboratory (Burkholder et al., 2015) for providing comprehensive absorption crosssection data for atmospheric compounds, which can be found at http://jpldataeval.jpl.nasa.gov.

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