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1 **A cascade of magmatic events during the assembly and eruption**
2 **of a super-sized magma body**

3

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23

24 **Abstract**

25 We use comprehensive geochemical and petrological records from whole-rock samples,
26 crystals, matrix glasses and melt inclusions to derive an integrated picture of the generation,
27 accumulation and evacuation of 530 km³ of crystal-poor rhyolite in the 25.4 ka Oruanui
28 supereruption (New Zealand). New data from plagioclase, orthopyroxene, amphibole, quartz,
29 Fe-Ti oxides, matrix glasses, and plagioclase- and quartz-hosted melt inclusions, in samples
30 spanning different phases of the eruption, are integrated with existing data to build a history
31 of the magma system prior to and during eruption. A thermally and compositionally zoned,
32 parental crystal-rich (mush) body was developed during two periods of intensive
33 crystallisation, 70 and 10-15 kyr before the eruption. The mush top was quartz-bearing and as
34 shallow as ~3.5 km deep, and the roots quartz-free and extending to >10 km depth. Less than
35 600 yr prior to the eruption, extraction of large volumes of ~840 °C low-silica rhyolite melt
36 with some crystal cargo (between 1 and 10%), began from this mush to form a melt-dominant
37 (eruptible) body that eventually extended from 3.5-6 km depth. Crystals from all levels of the
38 mush were entrained into the eruptible magma, as seen in mineral zonation and amphibole
39 model pressures. Rapid translation of crystals from the mush to the eruptible magma is
40 reflected in textural and compositional diversity in crystal cores and melt inclusion
41 compositions, versus uniformity in the outermost rims. Prior to eruption the assembled
42 eruptible magma body was not thermally or compositionally zoned and at temperatures of
43 ~790 °C, reflecting rapid cooling from the ~840 °C low-silica rhyolite feedstock magma. A
44 subordinate but significant volume (3-5 km³) of contrasting tholeiitic and calc-alkaline mafic
45 material was co-erupted with the dominant rhyolite. These mafic clasts host crystals with
46 compositions which demonstrate that there was some limited pre-eruptive physical
47 interaction of mafic magmas with the mush and melt-dominant body. However, the mafic
48 magmas do not appear to have triggered the eruption or controlled magmatic temperatures in
49 the erupted rhyolite. Integration of textural and compositional data from all available crystal
50 types, across all dominant and subordinate magmatic components, allows the history of the
51 Oruanui magma body to be reconstructed over a wide range of temporal scales using multiple
52 techniques. This history spans the tens of millennia required to grow the parental magma
53 system (U-Th disequilibrium dating in zircon), through the centuries and decades required to
54 assemble the eruptible magma body (textural and diffusion modelling in orthopyroxene), to
55 the months, days, hours and minutes over which individual phases of the eruption occurred,
56 identified through field observations tied to diffusion modelling in magnetite, olivine, quartz

57 and feldspar. Tectonic processes, rather than any inherent characteristics of the magmatic
58 system, were a principal factor acting to drive the rapid accumulation of magma and control
59 its release episodically during the eruption. This work highlights the richness of information
60 that can be gained by integrating multiple lines of petrologic evidence into a holistic timeline
61 of field-verifiable processes.

62

63

64 **Keywords**

65 Oruanui eruption, Taupo volcano, rhyolite, magma chamber, magma accumulation, eruption
66 timings

67

68 **Introduction**

69 **The nature and generation of large silicic magma bodies**

70 The study of large silicic eruptions and their parental magmatic systems offers many
71 challenges. One central issue is determining whether the processes and timescales involved in
72 the generation of small volumes of evolved magma are the same as for the accumulation of
73 vast magma bodies that feed supereruptions (i.e. $>10^{15}$ kg or ~ 450 km³ magma: e.g. Self
74 2006). Do large bodies of eruptible magma take longer to be generated than smaller ones (e.g.
75 Smith 1979; Shaw 1985; Reid 2008; de Silva and Gregg 2014), or do they accumulate on
76 similar timescales, driven by processes with faster rates (e.g. Wilson and Charlier 2009;
77 Gualda et al. 2012a; Wotzlaw et al. 2015)? By extension, questions arise as to how such vast
78 volumes of melt-rich magma can accumulate without erupting, what controls the lifetime of
79 eruptible magma bodies and, what triggers eruption onset (e.g. Gregg et al. 2012; Gelman et
80 al. 2013; Caricchi et al. 2014a; Malfait et al. 2014; Myers et al. 2016)?

81 Answers to these questions are sought either by numerical modelling (e.g. Jellinek
82 and DePaolo 2003; Bachmann and Bergantz 2004; Huber et al. 2011, 2012) or geochemical
83 and petrological investigation of eruption products (e.g. Girard and Stix 2010; Gualda et al.
84 2012a); the latter approach having the advantage of directly linking to hard evidence in the
85 field (e.g. Wilson et al. 2006; Wilson and Charlier 2009; Allan et al. 2012, 2013; Bain et al.
86 2013). Petrological studies can be controversial (particularly in the case of the Bishop Tuff),
87 as it is essential to interpretations of petrological data to determine if crystals grew together *in*
88 *situ* over a definable range of intensive variables (Hildreth 1979; Gualda et al. 2012a, b;
89 Evans and Bachmann 2013; Chamberlain et al. 2015), or if they are aggregated from a wide

90 range of sources (e.g. Eichelberger 1978; Gamble et al. 1999; Dungan and Davidson 2004;
91 Davidson et al. 2005). Furthermore, it is desirable for multiple co-erupted crystal phases to be
92 considered together (e.g. Hildreth 1979; Girard and Stix 2010; Chamberlain et al. 2014a,
93 2015). Although useful information can be recovered from studies of single phases (e.g.
94 Bishop Tuff quartz – Peppard et al. 2001; Wark et al. 2007; Thomas et al. 2010; Gualda et al.
95 2012a, Pamukcu et al. 2012, 2016; Gualda and Sutton 2016) there is a risk of introducing
96 unintended systematic bias.

97

98 **The crystal archive in crystal-poor ignimbrites**

99 The crystal mush model (e.g. Brophy 1991; Hildreth 2004; Bachmann and Bergantz 2004,
100 2008; Bachmann and Huber 2016) provides a widely utilised framework with which to
101 explore large crustal bodies of silicic magma. The mush itself has generally been envisaged
102 as a largely un-eruptible body of viscous material (Smith 1979), and numerical modelling has
103 focussed on determining how mush bodies might ‘defrost’ and/or remobilise (Burgisser and
104 Bergantz 2011; Huber et al. 2010, 2011). Mafic magma is often implicated in supplying the
105 heat and volatiles required to mobilise the mush to drive voluminous eruptions of crystal-rich
106 magma (e.g. Parmigiani et al. 2014). However, the origins and rates of accumulation of
107 crystal-poor rhyolites, which are generally thought have been extracted from mush zones, are
108 potentially widely diverse (Bachmann and Bergantz 2008; Bachmann and Huber 2016 for
109 overviews).

110 Two considerations from the crystal mush model pertain to the origins of crystal-poor
111 rhyolites. (1) Large volumes of evolved melt may be available for extraction from a mush
112 zone at any given point in time. (2) If crystal-poor rhyolite does originate from a crystal-rich
113 reservoir, crystal-melt segregation on a massive scale is required at some stage prior to
114 eruption. It is not yet known if such large scale segregation leaves a tangible record of this
115 process in the small fraction of crystals that are actually erupted and, if so, how such records
116 could be used to temporally and physically constrain these events. A further complication is
117 that any signals recorded in crystals of events related to melt body generation must be
118 deconvoluted from subsequent growth in the eruptible magma. This last point raises
119 questions of which crystals are phenocrystic (grown in the melt with which they were
120 erupted), antecrystic (derived from consanguineous parental sources), or xenocrystic (foreign
121 to the magma in which they are found) (Bacon and Lowenstern 2005; Charlier et al. 2005). It
122 is also possible that a given crystal may contain domains attributable to all three origins (e.g.
123 Charlier et al. 2008).

124 In this paper we document the timing and mechanisms of pre- and syn-eruptive
125 magmatic processes for the Oruanui eruption (Taupo, New Zealand). We integrate new and
126 published major- and trace-element analyses of whole-rock samples, matrix glasses, crystals,
127 and quartz- and plagioclase-hosted melt inclusions (see Table S.1), for both the
128 volumetrically dominant rhyolites and subordinate co-erupted mafic magmas. We explore the
129 textural and compositional relationships within and between crystal phases, to provide a
130 summary of the magmatic processes and their associated timescales that operated during the
131 establishment, assembly and eruption of the super-sized Oruanui magma body.

132

133 **The Oruanui eruption**

134 **Eruption summary**

135 The Oruanui eruption at $25,360 \pm 160$ (2 sd) cal yr BP (Vandergoes et al. 2013) from Taupo
136 volcano in the central North Island of New Zealand (Fig. 1) is the youngest supereruption on
137 Earth, discharging ~ 530 km³ of moderate- to high-silica rhyolite magma with 3 to 13 wt %
138 crystals (Wilson 2001; Wilson et al. 2006). Ten eruptive phases are identified, based on
139 bedding and grading characteristics in the fall deposits. The eruption was episodic, with
140 several time breaks varying in duration from probably only hours (i.e. long enough for the
141 eruption plume to have dispersed and ash to wholly settle out before the next phase began), to
142 a period of months between phases 1 and 2 during which time reworking and bioturbation of
143 the earlier ash occurred (Wilson 2001). Extensive interaction of magma with lake water
144 during the eruption caused rapid chilling, extensive fragmentation and produced extremely
145 widely dispersed, fine-grained fall deposits and a large, non-welded ignimbrite (Self and
146 Sparks 1978; Self 1983; Wilson 2001; Van Eaton et al. 2012; Van Eaton and Wilson 2013).
147 The eruption products are fresh, glassy and entirely non-welded, although pumices large
148 enough for detailed study are largely restricted to the ignimbrite. The interbedded nature of
149 the fall deposits and ignimbrite, and spikes in the abundance of juvenile mafic material, allow
150 the eruption phase boundaries to be linked between the fall and flow deposits (Wilson 2001),
151 such that clasts sampled from the ignimbrite can be placed within a time-stratigraphic
152 context.

153

154 **Juvenile products**

155 Juvenile products of the Oruanui eruption fall into seven categories (five felsic, two mafic;
156 Table 1). By far the most voluminous, comprising ~98% of the total erupted volume, is
157 highly vesicular white pumices of high-SiO₂ rhyolite (HSR hereafter), defined as having
158 whole-rock SiO₂ >74 wt%. HSR clasts characteristically contain 3-13 % crystals, have quartz
159 forming 9-21 % of the crystal fraction and mostly have more orthopyroxene than amphibole
160 (opx 45-82% to amph 55-18%: Sutton 1995 in Wilson et al. 2006). Sparse (~0.5 %), white,
161 low-SiO₂ rhyolite pumices (LSR, whole-rock compositions <74 wt% SiO₂) lack quartz and
162 have more amphibole than orthopyroxene (opx 29-42% to amph 71-58%; Ibid.). Wilson et al.
163 (2006) originally considered HSR and LSR to represent a compositional continuum, but
164 Allan et al. (2013) inferred that they were separate entities (see also Results section below)
165 that only encountered each other syn-eruptively. We focus on these two of the rhyolite types
166 in this paper.

167 Three additional, volumetrically minor (<< 1 %) but genetically significant Oruanui
168 rhyolitic components have previously been identified, and are summarized here for
169 completeness. The first is pumices showing a white/pale grey streaky nature, and which fall
170 along a mixing trend between the HSR and more mafic magma (Sutton 1995; Wilson et al.
171 2006, figure 18). The second is pumices that show enrichment in MgO for a given SiO₂, and
172 have been similarly attributed to mixing of HSR and more mafic magma but the clasts show
173 no signs of mingling in hand specimen (Sutton 1995; Wilson et al. 2006, figure 18). The third
174 is a denser, more crystal-rich, biotite-bearing rhyolite pumice, identified as a 'foreign'
175 magma sourced from the nearby but wholly chemically and isotopically independent 'NE
176 dome' system (Sutton et al. 1995; Fig. 1). This biotite-bearing magma is interpreted to have
177 been fed laterally by diking from its parental magma system into the Oruanui conduit at the
178 onset of eruption and throughout phases 1 and 2 (Allan et al. 2012), but it has not been
179 identified in later eruption products.

180 Clasts of juvenile mafic magma from two physically and chemically distinct lineages
181 also occur within the Oruanui deposits: an olivine-bearing calc-alkaline group and a tholeiitic
182 group in which olivine is very rare. Based on detailed componentry of ash-grade material
183 from the eruption deposits (Wilson 2001), the two mafic lineages together are estimated to
184 represent 3-5 km³ of magma. The relative proportions of the calc-alkaline to tholeiitic end-
185 members are not constrained, although subequal numbers of clasts from each type have been
186 collected. The mafic clasts are inferred to be juvenile on the basis of cauliform or crenulated
187 margins, indicating that they were hot and plastically deforming until quenched on eruption
188 (Wilson 2001; Rooyakkers 2015 and manuscript in preparation). The two mafic groups are

189 compositionally distinguished on plots of whole-rock SiO₂ versus FeO_T/MgO with both calc-
190 alkaline and tholeiitic groups being almost entirely contained within the appropriate fields as
191 defined by Miyashiro (1974).

192

193 **Samples studied**

194 For this study, we collected HSR pumices from early-erupted fall deposits (phases 1 and 2);
195 and from early (phases 2 and 3), mid- (phase 7) and late-erupted (phase 10) ignimbrite within
196 the stratigraphic framework of Wilson (2001). Wilson et al. (2006) showed that, although
197 there was notable compositional variability within the Oruanui rhyolites, the deposits
198 themselves were not systematically zoned with respect to stratigraphic height (Fig. S.1). The
199 melt-dominant magma body, accordingly, was inferred to have undergone thorough mixing at
200 some stage prior to eruption (Wilson et al. 2006). The LSR pumices were sampled from
201 phase 3 ignimbrite at a single location (at the Hinemaiaia C Dam in Lake Taupo Forest:
202 locality 1086 of Wilson 2001), where LSR material was previously observed to be most
203 abundant.

204

205 **Analytical techniques**

206 **Whole rock analytical methods**

207 To expand the data set in Wilson et al. (2006), single clasts of rhyolitic (Allan et al. 2012,
208 2013, this study) and mafic compositions (this study) were analysed for their bulk chemistries
209 by X-ray fluorescence (XRF) and solution ICP-MS techniques. Clasts were cleaned to
210 remove any adhering matrix, soaked overnight in de-ionised water, and placed in an oven at
211 110 °C until dry. The mafic clasts, which commonly contained thin adhering rinds of white
212 rhyolite, were coarsely crushed using an agate mortar and pestle, so that fragments of rhyolite
213 could be picked out prior to powdering to ensure characterisation of the mafic end-member.
214 The clasts were then reduced to powder using an agate mill. Whole-rock major element oxide
215 compositions were determined by XRF by either Spectrachem Analytical (CRL Energy LTD,
216 Wellington, New Zealand) or the late Dr John Watson at the Open University, UK.
217 International rock standards BHVO-2 (USGS, Hawaiian basalt), BCR-2 (USGS, Columbia
218 River basalt), and JR1 (GSJ, rhyolite) were analysed during the same sessions at
219 Spectrachem, and WSE (Whin Sill dolerite) and OU-3 (Nanhoron microgranite) at the Open
220 University. Data for the first three standards are summarised in the supplementary material of
221 Allan et al. (2013), and for the two Open University standards in the supplementary material

222 (Table S.2) for this paper. Overall, relative differences from accepted values are BHVO-2: <2
223 % except for MnO (-11%) and P₂O₅ (11%); BCR-2: <2% except for MnO (-10%), Na₂O (-
224 2.5%) and P₂O₅ (5.7%); JR-1: <2 % except for TiO₂ (-3.7%), MnO (-11.5%), MgO (23%)
225 and K₂O (5.4%); WSE: <1%; OU-3: < 1% except for MnO (1.7%), CaO (4.5%) and Na₂O
226 (1.6%). Overall relative 2 sd precisions are better than 1 % for SiO₂, Al₂O₃ and Fe₂O₃, 1-2%
227 for MnO, MgO, CaO, Na₂O and P₂O₅, and 3-5% for K₂O and TiO₂.

228 Whole-rock trace element compositions were determined using an Agilent 7500CS
229 ICP-MS and Thermo Scientific Element2 sector-field ICPMS at Victoria University of
230 Wellington. Sample and standard powders were digested using conventional HF + HNO₃
231 methods. Calculation of trace element concentrations were determined from raw count rates,
232 after correction for off-peak zeros, by reference to the known elemental concentrations of the
233 BHVO-2 whole-rock standard, and utilising ⁴³Ca as the internal standard with the CaO
234 content of samples previously determined to ~ 1% by XRF. Standards BCR-2, BHVO-2 and
235 JR-1 were analysed under identical conditions and during the same analytical sessions as the
236 Oruanui rhyolitic (Allan et al. 2012) and mafic samples (this study). Data for these standards
237 are available in the online supplementary information of Allan et al. (2012). Overall, relative
238 differences from accepted values are BHVO-2: <1% except for Zn (-5.7%), Tm (1.1%) and
239 Lu (1.4%); BCR-2: <5%, except for Cr (-16%) and Mo (25%); JR-1: <5% except for Yb
240 (5.2%), Er (6.6%), Zr (-7.7%), Ti (-8.6%), Gd (8.6%), Ho (9.2%), Sc (14%), V (-29%), Cr (-
241 40%), Mo (-75%), Zn (78%) and Cu (150%). Average % 2 sd precisions based on analysis of
242 whole rock standards are < 3% (MgO, TiO₂, V, Ga, Rb, Sr, Y, Zr, Nb, La); 3 to 5% (MnO,
243 Ba, Ce, Pr, Nd, Sm, Eu, Gd, Tb, Dy, Ho, Er, Tm, Yb, Lu, Hf, Th, U); 5 to 10% (Cu, Pb); 10
244 to 15% (Sc, Mo, Cs); 15 to 20% (Zn).

245

246 **In-situ analytical methods**

247 Crystals were liberated from the host felsic clasts by gently crushing in an agate mortar and
248 pestle, then hand-picked from the appropriate size fraction and mounted in epoxy resin
249 blocks. Analysis of crystals from the mafic clasts was approached in two ways. (1) Crystals
250 were handpicked from the 0.5 to 1.0 mm size fraction of crushed material and mounted in
251 epoxy resin blocks. (2) Clasts were thin sectioned so that mineral phases at sizes <0.5 mm
252 could be imaged and analysed within their host groundmasses. All *in situ* analyses of glass
253 and mineral phases considered here (Allan et al. 2012, 2013, and new data) were undertaken
254 using a JEOL JXA-8230 electron probe microanalyser (EPMA), and a New Wave 193 nm
255 laser ablation system coupled to an Agilent 7500CS ICP-MS at Victoria University of

256 Wellington. Glass analyses by EPMA were conducted at an accelerating voltage of 15 kV
257 and current of 8 nA with the beam de-focused to 10 μm and count times for Na reduced to
258 minimise the effects of alkali-loss. Mineral analyses by EPMA were conducted at 12 nA
259 under a focused electron beam ($\sim 2 \mu\text{m}$). Prior to *in-situ* analysis of the HSR and LSR mineral
260 phases, high resolution back-scattered electron (BSE) images of each crystal were obtained.
261 These images were used to note crystal textures and zonation features and infer, where
262 possible, a crystal stratigraphy to ensure that at least the inferred oldest and youngest, and
263 most- and least-evolved, domains of each crystal were targeted for analysis. Standardisation
264 was determined using a series of natural and synthetic standards, chosen to match as closely
265 as possible the matrix composition of the material being analysed. Secondary mineral
266 standards (Engels amphibole, Johnstown hypersthene, NMNH 115900 plagioclase, Kakanui
267 augite, Springwater olivine) and glass standards (ATHO-G, VG-A99, VG-568) were
268 analysed throughout the analytical sessions to monitor signal stability and inform estimates of
269 precision and accuracy.

