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Mantle anisotropy beneath the Earth's mid-ocean ridges

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Abstract

Observations of seismic anisotropy at oceanic spreading centres offer insights into mid-ocean ridge processes and the formation of new plates. Here, remote observations of seismic anisotropy beneath mid-ocean ridges are made using measurements of source-side shear wave splitting. Over 100 high-quality measurements are made using earthquakes that occur near mid-ocean ridges and transform faults, but are observed at teleseismic distances. In general, for off-axis ridge events, the polarisation of fast shear waves, ϕ'' , is approximately parallel to the spreading direction. Nearer the ridge ($\lesssim 50$ km), ϕ'' becomes more scattered and is often ridge-parallel. Delay times, δt , tend to increase from <1 s near the ridge axis to ~ 3 s further away. Slow-spreading regions (Gakkel and Southwest Indian Ridges) show smaller amounts of splitting than faster spreading centres. At transform zones, the pattern is more complex. Coverage beneath the East Pacific Rise is especially good, and we observe a systematic increase in delay times in S wave splitting measurements compared to previous SKS splitting observations made at ocean-bottom seismometers. One compatible explanation is the presence of horizontally-aligned, connected layers of melt at depth; this is also compatible with other observations of the 'LAB' discontinuity and surface-wave derived measurements of radial anisotropy.

Key words: mid-ocean ridges, seismic anisotropy, LAB, mantle dynamics, LPO

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1. Introduction

Although it is well known that mid-ocean ridges (MORs) mark sites where oceanic lithosphere is created, there is still considerable uncertainty about mantle processes near ridges and how melt is extracted to form new crust. It has been long understood that viscous shearing leads to the lattice-preferred orientation (LPO) of mantle minerals at spreading centres (e.g., Hess, 1964; Blackman et al., 1996; Tommasi et al., 1999). Additionally, upwelling and decompression lead to melt generation, and shearing and strain partitioning can cause melt segregation (Phipps Morgan, 1987; Holtzman and Kendall, 2010). Both effects can impart a significant anisotropic signature on seismic waves, measurements of which can be therefore used to probe the dynamics of the Earth's upper mantle (UM) beneath ridges.

Measurements of two orthogonally polarised and independent shear waves (i.e., shear wave splitting) are the most unambiguous observation of anisotropy, and are now routinely made in continental regions, or on oceanic islands (for reviews, see for instance Savage, 1999; Long and Silver, 2009). With UM anisotropy, the orientations of fast shear waves, as derived from splitting measurements, are usually interpreted in terms of LPO in peridotites, where olivine a-axes align roughly parallel to mantle flow directions (e.g., Mainprice, 2007). The delay time between the fast and slow shear-waves is proportional to the magnitude of the anisotropy and the extent of the anisotropic region.

Whilst subduction zones and orogens are well sampled, MORs have not been routinely investigated because of significant logistical problems with placing seismometers on the seafloor. Experiments using ocean-bottom seismometers (OBSs) (Blackman et al., 1993, 1995b; Wolfe and Solomon, 1998; Hung and Forsyth, 1999; Barclay and Toomey, 2003; Harmon et al., 2004) have provided vital insights into MOR processes, though there are still very few observations of shear wave splitting at MORs. Using teleseismic phases (e.g., SKS), these few studies generally reveal fast shear wave polarisations parallel to the direction of plate spreading, with increasing values in delay times moving away from the ridge axis (Wolfe and Solomon, 1998; Hung and Forsyth, 1999; Harmon et al., 2004). These observations are consistent with interpretations of olivine LPO as originally proposed by Hess (1964) (based on

29 observations of P-wave anisotropy) and as modelled by Blackman et al. (1996). In contrast,
30 shallow earthquakes measured within the axial valley show a fast shear-wave orientations
31 in the crust that are parallel to the ridge axis, which are attributed to aligned cracks and
32 layered intrusions of volcanic material (Barclay and Toomey, 2003). Blackman et al. (1996,
33 1995a, 1993) explained the early arrival of P-waves across the southern Mid-Atlantic Ridge
34 in terms of the vertical alignment on olivine a-axes in a mantle wedge beneath the ridge axis.
35 Subsequent modelling has suggested that the vertical alignment of melt in films, pockets or
36 bands would also be very effective in generating shear-wave splitting in near-vertically ar-
37 riving teleseismic phases (e.g., Kendall, 1994; Blackman and Kendall, 1997; Holtzman and
38 Kendall, 2010), and would also predict ridge-parallel fast shear-wave polarisations.

39 Previous studies of anisotropy beneath MORs in a global context have been undertaken
40 using surface waves to infer azimuthal anisotropy (see e.g., Becker et al., 2007). Debayle
41 et al. (2005), for instance, show that beneath MORs, fast orientations are generally similar
42 to the spreading direction, however the behaviour beneath transform zones is more complex
43 and such surface wave studies are limited in their horizontal resolution. It is also the case
44 that even for the simpler case of global inversions for radial anisotropy in the UM, *a priori*
45 corrections for the crust have a strong effect on the results of such inversions (Ferreira et al.,
46 2010). Hence whilst this should be less of a problem in the region of MORs, where the crust
47 is simple, caution in directly interpreting such results is still advisable. In a more localised
48 study Gaherty (2001) and Delorey et al. (2007) mapped vertical and lateral variations in
49 anisotropy beneath the Reykjanes Ridge. Using sources on the Gibbs fracture zone and
50 receivers on Iceland, differences in Love and Rayleigh wave arrival times revealed faster
51 vertically-polarised Rayleigh waves than horizontally-polarised Loves waves near the ridge
52 axis and at depths less than 100 km. This observation is consistent with either the vertical
53 alignment of olivine a-axes or a melt-induced anisotropy, but Holtzman and Kendall (2010)
54 argue that the latter is more likely.

55 In this study we evaluate MOR anisotropy using measurements of shear wave splitting
56 which occur beneath the earthquake, rather than the receiver, using direct S waves—a
57 technique often termed ‘source-side splitting’ (e.g., Schoenecker et al., 1997; Nowacki et al.,

58 2010; Foley and Long, 2011). Using seismic stations with well-characterised anisotropy in
59 the UM beneath the receivers, we can remove the effect of the splitting on the receiver side
60 and measure only that which occurs beneath the source. We then attempt to interpret these
61 observations in the context of previous observations and proposed mechanisms for anisotropy
62 beneath a MOR.

63 **2. Methods and data**

64 *2.1. Shear wave splitting*

65 We aim to measure the seismic anisotropy beneath MORs around the world using the
66 primary observable it produces, shear wave splitting. We use the ‘minimum eigenvalue’
67 technique of Teanby et al. (2004) (which is an extension of that of Silver and Chan (1991)),
68 which removes splitting by effectively maximising the linearity of the horizontal particle
69 motion for a given pair of splitting parameters: the fast direction, ϕ , and the delay between
70 the fast and slow waves, δt . Where measurements are available for an event at more than one
71 station within an azimuthal range of 15° , we use the method of Wolfe and Silver (1998) to
72 stack the small eigenvalue (λ_2) surfaces, with a backazimuth-independent implementation.
73 This significantly reduces the errors when for some stations the measurement is very near
74 null, as the initial polarisation is close to the fast direction beneath the event.

75 In this study, we make the common assumption that the lower mantle above D'' is not
76 significantly anisotropic: despite some evidence of its presence in the uppermost lower mantle
77 (Wookey and Kendall, 2004), several studies support this assumption (e.g., Meade et al.,
78 1995; Montagner and Kennett, 1996; Panning and Romanowicz, 2006; Kustowski et al.,
79 2008). Hence we can infer that any splitting is caused by anisotropy in the UM beneath the
80 source and receiver. If we have prior knowledge of splitting in the UM beneath the receiver,
81 we may correct for this and analyse the S phase, retrieving the splitting caused by anisotropy
82 beneath the source. We interpret the fast direction of the receiver-corrected signal simply
83 by considering the fast orientation at the source, $\phi'' = \text{azimuth} + \text{backazimuth} - \phi$. This
84 simple geometric relationship is true for rays which are vertically incident at the surface,

85 but is only less accurate by a few degrees than a fully slowness-dependent expression, for
86 the range of slownesses in this study. This error is generally less than the uncertainty in the
87 method.

88 *2.2. SKS UM splitting corrections*

89 Seismic anisotropy in the continental UM (where our stations are located) appears to be
90 ubiquitous, and is typically measured using phases such as SKS, PKS and SKKS; SKS is
91 the most commonly used. It converts from a compressional to an S wave upon exiting the
92 outer core, so begins its ascent through the mantle with no splitting present. It is polarised
93 radially, hence it is also polarised parallel to the backazimuth at the receiver. SKS also
94 propagates steeply through D'', which is known to be anisotropic in various places in the
95 lower mantle (see reviews by Kendall and Silver, 1998, 2000; Lay et al., 1998; Nowacki et al.,
96 2011). However, we assume that any contribution to splitting in the phase along this section
97 is minor, as it has spent relatively little time in D''. Studies on a global scale support this
98 approximation (Niu and Perez, 2004; Restivo and Helffrich, 2006), though any strong effects
99 should be visible and display backazimuthal variation in splitting parameters (Hall et al.,
100 2004).