270 Trace element analyses by LA-ICP-MS included a minor isotope of a major element
271 (typically ^{29}Si or ^{43}Ca), previously determined by EPMA for use as an internal standard
272 (Pearce et al. 1996). For analysis of the ferromagnesian mineral phases, the USGS basaltic
273 glass standard, BHVO-2G, was used as the calibration standard, whereas NIST612 was used
274 for analysis of plagioclase and melt inclusions. The ICP-MS was tuned by adjusting the
275 positions of the torch, lenses, and the inflow of ultra-pure He while rastering at 2 $\mu\text{m}/\text{sec}$
276 across the calibration standard. Data for samples and standards were collected under identical
277 conditions during 60 s acquisitions (plus 60 s for backgrounds), under a static laser beam of
278 25 to 35 μm diameter, and pulsed at 5 Hz. Analytical data were reduced using the Iolite
279 software package using the trace element (Internal Standard) data reduction scheme
280 (Hellstrom et al. 2008). The raw time-resolved signal from each analysis was scrutinised to
281 identify and remove any data compromised by the accidental ablation of contaminant (non-
282 targeted) phases, before final trace element contents were calculated. Estimates of precision
283 and accuracy of LA-ICP-MS data were derived from repeat analyses of the glass standards
284 BHVO-2 (USGS Hawaiian basalt), BCR-2G (USGS, Columbia River basalt) and ATHO-G
285 (MPI-DING, Icelandic rhyolite) and are given in supplementary material Table S.3. Overall,
286 relative differences from accepted values are BHVO-2G: <5% except for Sc (-5.7%), Ni
287 (5.6%), Cu (36%), Zn (-9.0%), Tm (-7.1%) and W (-9.0%); BCR-2G: <5%, except for Sc
288 (5.4%), Mn (6.6%), Ni (-7.4%), Cu (-21%), Zn (11%) and W (18%); ATHO-G: <10%

289 except for Mg (17%), Sc (140%), Ti (13%), V (-12%), Cr (-51%), Cu (-15%), Ga (-19%) and
290 Cu (17%). Average % 2 sd precisions based on analysis of glass standards are

- 291 • BHVO-2G: <10% (Li, Mg, Sc, Ti, V, Mn, Co, Zn, Ga, Rb, Sr, Y, Zr, Nb, Ba, La, Ce,
292 Pr, Nd, Eu, Tb, Dy, Hf, Ta, Th), 10-20% (Ni, Sm, Gd, Ho, Er, Tm, Yb, Pb, U) and
293 >20% (Cu, Cs, Lu, W);
- 294 • BCR-2G: <10% (Li, Mg, Sc, Ti, V, Mn, Co, Zn, Ga, Rb, Sr, Y, Zr, Nb, Ba, La, Ce,
295 Pr, Nd, Eu, Tb, Dy, Ho, Er, Hf, Th, U), 10-20% (Cr, Ni, Sm, Gd, Tm, Yb, Lu, Ta, Pb)
296 and >20% (Cu, Cs, W);
- 297 • ATHO-G: <10% (Li, Mg, Mn, Ga, Rb, Sr, Y, Zr, Nb, Ba, La, Ce, Pr, Nd, Tb, Dy, Ho,
298 Er, Yb, Hf, Pb, Th, U), 10-20% (CaO, Sc, Ti, V, Zn, Sm, Eu, Gd, Tm, Lu, Ta, W) and
299 >20% (B, Cr, Co, Ni, Cu, Cs).

300

301 **Results**

302 **Oruanui rhyolite compositions**

303 *Whole rock and glass chemistry*

304 Whole-rock rhyolite compositional data are presented in Wilson et al. (2006) and Allan et al.
305 (2013), and a representative suite of the latter data is given in Table 2. A compilation of all
306 existing and new data for the juvenile mafic clasts is in Electronic Appendix 1 and a
307 representative suite of new data given in Table 3. HSR pumices have relatively uniform glass
308 compositions, but all LSR clasts have bimodal glass compositions (Electronic Appendix 2;
309 Fig. 2a). The more evolved of these is compositionally identical to the HSR matrix glass, and
310 glass from the widespread Oruanui fall deposits (Lowe et al. 2008; Allan et al. 2008;
311 Vandergoes et al. 2013; Table S.4). The presence of the two, non-hybridized glass
312 populations has been interpreted to reflect latest-stage entrainment of HSR melt into LSR
313 magma, immediately prior to quenching (Allan et al. 2013). Quartz-hosted melt inclusion
314 data presented here (Electronic Appendix 2) confirm observations by Liu et al. (2006) and
315 Bégué et al. (2015) that inclusions in HSR samples are more evolved than the groundmass
316 glass (Fig. 2b). Although there is some overlap, these data indicate that for at least part of
317 their history most Oruanui quartz crystals grew from (and trapped) melts that were more
318 evolved than the carrier melts in which they were erupted (i.e., quartz is largely xenocrystic
319 relative to the groundmass pumice glass). The LSR glass data from both plagioclase-hosted
320 melt inclusions and groundmass glass forms two trends when Rb/Sr is plotted against Ti

321 content., one with a steeper, higher Ti array, and the other a shallower, lower Ti array (Fig.
322 2d) along trends towards less-evolved compositions.

323

324 *Textural linkages between crystal phases in the HSR*

325 Backscatter electron (BSE) images of amphibole, pyroxene, plagioclase and Fe-Ti
326 oxides, and cathodoluminescence [CL] images of quartz, reveal textural and zonation features
327 in the main crystal phases, particularly in those of the volumetrically dominant HSR. Allan et
328 al. (2013) noted that ~90% of HSR orthopyroxene showed distinctive core-rim relationships,
329 in which a more Mg- and inclusion-rich core was surrounded by 40-400 μm thick inclusion-
330 poor rim that was normally zoned with a consistent $\text{En}_{46\pm 2}$ composition in its outermost parts
331 (Fig. 3a). Zonation in individual crystals varies due to combined effects of dissolution and
332 recrystallisation, unequal kinetic/growth zonation and overprinting from Fe-Mg interdiffusion
333 (Allan et al. 2013). The remaining ~10% of HSR orthopyroxenes tended to be inclusion-poor
334 and lacked any significant core-rim zonation.

335 BSE images of individual HSR plagioclase crystals show that they record a
336 complementary suite of textures to the orthopyroxenes. In total, 89% of the HSR plagioclase
337 crystals ($n = 168$) also have prominent core-rim zonation (Fig. 3b). The cores are typically
338 truncated by at least one episode of resorption and are surrounded by subhedral to euhedral
339 overgrowths that are normally zoned towards an outermost composition of $\text{An}_{38\pm 2}$. Of these
340 zoned crystals, 48% have cores that are patchy/sieve textured, and 41% show significant
341 resorption but no evidence of sieve texturing (Fig. 3b); and they can be further subdivided
342 into those with high An ($>\text{An}_{70}$) cores, and those with prominent blue-grey cores (e.g.
343 Charlier et al. 2008). The remaining 11% of HSR plagioclase are either nominally unzoned
344 (An contents varying by <3 mol%, and accounting for ~7% of all HSR plagioclase), or
345 showed normal, oscillatory zonation with no distinction between core- or rim-zones.

346 The HSR quartz textures in CL images indicate a history that is not directly
347 comparable to those of the plagioclase and orthopyroxene (Fig. 4). Brightness of quartz under
348 CL scales with Ti concentration (Peppard et al. 2001; Wark et al. 2007; Matthews et al.
349 2012). The causes of varying or oscillating concentrations of Ti in magmatic quartz are
350 debated around fluctuations in temperature (e.g. Wark and Watson 2006; Wark et al. 2007),
351 pressure (Thomas et al. 2010, 2015; Huang and Audétat 2012), and/or the combined
352 influences of temperature, pressure and melt compositions (Wilson et al. 2012). Regardless
353 of the ultimate controls on CL ($\approx\text{Ti}$) zonation in quartz, the Oruanui HSR quartz crystals do

354 not follow an overarching zonation pattern. Some quartz crystals have darker cores and
355 brighter rims, and others show the opposite. Many crystals are multiply-zoned with many
356 bright and dark bands, yet some are devoid of any CL zonation at all. In direct contrast to the
357 plagioclase and orthopyroxene, the outermost quartz rims do not appear to converge on a
358 consistent zonation pattern or composition. This is consistent, however, with the analytical
359 data in Liu et al. (2006) and Wilson et al. (2012), where quartz rim analyses show roughly the
360 same mean (~85 ppm) and as wide a range of Ti concentrations (~50 to 120 ppm) as the
361 dataset as a whole, implying that mixing in of quartz crystals from a variety of sources
362 occurred.

363

364 *Compositional characteristics of crystals in the HSR and LSR*

365 A new suite of mineral analyses has been undertaken for this work to supplement the data in
366 Wilson et al. (2006) and provide the basis for our revised interpretations. Full data sets are
367 presented in Electronic Appendices.

368 *Plagioclase.* HSR plagioclase crystals (Electronic Appendix 3) have a wide range of
369 anorthite contents from An₂₆ to An₈₈, mostly reflecting diverse crystal core compositions
370 (Fig. 5a). The majority of crystal rims, however, converge on a narrow range of An_{38±2}. The
371 majority of crystal cores and interiors have compositions that are less evolved (57% are
372 An_{>42}), and a minority that are more evolved (12% are An₂₆₋₃₅) than the rims. Texturally
373 homogeneous grains (unzoned in BSE imagery) cluster in their composition around those of
374 the rim compositions of the other grains. The LSR plagioclase crystals (Fig. 5b) exhibit a
375 more restricted range of compositions, the majority (94%) ranging between An₃₈ and An₆₀.
376 The outermost rims of LSR plagioclase are relatively restricted in composition with 91 % of
377 analyses in the range An_{44±6}. The LSR plagioclases show steep positive or negative trends
378 between An content and Sr (Fig. 5c) and Ba (Fig. 5d). The same parameters for the HSR
379 crystals show two limiting trends in the cores and interiors data that converge around the
380 compositions of the outermost rims.

381 *Orthopyroxene.* Major and trace element compositional data for both HSR and LSR
382 orthopyroxenes (after Allan et al. 2013: Electronic Appendix 4) show similar ranges (Fig. 5e-
383 h), with the biggest difference being that the outermost HSR rims grains show a very
384 restricted compositional range (95% are En_{46±2}), whereas the outermost LSR rims are
385 generally more variable. HSR orthopyroxene rims have very uniform Al contents, whereas
386 HSR cores and interiors have a significant tail-off to higher Al values. The LSR

387 orthopyroxenes show a similar compositional range to the HSR crystals, but in general a
388 greater proportion of them are in the higher Al range. Ti and Al trend positively with En
389 content, whereas Mn and Zn are anti-correlated with Al (Fig. 5e-h). Overall, there are no
390 major compositional differences between the fields of HSR and LSR orthopyroxenes.

391 *Amphibole.* The majority of HSR outermost rim compositions (Electronic Appendices
392 5 and 6) are tightly clustered (47 ± 1 wt% SiO₂, 1.4 ± 0.2 wt% TiO₂, 7 ± 1 wt% Al₂O₃, 12 ± 1
393 wt% MgO) (Fig. 6). The data from HSR cores plus interiors and LSR amphiboles overlap this
394 range but also extend to higher concentrations of TiO₂ and MgO (Allan et al. 2013). The
395 amphibole Eu/Eu* ratio was used by Allan et al. (2013) as a first-order proxy for melt
396 composition, with higher Eu/Eu* values indicating growth from lesser evolved melts and vice
397 versa. Allan et al. (2013, their figure 3) reported that concentrations of Mn and Zn show an
398 initial increase with decreasing Eu/Eu* but then their abundance decreases below Eu/Eu* of
399 ~0.45. Similar noteworthy inflections in the abundances of Eu (going from static to
400 decreasing abundances: Fig. 6b), and Ni and Co (going from static to increasing abundances:
401 Fig. 6c,d) are also evident around this same hinge point. Eu/Eu* ratios of the amphiboles are
402 linearly and positively correlated (R^2 of 0.83: Allan et al. 2013) with the apparent pressures
403 of amphibole crystallisation derived from the calibration of Ridolfi et al. (2010). The
404 abundances of elements such as Mn, Eu, Ni and Co when plotted against model pressures
405 suggest that changes in amphibole chemistry occurred once the amphiboles crystallised at
406 model pressures of <140-160 MPa, equivalent to depths of ~6 km.

407

408 *Intensive variables*

409 Estimates of Oruanui magma intensive variables and water contents have been made through
410 a variety of measurements and models. The results of preferred thermometric models, i.e.
411 those that gave realistic, consistent and reproducible temperature estimates, are summarised
412 in Table 4 and Fig. S.2. Analytical data for Fe-Ti oxides are given in Electronic Appendix 7.
413 Application of the Ghiorso and Evans (2008) model (using the EPMA correction factors from
414 Evans et al. 2006) to equilibrium pairs (after Bacon and Hirschmann 1988) yielded mean
415 temperatures of 790 °C and 840 °C for the HSR and LSR magmas, respectively. The same
416 Fe-Ti oxide pairings applied to the Sauerzapf et al. (2008) model yield temperatures ~15 to
417 25 °C higher. Despite the offset between the two Fe-Ti oxide models, both indicate a ~40 to
418 50 °C difference between the final pre-eruptive temperatures of the HSR and LSR magmas.

419 The compositions of outermost amphibole rims for the HSR and LSR magmas, using
420 the Si* parameter of Ridolfi et al. (2010, equation 1), yielded mean temperatures of 784 °C
421 and 843 °C, respectively. However, both rhyolite groups are skewed by a subordinate
422 component of crystal rims that tail off to higher temperatures. Prominent model temperature
423 modes centre on ~ 770-780 °C (HSR rims) and ~830 °C (LSR rims: Table 4; Fig. S.2). The
424 plagioclase-melt and orthopyroxene-melt thermometers of Putirka (2008, equations 24a and
425 28a, respectively), were applied to the compositions of outermost crystal rims (Allan et al.
426 2013) and matrix glass. Plagioclase-melt and orthopyroxene-melt thermometry yielded mean
427 estimates of 787 °C and 785 °C, respectively, for HSR samples, and mean estimates of 831
428 °C and 830 °C, respectively, for LSR samples.

429 Estimates of the final pressures for storage and crystallisation of the Oruanui rhyolites
430 were determined using the total-Al based formulation of Ridolfi et al. (2010; their equation 4)
431 applied to outermost amphibole rim compositions (Fig. S.3). Based on these compositions
432 only, HSR samples yielded a mean model pressure of 113 MPa (range 88 to 185 MPa), with a
433 prominent mode between 90 and 120 MPa. The mean model pressure for LSR amphibole
434 rims was 174 MPa (range 114 to 243 MPa). As with the HSR amphiboles, the mode of the
435 estimates is slightly lower than the mean. These amphibole-derived pressures for the HSR are
436 consistent with those estimated from concentrations of H₂O and CO₂ in trapped quartz-hosted
437 melt inclusions (Liu et al. 2006) which have a mean pressure of 138 MPa (range 94-188
438 MPa). Broadly similar pressure estimates were also presented by Bégué et al. (2014), using
439 rhyolite-MELTS compositions from Oruanui [HSR] quartz-hosted melt inclusions.

440 Estimates of the final oxygen fugacity for the HSR and LSR magmas were calculated
441 using amphibole rim compositions (Ridolfi et al. 2010; their equation 2), and Fe-Ti oxide
442 equilibrium pairs (Ghiorso and Evans 2008; Sauerzapf et al. 2008) (Table 5). Use of the
443 Ghiorso and Evans (2008) model for Fe-Ti oxides in HSR clasts yields a mean estimate of -
444 0.2 NNO, compared to a mean of +0.47 NNO when the Sauerzapf et al. (2008) model is used.
445 The Ridolfi et al. (2010) formulation applied to amphibole rims yields a mean value of +0.9
446 NNO. Mean oxygen fugacity calculated for the LSR samples were as follows: -0.1 NNO
447 (Ghiorso and Evans 2008, model); +0.4 NNO (Sauerzapf et al. 2008); +0.5 NNO (Ridolfi et
448 al. 2010).

449 Measurements and estimates of the mean water content of the Oruanui HSR magma
450 vary from 4.5 wt% to 6.2 wt% (Table 5). The lowest values come from FTIR measurements
451 of H₂O in quartz-hosted melt inclusions (Liu et al. 2006). The higher estimates are inferred
452 from amphibole rim compositions (Ridolfi et al. 2010; their equation 3), empirical modelling

453 of H₂O solubility applied to Oruanui HSR compositions (Moore et al. 1998) and plagioclase-
454 melt equilibrium models (Housh and Luhr 1991; Waters and Lange 2015).

455

456 **Oruanui mafic compositions**

457 *Whole rock chemistry*

458 Mafic clasts analysed in here and previously (Wilson et al. 2006; Rooyakkers 2015) range
459 between 50-65 % SiO₂, in distinct tholeiitic (higher Fe, Ti) and calc alkaline (lower Fe, Ti)
460 groups (Fig. 7a; Table 3). New trace element data show that many of the mafic clasts notably
461 have Eu/Eu* values around or greater than 1 (Fig. 7b).

462

463 *Compositional characteristics of crystals in mafic clasts*

464 As previously stated, *in-situ* analysis of crystal phases from the mafic clasts was approached
465 in two ways, either on hand-picked separates, or from thin sections of slabbed material. In
466 general the sectioned samples contain greater proportions of smaller-sized and more primitive
467 crystals (higher En orthopyroxene, lower SiO₂ amphibole, higher An plagioclase). The
468 distribution of the data sets for sectioned clasts (analysed by EPMA only) versus handpicked
469 separates (EPMA + LA-ICP-MS) in Fig. 8 highlights that the trace element dataset for
470 crystals from the mafic clasts may under-represent the abundance of less-evolved crystals as
471 an artefact of the crystal size fraction used for LA-ICP-MS analysis. Plagioclase data appear
472 to be most affected by this apparent sampling bias, with no plagioclase from mafic clasts with
473 FeO_t contents greater than 0.5 wt % having been analysed by LA-ICP-MS (Fig. 8b). It is also
474 apparent that most of the major element compositional ranges in the larger crystals fall
475 entirely within the compositional fields defined by the corresponding crystal species in the
476 HSR and LSR rhyolites. Linkages between plagioclase textures and major element
477 chemistries show several features (Fig. 9). The majority of analyses that returned high Fe
478 contents at a given An content were almost exclusively analyses on groundmass plagioclase
479 (microlites) or from very thin (<5 µm) quench crystallisation rims on larger crystals (see also
480 Rooyakkers 2015 and manuscript in preparation). Of the plagioclase crystals that are
481 consistent with derivation from the rhyolites (i.e. overlapping the compositional fields for
482 HSR and LSR crystals), most show thin rims that increase markedly in BSE brightness (i.e.
483 higher An) reflecting their late-stage immersion in hotter, more mafic melt compositions
484 (e.g. CA_P919 plag 16 in Fig. 9). Others have thicker (>100 µm) rim zones of mottled high

485 An surrounding relatively featureless more sodic cores (e.g. CA_P987 plag 2 in Fig. 9),
486 possibly indicating a more prolonged exposure to mafic magma.

487 Olivines from calc-alkaline clasts are almost entirely unzoned with respect to BSE
488 imagery and major element composition, and returned uniform compositions of $Fo_{85\pm 1}$
489 (Electronic Appendix 8). However, the outermost rims of these olivines had very thin (<5
490 μm) BSE-brighter, more Fe-rich rims of Fo_{74} . Olivine is rare in the tholeiitic clasts but three
491 examples gave compositions between Fo_{75} and Fo_{65} . Rare clinopyroxene crystals ranged from
492 En_{34} to En_{49} and Wo_{32} to Wo_{44} (Electronic Appendix 8). The trace element compositions of
493 plagioclase, orthopyroxene and amphibole from the mafic clasts fall almost entirely within
494 the corresponding HSR and LSR crystal fields making identification of crystals considered to
495 be phenocrystic in the mafic magmas difficult (Figs. 10 and S.4). Notable exceptions to this
496 in the calc-alkaline clasts are the very high-An plagioclase (An_{85-95}), and in the tholeiitic
497 clasts the high-En orthopyroxenes and some of the least evolved amphiboles ($Eu/Eu^* > \sim 0.8$).
498

499 *Intensive variables*

500 Constraining magma intensive variables in the Oruanui mafic magmas is challenging because
501 the compositional and textural data indicates that many of the crystals in the mafic clasts are
502 not equilibrium phases. Of the models and phases available to constrain pre-eruptive
503 temperatures, Fe-Ti oxides may be considered the most appropriate because they are known
504 to re-equilibrate very rapidly at magmatic temperatures (typically days to weeks; e.g. Freer
505 and Hauptman 1978; Hammond and Taylor 1982; Venezky and Rutherford 1999).
506 Temperatures calculated from equilibrium pairs (after Bacon and Hirschmann 1988) of Fe-Ti
507 oxides from the mafic samples return bimodal temperature estimates centred around $1000\text{ }^{\circ}C$
508 and $840\text{ }^{\circ}C$ (Fig. 11). This bimodality suggests that there are Fe-Ti oxides from both the
509 mafic and rhyolitic end-members in the mafic clasts. We do not use the Ridolfi et al. (2010)
510 thermobarometric formulations for the mafic amphiboles because the larger crystals are
511 clearly inherited from the rhyolite and the groundmass amphiboles are inferred to have grown
512 under disequilibrium conditions during rapid crystallisation (Rooyackers 2015). Two-
513 pyroxene thermometry has been applied by pairing together clinopyroxene (inferred to be
514 phenocrystic) and the most primitive orthopyroxenes (i.e. those with compositions that are
515 only found in the mafic clasts). When only pairings that pass the Fe-Mg exchange
516 equilibrium test of Putirka (2008) (i.e. $K_D = 1.09 \pm 0.14$) are used, average temperatures of

517 ~1010 °C or 1038 °C are calculated using the Brey and Kohler (1990) and Putirka (2008)
518 models, respectively (n=19: Fig. 11b).

519

520 **Discussion**

521 The diversity of crystal textures and compositions presented in this study highlights the wide
522 range in magma compositions and intensive conditions within the pre-eruptive Oruanui
523 magmatic system. Wilson et al. (2006) demonstrated that although there was significant
524 whole-rock compositional diversity in the Oruanui rhyolitic clasts, there were no systematic
525 changes in composition with respect to eruption phase (i.e. stratigraphic height: Fig. S.1), and
526 therefore no systematic zonation within the immediately pre-eruptive Oruanui melt-dominant
527 magma body. This is confirmed by HSR crystal rims that converge to uniform compositions
528 for each eruptive phase, whether the earliest (phase 1) or latest (phase 10) or vented from
529 different areas within what became the Oruanui caldera (Fig. S.5). Below we consider a
530 sequence of events that led to this final uniformity from what was originally a diverse magma
531 reservoir, and examine factors that governed the development, storage and evacuation of the
532 530 km³ Oruanui melt dominant magma body.