101 Because we wish to remove UM anisotropy from the S phase, we choose seismic stations
102 which have many SKS splitting measurements along a variety of backazimuths. If dipping
103 or multiple layers of anisotropy exist beneath the station, then we expect the results to show
104 a 90° or 180° periodicity to the measurements of ϕ and δt (Silver and Savage, 1994). We do
105 not use stations which exhibit such measurements, as complicated UM anisotropy beneath
106 the receiver is difficult to infer uniquely, and therefore we cannot confidently remove its
107 effects on direct S phases, as they will be arbitrarily polarised compared to the backazimuth,
108 depending on the source mechanism and anisotropic fabric they have encountered near the
109 source. Stations which exhibit backazimuthal variation in SKS splitting may also do so
110 because of laterally heterogeneous anisotropy beneath them. We also avoid using such
111 stations for similar reasons. Our approach is slightly different from some authors, who opt
112 to use stations which appear to show no anisotropy beneath them (Foley and Long, 2011),

113 however these are rare and UM anisotropy appears to be the norm, rather than isotropy.

114 Some studies using surface waves (e.g., Gaherty, 2004) or combining long-period waves
115 with SKS splitting measurements (Yuan and Romanowicz, 2010) show evidence for multiple
116 layers of anisotropy beneath North America, including beneath stations which do not exhibit
117 backazimuthal variation in SKS splitting. It is possible that the same might also be true
118 beneath Ethiopia. It therefore may be that any complexity of anisotropy (e.g., multiple
119 layers) is not imaged in backazimuthal variations in splitting in SKS alone at our stations,
120 despite the apparent requirement for it in surface waves, thus breaking the assumption of
121 simple sub-station anisotropy. However, because our study is concerned only with shear wave
122 splitting, so long as the splitting experienced by SKS and S is similar enough, this should
123 not impact on the use of SKS splitting measurements as corrections here. The similarity
124 between splitting in SKS and S for a given anisotropy beneath the seismic stations used here
125 is therefore the critical assumption we make in this study.

126 In order to be confident of our measurements, we wish to make several for each MOR
127 event, and so we choose from sets of stations in North America and Ethiopia, where extensive
128 SKS splitting studies have been conducted (Ayele et al., 2004; Barruol et al., 1997; Evans
129 et al., 2006; Fouch et al., 2000; Kendall et al., 2005; Liu, 2009; Niu and Perez, 2004; J.O.S.
130 Hammond, pers. comm., 2010). As explained, we reject stations with apparently complicated
131 sub-station anisotropy. SKS measurements for two example stations used in this study are
132 shown in Supplementary Figure 1. The stations used in this study and the SKS splitting
133 parameters used as UM corrections are shown in Supplementary Figure 2.

134 We use these SKS-derived corrections and analyse the direct S phase from events beneath
135 MORs, applying the correction during the analysis. We note that even though reciprocity
136 must apply along the ray path (see, for example, Kendall et al., 1992), the splitting operators
137 are not commutative (Wolfe and Silver, 1998), so it is essential to make the corrections in
138 the correct order (see Wookey and Kendall, 2008; Wookey et al., 2005). As a further check
139 that the correction is valid, after the measurement we check that the source polarisation of S
140 matches that predicted by the event's focal mechanism (see section 2.4). This helps mitigate
141 against the possibility that the S phase we analyse is contaminated by depth phases (sS and

142 pS), as these will generally alter the apparent source polarisation of the combined phase to
 143 be different to that expected from the CMT solution. A difference in the measured source
 144 polarisation may also occur due to the application of an incorrect receiver correction in the
 145 analysis (see below), which also leads us to reject measurements where the two are not in
 146 agreement within 15° .

147 2.3. Testing the use of receiver corrections

148 Whilst we make every effort to ensure that we use seismic stations which have very well-
 149 characterised anisotropy beneath them, some error will be present in the measurement. Part
 150 of the difference will result because of the different slownesses between the S waves we study
 151 and the SKS phases used to make the splitting measurements we use as station corrections,
 152 but the difference is usually negligible in ϕ and very small in δt (see discussion in Nowacki
 153 et al., 2011). The majority of the error therefore likely comes from the assumption that
 154 the anisotropy is simple beneath the station, and that the SKS splitting measurements are
 155 accurate.

156 We conduct synthetic tests to determine how large the uncertainty in the measured
 157 source splitting parameters are when an ‘incorrect’ receiver correction is used. We apply a
 158 known initial amount of splitting (the ‘source-side’ splitting, $\phi_s''^{\text{true}}, \delta t_s^{\text{true}}$) to a synthetic wave
 159 of dominant frequency 0.1 Hz, then a known receiver-side splitting, $\phi_r^{\text{true}}, \delta t_r^{\text{true}}$. We then
 160 analyse the splitting in the wave with a range of receiver corrections ($\phi_r^{\text{trial}}, \delta t_r^{\text{trial}}$) to obtain
 161 the ‘observed’ splitting parameters at the source ($\phi_s''^{\text{trial}}, \delta t_s^{\text{trial}}$) and compare the known and
 162 measured source-side splitting. The procedure can be repeated for any combination of true
 163 source and receiver splitting operators, and all receiver ‘corrections’.

164 Supplementary Figure 3 shows the difference between the true and measured splitting
 165 parameters where $\phi_s''^{\text{true}} = 20^\circ$, $\delta t_s^{\text{true}} = 1.0$ s, and $\phi_r^{\text{true}} = 0^\circ$, $\delta t_r^{\text{true}} = 1.0$ s. The difference in
 166 fast orientation, $\Delta(\phi'') = \text{abs}(\phi_s''^{\text{trial}} - \phi_s''^{\text{true}})$, is within about 15° whilst the trial receiver
 167 correction is within about 40° and 0.4 s of the true receiver splitting parameters. In these
 168 limits, the difference in source delay time, $\Delta(\delta t) = \text{abs}(\delta t_s^{\text{trial}} - \delta t_s^{\text{true}})$, is up to 0.6 s. Consis-
 169 tent with previous tests using real data (Russo and Mocanu, 2009), we find that errors in δt_r

170 appear to cause the largest uncertainty in the ‘observed’ source-side splitting parameters.
171 Supplementary Figure 4 shows the case when $\phi_s^{\text{true}} = 45^\circ$, $\delta t_s^{\text{true}} = 1.0$ s.

172 We also show in Supplementary Figures 3 and 4 the difference between the known and
173 measured source polarisation for a range of ϕ_r^{trial} and $\delta t_r^{\text{trial}}$. The initial polarisation is 0° in
174 both cases. Again, the difference in the true and trial receiver delay times plays a large rôle,
175 and when the ‘observed’ source-side splitting parameters are most inaccurate, the source
176 polarisation is often incorrect by about $10\text{--}20^\circ$. Hence the use of the source polarisation as
177 a diagnostic of the quality of the result is important and helpful.

178 Finally, manual inspection of the results indicates that in several instances the ‘observed’
179 source splitting parameters would be classified as null events, especially where the delay times
180 are large as shown in Supplementary Figures 3 and 4. This also highlights the strength of
181 using manual inspection or an automated null-classifying scheme to maintain the integrity
182 of measurements (Wuestefeld et al., 2010). When all of these diagnostics are included, and
183 the receiver corrections are within an acceptable uncertainty range of within about 20° for
184 the fast direction and 0.4 s for the delay time, we can be confident that the source-side shear
185 wave splitting measurement is a true reflection of the splitting which has affected the wave
186 in the source anisotropic region.

187 *2.4. Event locations and focal mechanisms*

188 In order to make inferences about anisotropy beneath MORs, it is obviously important
189 to accurately know the earthquake location. Because MOR events typically have large
190 uncertainties on their locations in time and space, where possible (for events before 2008)
191 we take these parameters from the ISC’s relocations using the EHB algorithm (Engdahl et al.,
192 1998). The published horizontal uncertainty in the standard ISC locations is approximately
193 20 km; for the EHB locations in this study, the average uncertainty is 7 km.

194 The location of an event—whether beneath a ridge segment or transform zone—may
195 affect the type of anisotropy we expect, hence each event was assigned to one of these
196 categories based on its location relative to the bathymetry (Smith and Sandwell, 1997), and
197 in part its focal mechanism. These were taken from the Global CMT catalogue. Where

198 there was ambiguity from bathymetry, the event was classified as being located on a ridge
199 if the focal mechanism was mainly dip-slip, and as on a transform if mainly strike-slip.

200 2.5. Dataset

201 We consider events of $M > 5.0$, depth ≤ 35 km, in the epicentral distance range $55^\circ \leq$
202 $\Delta \leq 82^\circ$, which are located on the East Pacific Rise (EPR), Mid-Atlantic Ridge (MAR),
203 Gakkel Ridge, and the Southwest and Southeast Indian Ridges (IRs) (Figure 1). At distances
204 less than $\sim 55^\circ$, the difference in incidence angle between SKS and S becomes large enough
205 that the vertical-incidence approximation may no longer be appropriate, and increases the
206 possibility that an SKS-correction for UM anisotropy is inaccurate; beyond $\sim 82^\circ$, the S
207 phase interferes with ScS, or there may be a triplication due to the presence of the D'' layer,
208 contaminating the S signal in the splitting analysis. We of course also wish to avoid D''-
209 traversing rays due to the anisotropy present there. Events deeper than 35 km are unlikely
210 to occur near MORs, and such depths may indicate a poor event location. The seismograms
211 were band-pass filtered between 0.001 and 0.3 Hz.

212 After selection, over 2000 events matched the criteria between 1979 and 2009, according
213 to the USGS National Earthquake Information Center (NEIC) and International Seismolog-
214 ical Centre (ISC) catalogues. Due mainly to signal-to-noise requirements, ~ 400 events were
215 retained for analysis, leaving ~ 820 event-station pairs.

216 During analysis, we apply a strict set of criteria to select the optimum splitting results.
217 Only non-null results which meet the following are retained: (i) acceptable signal-to-noise
218 ratio on both horizontal components; (ii) clear elliptical particle motion before analysis; (iii)
219 clear linearisation of particle motion when corrected; (iv) measured source polarisation is
220 within 15° of the CMT-predicted source polarisation; (v) clear minimum on the λ_2 surface.
221 A quality of 1 (excellent) to 4 (very poor) is assigned manually to each measurement. Null
222 measurements are retained, provided the signal-to-noise ratio is adequate and particle motion
223 is clearly linear before analysis, but after correction for receiver anisotropy.

224 Following analysis, 350 measurements of splitting of ‘fair’ (3) quality or better beneath
225 67 events comprise the dataset. Of these, 122 are of quality ‘good’ (2) or better. There are

226 189 null measurements. The events have magnitude range $4.4 \leq M_b \leq 6.7$, and depth range
227 0–33 km.

228 **3. Results**

229 *3.1. East Pacific Rise*

230 The EPR is the best-sampled MOR segment in this work. Our results agree excellently
231 with SKS splitting results from ocean-bottom seismometers (OBSs) deployed as part of the
232 MELT and GLIMPSE projects (Wolfe and Solomon, 1998; Harmon et al., 2004) (orange
233 bars, Figure 2). Here and in the OBS experiments, ϕ'' or ϕ_{SKS} is approximately parallel to
234 the spreading direction, with δt varying from 1–3 s, depending on distance from the ridge
235 axis. Figure 3 shows the variation of splitting parameters with distance for results on the
236 EPR which are classified as ‘ridge’ events, alongside the MELT and GLIMPSE data.

237 Away from the straightest segments of the EPR, where frequent fracture zones offset
238 the ridge axis, the pattern of observed splitting is different. There is no clear spreading
239 direction-parallel trend to ϕ'' , and the change in δt is also complicated. At about -5°
240 latitude, for example, ϕ'' seems to change over a short distance by $\sim 70^\circ$ from spreading
241 direction-parallel to transform zone-perpendicular. Similarly, the pattern of ϕ'' and δt east
242 of the Pacific–Nazca–Antarctic triple-junction is also complex, with a variation of ϕ'' from
243 parallel to perpendicular to the Challenger Fracture Zone (at about -35° latitude).

244 *3.2. Mid-Atlantic Ridge*

245 Events which produced ‘good’ source-side splitting measurements were limited to lati-
246 tudes between -40° and 15° . Very few events of sufficient magnitude are reported in the
247 catalogues along the Reykjanes Ridge, and no ‘good’ measurements could be made north
248 of the equator. We note that few measured earthquakes occur along clear, linear ridge seg-
249 ments along the MAR, and most seismicity for which we have results is located instead on
250 the transform zones. Nonetheless, the few events clearly beneath ridges (e.g., the stack at
251 -30° latitude) do seem to show spreading direction-parallel ϕ'' . This agrees approximately

252 with SKS splitting measurements made at ASCN (Butt Crater, Ascension Island; Wolfe and
253 Silver, 1998) and SHEL (Horse Pasture, St. Helena; Behn et al., 2004).

254 Along transform zones, about half the results show ϕ'' close to the spreading direction,
255 whilst many show large ($\sim 2.5\text{--}3$ s) δt and ϕ'' roughly perpendicular to the strike of the
256 transform. The dependence of splitting parameters upon distance along the transform zone,
257 away from the nearest ridge segment, is shown in Supplementary Figure 5. The pattern
258 shows considerable variation near the ends of the transform zones, close to the ridge axes,
259 perhaps related to the complex tectonic environment and resultant shearing and melt pro-
260 duction. However, there is a decrease in the maximum delay time as distance from the ridge
261 axis increases, possibly indicating a reduced contribution from a mechanism of anisotropy
262 arising due to melt or other sub-axial process.

263 Two events on the MAR gave results at stations in both North America and Ethiopia. In
264 this case, we may examine the azimuthal dependence of the splitting. Figure 5 shows equal-
265 area lower-hemisphere stereoplots of the splitting parameters, which are notably different
266 along the two different azimuths. Splitting measured at North American stations for both
267 events has smaller δt (stacked splitting parameters: $\phi'' = (25 \pm 4)^\circ$, $\delta t = (1.9 \pm 0.1)$ s), whereas
268 δt is larger when measured along the other azimuth at Ethiopian stations ($\phi'' = (76 \pm 2)^\circ$,
269 $\delta t = (2.6 \pm 0.1)$ s). At this limited range of slownesses, there is not much variation in the
270 angle away from the vertical for the rays, so the differences primarily arise due to azimuth.
271 The Fresnel zones of the two rays of period 20 s stop overlapping significantly when deeper
272 than ~ 200 km, so if heterogeneity were the cause, then the majority of the anisotropy would
273 need to be present below this.

274 3.3. Gakkel Ridge

275 Ten results from events on the Gakkel Ridge were of ‘good’ or better quality, with an
276 equal number of null results. The splitting parameters are shown in Figure 6A. It is notable
277 that most results show a small amount of splitting ($\langle \delta t \rangle = 1.1$ s), and there is a higher
278 proportion of null results than in other regions. The splitting that is present is often ridge-
279 parallel. The spreading rate predicted by NUVEL-1A (DeMets et al., 1994) increases from

280 ~ 6 to 18 mm a^{-1} from right to left in Figure 6A, however there is no clear corresponding
281 trend in the amount of splitting. There is also no obvious systematic variation of parameters
282 for the cluster of events furthest north (rightmost in Figure 6A, circled) with azimuth. A
283 lower amount of splitting beneath such extremely slow-spreading ridges might be related to
284 reduced melt production caused by slow exhumation of material and a consequently small
285 amount of adiabatic decompression melting. If this is the case, the dominant contribution
286 to seismic anisotropy at teleseismic distances would then be from LPO, yet the axis-parallel
287 fast orientations we observe are hard to explain via mineral alignment.

288 *3.4. Southwest and Southeast Indian Ridges*

289 Beneath events on the SWIR and SEIR, 43 individual results, allowing three stacked
290 results, and seven null measurements were made. These are shown in Figure 6B. Again, the
291 pattern is complicated, and few events lie on ridge segments: most are on transforms. The
292 Southwest Indian Ridge shows some of the most oblique spreading of any MOR, so it may
293 help distinguish between processes which lead to anisotropy which is ridge-perpendicular
294 or spreading direction parallel. However, there are insufficient large earthquakes to make
295 any strong inferences from source-side splitting. Interestingly, all measurements made from
296 beneath the ridge segment at longitude 20° appear to be null. This might result from the
297 absence of anisotropy in the region, but it may also occur if the source polarisation is parallel
298 or perpendicular the local fast orientation of some anisotropy. With only one azimuth of
299 measurements and no other events with different source polarisations, it is not possible to
300 distinguish these scenarios.

301 **4. Interpretation and discussion**

302 In interpreting our results, it is difficult to draw firm conclusions about the behaviour of
303 MORs in general because of poor sampling, arising from the lack of stations outside USA
304 and Ethiopia with comprehensive studies published on the backazimuthal variation of SKS
305 splitting parameters. For the purposes of studying relatively small-magnitude earthquakes
306 at teleseismic distances such as is done here, networks of stations with such measurements

307 are necessary to allow stacking of data, especially when fast directions are near the source
308 polarisation. This limitation also means that comparisons between fast and slow ridges are
309 hard to make. However, our measurements do suggest that splitting near the ridge axis is
310 greater beneath fast-spreading ridges (full-rate $> 100 \text{ mm a}^{-1}$) than slow-spreading ones.

311 Our measurements of splitting show ϕ'' to be very similar to the fast orientations observed
312 by regional SKS studies near MORs, but these are extremely limited in coverage because
313 of the practical difficulties in operating such OBS sites. Surface wave studies examining
314 azimuthal anisotropy globally (e.g., Debayle et al., 2005) can provide better coverage near
315 MORs, but limited horizontal resolution means changes over relatively small distances (up
316 to few tens of kilometres and less) cannot be imaged well. Such global measurements tend
317 to show fast orientations approximately parallel to the spreading direction, but this can vary
318 by up to 45° in some places, notably near large fracture zones.