533

534 **Generation of the Oruanui melt-dominant magma body from a zoned crystal mush**

535 On the basis of amphibole and orthopyroxene textural and chemical data, combined with P–
536 T–H₂O constraints, Allan et al. (2013) proposed that the Oruanui HSR magma body had
537 accumulated following rapid decompression and extraction of large volumes of melt (plus
538 entrained crystals) from a crystal-rich mush-like source body. The significant decompression
539 event is reflected in the *in-situ* chemistry of amphibole cores and interior zones which were
540 interpreted to record a chemical signature of the dissolution-recrystallisation history that is
541 texturally evident in orthopyroxene crystals (Fig. 3a). Orthopyroxene destabilisation occurred
542 during decompression from as much as 270 MPa to ~140 MPa, and is recorded in the
543 amphiboles as increasing levels of Mn and Zn, two elements notably enriched in the
544 orthopyroxenes and inferred to have been taken up in the amphiboles as the orthopyroxenes
545 dissolved (Fig. 6a). Decompression to less than ~140 MPa saw the re-stabilisation and re-
546 growth of orthopyroxene rims and a decrease in the abundance of Mn and Zn in the co-
547 crystallising amphiboles.

548 New data presented here for plagioclase crystals further supports the interpretation of
549 rapid extraction and melt accumulation. The significant resorption ± sieve texturing of ~90%

550 of the plagioclase cores, followed by overgrowths of rim-zones (Fig. 3b) are inferred to
551 represent a disequilibrium-then-regrowth history parallel to that experienced by 89% of the
552 orthopyroxenes (Fig. 3). The presence of resorption zones and raddled cores in the feldspars
553 could represent a decompression event or interaction with hotter, less evolved melts within
554 the mush zone, including those parental to the erupted mafic clasts. Experimental studies on
555 andesitic compositions have shown that sieve-textures similar to that observed in the HSR
556 plagioclase cores can form in response to rapid decompression under fluid-undersaturated
557 conditions (Nelson and Montana 1992). Detailed consideration of the blue-grey-cored
558 plagioclases by Charlier et al. (2008; Fig. 3b) showed that these cores had been partially
559 resorbed, then interacted with hotter melts (including those of metasedimentary crustal
560 origin) to yield dusty textures (cf. Tsuchiyama 1985) in some cases before overgrowth of
561 rims zoned out to the common $An_{38\pm 2}$ range. A general origin for the raddled cores through
562 rapid growth is unlikely because these cores span a wide variety of compositions, indicating
563 that their original growth occurred under a variety of conditions, and the inclusions cross-cut
564 the growth zones, indicating that they are a superimposed feature (cf. Nelson and Montana
565 1992). The notable decrease in the abundance of Eu in the amphiboles at ~140 MPa is
566 commensurate with the inflection of Mn concentrations (Fig. 6a, b) that Allan et al. (2013)
567 interpreted to reflect the re-stabilisation and re-growth of orthopyroxene upon reaching the
568 accumulating melt-dominant magma body. The decrease in Eu content of the amphiboles
569 may similarly reflect the restabilisation and crystallisation of plagioclase in the melt-
570 dominant magma body following the decompression experienced by the host magma parcel.

571 A common magmatic history can therefore be inferred for the majority of plagioclase,
572 orthopyroxene and amphibole crystals in the HSR clasts: derivation from a deeper, hotter and
573 less evolved source mush, followed by entrainment and crystallisation of rim-zones in the
574 cooler, shallower and more evolved melt-dominant magma body. The earlier history of HSR
575 quartz crystals, however, contrasts with that inferred for these other main phases, with the
576 trapped melt inclusions recording more evolved compositions than the final (HSR) melt-
577 dominant magma body. Crystallisation of the quartz rim-zones took place in the melt-
578 dominant magma body, along with plagioclase, orthopyroxene and amphibole, as shown by
579 the close correspondence between compositions of the host melt (i.e. groundmass glass) and
580 those of re-entrant inclusions (Liu et al. 2006).

581 Allan et al. (2013) highlighted the significant compositional overlap between the
582 earlier-grown portions of the HSR crystals (i.e. cores plus interiors) and whole crystals from
583 the LSR, in addition to overlapping P-T conditions indicated by amphibole compositions.

584 They argued that the LSR magma, seen as discrete clasts ejected during the eruption in only
585 minor proportions, in fact approximated the initial composition of the voluminous crystal-
586 poor magmas extracted from the crystal mush/source zone. The hundreds of cubic kilometres
587 of this magma that were extracted then, following crystallisation of the rim-zones on many
588 crystals and associated cooling of the order of 40-50 °C, culminated in the final HSR
589 composition. This is consistent with the findings of Wilson et al. (2006), that the HSR
590 compositions could be generated from LSR compositions following ~ 28% fractionation of
591 the crystal phases recorded in the pumices. In this light, the LSR pumices that were sampled
592 during the eruption itself are considered to represent isolated bodies of crystal-poor rhyolite
593 magma that existed in the crystal-mush source reservoir, and that were syn-eruptively tapped
594 during evacuation of the HSR magma body (cf. Cashman and Giordano 2014). Allan et al.
595 (2013) noted that all LSR clasts examined were quartz-free and that quartz may have only
596 begun to grow in the melt-dominant magma body after the melt had evolved to SiO₂-
597 saturated conditions. The short (decades to centuries) timescales proposed from melt
598 inclusion faceting in the Oruanui quartz by Pamukcu et al. (2015) thus only reflect a part of
599 the physical and chemical evolution of the magma body. In light of the expanded crystal-
600 specific dataset considered here, we review these approaches in order to more fully explain
601 some of the features of quartz and the other crystal phases discussed above.

602

603 **A revised model for the pre-eruptive development of the Oruanui magma body**

604 The diversity of compositions and inferred P-T conditions determined from the crystal cores
605 and interiors reflect earlier derivation of the crystals from a thermally and compositionally
606 zoned crystal-mush source. During the processes of crystal-liquid separation, large volumes
607 of melt plus entrained crystals were carried upwards into the melt-dominant magma body that
608 was forming at shallow crustal levels (~90-140 MPa, or ~3.5-6 km depth). Convection and
609 stirring of the accumulating melt-rich body destroyed any compositional stratification (cf.
610 Bishop Tuff: Hildreth and Wilson 2007; Chamberlain et al. 2015) and resulted in melt
611 homogenization that generated compositionally uniform crystal rims. Quartz-hosted melt
612 inclusions (indicating crystal growth from more evolved melts during the formation of the
613 crystal interiors, to slightly less-evolved conditions in the main melt-dominant magma body:
614 Fig. 4; Liu et al. 2006, also Bégué et al. 2015) indicate that quartz was only stable in the
615 uppermost layers of the mush body prior to melt segregation. The interstitial melt in this part
616 of the mush was highly evolved (on the basis of Rb/Sr ratios) and the modal values from

617 volatile saturation pressures (Liu et al. 2006) indicate that this zone was at pressures of ~120
618 to 150 MPa (~4.5-5.5 km depth) but not greater than ~190 MPa (>7 km depth).
619 Compositional characteristics of melt inclusions in sparse CL-dark quartz cores observed by
620 Liu et al. (2006) and age characteristics of inherited zircons in post-Oruanui eruptives
621 (Charlier et al. 2010; Barker et al. 2014) are consistent with at least part of this mush being an
622 intrusive remnant of the 350 ka Whakamaru magma systems (Brown et al. 1998; Matthews et
623 al. 2012). It thus seems likely that some of the plagioclase (e.g., those with compositions
624 notably more evolved than the outermost rims), amphibole and orthopyroxene co-crystallised
625 with quartz in this upper part of the mush layer. Similarly, a small component of the
626 amphibole cores and interiors have low-Al compositions that may reflect their earlier growth
627 and crystallisation from evolved interstitial melt in the quartz-bearing portion of the mush.
628 Quartz is absent in the post-Oruanui rhyolites (Barker et al. 2015), which are typically 20-60
629 °C hotter than the Oruanui magma, supporting the inference that quartz was only stable in the
630 upper and cooler parts of the pre-Oruanui mush.

631 Downward thermal and compositional stratification of the mush below the quartz-
632 bearing zone is reflected in the compositions of plagioclase (increasing An content: Fig 5a-d)
633 and orthopyroxene (increasing En content: Fig 5e, f). The amphibole compositions reflect
634 increasing model temperatures with depth in the mush pile, reaching ~920 °C at pressures of
635 ~270 MPa or ~10 km depth, with corresponding compositional gradients reflected in co-
636 related Eu/Eu* values (Allan et al. 2013: Fig. 6). In addition, the plagioclase compositional
637 data from the HSR clasts show evidence for two contrasting sources for the An>50
638 plagioclase cores (Fig. 5). A subordinate population has somewhat depleted concentrations of
639 compatible elements (e.g. Sr, Mg; Figs. 5c) and elevated levels of incompatible elements (e.g.
640 Ba, Pb: Fig 5d) at values of An >50. Although these particular cores have bulk compositions
641 that are consistent with growth from (qualitatively) hotter and/or drier melts, their trace
642 element characteristics indicate that those melts were comparatively more evolved. In
643 contrast, the dominant population of plagioclase crystals (both cores and rims) show higher
644 Sr, Mg and lower Ba, Pb contents and are thus inferred to have crystallised from contrasting
645 melts. On the basis of these characteristics we suggest that the high-An, low Sr (and Mg)
646 plagioclase cores grew from melts with a significant (tens of percent) crustal component,
647 although there is a diversity of possible crustal compositions available (Ewart and Stipp
648 1968; Charlier et al. 2010; Price et al. 2015). In contrast, the low Ba (and Pb) parts of the
649 crystals are inferred to have grown from more broadly intermediate melts, with lower crustal

650 contributions. This inference is in line with evidence from the blue-grey cored Oruanui
651 plagioclase crystals (Charlier et al. 2008), in which high An overgrowths around the cores
652 were interpreted to have grown from almost ‘pure’ crustal melts with accompanying high
653 $^{87}\text{Sr}/^{86}\text{Sr}$ values (see also Charlier et al. 2010). Such findings not only highlight the wide
654 variety of source melt compositions and the open-system nature of the Oruanui reservoir, but
655 also a heterogeneity of crustal sources and degrees of assimilation within different regions of
656 the pre-eruptive mush body.

657

658 **The role of mafic magma in the Oruanui eruption**

659 In applications of the crystal mush model to the development of eruptible material,
660 importance is often placed on the role that mafic magmas play in providing heat and volatiles
661 to defrost and remobilise the mush (e.g. Bachmann and Bergantz 2006; Burgisser and
662 Bergantz 2011; Huber et al. 2011; Karlstrom et al. 2012; Parmigiani et al. 2014; Bachmann
663 and Huber 2016). Elevated rates of mafic intrusion have also been suggested as a mechanism
664 for pre-eruptive heating and priming in post-Oruanui eruptions from Taupo (Barker et al.
665 2016). The involvement of two distinct, mafic magmas in the Oruanui eruption therefore
666 raises questions as to what role these mafic magmas had in mush reactivation or initiating the
667 eruption. The estimated total volume of mafic magma erupted ($\sim 3\text{-}5\text{ km}^3$) is significantly
668 larger than any other known eruption of purely mafic material in the TVZ (Wilson 2001).
669 However, the ubiquitous down-temperature signatures in the rim-zones of the main crystal
670 phases in the HSR (e.g. decreasing An in plagioclase, decreasing En in orthopyroxene)
671 precludes the notion that injection of mafic material triggered the eruption through mafic
672 recharge-driven heating (Fig. 5; cf. Barker et al. 2016). Wilson et al. (2006) suggested that
673 some tholeiitic mafic magma may have ponded at the base of the primed Oruanui melt-
674 dominant magma body in the weeks to months prior to eruption, and this is supported by
675 evidence presented here for interaction between tholeiitic magma and the LSR and (but to a
676 lesser extent) HSR magmas (Figs. 12 and 13). The resulting convection and stirring brought
677 about by the hot mafic floor to the melt-dominant magma body was used by Wilson et al.
678 (2006) to explain the homogenisation of any significant pre-existing compositional zonation.
679 Here, we consider how such processes would be reflected in the zonation of minerals and in
680 the textures in the mafic clasts.

681 Chemical analysis of the macroscopic crystals from individual mafic clasts (Figs. 8
682 and Fig. S.4) shows that many of these crystals are identical in composition to those found in

683 HSR and LSR clasts and hence inferred to be inherited. The positive Eu/Eu^* values in almost
684 all of the tholeiitic clasts (Fig. 7) is also indicative that this magma has accumulated or
685 assimilated significant amounts of plagioclase. One of the two trends observed in the trace
686 element glass compositions of LSR plagioclase melt inclusions and selvages (Fig. 2b) can
687 be explained by binary mixing between LSR compositions and the least-evolved tholeiitic
688 magma (Fig. 12). Trends within the mafic whole-rock data can be re-visited in the light of the
689 inference that the mafic magmas accumulated a significant component of ‘rhyolitic’ crystals.
690 Three similar, but subtly different trends are highlighted within the tholeiitic clasts (Fig. 13)
691 suggesting that the overall tholeiitic group itself may comprise at least three discrete batches
692 of magma. The overall patterns within the whole rock data trend towards LSR compositions,
693 with the exception of trend 3 on Fig. 13 (four samples) that plots on a compositional
694 trajectory towards HSR compositions. The whole-rock compositional trends within the mafic
695 magma suites, particularly the tholeiitic group, and the apparent mixing trend seen in the
696 glass data of LSR samples indicate that some mafic magma interacted with rhyolitic
697 compositions for an extended period before the clasts were quenched or the eruption
698 commenced.

699 In addition, a particular feature of the amphibole trace element data is the unusual
700 increase in Ni and Co abundance with decreasing Eu/Eu^* and model crystallisation pressure
701 (Fig. 6c, d). These two elements occur at very low levels in the HSR rhyolite in general
702 (Table 2), although the MgO-enriched pumices reported by Wilson et al. (2006: see Table 1)
703 also have slightly enhanced levels of Ni and Co, barely discernible in their XRF data. Similar
704 results to those of Wilson et al. (2006) were reported by Kaneko et al. (2015), also from
705 whole-rock XRF data. Mafic recharge is unlikely, however, to have caused the increased Ni
706 and Co levels widely distributed in the amphiboles (i.e. those crystals from HSR pumices that
707 do not show higher MgO values) because the HSR amphiboles (plus orthopyroxene and
708 plagioclase) all are zoned outwards to compositions reflecting cooler and more evolved melt
709 compositions. We infer that the elevated amphibole Ni and Co levels reflect the dissolution of
710 a high Ni-Co phase into the HSR magma, with these elements then being sequestered into the
711 amphiboles. Of the other crystal phases in the Oruanui eruptives only the Fo_{85} olivines found
712 in the calc-alkaline mafic magmas are strongly enriched in these elements (Ni ~750 ppm; Co
713 ~180 ppm: Electronic Appendix 8). We infer that at some pre-eruptive stage(s), olivine-
714 bearing, relatively primitive mafic magma reached either the melt-dominant body or the
715 upper, evolved levels of the crystal mush zone and the olivines reacted out. Due to the
716 relatively low levels of other trace-elements in these olivines (see Electronic Appendix 8),

717 their dissolution into a large rhyolitic magma body would leave little trace, with Fe-Mg
718 contents being readily taken into ferromagnesian minerals in the rhyolitic assemblage or
719 slightly increasing the MgO content of a small fraction of the HSR. The elevated Ni and Co
720 contents in the HSR amphibole rims may thus be cryptic evidence for a role olivine-bearing
721 (i.e. calc-alkaline) mafic magma in the earlier assembly of the Oruanui melt-dominant
722 magma body than can be inferred from any other information at present.

723

724 **Diffusive timescales in the Oruanui magmatic system**

725 Elemental diffusion in volcanic crystals can be used to provide information about the
726 timescales of sub-surface processes (Costa et al. 2008, for review), and here we highlight
727 three mineral species in particular to place a range of temporal constraints on pre- and syn-
728 eruptive processes in the Oruanui magmatic system.

729

730 *Fe-Mg diffusion timescales in orthopyroxene*

731 Within the Oruanui magmas pre-eruptive timescales on the order of decades to millennia
732 have been obtained from modelling of Fe-Mg diffusion in orthopyroxenes. Allan et al. (2013)
733 modelled the prominent main core-rim boundaries in HSR orthopyroxenes (Fig. 3a), a feature
734 that they interpreted to represent the physical arrival of individual orthopyroxene crystals
735 along with their host LSR melt from the underlying mush into the growing melt-dominant
736 magma body. With allowance for maximum uncertainties in model parameters (especially
737 temperature and oxygen fugacity: Chamberlain et al. 2014a), it was inferred that the crystal-
738 liquid segregation process that led to the generation of the melt-dominant magma body did
739 not begin until only ~1600 years prior to the eruption, consistent with previous inferences
740 from zircon model-age spectra (Wilson and Charlier 2009). Probability density function
741 analysis indicated that the peak ‘extraction age’ was ~230 years earlier and that there was an
742 apparent period of stasis ~60 years just before the eruption.

743 With revisions to magmatic temperatures (Barker et al. 2015; this paper), and the
744 publication of a new model for Fe-Mg interdiffusion in magnesian olivine (Dohmen et al.
745 2016), we have revisited the timescale estimates of Allan et al. (2013). An extended
746 discussion is given in the supplementary material, but the various changes to parameters
747 (Table S.5), combined with the correction of a numerical error, leads to a reduction in
748 timescales of a factor of three (Fig. S.6). This migrates the peak extraction time to 80 years
749 pre-eruption with the inferred pre-eruption stasis period of 20 years (Figs. S.7 and S.8) but

750 does not substantively change the conclusion that mobilisation is geologically rapid, and by
751 implication involving the upwards transport of $>1 \text{ km}^3/\text{y}$ of magma). These timescales are
752 similar to those proposed from the faceting of Oruanui quartz-hosted melt inclusions
753 (Pamukcu et al. 2015), and both timescales may reflect the same process of rapid assembly of
754 the melt-dominant body, although the mineral species concerned come from different
755 domains within the mush.

756

757 *Fe-Ti diffusion timescales from magnetite*

758 BSE imagery of Fe-Ti oxide crystals extracted from the mafic clasts shows that many
759 magnetite crystals in these samples are compositionally zoned (Figs. 14 and S.9). Fe-Ti
760 oxides, and magnetite in particular, are known to rapidly re-equilibrate in response to
761 fluctuations in magmatic conditions (Freer and Hauptman 1978; Aragon et al. 1984;
762 Nakamura 1995; Venezky and Rutherford 1999; Coombs et al. 2000; Devine et al. 2003).
763 The preservation of Fe-Ti zonation in magnetites in the Oruanui mafic clasts, therefore, is
764 taken as a record of a process (or processes) operating shortly before quenching on eruption.
765 Using ImageJ (<http://rsb.info.nih.gov/ij/>), we retrieved spatially-resolved transects across the
766 BSE images and compared them to the chemical composition along the same transects as
767 measured by EPMA (e.g. Fig. 14). The greyscale values from the BSE images strongly
768 reflected changes in the Fe-Ti content of the magnetite. Although zonation in some minor
769 element oxides was also evident (e.g. Al_2O_3 and MgO), variations in the concentrations of
770 these elements appeared to have little or no influence over the greyscale zonation, allowing
771 the BSE intensity to be calibrated for Ti content. Using the changing greyscale intensity to
772 infer the changing Fe-Ti contents of the zoned crystals, we used the BSE images to recover
773 information about Fe-Ti zonation at a higher spatial resolution than that possible by direct
774 EPMA spot analyses. We modelled Fe-Ti interdiffusion in magnetite crystals to calculate the
775 timescale(s) over which mafic-felsic mingling occurred prior to quenching upon eruption
776 (Fig. 14). In order to determine the magnetite Fe-Ti diffusion coefficient in the Oruanui
777 system, we used the experimentally derived parameters of Aragon et al. (1984), and
778 conducted our modelling at conditions of $820 \text{ }^\circ\text{C}$ and an oxygen fugacity of NNO. These
779 conditions correspond to the lowest temperatures and oxygen fugacity conditions inferred
780 from the cores of equilibrium Fe-Ti oxide pairings within the mafic clasts (Fig. 11) and
781 should yield a maximum estimate. On this basis, the timescales obtained are between 6.7 and
782 37 hours ($n = 18$), but mostly <18 hours, and are interpreted to reflect syn-eruptive mixing

783 processes prior to quenching at the surface. Alternative modelling using diffusivity from
784 Freer and Hauptmann (1978) at the same conditions of T and fO_2 yields timescales a factor of
785 ~ 3 shorter than the Aragon et al. (1984) formulation, but this does not change the conclusion
786 that the Fe-Ti oxide crystals record magma mixing processes occurring just hours before the
787 magma was quenched. The samples used for modelling were from pumices in phase 7
788 ignimbrite, the coeval powerfully dispersed fall unit of which contains the greatest
789 proportions of mafic material (up to 4 % of the juvenile fraction: Wilson 2001). Fall deposits
790 of phases 6 and 7 are separated by a short time gap during which the eruption plume fully
791 dissipated, but no erosion occurred (Wilson 2001). These factors suggest that the Fe-Ti oxide
792 zonation reported here is recording syn-eruptive timescales related to inputs of fresh mafic
793 magma during the time break between eruption phases 6 and 7.