319 *4.1. Doldrums FZ observations*

320 The multi-azimuth observations beneath the Doldrums Fracture Zone (FZ) in section
321 3.2 are interesting (Figure 5). The observation of this azimuthal dependence in splitting
322 parameters appears to be a robust feature: tests requiring ϕ'' to be the same for stations in
323 Ethiopia and North America show the splitting experienced by the direct S wave beneath
324 both sets of stations would have to be different by around 45° to that observed in SKS
325 waves and for which we correct. Such a strong incidence dependence in splitting beneath
326 the receiver would in all likelihood appear as complexity in SKS splitting observations, and
327 here we deliberately avoid stations where this is the case. Whilst the limited number of data
328 prevents detailed analysis, we can speculate on the likely causes of the observed pattern.

329 With two azimuths of observations, we can seek to define an hexagonal symmetry, ori-
330 ented arbitrarily (called ‘tilted transverse isotropy’, or TTI). If we assume Thomsen’s (1986)
331 anisotropic parameters $\delta = \epsilon$ (the case of elliptical anisotropy), we can use the two azimuths
332 of observations to find the plane of isotropy, or axis of symmetry, by simple trigonometry
333 (Nowacki et al., 2011). This dips shallowly to the southwest, as shown by the dashed line
334 in Figure 7. For TTI derived from aligned material, for instance, this would correspond to

335 penny-shaped inclusions having their short axis aligned about 35° from the vertical. This
336 is in some sense similar to the orientations predicted by simulations (Weatherley and Katz,
337 2010), which suggest melt should be focussed along northwest–southeast flow lines for a
338 transform in this orientation. Intriguingly, it also would be consistent with the suggestion
339 of van Wijk and Blackman (2005), who speculate that the transform fault itself would dip
340 towards the ridge segment near the ends of the transform.

341 Another likely contributor to seismic anisotropy in the FZ would be the alignment of
342 olivine in response to flow. Natural samples and deformation experiments show that the
343 dominant way in which olivine develops an LPO is by slip along [100] (a-direction), on
344 $\{0kl\}$ or (010) (b-planes), known as D- and A-type olivine respectively. We examine the
345 possibility that the observed anisotropy at the Doldrums FZ is caused by olivine LPO by
346 using the method described by Wookey and Kendall (2008) and Nowacki et al. (2010). We
347 use the single-crystal elastic constants of olivine (Abramson et al., 1997) and mix them in
348 all proportions with an isotropic average. We then rotate these constants to all possible
349 orientations and compute the shear wave splitting accrued over a 200 km thick layer for the
350 two raypaths observed, and plot the orientations producing splitting compatible with the
351 observations. Figure 7 shows the compatible orientations and degree of alignment as the a-
352 and b-axes of the aligned olivine on a lower-hemisphere equal-area projection.

353 Compatible orientations of the a-axes are northwest–southeast, with glide planes dipping
354 north, northwest or west, all shallowly. This direction of shear is not parallel to the spreading
355 direction or FZ strike, or to the absolute plate motion (APM) (Figure 4). It is consistent
356 with some dynamics models of ridge transforms (e.g., van Wijk and Blackman, 2005; Sparks
357 et al., 1993; Phipps Morgan and Forsyth, 1988), though such models also predict low strain
358 rates in such regions and relate to shorter FZs. This interpretation is, however, inconsistent
359 with the expected behaviour for a pure strike-slip fault, where strains are sufficiently high
360 that olivine [100] directions should be parallel to the FZ strike. This incongruity might
361 support the hypothesis that SPO due to melt or another material is the cause: if olivine [100]
362 directions were parallel with APM or FZ strike, but some other anisotropy were overprinted,
363 then we would not necessarily retrieve the olivine orientations with this method. Equally,

364 such conditions as are present beneath the FZ might lead to the dominance of a different
365 slip system in olivine, in which case flow might still be parallel to the FZ, but not olivine
366 [100] axes. Whilst the uncertainty in our measurements is not insignificant, even with more
367 relaxed constraints on the orientations the picture is much the same (Supplementary Figure
368 6). We of course neglect other anisotropic phases in this approach, but would expect this to
369 require a stronger texturing in the olivine itself to match observations.

370 Finally, heterogeneity between the two raypaths away from the source may also be impor-
371 tant, as paths to Ethiopia spend considerable distance close to the FZ, whilst paths to North
372 America travel away from this region. This would have to be true at depths of ~ 200 km, as
373 the Fresnel zones of waves of this period overlap significantly until that depth. Hence the
374 signal may reflect different anisotropies along these two paths, perhaps with the addition of
375 a common anisotropic region immediately below the source. However, other studies (e.g.,
376) suggest that such strong anisotropy at depths is unlikely, and we prefer to interpret the
377 observations in terms of a mixture of olivine alignment and SPO

378 *4.2. Spreading rate and strength of anisotropy*

379 Dynamic models of MOR accretion have predicted the amount and orientation of shear
380 wave splitting near spreading ridges on the basis of LPO of pure olivine (Blackman et al.,
381 1996), LPO of olivine and enstatite (Blackman et al., 2002; Blackman and Kendall, 2002;
382 Nippres et al., 2007), and combined olivine LPO with the effect of oriented melt pockets
383 (Blackman and Kendall, 1997). In such models, the spreading rate controls the behaviour
384 of upwelling beneath the ridge axis and the shape of the melt-rich region, and hence the
385 orientation and amount of shear wave splitting observed at the surface. In Blackman and
386 Kendall's (2002) simulations, slow spreading ridges (full-rate ~ 40 mm a⁻¹) show significant
387 ridge-parallel splitting within ~ 20 km from the axis because of the requirement that buoyant
388 flow beneath the ridge axis supplies the upwelling material in a small region. Fast-spreading
389 (full-rate ~ 140 mm a⁻¹) ridges, by contrast, do not focus material so efficiently towards the
390 centre and should not produce much observable difference in splitting times between the
391 ridge axis and at distance (> 50 km). Both cases show ~ 0.5 – 1 s of splitting away from

392 the axis. Hence their model predicts there should be little observable difference in splitting
393 in SKS for fast ridges at the axis compared to at distance (>50 km). This agrees with
394 SKS measurements at the EPR (Wolfe and Solomon, 1998; Harmon et al., 2004), where
395 full-spreading rates are >60 mm a $^{-1}$, and on average ~ 150 mm a $^{-1}$.

396 Figure 8 shows δt versus spreading rate for measurements made beneath ridge events,
397 with filled circles indicating those within 50 km of the ridge axis. The above models would
398 predict a negative trend in the near-axis data, with larger splitting times observed at the
399 slowest spreading centres; however, there is no strong trend apparent in our results and the
400 data suggest a weakly positive correlation if any. (Weighted least-squares linear regression
401 for results <50 km from axis gives $R^2 = 0.42$.)

402 Whilst most authors predict increased splitting at the slowest MORs, some observa-
403 tions suggest that anisotropy away from the ridge increases with palaeo-spreading rate. P
404 wave anisotropy in the shallow lithosphere beneath the northwest Atlantic (spreading full-
405 rate ~ 20 mm a $^{-1}$) is significantly less at $\sim 3\%$ (Gaherty et al., 2004) than that observed
406 at present-day fast-spreading sites near the East Pacific Rise ($\sim 6\%$, rate ~ 100 mm a $^{-1}$)
407 (Dunn and Toomey, 1997) and old lithosphere in the western Pacific ($\sim 6\%$, palaeo-rate
408 ~ 60 mm a $^{-1}$) (Shearer and Orcutt, 1986). Such observations constrain the anisotropy in
409 the uppermost mantle, hence probably reflect the effect of ‘frozen-in’ olivine LPO and pro-
410 cesses contemporaneous with lithosphere creation. Gaherty et al. (2004) suggest a spreading
411 rate dependence could be due to slower ridges accommodating more deformation by brittle
412 failure in the crust, leading to reduced LPO in the uppermost mantle. However, our shear
413 wave splitting measurements integrate anisotropy over the complete ray path in the upper
414 mantle, and it is not clear that this effect could cause the change in δt we observe, given the
415 thickness of the brittle crust.

416 It is also important to note that this discussion ignores any azimuthal dependence on
417 splitting parameters. With stations only in North America and Ethiopia, there may be an
418 azimuthal bias between ridges of different spreading rates, which could account for some of
419 the variability we observe.

4.3. Splitting in S and SKS at the EPR

Figure 3 indicates that the amount of shear wave splitting in direct S increases away from the ridge more quickly than that in SKS phases. Assuming that the splitting is accrued over a horizontal layer of constant anisotropy, this difference cannot be accounted for simply by the difference in incidence angle between the phases (which would predict a difference in δt of $\lesssim 0.3$ s for the largest δt_{SKS}). Accruing splitting operators through a model of LPO development at the EPR (Blackman and Kendall, 2002) along the raypaths of the S and SKS phases also indicates that the contrasting azimuths and incidences of the waves are insufficient to produce the observed discrepancy.

One possible explanation might be along-ridge variability in the strength and style of mantle anisotropy, leading to varying splitting dependent on the location along ridge segments. This would imply a sampling bias, whereby events nearest the ridge are in weak anisotropy regions, whilst those furthest are in strong anisotropy regions. This correlation arising by chance or through some earthquake mechanism seems unlikely. More likely may be the influence of azimuth and large-scale heterogeneity. SKS waves travelling to the MELT and GLIMPSE OBSs are along a backazimuth of $\sim 280^\circ$, approximately ridge-perpendicular, whilst S waves in this study travel along an azimuth of $\sim 20^\circ$, closely parallel to the ridge. Hence the S waves may be more sensitive to ridge structure for events near the ridge, and less so at distance, magnifying the contrast in structure between the on- and off-axis mantle.

Another simple explanation may be that current LPO models do not include effects of short-wavelength segregation of material such as that observed in experiments which deform partially molten olivine aggregates (Holtzman et al., 2003a,b), and in numerical experiments incorporating porosity and strain rate-dependent viscosity (Katz et al., 2006). These observations predict lenses of melt at MORs will be aligned approximately with long axes of the strain ellipses expected for corner flow, forming bands which dip away from the ridge axis, becoming approximately horizontal beyond about 50 km from the centre. The alignment of seismically distinct material on scales shorter than the seismic wavelength would lead to a shape-preferred orientation (SPO). Such an SPO with inclusions of uniaxial symmetry would lead to transverse isotropy (TI), where the seismic velocities and amount

449 of splitting vary only away from the axis of rotational symmetry (e.g., Hudson, 1980). This
450 would mean SKS waves travelling perpendicular to the axis (such as those measured by
451 the MELT and GLIMPSE OBSs) would not be split due being polarised in the sagittal
452 plane, whilst near the ridge fast orientations for S waves would be between ridge-parallel
453 and ridge-perpendicular, depending on the dip of the fabric. Further from the ridge, where
454 bands are horizontal, SKS would be unsplit in any azimuth; S waves travelling parallel to
455 the ridge would have ridge-perpendicular fast orientations. This mechanism would increase
456 delay times in S over SKS away from the axis, as we observe at the EPR. Aligned inclusions
457 leading to SPO could also explain surface wave observations that $V_{SH} > V_{SV}$ beneath the
458 Pacific (e.g., Ekström and Dziewoński, 1998; Nettles and Dziewoński, 2008), and receiver
459 function observations of a suboceanic seismic discontinuity at depths of 50 to 150 km (the
460 so-called ‘LAB’ discontinuity; Rychert and Shearer, 2009; Kawakatsu et al., 2009; Rychert
461 et al., 2010; Kumar and Kawakatsu, 2011).

462 Several studies show that the structure beneath the EPR is asymmetric (e.g., Conder,
463 2007; Harmon et al., 2004; Podolefsky et al., 2004; Wolfe and Solomon, 1998), hence this
464 might play some part in the observed difference in splitting times between S and SKS phases.
465 However, our data also sample both sides of the ridge, so presumably are also affected by the
466 same asymmetry, yet still consistently show larger splitting in S than SKS. Blackman and
467 Kendall (1997) calculate the splitting times for vertical-incidence shear waves on a suite of
468 asymmetric model of MOR LPO development, as for the EPR, and for the best-fitting case
469 predict splitting times for SKS of up to 2–3 s on the Pacific plate, and up to 1.5 s on the
470 Nazca plate. This asymmetry is observed to a lesser extent in the data (Figure 3). However,
471 again the effect of azimuth is not tested and the path of the S waves (ridge-parallel) may
472 negate some of the effects on δt expected for an asymmetric spreading centre.

473 4.4. *Plate motion and mantle flow*

474 Several authors interpret shear wave splitting measurements in terms of the alignment
475 of olivine due to shearing of the asthenosphere by the relative motion of the lithosphere
476 above it (e.g., Tommasi, 1998; Conrad et al., 2007). In this study, it is hard to discriminate

477 between this and spreading processes for the EPR, MAR and SWIR because the directions of
478 spreading and APM are very similar (Figures 2, 4 and 6). APM is also small ($<10 \text{ mm a}^{-1}$)
479 at the SWIR. Beneath our measurements on the SEIR, the APM is eastwards, but no fast
480 orientations are parallel to this. At the eastern Gakkel ridge, plate motion is approximately
481 parallel to the ridge, and is faster than the spreading rate ($|\text{APM}| \approx 12 \text{ mm a}^{-1}$). It may be,
482 therefore, that the ridge-parallel ϕ'' reflects the shearing of the North American and Eurasian
483 plates over the asthenosphere. Another interpretation would be that because spreading rate
484 varies along the ridge, mantle material again flows parallel to the axis, again leading to
485 olivine LPO with [100] directions which are compatible with our observations. In both
486 cases, however, the mechanisms which cause this at spreading centres themselves is still
487 unclear. Conrad et al. (2007) point out that parallelism between APM direction and olivine
488 a-axes—assuming LPO in olivine to be the cause of the observed anisotropy—may not be
489 a good assumption beneath MORs, because of the complicated combination of radial and
490 horizontal flow present there. One might also expect along-axis flow of material due to the
491 fact that spreading rate varies along the ridge, which might also explain ridge-parallel fast
492 orientations.

493 A more involved explanation of upper mantle shear wave splitting results can be invoked
494 by considering more complex flow regimes. Behn et al. (2004), for instance, combine plate
495 motion models with a model of mantle flow derived from seismic tomography, and compare
496 SKS splitting observations with the fast orientations predicted by the flow models. In
497 this case, the authors conclude that plate spreading directions adequately describe SKS
498 fast orientations on oceanic plates within 500 km of the ridge. In this study, only one
499 measurement is further than this from the spreading centre (westernmost stack in Figure
500 2), and here the fossil spreading direction and APM are the same within a few degrees.
501 This might serve to reinforce the fossil anisotropy in the lithosphere, but δt here is modest
502 ($(1.7 \pm 0.5) \text{ s}$).

503 Where inferred mantle flow beneath Africa might affect our measurements, on the SWIR,
504 we see several fast orientations parallel to the spreading direction, however one stacked
505 measurement (Figure 6) at $\sim 32^\circ \text{E}$ is perpendicular to spreading. This direction does not

506 correlate with APM. Whilst more rigorous modelling of deeper flow could be attempted (e.g.,
507 Forte et al., 2010), this is beyond the scope of this study, and may be more appropriate when
508 further data are presented.

509 **5. Conclusions**

510 We present measurements made using the source-side shear wave splitting technique of
511 upper mantle anisotropy beneath mid-ocean ridges around the world. We correct for the
512 UM on the receiver side for seismic stations where the anisotropy beneath is very well char-
513 acterised, and can resolve the source anisotropy, subject to a series of rigorous tests. With
514 122 new observations, the presented dataset adds significantly to the current knowledge of
515 anisotropy beneath MORs. There is no strong trend that corroborates the prediction of more
516 splitting beneath slow-spreading ridges, and it may be true that more splitting is present at
517 fast-spreading ones. For the EPR, comparisons with previous SKS splitting measurements
518 show more splitting in S away from the ridge. We suggest that TI dipping away from the
519 ridge axis, becoming horizontal at distance, is compatible with our observations, in addition
520 to LPO development. This would be consistent with other observations of anisotropy and
521 a seismic discontinuity beneath the oceans. We find anisotropy at MORs appears to be
522 dominated by ridge processes, rather than plate motion over the asthenosphere. As further
523 rigorous study of UM anisotropy using SKS phases becomes routine, more stations can be
524 used to measure the seismic shear wave splitting beneath MORs and other remote parts of
525 the Earth where earthquakes occur, and hence our understanding of mantle dynamics in
526 these regions will be vastly improved.

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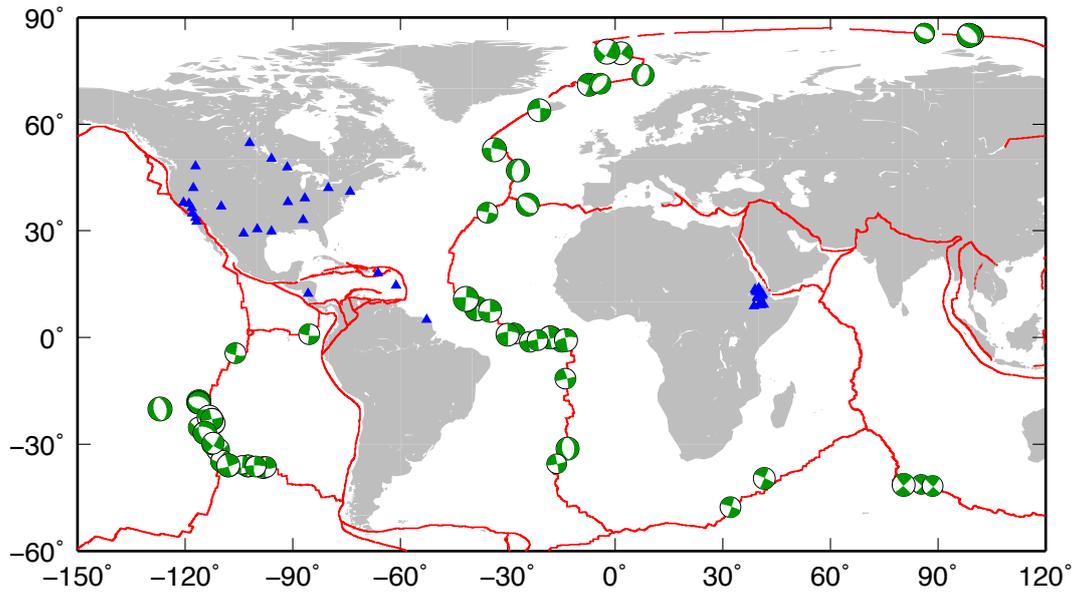


Figure 1: Location of events used in this study with plate boundaries of Bird (2003). Lower-hemisphere focal mechanisms are the best-fitting double-couple solutions as given by the Global CMT project. Blue triangles are seismic stations. Magnitude range ($4.4 \leq M_b \leq 6.7$) shown by size of hemispheres.