794

795 *Fe-Mg diffusion timescales from olivine*

796 Despite the abundance of olivine in the calc-alkaline mafic clasts, very few crystals were
797 suitable for Fe-Mg diffusion modelling. All of the large (1–2 mm) olivine crystals in the calc-
798 alkaline samples have very thin ($< 5 \mu\text{m}$) bright rims in the BSE images reflecting a higher
799 Fe-content. However, closer inspection of the greyscale zonation patterns of the rims (which
800 correlate directly to the changes in the Fe-Mg content of the olivine, e.g. Martin et al. 2008)
801 suggests that these zonation features are not solely diffusion controlled, but reflect a quench
802 crystallisation/growth feature of the olivines accompanying mafic-felsic interaction, indicated
803 by regions of linear compositional gradient in the crystal rim and the lack of a single rim
804 plateau value. Although the rim zonation was not suitable for more rigorous Fe-Mg diffusion
805 modelling, examination of the interaction between the homogeneous cores and the growing
806 rim zone does impart some constraints. Typically we see a very abrupt change between the
807 core and the steadily increasing rim, with only a very narrow region of curvature between the
808 two, which would be the natural consequence of diffusion. Based on the uniformly small
809 length-scales of these regions ($< 5 \mu\text{m}$), we can estimate a maximum timescale over which
810 these regions of abrupt curvature could be preserved. Based on a temperature of $\sim 820 \text{ }^\circ\text{C}$ and
811 NNO oxygen fugacity, the olivine rim zonations are consistent with an event that occurred at
812 most no more than a few weeks before quenching, consistent with short timescales inferred
813 from experimental studies on olivine mixing with silicic melts (e.g. Coombs and Gardner
814 2004). Although more rigorous constraints are not available we suggest that it is possible that

815 the quench rims on the olivines formed during the same event recorded by the zoned
816 magnetite crystals.

817

818 **A cascade of events and their associated timescales as inferred from the Oruanui crystal** 819 **archive**

820 By combining our new results with previously published studies, an overview of the
821 cascading timescales of pre- and syn-eruptive processes culminating in the Oruanui eruption
822 is presented (Fig. 15). From these studies, it is clear that the overall history of the magmatic
823 system that gave rise to the Oruanui eruption (and its precursory leaks: Sutton et al. 1995;
824 Wilson and Charlier 2009) is not encapsulated in any one crystal phase or manifested on any
825 one time scale.

826 At one extreme, the timescale for the growth of the Oruanui mush zone, with its
827 distinctive chemical and isotopic compositional characteristics and zircon model-age spectra
828 goes back around 40 kyr (Sutton et al. 1995; Wilson et al. 2006; Wilson and Charlier 2009).
829 A model that used Oruanui zircon model-ages (Wilson and Charlier 2009) in isolation to infer
830 a magma chamber lifetime of 200 kyr (Caricchi et al. 2014b) is flawed, for three reasons.
831 First, the model-age spectrum used only represents those grains (or parts thereof) that yielded
832 model ages 1 sd or more below the equiline in the U-Th disequilibrium system. The earliest
833 ages presented in Wilson and Charlier (2009) are thus not indicative of the oldest ages
834 present in the zircon suite. Second, a given age estimate from part of a zircon gives no
835 indication in itself as to whether that age reflects material that is foreign and inherited
836 through recycling (xenocrystic: cf. Charlier et al. 2005, 2010; Barker et al. 2014),
837 consanguineous and inherited (antecrystic), or phenocrystic and growing in the magma body
838 that actually erupted. There is thus no single zircon age determination (regardless of its
839 precision) that can be uniquely defined as representing the onset of the Oruanui magmatic
840 system. Third, during the ~200 kyr period prior to the Oruanui event, large volumes of
841 volcanic products attest to the presence of compositionally unrelated magma systems (such as
842 the NE dome system) active in the Taupo area (Sutton et al. 1995; Leonard et al. 2010;
843 Cattell et al. 2016). The onset of the Oruanui magma system in terms of surface eruptives is
844 around 60 ka (Sutton et al. 1995), and the commonality of a ~86-96 ka peak in zircon model
845 ages for the Oruanui and earlier geochemically linked rhyolites (Charlier et al. 2005; Wilson
846 and Charlier 2009) suggests a common source that experienced an enhanced episode of
847 cooling and crystallisation at that time. The dominant younger peak of zircon model ages in

848 the Oruanui magma (37-41 ka across three samples: Wilson and Charlier 2009) also
849 represents a peak crystallisation in prior to formation of the Oruanui melt-dominant body at a
850 time when surface volcanism was inactive in the Taupo area (Wilson et al. 2009).

851 Extraction of melt plus entrained crystals to form the melt-dominant body only
852 occurred within the ~3000 year time gap between the Oruanui (25.4 ka) and its precursor
853 Okaia event at 28.6 ka from a geographically overlapping vent (Allan et al. 2013; Lowe et al.
854 2013) on the basis of the greatly contrasting zircon model-age spectra between the two suites
855 of products (Charlier et al. 2005; Wilson and Charlier 2009). This time gap is only bracketed
856 by the zircon data, but quantified in detail by the textural information and diffusion-profile
857 model age information from Oruanui pyroxenes to have been mostly accomplished in
858 decades to centuries (Allan et al. 2013; this paper). Faceting of quartz-hosted melt inclusions
859 (that were avoided in the work by Liu et al. 2006) also suggests a centennial timescale for the
860 lifetime of quartz growth in the melt-dominant body (Pamukcu et al. 2015).

861 Over years to days, interactions between the mafic and rhyolitic magmas are likely to
862 become important. In many examples globally it has been inferred that injection of mafic
863 magma into a body of evolved magma may precede eruption by days to months (e.g. Sparks
864 et al. 1977; Pallister et al. 1992; Snyder 2002; Leonard et al. 2002). In the case of the
865 Oruanui, however, evidence from the deposits of phases 1 and 2 also implies strongly that
866 external tectonic forces played a central role in starting and modulating the eruption during
867 these phases over comparable time periods (Allan et al. 2012). The evidence for the ‘foreign’
868 NE-dome-type rhyolitic magma migrating at shallow levels into the Oruanui conduit again
869 implies strongly that the onset and initial stages of the eruption were not controlled by
870 properties (overpressure, volatile saturation, etc.) intrinsic to the magma chamber itself (cf.
871 Blake 1984; Foroozan et al. 2011; Gregg et al. 2012). What remains uncertain is whether the
872 large volumes of mafic magmas involved in the Oruanui eruption were opportunistic in
873 exploiting a changing stress field to rise into the mush zone and base of the melt-dominant
874 body, or whether the mafic magmas were active participants in controlling the crustal stress
875 field (Rowland et al. 2010). Evidence given above implies that some mafic magmas
876 interacted with the mush zone or the melt-dominant body prior to the eruption and that
877 multiple batches of magma were involved. However, these interactions were not to an extent
878 that led to any widespread heating signal in the compositions of crystals growing in the melt-
879 dominant body and thus were at best, minor. Later in the eruption sequence, diffusive time
880 estimates from Fe-Ti oxides suggest that the sharply defined onset of phase 7 coincided with
881 intra-eruptive injection of a batch of mafic magma. Such injections resulted in enhanced

882 juvenile mafic proportions, and may have also contributed to the marked escalations of
883 eruptive vigour into phases 3 (when LSR clasts also peak in abundance), 7 and 9 (Wilson
884 2001).

885 At the other extreme, the shortest timescales that can be quantified here are limited by
886 the ability of eruption products to preserve changing conditions in the crystal record. For the
887 Oruanui eruption, the limits are in those timescales that are captured by fast-diffusing
888 elements or species that can react to changes in physical conditions associated with magma
889 rise during eruptive activity. Water and CO₂ gradients in melt pockets in quartz (Liu et al.
890 2007) record rise rates of 5-35 cm/s as the magma started to actively vesiculate, and Li
891 gradients in quartz and feldspar record transient processes as the magma ascended at metres
892 to tens of metres per second (Charlier et al. 2012).

893

894 **Magmatism at Taupo: typical or unique for large silicic systems?**

895 Hildreth and Wilson (2007) put forward a model for the Bishop Tuff magma chamber in
896 which gradients in melt chemistry, crystal content and temperature were established and
897 maintained through incremental accumulation from a source mush reservoir. Subsequent
898 work on mineral-specific aspects of the Bishop record has reinforced that concept
899 (Chamberlain et al. 2015; Evans et al. 2016). This model is broadly similar to that for the
900 Oruanui proposed by Allan et al. (2013) and re-emphasised here. However, several features
901 of the two deposits are in stark contrast to one another. For example, the Bishop Tuff magma
902 was compositionally zoned at the time of eruption such that first erupted compositions were
903 the most evolved, while less-evolved and more crystal-rich compositions became more
904 prominent as the eruption progressed (Hildreth 1979; Hildreth and Wilson 2007). There thus
905 cannot have been significant large-scale convection operating in the Bishop magma chamber.
906 In contrast, the Oruanui magma, although compositionally diverse, was not zoned in any
907 systematic way (Fig. S.1). This feature and the diversity of mineral and melt inclusion
908 compositions and textures in any one clast (Liu et al. 2006; Wilson et al. 2006, 2012; Wilson
909 and Charlier 2009; this paper) imply that the Oruanui magma body was vigorously
910 convecting.

911 Early-erupted crystals in the Bishop magma lack prominent zonation and lack
912 evidence for any significant and widespread dissolution/disequilibrium features (Hildreth
913 1979; Chamberlain et al. 2015), leading Hildreth (1979) to conclude that the Bishop Tuff
914 magma underwent largely *in situ* crystallisation and that the crystal cargo was phenocrystic

915 (in the sense used in this paper). Later-erupted Bishop crystals do, however, show evidence
916 for influx of a contrasting magma, growth of compositionally contrasting rims on quartz,
917 feldspar and zircons, and introduction of pyroxenes (Peppard et al. 2001; Wark et al. 2007;
918 Reid et al. 2011; Gualda et al. 2012a; Chamberlain et al. 2014a,b, 2015; Gualda and Sutton
919 2016). Oruanui HSR crystals often have two, readily identifiable and distinct periods of
920 growth. In the orthopyroxene and plagioclase, the earlier and later histories are readily
921 distinguished on the basis of resorption horizons \pm sieved texturing or patchy zonation in the
922 cores (Fig. 3), features which are absent in the crystals of the Bishop Tuff. The earlier
923 growth stage is inferred to have occurred within the thermally and compositionally zoned
924 mush body, and the other is reflected in the rim-zones crystallised within the melt-dominant
925 magma body. This melt-dominant body, stirred by convective mixing, also evolved over time
926 (decades), as seen in the decreases in An proportions in plagioclase rims to An₃₈ and En in
927 the orthopyroxene rims to En₄₆.

928 The question then arises, that if both magma bodies accumulated following broadly
929 similar mechanisms, what factors led to the strongly contrasting characteristics of the melt-
930 dominant magma bodies and their crystal cargoes? We suggest that the answer lies in the
931 process(es) that ultimately drive the crystal-liquid segregation and the rates at which large-
932 scale melt extraction process take place. In the Oruanui, physical extraction from the mush
933 and assembly of the melt-dominant magma body was extremely rapid, at rates exceeding >1
934 km³/y (Wilson and Charlier 2009; Allan et al. 2013; this paper). Similarly rapid rates of melt
935 extraction have also been demonstrated to have occurred prior to younger eruptions from
936 Taupo (Barker et al. 2016). Allan et al. (2013) suggested that magma-assisted rifting
937 processes prior to the Oruanui eruption may have acted to enhance the melt extraction
938 process by creating gashes through the mush body that acted as focussing channels for melt
939 (and crystals) to migrate upwards and accumulate, as compared with more passive processes
940 (e.g. hindered settling, gas filter pressing) that are generally assumed to drive melt extraction
941 (Bachmann and Huber 2016 for overview). The vigour with which melt and crystals from
942 contrasting compositions, depths and temperatures were brought together in the shallow
943 crustal holding chamber, and the inferred heat loss associated with the 40-50 °C cooling
944 (Allan et al. 2013; this paper), likely resulted in convection and stirring that acted to destroy
945 any chamber-wide compositional and thermal gradients during formation of the melt-
946 dominant magma body. Any early increments of mafic magma that reached the floor of the
947 melt-dominant body may also have served to intensify the mixing processes through
948 providing a source of heat and volatiles (e.g. Sparks et al. 1984). However, such increments

949 were not big enough in volume to counteract the overall cooling of the melt-dominant HSR
950 body, as reflected in the ubiquitous down-temperature signals in compositions of the major
951 crystal phases.

952 In contrast, no mineral textures that could represent the contrasts between growth in
953 the crystal-dominant mush versus that in the melt-dominant body above have been clearly
954 identified in the Bishop eruption products (particularly the early-erupted material). A
955 maximum timescale for accumulation of the Bishop melt-dominant body has been suggested
956 as 70-80 kyr based on zircon textural and age relationships (Chamberlain et al. 2014b). Much
957 shorter timescales proposed from diffusive profiling across zonation boundaries in quartz
958 (Gualda et al. 2012a; Gualda and Sutton 2016) and melt inclusion faceting (Pamucku et al.
959 2015) are interpreted to reflect only the later stages of mixing events that gave rise to zoned
960 crystals (quartz, sanidine, zircon) in the later-erupted, deeper-derived parts of the melt-
961 dominant body (Chamberlain et al. 2014a, 2015). It seems necessary that in order to maintain
962 the apparent equilibrium relationships between crystals and melt, and the vertical
963 compositional and thermal stratification, the fluid dynamical processes involved in the melt-
964 dominant magma bodies must have been much more sluggish in the Bishop than the Oruanui.

965 Identifying minerals, or portions thereof, as phenocrystic, antecrystic or xenocrystic in
966 the Oruanui pumices is challenging. The clearest xenocrystic components are older zircons
967 found as whole grains or cores, and identifiable as such from their age spectra (Charlier et al.
968 2005, 2010), quartz cores with contrasting entrapped melt inclusion compositions (Liu et al.
969 2006) and cloudy feldspar cores with contrasting Sr-isotopic systematics (Charlier et al.
970 2008). The sieve-textured or resorbed cores in the Oruanui HSR plagioclase and
971 orthopyroxenes (Fig. 3) could be characterised as antecrystic, with the interiors having grown
972 in a parental, mush-like reservoir that was, however, physically, compositionally and
973 temporally distinct from the melt-dominant magma body that ultimately erupted. The zircons
974 defining the ~95 ka and 41-37 ka model-age peaks in the Oruanui and its precursor eruptions
975 are also antecrystic on this basis. The phenocryst population, *sensu stricto*, present in the
976 Oruanui pumices is thus quite restricted. It includes the rim-zones of those crystals (quartz,
977 feldspar, orthopyroxene) that have resorbed or non-equilibrium cores, together with the
978 sparse grains of the same minerals that have homogeneous textures and compositions
979 appropriate to the melt-dominant body (e.g. crystallisation model temperatures diminishing
980 from ~840 to ~790 °C). The Fe-Ti oxides are phenocrystic, regardless of their antecedents,
981 because any inheritance patterns are erased by rapid diffusion. The Fe-Ti oxides therefore
982 remained in equilibrium with conditions in the melt-dominant body (and the rim

983 compositions of other crystal phases) for all except the shortest timescales (hours to days)
984 associated with syn-eruptive magma mixing (Fig. 14). Amphiboles include antecrystic cores
985 and phenocrystic rims (plus subordinate phenocrystic grains), but these are only identifiable
986 by chemical analysis, particularly of trace elements (Figs. 6 and S.5), and are not
987 distinguishable on a textural basis. The bulk crystal population in the Oruanui system is thus
988 not amenable to equilibrium modelling (cf. Bégué et al. 2014) as the majority of the crystal
989 mass is inherited and the host melt temperature decreasing rapidly, with corresponding
990 changes in the outwards crystallising compositions.

991 Wilson et al. (2006) presented a scaled, schematic cross section of the Oruanui
992 magmatic system immediately prior to eruption. This cross section is revised and updated
993 here in Fig. 16 in light of new details from Allan et al. (2012, 2013) and the findings of this
994 study. Integration of the existing information and new datasets presented here highlight the
995 following details:

996 1) The realisation that tectonic rifting-related processes played an integral part in
997 triggering the onset of the Oruanui eruption. A body of biotite-bearing rhyolite magma was
998 present in the shallow crust beneath the NE dome system (Figs. 1 and 16; Sutton et al. 1995;
999 Wilson and Charlier 2009). Syn-eruptive rifting processes resulted in the southerly
1000 propagation of NE-dome-type rhyolite magma laterally, to intercept Oruanui HSR magma in
1001 the conduit during phases 1 and 2 of the eruption (Allan et al. 2012). Comparable tectonic
1002 controls may well have been important in other large, caldera-forming events (e.g. Myers et
1003 al. 2016), but the evidence more subtle and not involving magmas of obviously contrasting
1004 visible or compositional characteristics.

1005 2) The approximate depth to the top of the melt-dominant magma body. Although
1006 similar to that proposed from volatile data (Liu et al. 2006), newer data suggest that its top
1007 may have been slightly shallower and lay between ~3.5 and 6.0 km depth (~90 to 140 MPa).
1008 These dimensions, inferred from thermobarometry applied to the outermost rims of
1009 amphibole crystals in the HSR, are consistent with the lack of thermal and compositional
1010 gradient within the body, and consistent with it being contained within Wilson's (2001) areas
1011 'A' (the structural caldera collapse area) and 'B' (the collapse collar: terms after Lipman
1012 1997) of the Oruanui caldera (Fig. 1).

1013 3) Recognition of the crystal diversity and a range of potential source regions.
1014 Approximately 90% of the crystal cargo in the HSR is inferred to have cores and interior
1015 domains that were inherited from a stratified crystal mush body prior to their confluence in
1016 the melt-dominant magma body. Quartz crystals are inferred to have originated from an

1017 upper, cooler, SiO₂-saturated and highly-evolved layer, whereas the majority of amphibole,
1018 orthopyroxene and plagioclase crystals largely originated from deeper, lesser-evolved levels.

1019 4) Origin of the LSR pumices. The LSR pumices were originally inferred by Wilson
1020 et al. (2006) to represent deeper and hotter parts of the melt dominant body. Here, these
1021 samples (as represented by individual pumices) are now interpreted to represent isolated
1022 bodies of crystal-poor rhyolite that were held within the crystal mush and were tapped syn-
1023 eruptively (e.g. Cashman and Giordano 2014). In addition, the LSR melt (and entrained
1024 crystals) is recognised by us as having also previously been the feedstock material for the
1025 melt-dominant body. These conclusions are based on the distinct differences between the
1026 HSR and LSR glass chemistries (Fig. 2) and limited degrees of hybridisation in the pumices
1027 (i.e. minor melt and crystal exchanges between HSR and LSR, as inferred by Allan et al.
1028 2013).

1029 4) Wilson et al. (2006) suggested that the Oruanui mafic magmas, in particular the
1030 tholeiitic magma, had ponded at the base of the melt-dominant magma body and initiated
1031 chamber wide stirring and homogenisation. Subsequent work (Allan et al. 2013; Rooyackers
1032 2015; this paper) has shown that the mafic magmas were mostly interacting with the mush, as
1033 indicated by their accumulated felsic crystal populations and whole-rock trajectories trending
1034 towards LSR (rather than HSR) compositions (Figs. 12 and 13). Commensurate with this
1035 revised interpretation, our data imply that vigorous convection was an inherent feature of the
1036 melt-dominant body from the moment of its inception, and not simply the result of late-stage
1037 introduction of mafic magmas.

1038

1039 **Conclusions**

1040 The assembly and eruption of the Oruanui magma body was a complex, but geologically
1041 short-lived, sequence of events. The work presented here and in other papers cited shows that
1042 there is a rich petrological record of these events in the compositions of melts and minerals,
1043 and in the textures of the mineral phases. The associated processes can be resolved into a
1044 cascade of events over timescales ranging from tens of thousands of years for the
1045 development of the deeper source mush zone to hundreds of seconds for the latest stages of
1046 magma rise before fragmentation and quenching. Key conclusions from the integration of
1047 previous and current Oruanui datasets are as follows.

1048 1. Detailed compositional and textural records within Oruanui HSR quartz,
1049 plagioclase, orthopyroxene and amphibole crystals highlight the transition from

1050 compositional diversity in their early histories, to uniformity immediately prior to eruption.
1051 Melt extraction from the mush zone entrained crystals from all levels of the stratified crystal
1052 mush body, resulting in the juxtaposition of a diverse range of crystals within the melt-
1053 dominant magma body. A general outwards-cooling trend reflected in most mineral
1054 compositions show that the melt-dominant body cooled by ~40-50 °C during its lifetime and
1055 at that rate would have cooled back to mush within a few centuries to millenia had evacuation
1056 not occurred. Our data support the concept that eruptible magma bodies are very short-lived
1057 when compared to the overall lifetime of the parental magmatic system (e.g. Glazner et al.
1058 2004; Barboni and Schoene 2014; Barker et al. 2016).

1059 2. The Oruanui mafic magmas record significant interactions with, or inheritance of,
1060 ‘rhyolite’ crystals (i.e. crystals with compositions appropriate for growth from rhyolite melts)
1061 from traversing through the mush zone during their eruption. Inheritance occurred over
1062 varying timescales, with reaction rims varying from only a few microns to >100 µm.
1063 However, ubiquitous down-temperature signals recorded in HSR crystals precludes the
1064 possibility that mafic magmas played any significant role in wholesale re-heating and/or re-
1065 mobilisation of the mush to cause eruption (cf. Bachmann and Dungan 2002; Huber et al.
1066 2011; Barker et al. 2016). Syn-eruptive interactions between the mafic magmas and the melt-
1067 dominant magma body are recorded in zoned magnetite and olivine crystals, down to
1068 timescales of hours to tens of hours. Such short-term interactions can be directly linked with
1069 inferred eruptive timings and variations in the erupted proportions of mafic materials (Wilson
1070 2001).