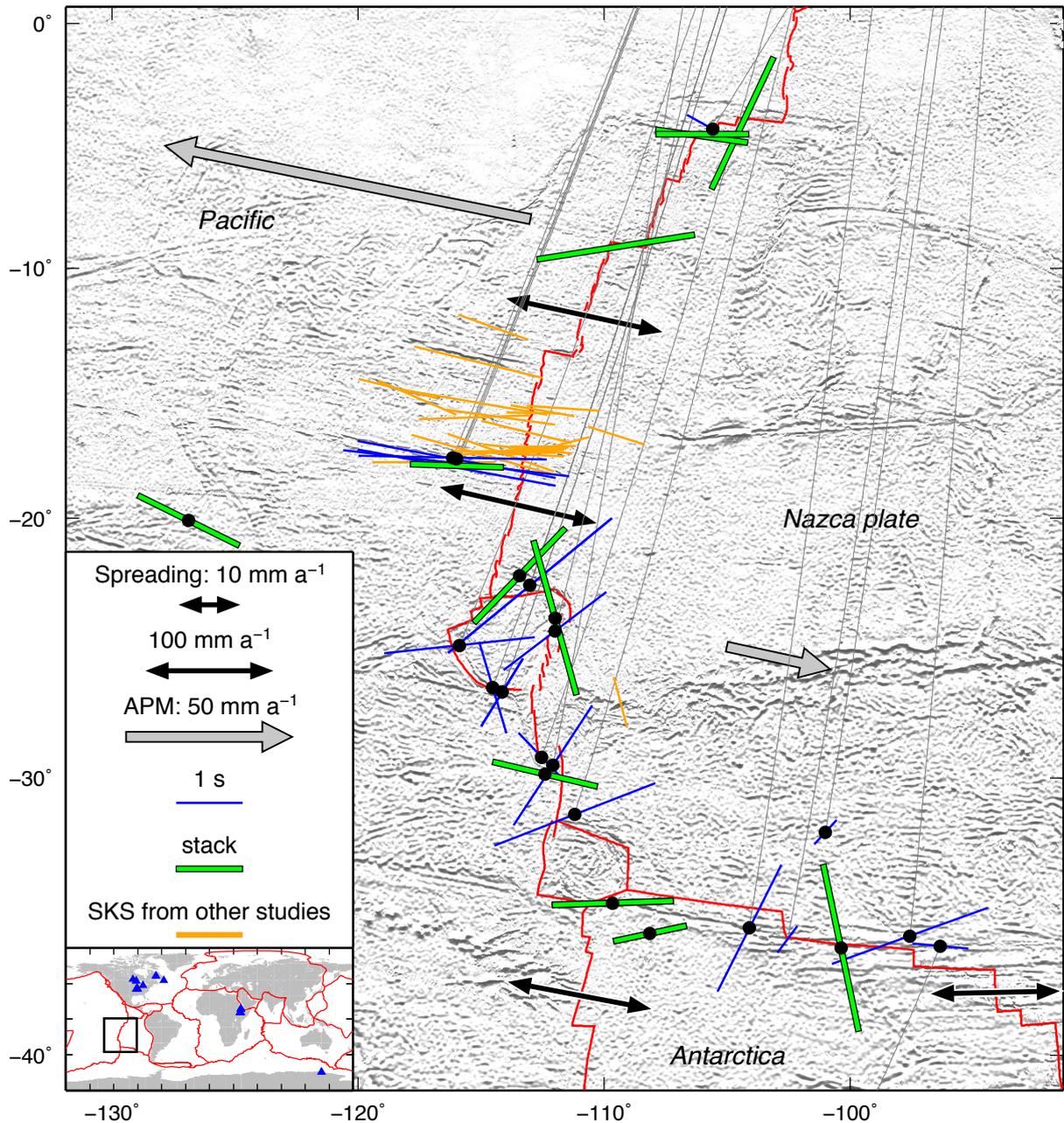


Figure 2: Source-side splitting beneath events on the EPR. Dots show earthquake locations, with bars indicating splitting parameters, where the orientation shows ϕ'' and the length δt , as in the legend. Blue bars are for single measurements; green for stacks. Orange bars show SKS splitting parameters from previous studies (Wolfe and Solomon, 1998; Harmon et al., 2004). Thin grey lines show raypaths to stations (blue triangles, inset map). Shading indicates bathymetry (Smith and Sandwell, 1997), and thin red lines are plate boundaries. Black double-headed arrows show base-10 logarithm of NUVEL-1A full spreading rates at selected locations along the ridge. Grey arrows show the absolute plate motion (APM) in the HS3 reference frame of the NUVEL-1A model (Gripp and Gordon, 2002). The legend indicates spreading and APM rates. Results include stacks from Nowacki et al. (2010).

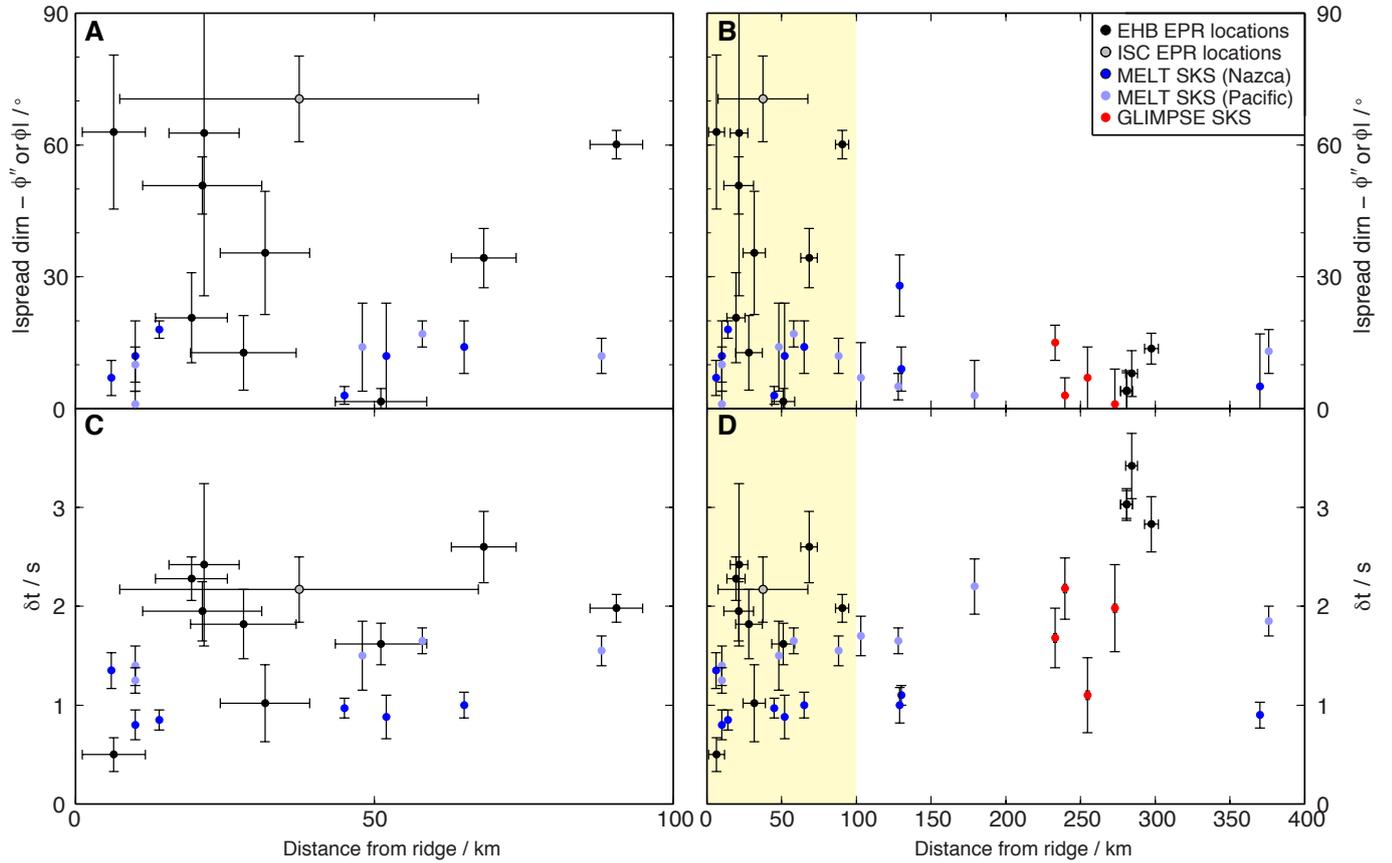


Figure 3: Variation of splitting parameters beneath ‘ridge’ events with distance away from the axial ridge at the EPR. Black and grey circles indicate respectively the EHB and standard ISC locations of the events in this study. Error bars show 95% confidence interval in splitting parameters and stated uncertainty in event locations. Coloured circles indicate MELT (blue) and GLIMPSE (red) SKS splitting parameters as shown in the legend, where MELT stations on the Pacific and Nazca plates are coloured lighter and darker respectively. All GLIMPSE stations are on the Pacific plate. Shaded part of panels on right shows regions shown by panels on left. A) and B) Modulus of difference in angle between ϕ'' or ϕ_{SKS} and the plate spreading direction (DeMets et al., 1994). C) and D) Splitting times for S or SKS.

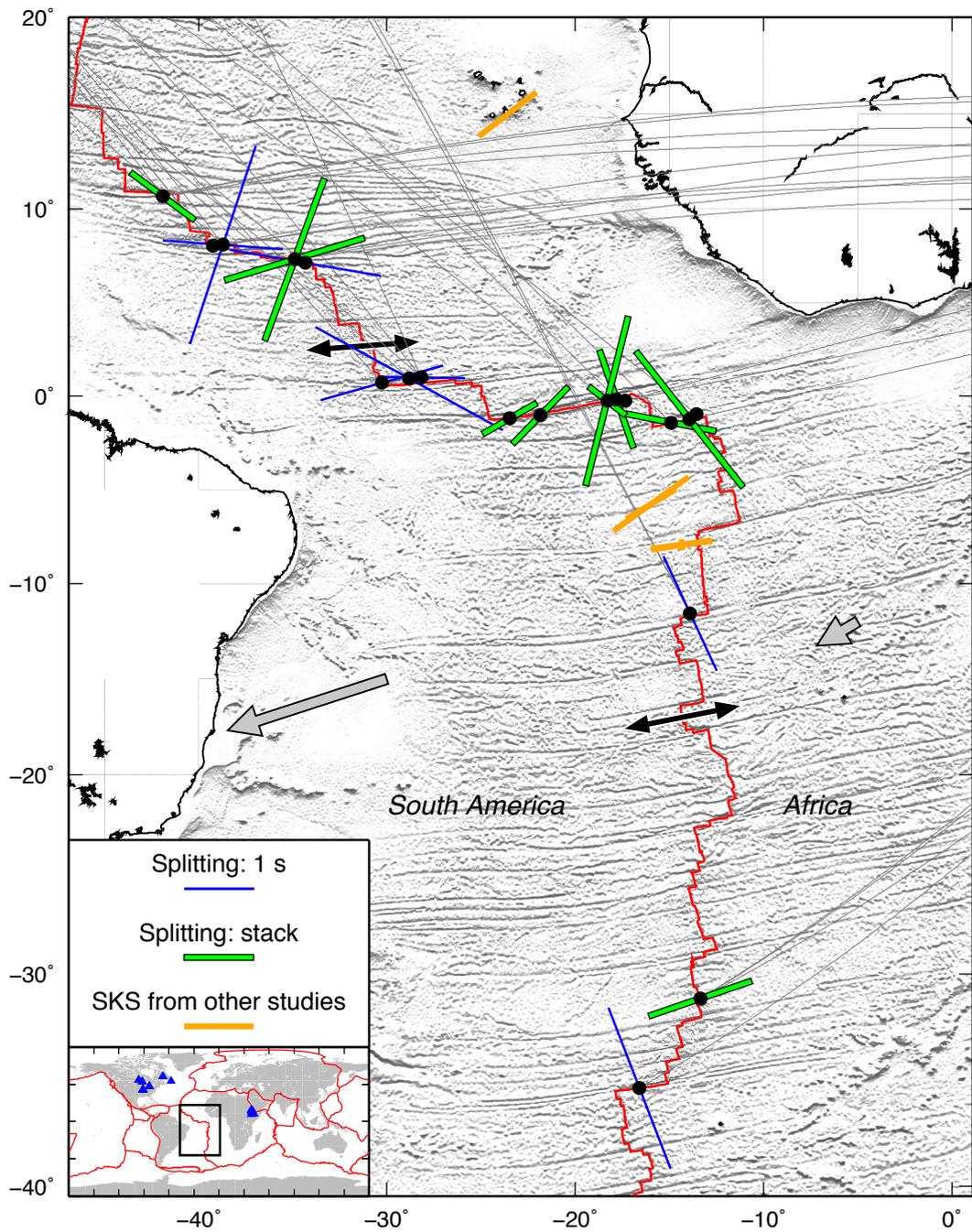


Figure 4: Splitting parameters beneath events on the MAR. Symbols as for Figure 2. Orange bars show SKS results of Wolfe and Silver (1998) and Behn et al. (2004). Note that some events are measured at stations in both North America and Ethiopia, in which case stacks of results for both directions are shown. Spreading rates and APM are of same scale as Figure 2.

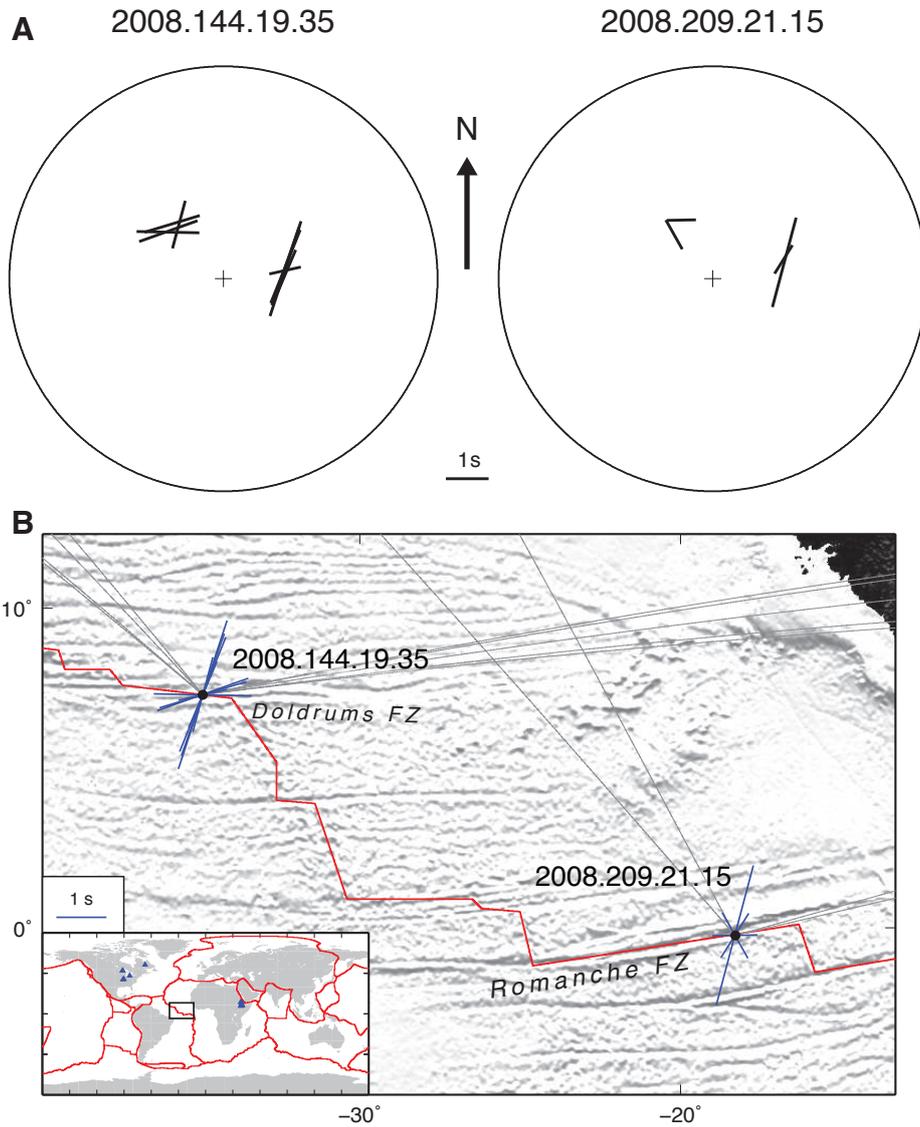


Figure 5: Lower-hemisphere diagrams for splitting parameters measured beneath two events on the MAR. A) Azimuthal and inclination-dependence of splitting parameters shown on equal-area lower-hemisphere projections. Average inclination of downgoing rays in top 150 km of IASP91 (Kennett and Engdahl, 1991) is shown by radial distance (with vertical at the centre). Azimuth corresponds to azimuth at the event. Bar orientation and length corresponds to ϕ'' and δt respectively, as per the scale, centre. The splitting times measured at Ethiopian stations (group on right of hemispheres) are larger for both events, and ϕ'' is also different. B) Location of events and individual splitting measurements shown at earthquake location. Bars correspond to splitting parameters as for previous figures, with delay time indicated by length as per the legend (left). Inset map show location of larger map by thick black box. The raypaths to Ethiopia run along the transform zones, whilst those to North America move away from the transforms.