1071 3. The disparate textural and compositional domains within single crystals,
1072 particularly HSR plagioclase and orthopyroxene, make complex the definitions of
1073 phenocrysts versus antecrysts for the Oruanui system. Despite the near-uniformity of rim
1074 compositions in the HSR crystals, the compositions and intensive parameters recorded by the
1075 crystal cores and interiors bear no direct relation to the melt-dominant body in which they
1076 were immersed at the time of eruption.

1077 4. The magmatic systems giving rise to voluminous, crystal-poor rhyolites can be
1078 considered to operate in general on two contrasting timescales: longer-term processes
1079 governing the broad compositional characteristics of the magma (i.e. fractional
1080 crystallisation, protolith assimilation, etc.) versus short-lived processes governing physical
1081 assembly of the melt dominant magma body that actually erupts. The major crystallisation
1082 associated with the Oruanui magmatic system is represented by a dominant peak of zircon
1083 model ages centred on ~37-41 ka (Charlier et al. 2005; Wilson and Charlier 2009), in accord

1084 with the U/Th model age of bulk separates of the main crystal phases of 33 ka (+18 ka/– 16
1085 ka: Charlier et al. 2005). Physical assembly of the melt-dominant magma body, as reflected
1086 in orthopyroxene zoning (Allan et al. 2013; this paper) and melt inclusion faceting (Pamukcu
1087 et al. 2015), occurred in <600 years, and mainly within the last ~100 years prior to eruption
1088 (i.e. 2-3 orders of magnitude faster than the processes of chemical evolution in the progenitor
1089 melts). On a global scale, such accumulation rates for eruptible magma are at the extreme,
1090 especially for such a large-sized eruption and appear to reflect the influence of tectonic
1091 processes acting on a mush system that held large quantities of melt. Only slightly smaller
1092 rates applied also to the youngest eruptions at Taupo (Barker et al. 2016). Both rate estimates
1093 are apparently orders of magnitude faster than those inferred for the comparably-sized Bishop
1094 Tuff system (Chamberlain et al. 2014b; Simon et al. 2014).

1095 5. The model inferred for the Oruanui magma system (extraction of melt plus some
1096 fraction of crystals from a crystal mush into a melt-dominant holding chamber) is similar to
1097 the model invoked for the Bishop Tuff (Hildreth and Wilson 2007; Chamberlain et al. 2015).
1098 Despite these broad similarities, the zonation characteristics of the deposits from these two
1099 eruptions are greatly different at whole-rock and single-crystal scales, reflecting different
1100 accumulation timescales for the respective melt-dominant bodies. The rapidity with which the
1101 Oruanui melt-dominant magma body accumulated, and the inferred rapid rates of heat loss,
1102 led to vigorous convection that destroyed any significant compositional and thermal zonation
1103 within the melt-dominant magma body prior to eruption.

1104 6. The timescales inferred either for the lifetime of the overall Oruanui magmatic
1105 system or accumulation of its ~530 km³ melt-dominant magma body are much shorter than
1106 generalised estimates based on numerical modelling for similar sized magma bodies (e.g.
1107 Bachmann and Bergantz 2004; Annen 2009; Gelman et al. 2013; Caricchi et al 2014b). The
1108 involvement of the mafic magmas, and the rapid accumulation of the melt-dominant magma
1109 body is inferred to be the consequence of an enhanced episode of (magma-assisted) rifting in
1110 the crystal mush-bearing crust beneath Taupo during the lead-in to the Oruanui eruption.
1111 Following Rowland et al. (2010) we infer that tectonism associated with the rifted arc setting
1112 served to be a major (dominant?) control on (a) the rapidity of accumulation of the Oruanui
1113 melt-dominant body, (b) the fact that it grew to such a large size without prematurely
1114 evacuating, and (c) its consequent release on eruption in spasmodic fashion. Such a strong
1115 control by a non-volcanic factor raises questions as to how tectonic processes might have
1116 acted to control the development and eruption of other large caldera-related silicic systems.

1117

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1127

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1467 **Figure Captions**

1468

1469 **Fig. 1** Key volcanic and tectonic features of the Taupo area, located in the central North
1470 Island of New Zealand (see inset: TVZ – Taupo Volcanic Zone). Caldera structure
1471 and vent areas for phases 1 to 4 of the Oruanui eruption are from Wilson (2001).
1472 The NE-SW alignment of ‘young’ faults (<20 ka: Langridge et al. 2016) form the
1473 axis of regional extension in the modern TVZ. Dashed line marks the boundary
1474 between Taupo (T) and Maroa (M) volcanoes (Sutton et al. 1995). The stippled
1475 grey region encompasses the NE Dome magmatic system and associated biotite-
1476 rhyolite domes erupted in the last ~50 kyr (triangles – after Sutton et al. 1995).
1477 Modified from Wilson and Charlier (2009).

1478 **Fig. 2** **a** CaO versus Fe₂O₃ and **b** MgO versus SiO₂ data from fragments of pumice matrix
1479 glass in Oruanui high-silica rhyolite (HSR) and low-silica rhyolite (LSR) clasts
1480 (see text for definitions). All LSR clasts contain a bimodal glass compositions, the
1481 most evolved of which corresponds to that of the HSR clasts. The lack of a
1482 mixing/mingling relationship between the two groups of compositions suggests
1483 that the two rhyolite magmas only encountered each other syn-eruptively. **c** Ti
1484 versus Rb/Sr of (whole-rock) HSR pumices, HSR groundmass glass, and quartz-
1485 hosted melt inclusions. The majority of quartz-hosted melt inclusions record more
1486 evolved compositions (higher Rb/Sr ratio and lower concentrations of compatible
1487 elements) than the groundmass glass. Data sources: HSR pumice (Allan et al. 2012
1488 and this paper); quartz-hosted melt inclusions (this paper); HSR groundmass glass
1489 from Allan et al. (2008) and Vandergoes et al. (2013). **d** Ti versus Rb/Sr of
1490 Oruanui rhyolite glasses. The HSR groundmass glass data here are from the distal
1491 fall deposits where single glass shards are large enough for analysis by LA-ICP-
1492 MS (>40 µm thick). The LSR glass values are from selvages on, and melt
1493 inclusions in, plagioclase crystals in LSR clasts. Note the two distinct trends in
1494 LSR glass compositions.

1495 **Fig. 3** **a** Pictorial summary of the textural and zonation characteristics of orthopyroxene
1496 from Oruanui HSR clasts. See text for descriptions. White scale bars are 100 µm.
1497 **b** Pictorial summary of the textural and zonation characteristics of plagioclase
1498 crystals from Oruanui HSR clasts. See text for descriptions. White scale bars are
1499 100 µm.

1500

1501 **Fig. 4** Representative examples of quartz zonation under cathodoluminescence (CL)
1502 imagery. **a** Multiply zoned crystals; **b** dark CL cores, bright CL rims; **c** bright CL
1503 cores, dark CL rims; **d** nominally unzoned. The black scale bars are 200 μm .

1504 **Fig. 5** Selected major element compositional data for plagioclase from **a** HSR and **b** LSR
1505 clasts, and **c, d** selected trace element data versus An content of plagioclase from
1506 HSR and LSR pumices. **e, f** Concentrations of Ti and Al versus mol % En for
1507 orthopyroxene from HSR and LSR clasts. **g, h** Concentrations of Mn and Zn versus
1508 Al content for orthopyroxenes from HSR and LSR clasts.

1509 **Fig. 6** Selected trace element concentration data for amphiboles from HSR and LSR
1510 clasts (symbols are as in Fig. 5). Data shown in panels **a, b** are plotted against the
1511 apparent crystallisation pressure as modelled by the Ridolfi et al. (2010) total-Al
1512 barometer and in panels **c, d** against the Eu/Eu* of the amphibole as a qualitative
1513 measure of evolution of the melt from which the amphibole grew. The origins of
1514 trends marked in each panel are discussed in the text.

1515 **Fig. 7** Selected whole-rock analytical data for single Oruanui mafic clasts, from this
1516 study, Rooyackers (2015) and Wilson et al. (2006). **a** Plot of FeO*/MgO versus
1517 SiO₂ to show the two compositional groupings (calc-alkaline and tholeiitic: Wilson
1518 et al. 2006, dividing line from Miyashiro 1974). **b** New ICP-MS derived values for
1519 the Eu/Eu* ratio (Table 3 and Electronic Appendix 1). Note, in the tholeiitic group
1520 in particular, the positive Eu anomalies (i.e. Eu/Eu* >1) of most clasts, indicative
1521 of plagioclase accumulation.

1522 **Fig. 8** Comparisons of selected major element compositions of crystals from the mafic
1523 versus rhyolitic clasts. The yellow symbols identify analyses for which in-situ trace
1524 element data are available, whereas the orange symbols show analyses for which
1525 there is no corresponding trace element data. The higher spatial resolution of
1526 EPMA versus LA-ICP-MS was the main limiting factor in determining which
1527 crystals could be analysed for their trace element contents.

1528 **Fig. 9** Figure to illustrate the textural and compositional linkages between plagioclase
1529 crystals in the mafic clasts. BSE images of selected crystals are shown here to
1530 highlight some of the linkages that can be made between textures and composition,
1531 represented by a plot of FeO_t versus An content. Analysis spot points in the crystal
1532 images are marked with coloured stars that correspond to the equivalent coloured
1533 symbols in the central plot. The portion of the mafic clast dataset that plots at
1534 higher FeO_t contents at a given value of An (black and grey filled circles) consists

1535 entirely of analyses of rapidly grown plagioclase, i.e. flash crystallised microlites
1536 in the groundmass and thin, quench-crystallised rims. The illustrated samples are
1537 labelled at the bottom of the images with CA and TH denoting crystals from calc-
1538 alkaline and tholeiitic clasts, respectively.

1539 **Fig. 10** Selected trace element compositional data of crystals from Oruanui juvenile mafic
1540 clasts. **a** Plagioclase. **b** Orthopyroxene from a representative evolved tholeiitic
1541 mafic clast (P574; SiO₂ = 62.4 wt.%). Note: only 35 of the 60 analyses by LA-ICP-
1542 MS on this sample included Al in the analytical scan. **c** Amphibole. The grey
1543 shaded regions in all panels show the corresponding compositional ranges from
1544 crystals from the HSR (blue) and LSR (green).

1545 **Fig. 11** **a** Fe-Ti oxide temperatures for equilibrium pairs (after Bacon and Hirschmann
1546 1988) extracted from Oruanui juvenile mafic clasts. Temperatures are calculated
1547 using the model of Ghiorso and Evans (2008). Note the bimodality of the
1548 temperature estimates. The lower temperature group is assumed to represent pairs
1549 of oxides that were entrained from the crystal mush pile and/or during interactions
1550 with the rhyolite body. **b** Mean (± 2 sd) temperature determinations from two-
1551 pyroxene thermometry for 19 equilibrium opx-cpx pairs (i.e. $K_D(\text{Fe-Mg}) = 1.09 \pm$
1552 0.14). BK90 - Brey and Kohler (1990); P08 - Putirka (2008, his eqn. 37). **c** Fe-Ti
1553 oxide temperatures in the mafic clasts (from **a**) versus oxygen fugacity, with the
1554 respective fields for the LSR and HSR plotted.

1555 **Fig. 12** Ti versus Rb/Sr data for rhyolitic glasses (as annotated) and juvenile mafic clasts
1556 (whole-rock). The steeper trend observed in the LSR plagioclase selvage and melt
1557 inclusion data are taken as indicating involvement of the tholeiitic mafic magma
1558 with the LSR glasses, and mixing controls on the tholeiitic whole-rock variations
1559 being with melts in the mush zone, not the melt-dominant body.

1560 **Fig. 13** Selected whole-rock analytical data for Oruanui felsic and mafic clasts. The
1561 tholeiitic group is split into three sub-groups on the basis of possible mixing trends.
1562 Chemical trends in the mafic groups overwhelmingly trend towards the LSR
1563 compositions and only 3 clasts ('trend 3') plausibly trend towards the
1564 volumetrically dominant HSR compositions.

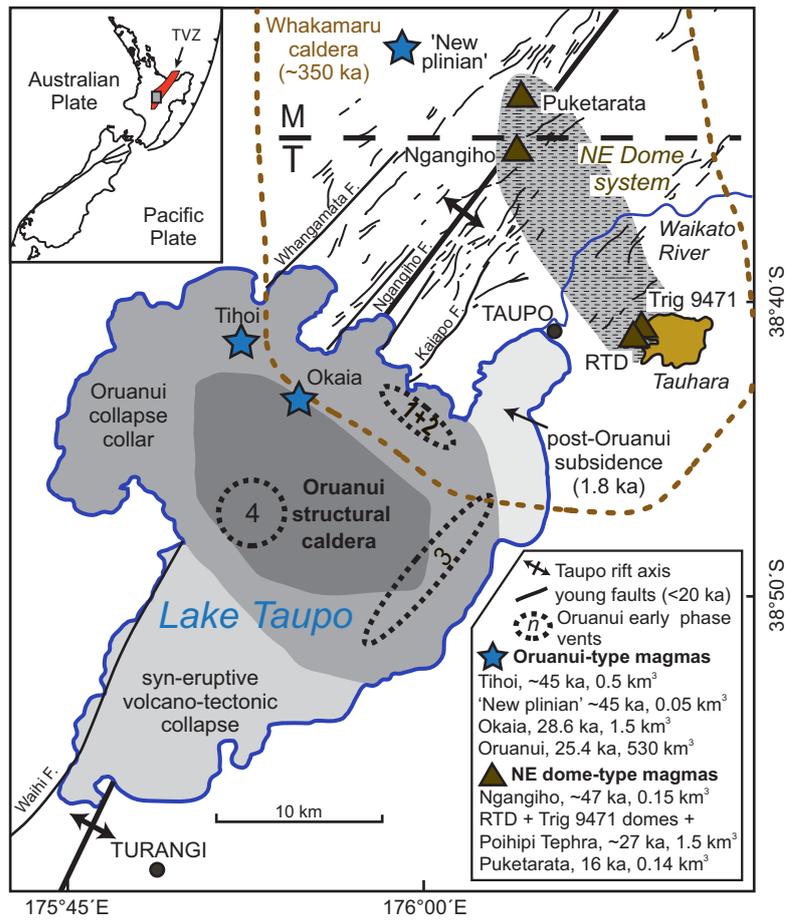
1565 **Fig. 14** An example of a Fe-Ti diffusion modelling of a zoned magnetite crystal from clast
1566 P987 (calc-alkaline). In panels **a** and **b** the greyscale zonation across the crystal is
1567 examined to check for symmetry of the profile. Note the 'flaring' effect at the very
1568 edges of the crystal. In panel **c** a higher resolution portion of the crystal is selected

1569 for modelling, and in panel **d** a model curve is fitted to the observed profile. In this
1570 example the crystal is interpreted to record a thermal perturbation in the final ~11
1571 hours prior to quenching on eruption.

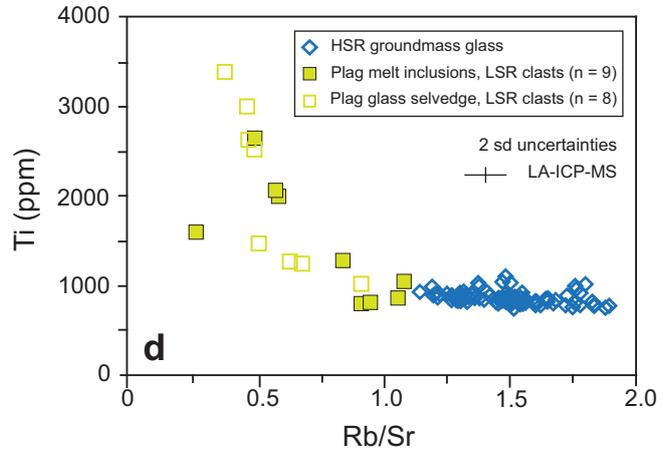
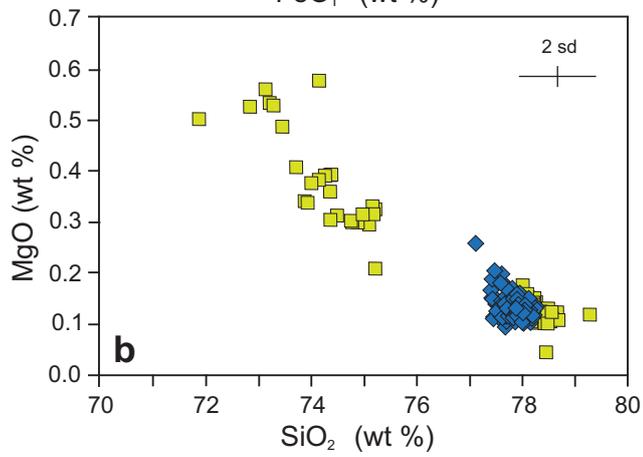
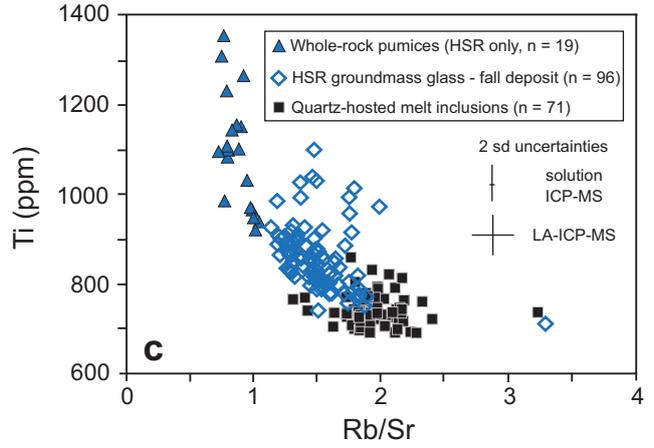
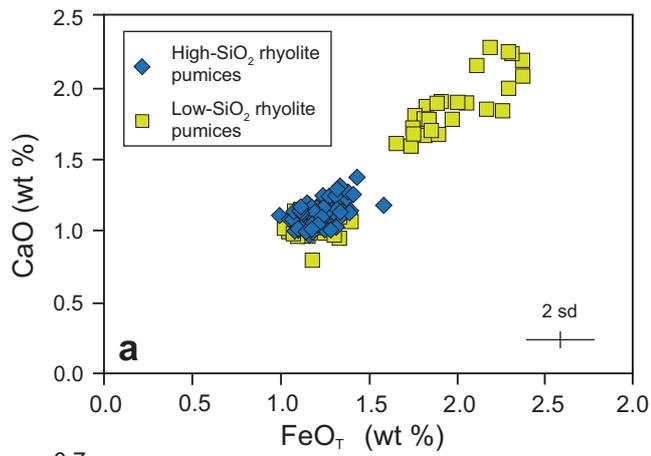
1572 **Fig. 15** Summary figure to show the time-scales of processes involved in the generation,
1573 accumulation and destruction of the Oruanui magma body as inferred from
1574 combined crystal-specific and field studies. Processes are numbered as follows.
1575 1. Compositional characteristic of melt inclusions in xenocrystic quartz cores (Liu
1576 et al. 2006), radiogenic fingerprinting of xenocrystic blue-grey-cored plagioclase
1577 (Charlier et al. 2008), U-Pb zircon age fingerprinting of ‘old’ zircons (Charlier et
1578 al. 2010).
1579 2. U-Th age spectra of zircons and crystal inheritance patterns with Oruanui
1580 precursor eruptions (Charlier et al. 2005; Wilson and Charlier 2009).
1581 3. Processes identified by reconciliation of textural, thermobarometric and
1582 compositional data from Opx and amphibole (Allan et al. 2013). Timescales
1583 constrained as outlined in point 4 and in the supplementary material.
1584 4. Fe-Mg diffusion timescales across the main core-rim boundaries in Opx (Allan
1585 et al. 2013). The equivalent boundary is texturally observed in plagioclase (this
1586 paper and Charlier et al. 2008).
1587 5. Fe-Mg diffusion timescales from re-crystallised streaky, discontinuous zones in
1588 Opx (Allan et al. 2013)
1589 6. Fe-Mg diffusion timescales from growth zones in Opx (Allan et al. 2013)
1590 7. Crystal inheritance pathways, textural observation and disequilibrium textures
1591 (this paper)
1592 8. Fe-Ti diffusion timescales in magnetite crystals (this paper)
1593 9. Compositional relationships in matrix glass and crystals (this paper)
1594 10. Compositional source fingerprinting and field constraints (Allan et al. 2012)
1595 11. Li diffusion timescales in plagioclase and quartz (Charlier et al. 2012) and H₂O
1596 and CO₂ diffusion timescales from melt pockets (Liu et al. 2007)

1597 **Fig. 16** Schematised scaled cross section through the Taupo area to show the sources,
1598 depths and pathways for the various crystal and melt components represented in
1599 the Oruanui eruption products.

1600



Allan et al. Figure 1

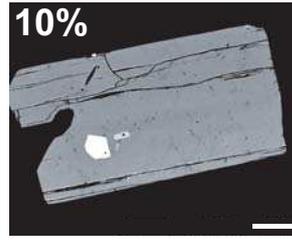
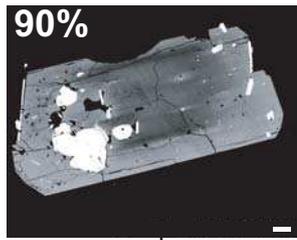


Allan et al. Figure 2

a Orthopyroxene in HSR clasts

Prominent zonation

Comparatively Mg-rich core, rich in melt and mineral inclusions, and variably modified by dissolution, recrystallisation and diffusion
40-400 µm thick rim-zones grade out to En_{46±2} in the outermost rims.