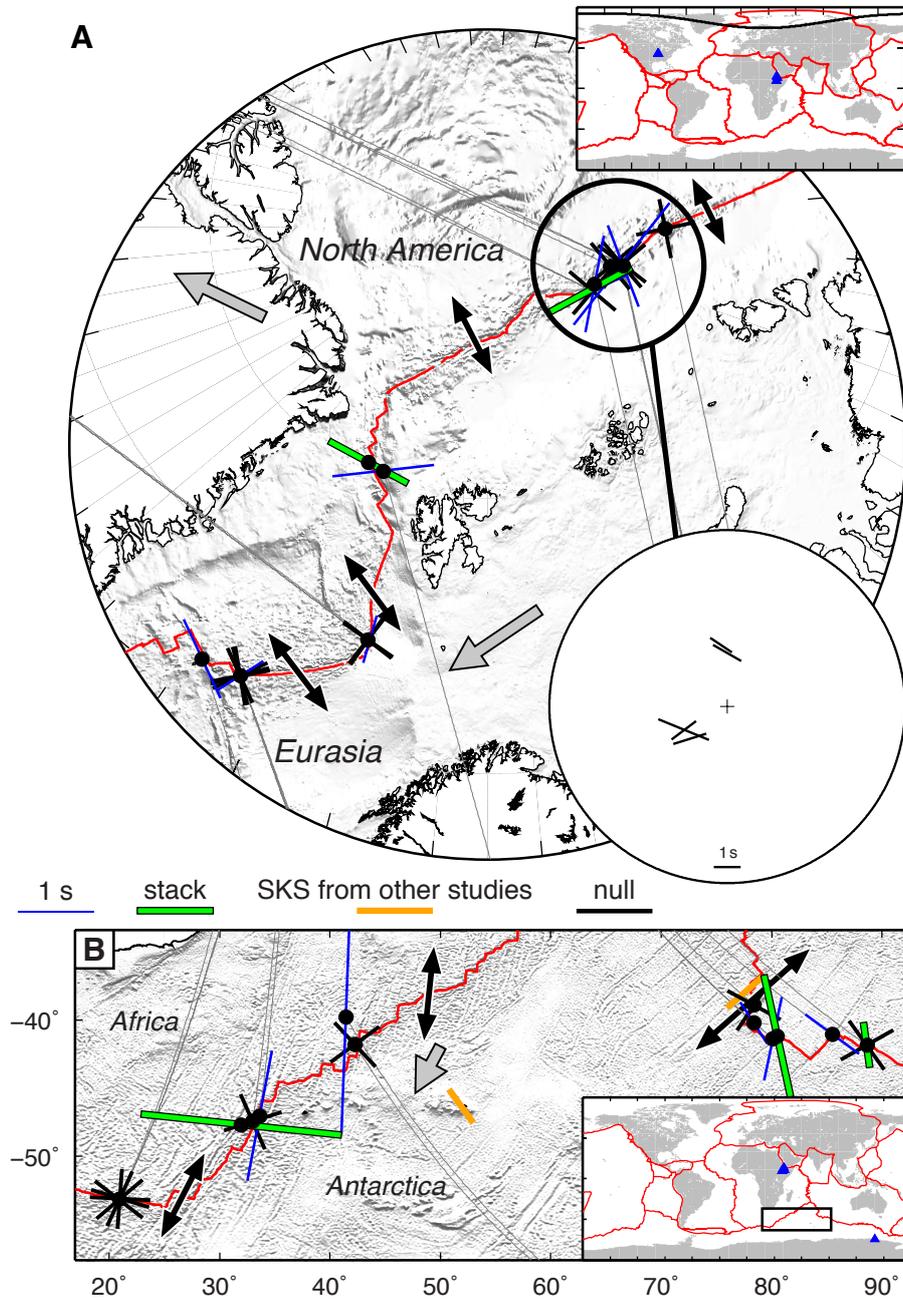


Figure 6: Splitting parameters beneath events on the Gakkel, Southeast and Southwest Indian Ridges. Symbols as for Figure 2, except null results are shown by black crosses with bars parallel to the null directions. A) Results for events on Gakkel Ridge. Thick black circle shows results included in lower hemisphere stereonet, inset lower right. Bars above the centre show measurements made at North American stations; those to the lower left show measurements at Ethiopian stations. Scale indicated at bottom. B) Results for events on Southwest and Southeast Indian Ridges. SKS splitting at CRZF (Base Alfred Faure, Crozet Islands) and AIS (île Nouvelle-Amsterdam; Behn et al., 2004) is shown by the orange bar. APM is less than 10 mm a^{-1} for African plate, and parallel to spreading direction at $\sim 65 \text{ mm a}^{-1}$ for Australian plate (northeast corner, not labelled).

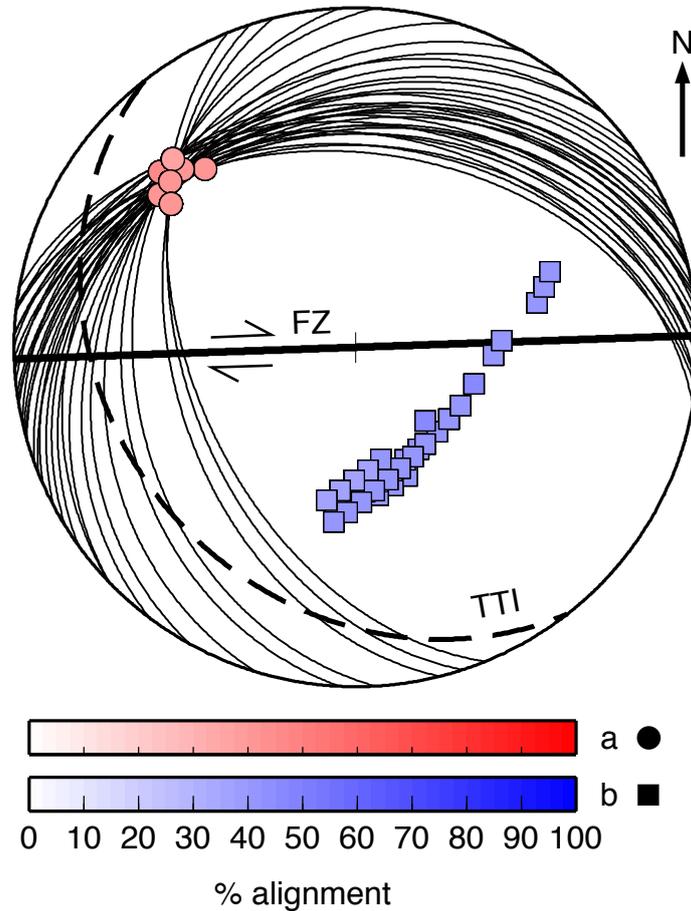


Figure 7: Orientations of olivine a- and b-axes compatible with observations from event 2008.144.19.35 on the Doldrums FZ, MAR. Lower-hemisphere equal area plot shows north upward and the vertical direction out of the page. Red circles (a-axes) and blue squares (b-axes) are shaded per the degree of alignment according to the scale below. Thick black solid line shows approximate strike of FZ and spreading direction, with strike-slip arrows indicating sense of shear. Thick dashed line is best-fitting plane of isotropy from fit of TTI to fast orientations. Thin solid lines are crystallographic slip planes (b-planes) for the case of ‘A-type’ olivine LPO.

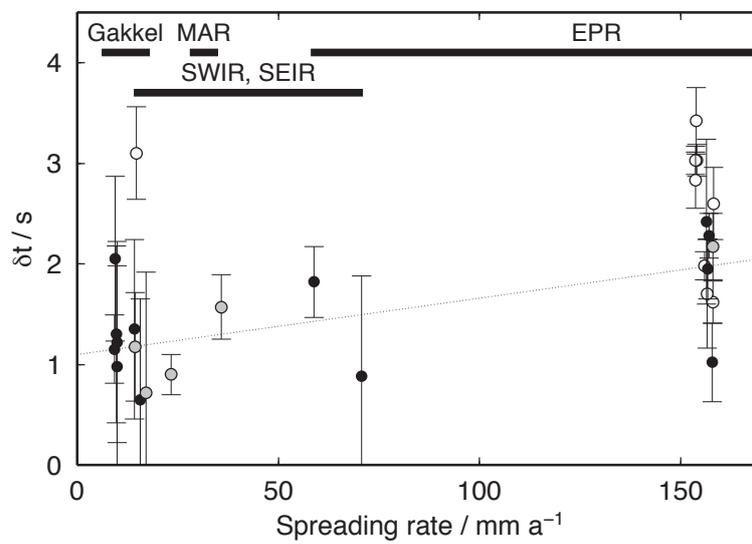


Figure 8: Splitting time versus spreading full-rate beneath all ‘ridge’ events. Filled circles show results less than 50 km from the ridge axis: black circles shows events with EHB locations; grey circles indicate ISC locations. Open circles indicate events >50 km from the axis (all EHB locations). Thick bars show range of spreading rates represented by events beneath each MOR in this study. Weighted linear fit to near-axis (filled circles) data is shown with thin dotted line.