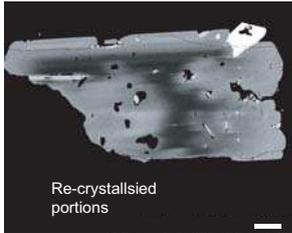
No core-rim zonation

Inclusion-poor crystals with no clear core-rim relationship.

No significant zonation in backscattered electron images.

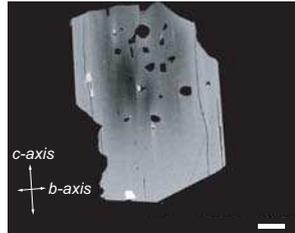
Zoned subgroups

Heavily recrystallised cores



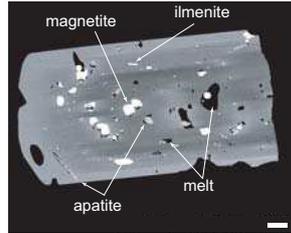
Feature seen to some extent in all 'cored' Opx. Typically evident around melt and mineral inclusions.

Streaky, discontinuous zonation parallel to c-axis



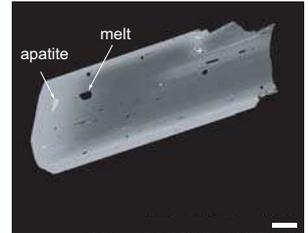
Feature seen in all 'cored' Opx but is most evident in examples with inclusion-rich cores.

Inclusion-rich cores



Common inclusions, in order of decreasing abundance, include magnetite, melt, apatite, and ilmenite, with very rare zircon.

Inclusion-poor cores



Common → Rare

89%

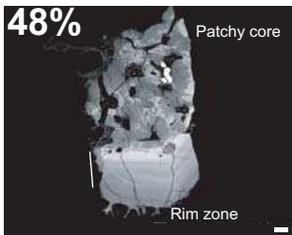
b Plagioclase in HSR clasts

11%

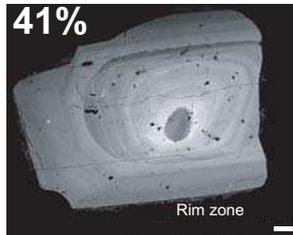
Prominent zonation

Patchy/sieve textured ± resorbed cores

Resorbed cores
NO patchy/sieve texture



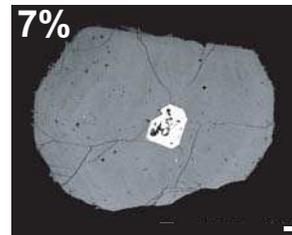
Normally zoned outward toward compositions of An_{38±2}



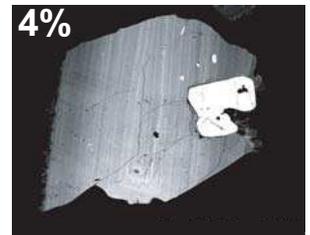
No overarching core-rim zonation

Nominally unzoned

Normal, oscillatory zoned



Lack prominent core-rim zonation



Sub-groups within 'cored' plagioclase

4.5% (of cored plag)
High-An (>An₇₀) cores

Sub-groups are not mutually exclusive

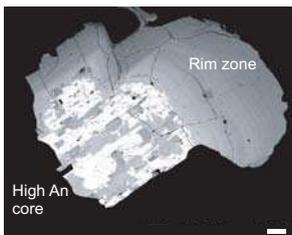
6% (of cored plag)
Blue-grey cored plagioclase

Patchy/sieve textured ± resorbed cores

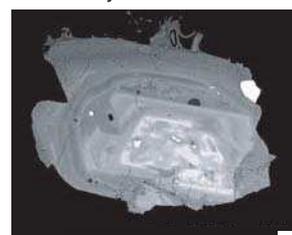
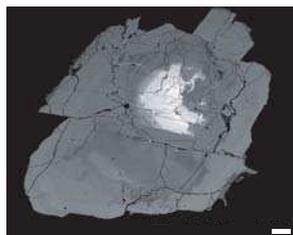
Resorbed cores
NO patchy/sieve texture

Patchy/sieve textured

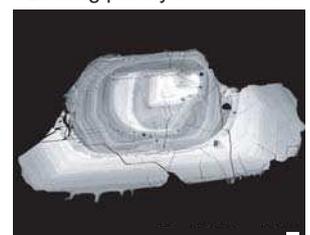
Lacking patchy/sieve texture

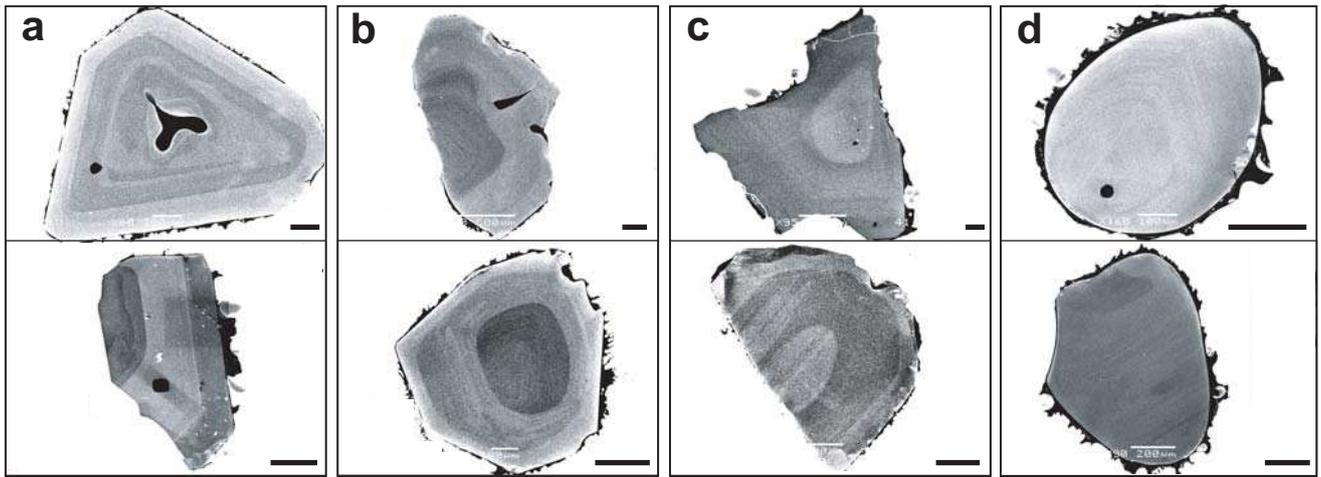


High An cores inherited from mafic magmas

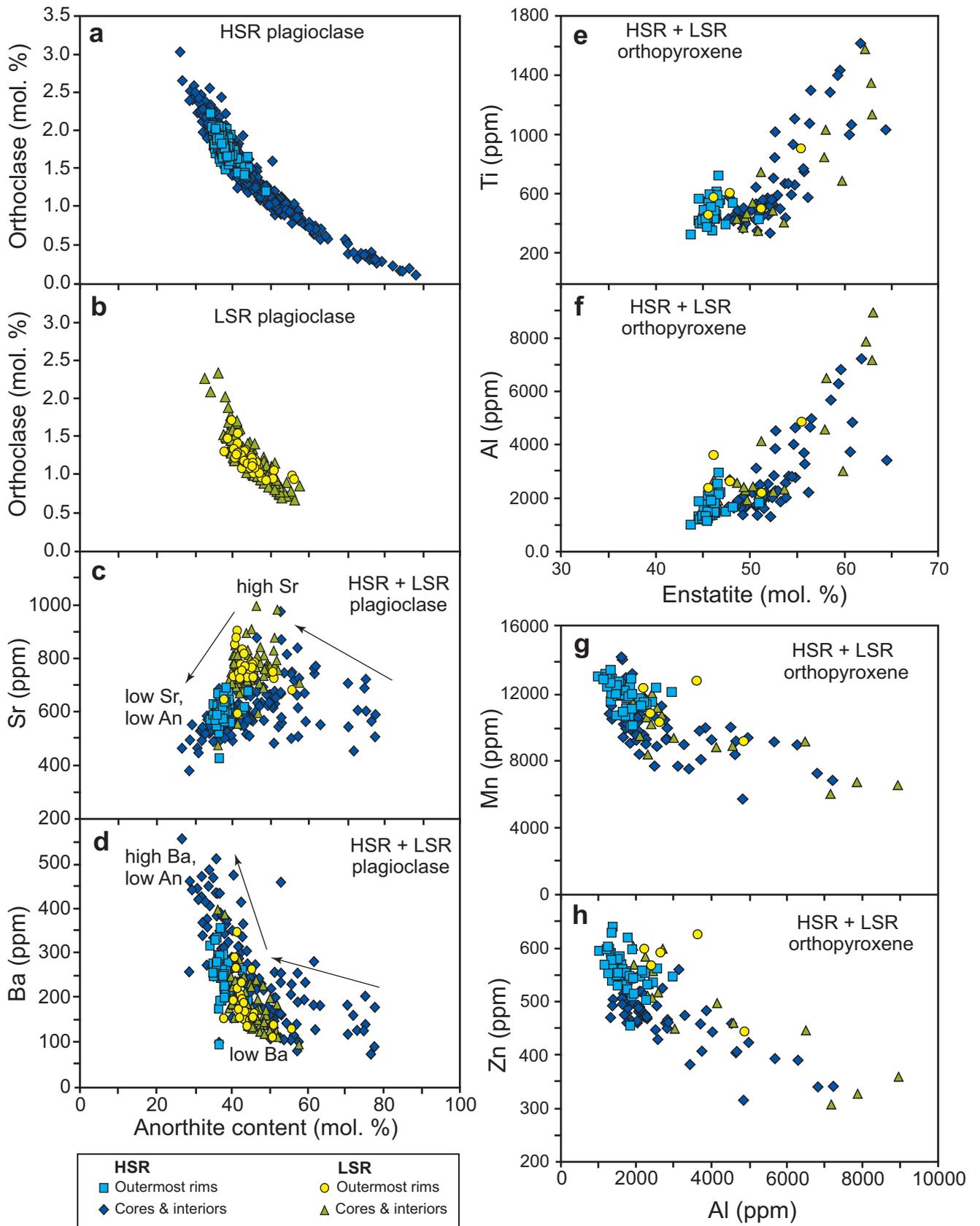


Blue-grey cored plagioclase interacting with greynacke melts

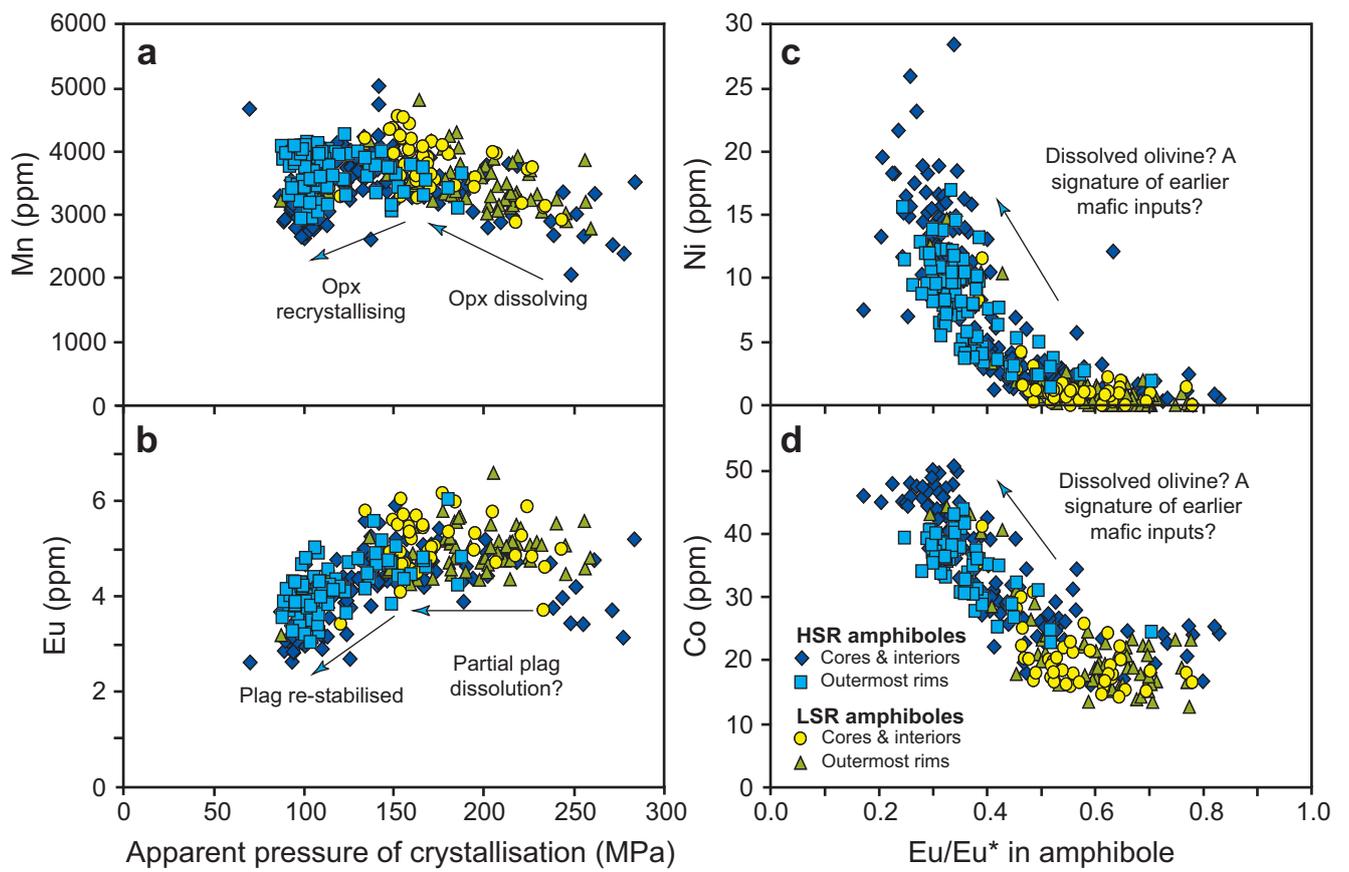




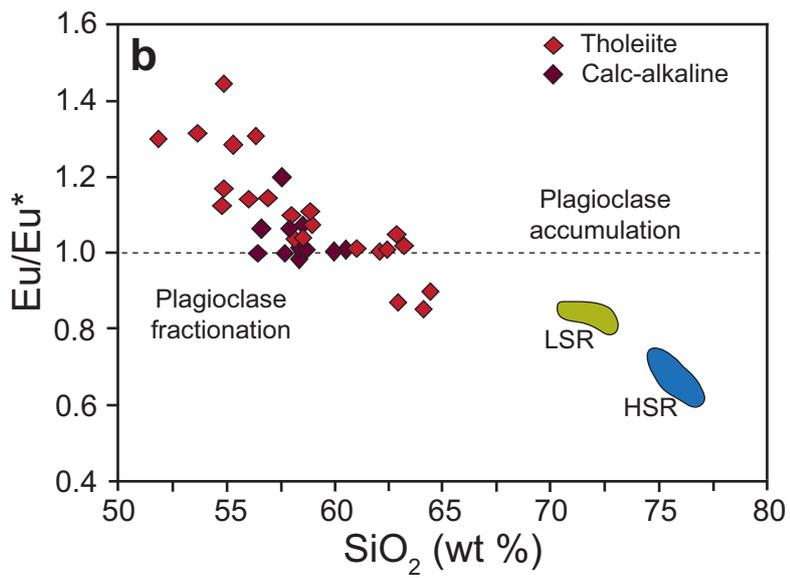
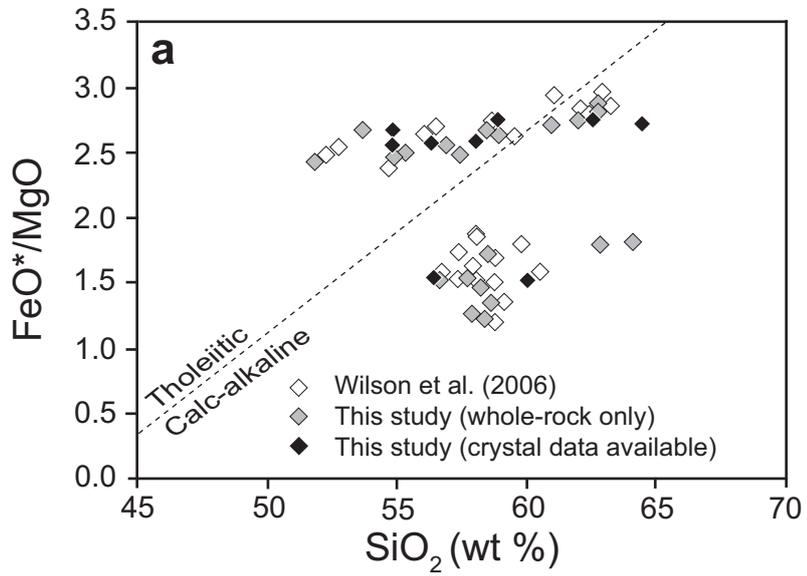
Allan et al. Figure 4



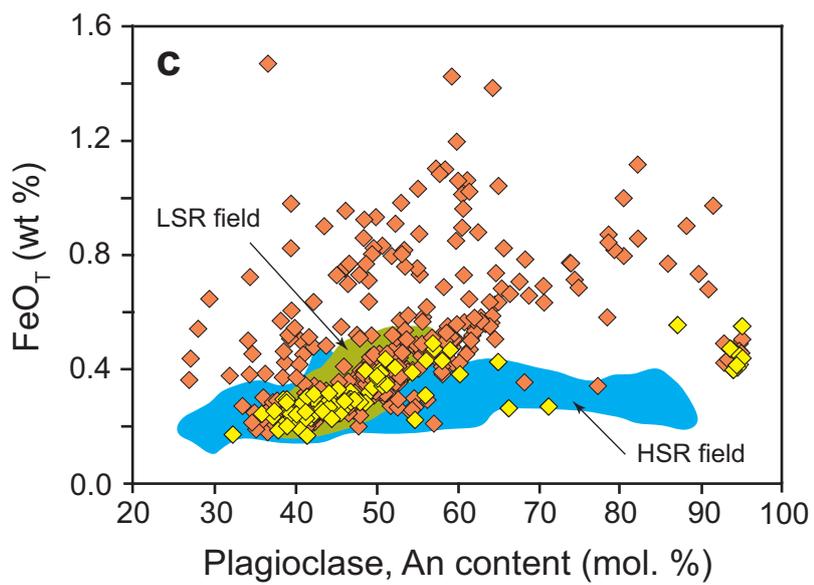
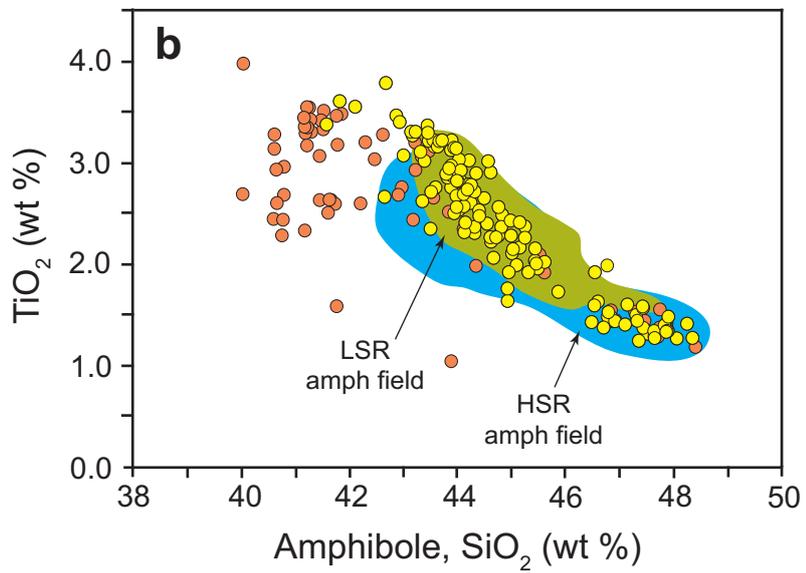
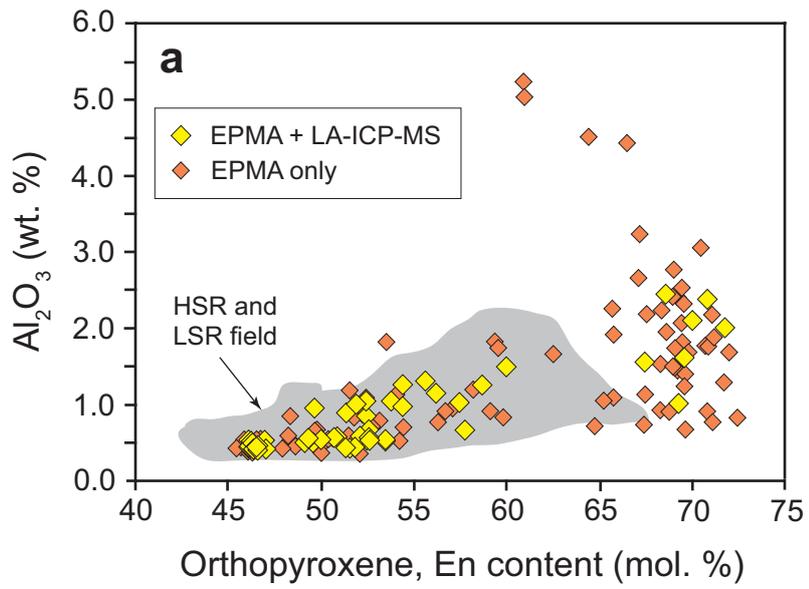
Allan et al. Figure 5



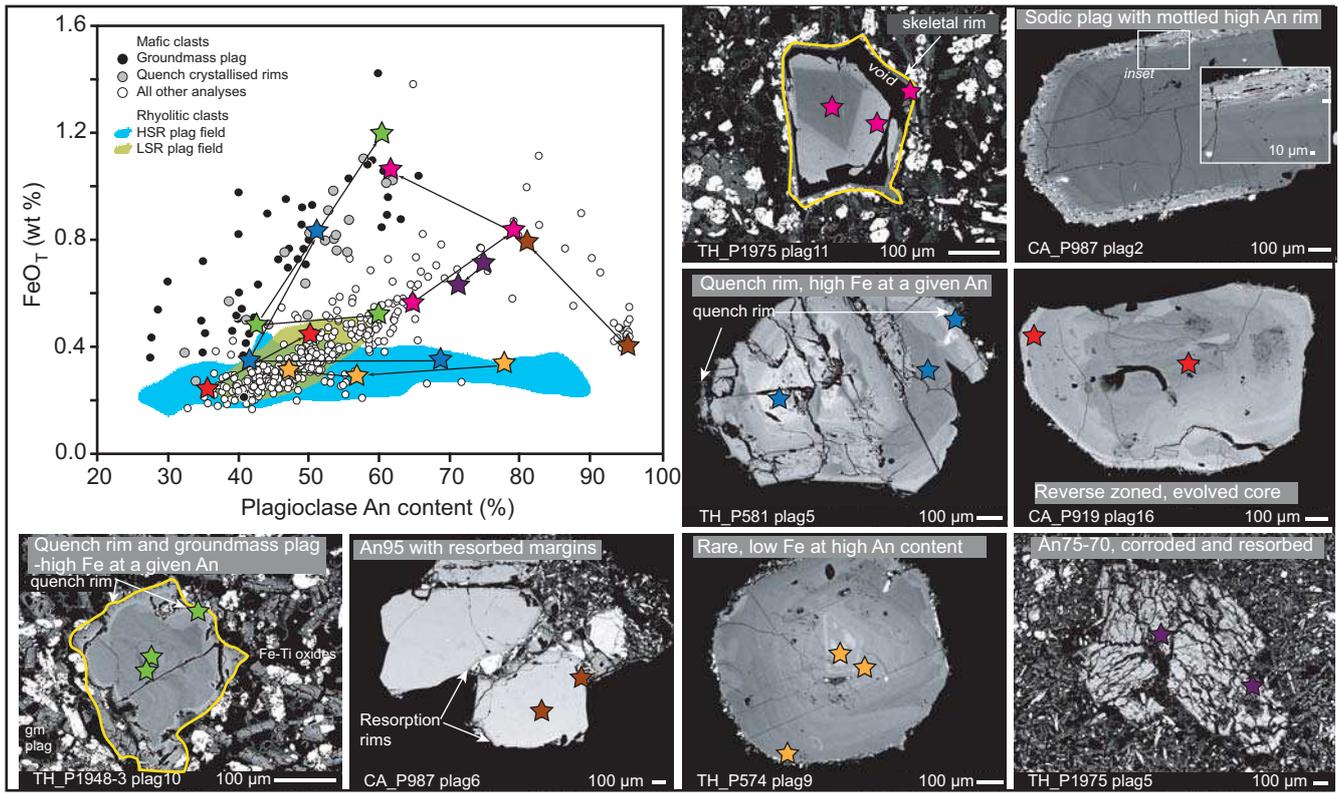
Allan et al. Figure 6



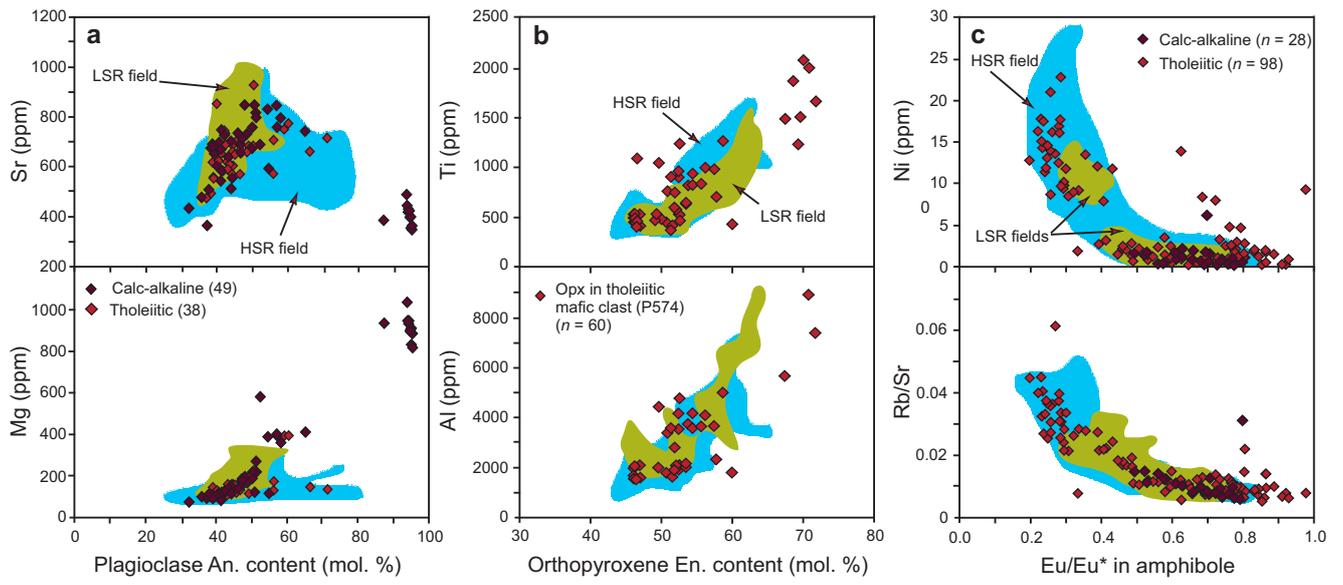
Allan et al. Figure 7



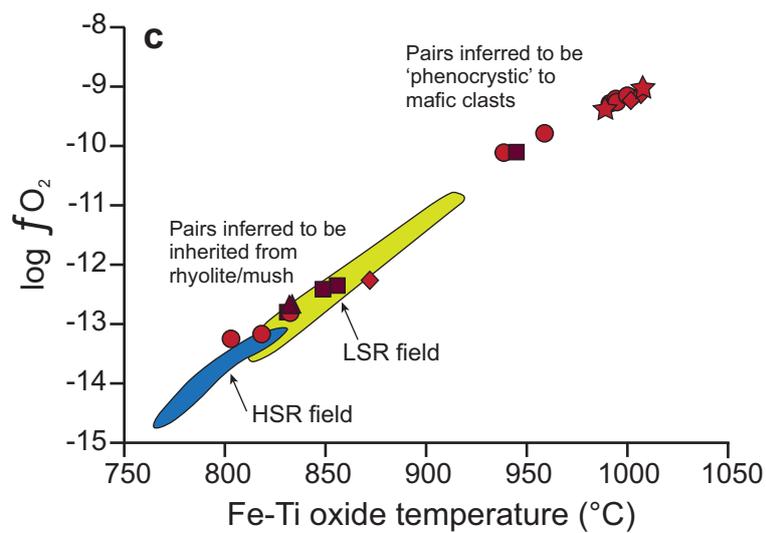
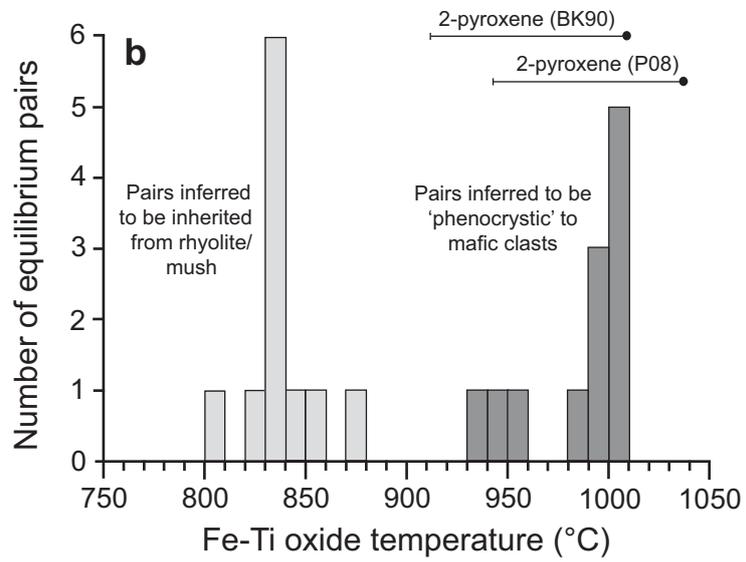
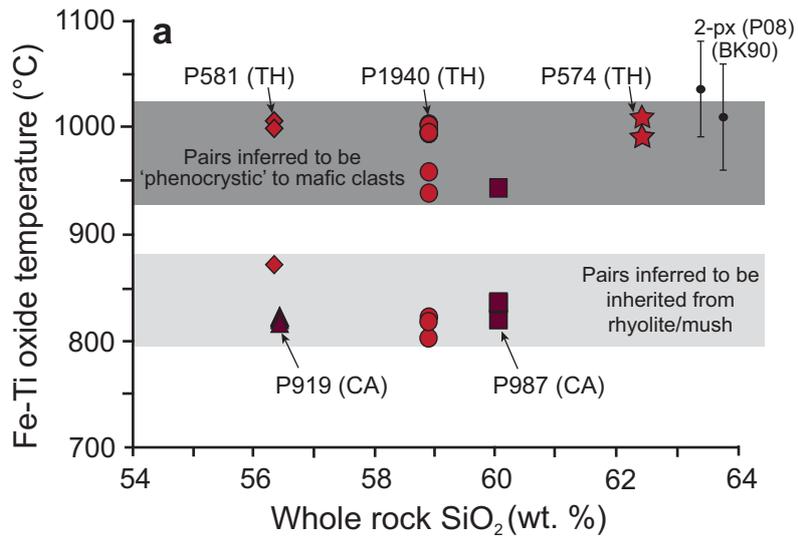
Allan et al Figure 8



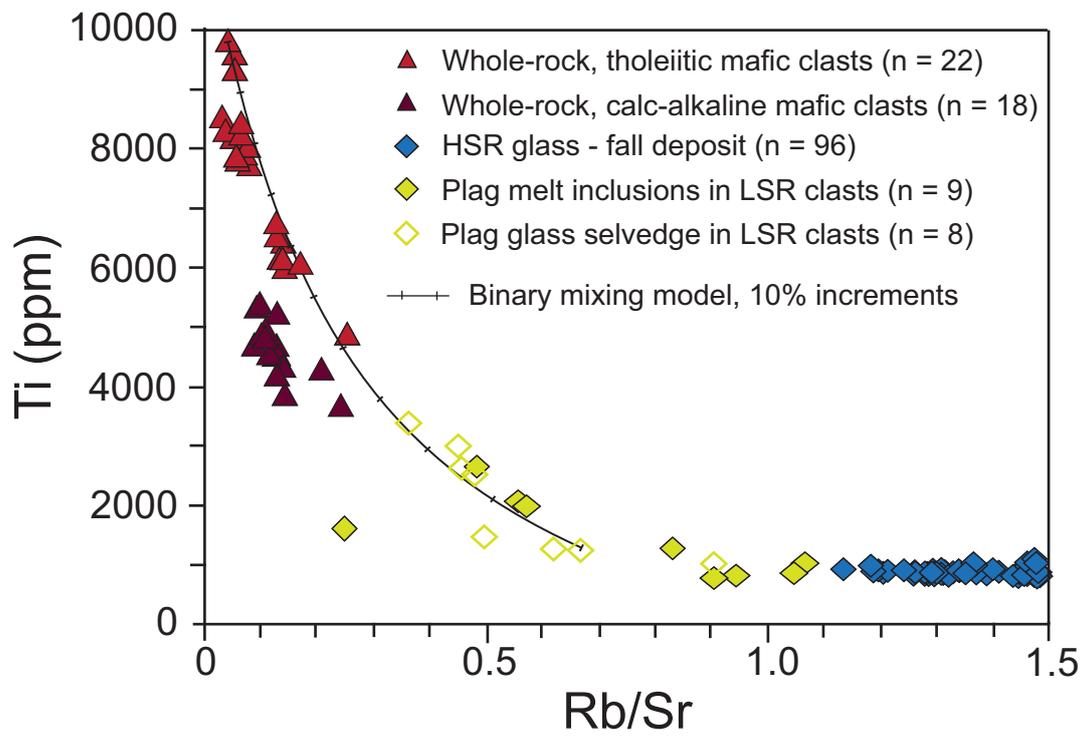
Allan et al. Figure 9



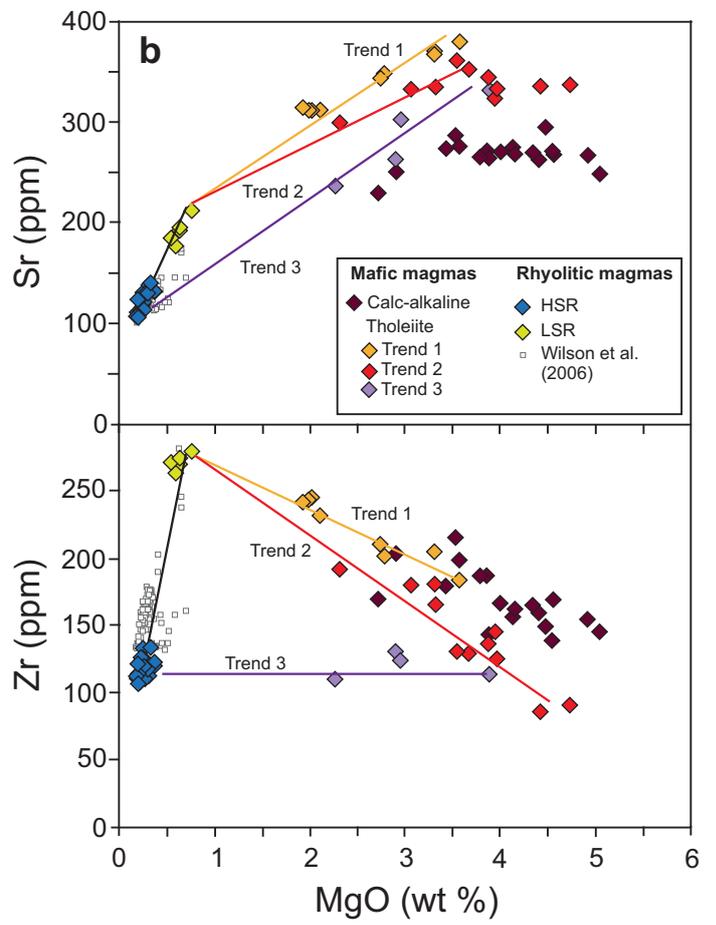
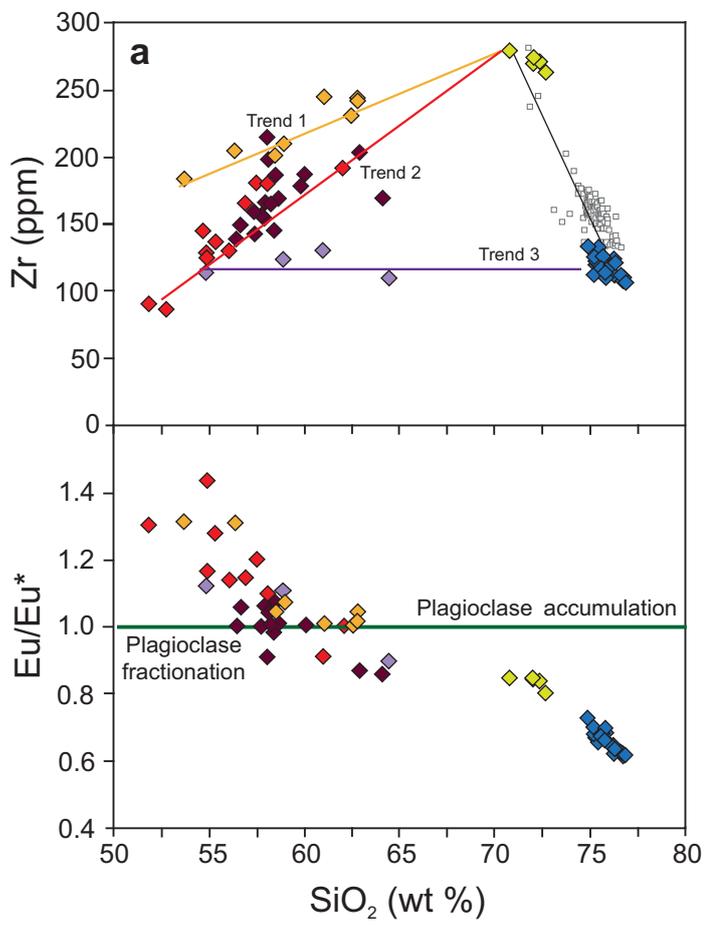
Allan et al. Figure 10



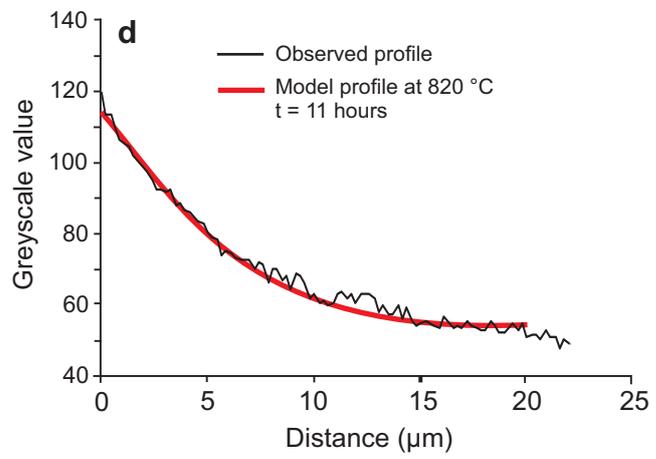
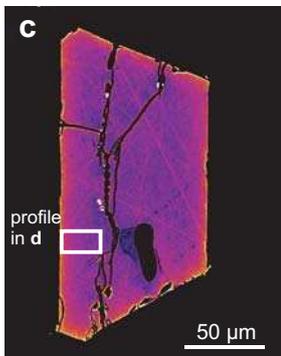
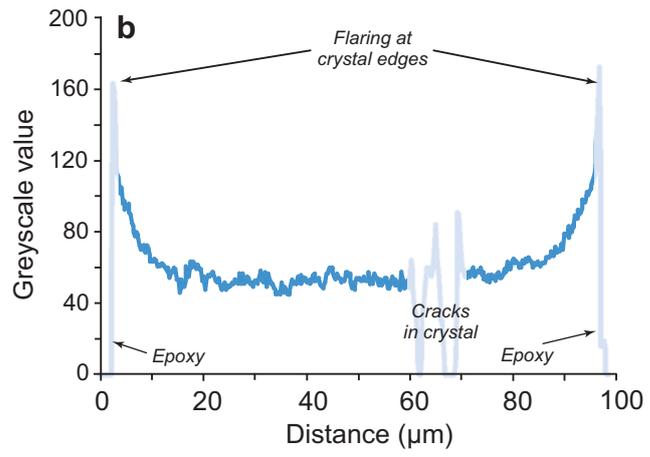
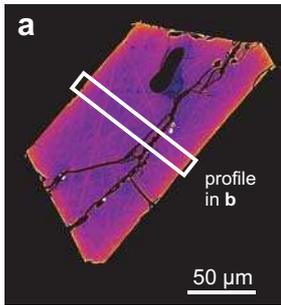
Allan et al. Fig. 11



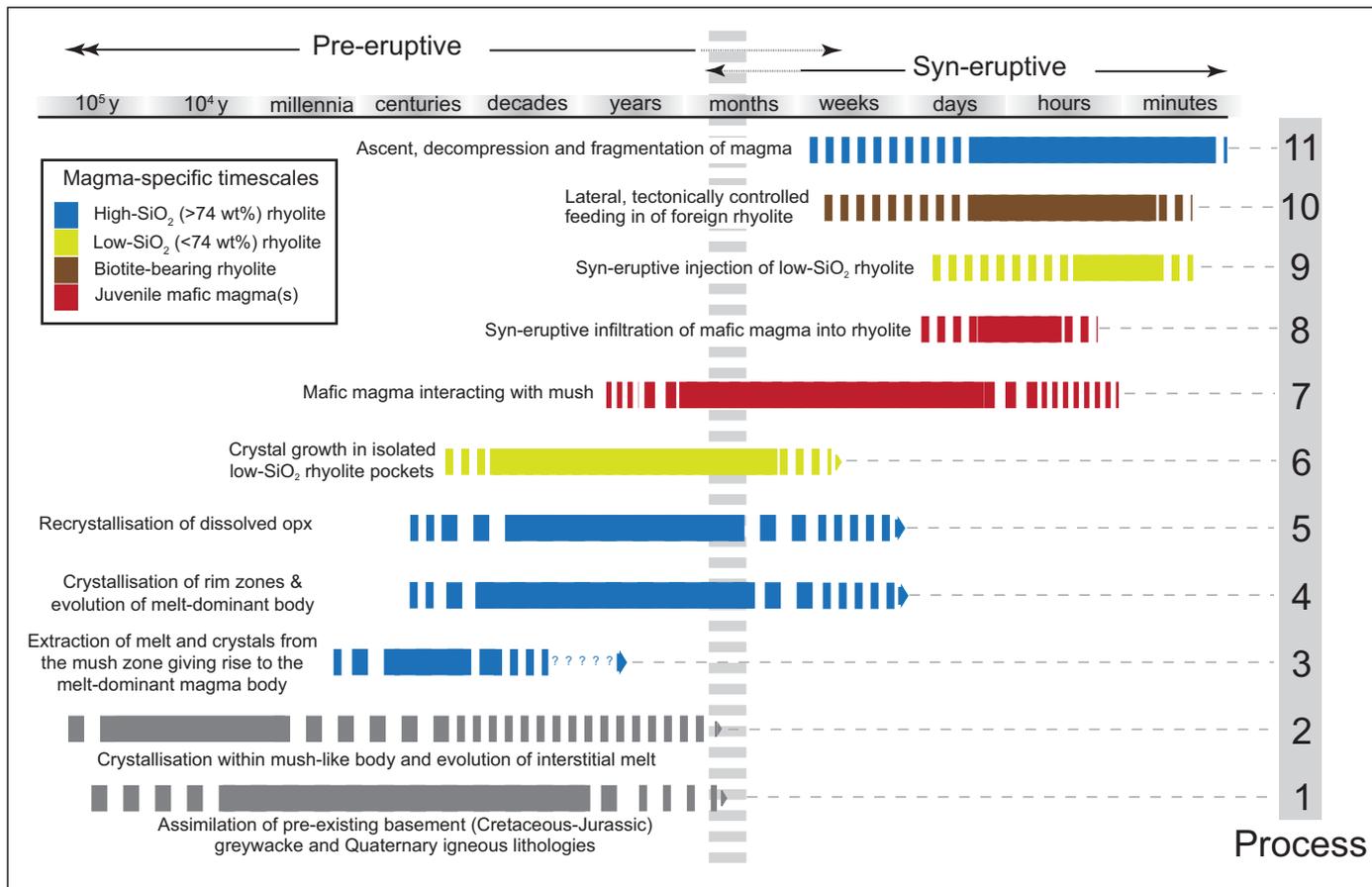
Allan et al. Figure 12



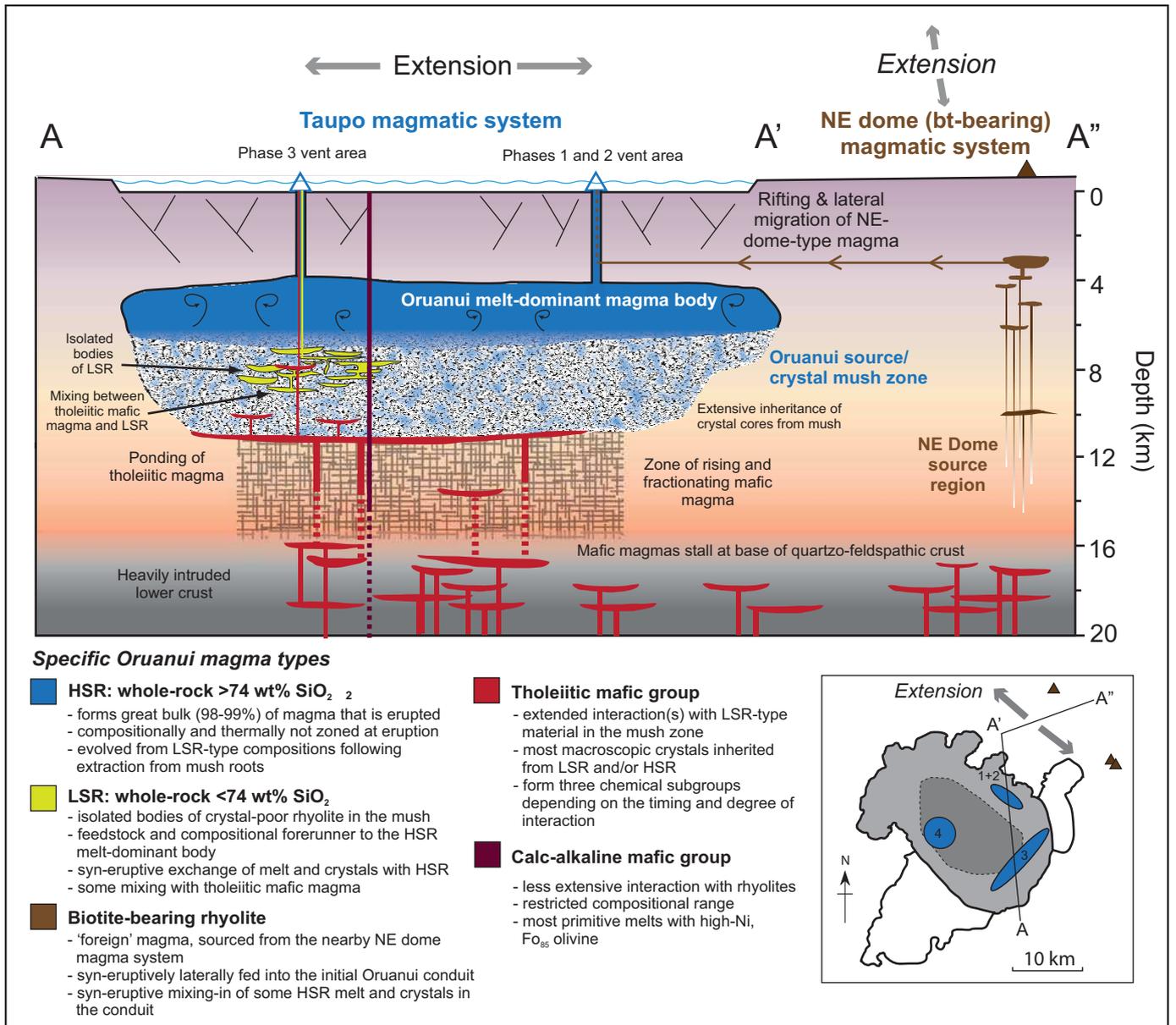
Allan et al. Figure 13



Allan et al. Figure 14



Allan et al. Figure 15



Allan et al. Figure 16

Table 1. Juvenile products of the Oruanui eruption.

Magma type	Features	Approx. magma volume	Abbreviations used in this paper	Key references
High-SiO ₂ rhyolite	Whole-rock SiO ₂ >74 wt%; 3-13% crystals; 64-73% plag, 15-21% qtz 7-8 % opx, 1.3-5% amph	525 km ³ (>98% of total)	HSR	Wilson et al. (2006); Liu et al. (2006); Allan et al. (2012, 2013)
Low-SiO ₂ rhyolite	Whole-rock SiO ₂ <74 wt%; 6% crystals; 67% plag, 0% qtz 8 % opx, 18% amph. Greatest abundance in deposits of phases 3 and 6.	0.5 km ³	LSR	Wilson et al. (2006); Allan et al. (2012, 2013)
Biotite-bearing	Biotite-bearing, 12-24% crystals, Rhyolite derived from an adjacent, independent magmatic system during phases 1 and 2	0.03 km ³	BtB	Sutton et al. (1995), Allan et al. (2012)
High-MgO rhyolite	Visually homogeneous pumices that fall along a trend of least evolved tholeiite mafic magma and HSR. Exceedingly rare as hand specimens	negl.	n/a	Wilson et al. (2006)
Streaky/mixed pumices	Streaky mixtures that fall along a trend of least evolved tholeiite mafic magma and HSR. Exceedingly rare as hand specimens	negl.	n/a	Wilson et al. (2006)
Juvenile mafic 1	Calc-alkaline trend, olivine common	1-3 km ³	CA	Wilson et al. (2006); Rooyakkers (2015); Rooyakkers et al. (in prep.)
Juvenile mafic 2	Tholeiitic trend, olivine rare	1-3 km ³	TH	Wilson et al. (2006); Rooyakkers (2015); Rooyakkers et al. (in prep.)

Note: plag = plagioclase, qtz = quartz, opx = orthopyroxene, amph = amphibole.

Table 2. Representative major (XRF) and trace element (ICP-MS) analyses of Oruanui rhyolitic clasts.

Sample no Eruption phase	P1957-3 1	P1959 1	P2110F 3	P1676 3	P2075 7	P2059 10	P1670 3	P2110C 3	P2110A 3	P2110B 3
SiO ₂	75.90	75.33	76.31	76.76	76.14	75.83	70.80	72.05	72.71	72.39
TiO ₂	0.19	0.22	0.18	0.16	0.20	0.17	0.47	0.40	0.37	0.38
Al ₂ O ₃	13.19	13.30	12.97	12.81	12.84	13.78	14.96	14.56	14.39	14.52
Fe ₂ O ₃	1.72	1.92	1.69	1.60	1.81	1.68	3.25	2.93	2.82	2.87
MnO	0.06	0.06	0.06	0.05	0.06	0.05	0.10	0.10	0.09	0.09
MgO	0.27	0.37	0.26	0.22	0.26	0.20	0.76	0.63	0.59	0.55
CaO	1.65	1.81	1.55	1.42	1.50	1.41	2.81	2.38	2.25	2.33
Na ₂ O	4.04	4.07	3.70	3.68	4.11	3.87	4.35	4.21	3.91	4.05
K ₂ O	2.98	2.90	3.23	3.27	3.04	2.99	2.35	2.63	2.77	2.73
P ₂ O ₅	0.02	0.03	0.04	0.03	0.04	0.02	0.13	0.10	0.09	0.10
LOI	2.82	3.41	3.44	3.36	2.76	3.39	3.91	4.14	4.63	4.47
SUM	96.80	96.15	95.80	96.62	96.63	95.96	95.68	95.14	94.58	94.63
FeO*/MgO	5.82	4.64	5.81	6.54	6.23	7.51	3.84	4.16	4.27	4.73
Li	31	28	13.7	13.5	36	29	16.6	15.5	26	26
Sc	5.7	6.8	5.9	5.3	6.7	6.0	11.9	10.2	10.6	10.1
Ti	1080	1350	1100	970	1270	1030	2820	2420	2300	2230
V	6.0	10.8	5.9	4.7	6.2	4.5	16.0	8.5	9.5	9.6
Cr	1.2	4.3	1.3	0.82	0.54	2.2	1.2	2.4	4.5	3.4
Co	1.3	2.0	1.8	1.2	1.2	1.0	-	-	-	-
Ni	-	-	-	-	1.5	1.3	-	-	-	-
Cu	4.5	4.6	3.1	3.1	3.1	4.1	2.9	4.2	5.8	4.6
Zn	56	62	63	58	46	45	76	79	66	65
Ga	15.2	15.6	15.3	14.8	14.8	15.2	17.8	17.5	17.3	17.1
Rb	102	102	108	110	105	103	72	76	83	79
Sr	129	133	122	113	114	108	213	193	178	186
Y	23	24	24	23	24	22	28	28	28	27
Zr	114	123	112	108	120	114	280	270	264	272
Nb	6.5	6.7	6.6	6.4	7.3	7.0	-	-	-	-
Cs	5.5	5.4	6.1	6.0	4.7	4.4	3.8	4.0	4.4	3.4
Ba	628	604	646	614	658	672	510	539	565	545
La	22	22	23	22	23	21	21	22	21	21
Ce	44	44	47	45	47	45	44	46	45	45
Pr	4.9	5.0	5.2	5.1	5.1	4.8	5.3	5.5	5.3	5.3
Nd	17.8	18.5	19.1	18.5	18.8	17.4	21	22	21	21
Sm	3.6	3.8	3.9	3.8	3.9	3.5	4.6	4.7	4.6	4.5
Eu	0.82	0.84	0.80	0.78	0.82	0.77	1.3	1.3	1.2	1.3
Gd	3.7	3.9	3.9	3.8	3.9	3.6	4.9	4.9	4.7	4.6
Tb	0.57	0.60	0.60	0.57	0.60	0.55	0.76	0.76	0.74	0.72
Dy	3.7	3.8	3.9	3.7	3.9	3.5	4.8	4.8	4.7	4.6
Ho	0.78	0.80	0.80	0.78	0.81	0.74	1.0	1.0	1.0	0.95
Er	2.4	2.5	2.4	2.4	2.5	2.2	3.0	3.0	3.0	2.9
Tm	0.36	0.38	0.37	0.36	0.38	0.34	0.45	0.45	0.45	0.44
Yb	2.5	2.6	2.6	2.6	2.6	2.4	3.0	3.0	3.1	3.0
Lu	0.39	0.40	0.40	0.39	0.39	0.36	0.47	0.47	0.47	0.46
Hf	3.5	3.6	3.4	3.4	3.6	3.6	6.4	6.5	6.4	6.3
Pb	12.7	12.6	13.2	13.1	13.1	13.7	9.9	10.8	11.4	11.0
Th	10.7	10.7	11.1	11.0	10.6	11.3	7.4	8.0	8.6	8.0
U	2.4	2.4	2.5	2.6	2.4	2.4	1.8	1.9	2.0	1.9
Rb/Sr	0.79	0.76	0.88	0.97	0.92	0.95	0.34	0.39	0.47	0.42
Eu/Eu*	0.68	0.67	0.62	0.62	0.65	0.66	0.85	0.84	0.80	0.84

Table 3. Representative major (XRF) and trace element (ICP-MS) analyses of Oruanui juvenile mafic clasts.

Sample no Group	P554 TH	P1975 TH	P581 TH	P1974 TH	P574 TH	P1976 TH	P1643 CA	P919 CA	P1664 CA	P987 CA
SiO ₂	52.27	54.83	56.52	58.05	62.43	64.47	56.66	56.74	57.90	60.53
TiO ₂	1.44	1.31	1.68	1.24	1.02	0.79	0.79	0.80	0.67	0.78
Al ₂ O ₃	16.92	16.54	17.67	16.71	16.59	14.89	17.84	17.25	17.74	16.61
Fe ₂ O ₃	12.60	11.04	9.60	8.81	6.30	6.86	7.56	7.83	6.91	6.53
MnO	0.19	0.17	0.18	0.16	0.15	0.12	0.14	0.15	0.12	0.14
MgO	4.57	3.89	3.20	3.07	2.03	2.27	4.47	4.45	4.92	3.70
CaO	8.13	7.62	5.98	6.53	5.23	4.94	8.31	8.69	7.44	6.59
Na ₂ O	3.28	3.51	4.30	4.06	4.46	3.70	3.15	3.06	3.16	3.67
K ₂ O	0.48	0.77	0.74	0.96	1.40	1.78	0.85	0.83	0.97	1.22
P ₂ O ₅	0.11	0.33	0.14	0.41	0.40	0.19	0.23	0.20	0.17	0.22
LOI	1.80	0.80	2.22	1.01	1.22	1.47	0.91	0.76	1.23	1.46
SUM	100.30	99.31	100.32	99.07	99.36	98.55	99.06	98.57	98.78	99.18
FeO*/MgO	2.48	2.55	2.70	2.58	2.79	2.72	1.52	1.58	1.26	1.58
Li	10.7	14.7	13.4	16.0	11.7	23.1	14.5	13.5	12.8	15.0
Sc	40	32	28	28	21	21	28	30	25	23
Ti	8550	7820	9610	7740	6440	4890	4710	4710	4200	4350
V	364	272	123	163	78	159	154	162	122	102
Cr	4.7	3.8	0.76	3.1	1.8	8.6	102	111	100	85
Co	28	22	9.6	13.7	7.5	14.1	21	21	22	15
Ni	3.6	2.3	2.0	1.7	1.9	2.0	21	18.4	41	20
Cu	12.9	14.3	6.0	6.8	4.8	13.5	14.6	14.5	13.8	8.6
Zn	106	94	107	93	83	69	74	75	81	74
Ga	21	20	22	20	20	17.9	18.2	17.9	17.6	18.0
Rb	10.6	19.7	20	27	44	60	26	24	35	38
Sr	338	332	367	333	312	237	295	271	267	271
Y	17.3	20	20	24	27	22	20	20	19.3	22
Zr	90	114	204	180	231	110	149	138	154	187
Nb	2.7	3.5	6.2	5.1	6.8	5.5	4.0	3.8	4.2	5.1
Cs	0.41	0.98	0.83	1.3	1.8	3.1	0.96	0.96	1.7	1.5
Ba	152	200	279	253	365	405	313	212	285	302
La	5.1	9.3	8.3	11.9	16.3	15.7	10.0	9.6	9.7	12.7
Ce	14.5	21	19.6	27	36	33	23	22	22	28
Pr	1.8	2.9	2.5	3.6	4.5	3.9	3.0	2.8	2.8	3.5
Nd	8.6	12.8	11.0	16.0	19.3	15.9	12.9	12.4	11.9	15.1
Sm	2.5	3.2	3.0	4.0	4.6	3.5	3.2	3.1	2.8	3.7
Eu	1.1	1.3	1.4	1.5	1.5	1.1	1.2	1.1	1.1	1.3
Gd	2.8	3.5	3.4	4.4	5.0	3.7	3.6	3.5	3.2	3.9
Tb	0.48	0.55	0.56	0.68	0.74	0.58	0.56	0.54	0.51	0.61
Dy	3.1	3.5	3.7	4.3	4.7	3.7	3.6	3.5	3.3	3.9
Ho	0.67	0.73	0.76	0.87	0.97	0.77	0.75	0.73	0.69	0.81
Er	1.9	2.1	2.2	2.5	2.8	2.3	2.2	2.1	2.0	2.4
Tm	0.27	0.31	0.32	0.36	0.40	0.34	0.32	0.31	0.30	0.34
Yb	1.8	2.0	2.2	2.4	2.7	2.3	2.1	2.1	2.0	2.3
Lu	0.27	0.29	0.32	0.36	0.41	0.35	0.32	0.32	0.31	0.35
Hf	2.3	2.7	4.6	4.0	5.2	3.0	3.4	3.2	3.5	4.2
Pb	3.4	6.4	8.2	4.9	7.0	14.4	4.0	4.1	5.2	5.9
Th	1.5	2.1	3.1	3.0	4.8	6.2	2.8	2.7	3.4	4.0
U	0.38	0.53	0.73	0.73	1.2	1.4	0.67	0.65	0.84	0.97
Rb/Sr	0.03	0.06	0.05	0.08	0.14	0.25	0.09	0.09	0.13	0.14
Eu/Eu*	1.30	1.13	1.31	1.10	1.01	0.90	1.06	1.00	1.07	1.01

Table 4. Summary of estimates for the pre-eruptive temperature of the Oruanui melt-dominant magma body yielded by various mineral equilibrium based techniques

Estimates of outermost rim crystallisation temperatures (T°C)	Number of estimates	Mean temp (°C)	Temp range (°C)	Modal values	Model details and/or parameters
HSR magma					
Plag-melt (rims)	158	787	783 - 797		Putirka (2008), eqn 24a, matrix glass and 5.5 wt% melt H ₂ O
Opx-melt (rims)	222	785	779 - 793		Putirka (2008), eqn 28a, matrix glass and 5.5 wt% melt H ₂ O
Amphibole compositions (rims)	121	784	751 - 851	770-780	Ridolfi et al. (2010) model
Fe-Ti oxides	378	796	765 - 836	790	Ghiorso and Evans (2008) model
LSR magma					
Plag-melt (rims)	47	831		820-840	Putirka (2008), eqn 24a, matrix glass and 6 wt% melt H ₂ O
Opx-melt (rims)	49	830		820-840	Putirka (2008), eqn 28a, matrix glass and 6 wt% melt H ₂ O
Amphibole compositions (rims)	59	837	783-903	820-850	Ridolfi et al. (2010) model
Fe-Ti oxides	61	840	812-910	840	Ghiorso and Evans (2008) model

Table 5. Summary of oxygen fugacity and melt H₂O estimates for the final pre-eruptive conditions in the Oruanui HSR and LSR magmas.

Estimates of oxygen fugacity (Δ NNO)	n	Mean (Δ NNO)	Range (Δ NNO)	Model details and/or parameters
HSR magma				
Fe-Ti oxides	379	-0.20	-0.52 to 0.03	Ghiorso and Evans (2008)
Amphibole rims	121	0.90	0.40 to 1.20	Ridolfi et al. (2010)
LSR magma				
Fe-Ti oxides	59	0.5	0.2-0.8	Ghiorso and Evans (2008)
Amphibole rims	59	0.5	0.2-0.8	Ridolfi et al. (2010)
Estimates of H ₂ O melt (wt %)		Mean (wt %)	Range	Model details and/or parameters
HSR magma				
Amphibole rims	121	5.5	5.0 - 6.2	Ridolfi et al. (2010) model
Quartz-host melt inclusions	66	4.5	3.8 - 5.1	Liu et al. (2006), H ₂ O in quartz-hosted melt inclusions measured by FTIR
Empirical H ₂ O solubility model	-	5.2	3.9 - 6.6	Moore et al. (1998), model run at 796 °C and 90-240 MPa, mean value at 150 MPa
Glass H ₂ O by difference	141	4.5	0.6 - 7.4	Difference from 100 wt% of raw EPMA totals for pumice matrix glass
Plag-melt equilibrium		4.6		Anorthite-method of Housh and Luhr (1991), reported in Wilson et al. (2006)
Plag-melt equilibrium		6.2		Albite-method of Housh and Luhr (1991), reported in Wilson et al. (2006)
Plag-melt equilibrium		5.5		Method of Waters and Lange (2015) with 790 °C and outer rim values of An _{38±2}
LSR magma				
Amphibole rims	59	6.0	5.6-6.5	Ridolfi et al. (2010) model
Plag-melt equilibrium		5.3		Method of Waters and Lange (2015) with 840 °C and outer rim values of An ₄₄