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2	The Reaction between CH <sub>3</sub> O <sub>2</sub> and OH Radicals: Product
3	Yields and Atmospheric Implications
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## 28 Abstract

29 The reaction between  $CH_3O_2$  and OH radicals has been shown to be fast and to play an appreciable 30 role for the removal of CH<sub>3</sub>O<sub>2</sub> radials in remote environments such as the marine boundary layer. Two 31 different experimental techniques have been used here to determine the products of this reaction. The 32 HO<sub>2</sub> yield has been obtained from simultaneous time-resolved measurements of the absolute concentration of CH<sub>3</sub>O<sub>2</sub>, OH and HO<sub>2</sub> radicals by cw-CRDS. The possible formation of a Criegee 33 34 intermediate has been measured by broadband cavity enhanced UV absorption. A yield of  $\phi_{HO_2}$ = 35  $(0.8\pm0.2)$  and an upper limit for  $\phi_{Criegee} = 0.05$  has been determined for this reaction, suggesting a minor 36 yield of methanol or stabilized trioxide as a product. The impact of this reaction on the composition of 37 the remote marine boundary layer has been determined by implementing these findings into a box 38 model utilizing the Master Chemical Mechanism v3.2, and constraining the model for conditions 39 found at the Cape Verde Atmospheric Observatory in the remote tropical Atlantic Ocean. Inclusion of the CH<sub>3</sub>O<sub>2</sub>+OH reaction into the model results in up to 30% decrease in the CH<sub>3</sub>O<sub>2</sub> radical 40 41 concentration while the HO<sub>2</sub> concentration increased by up to 15%. Production and destruction of O<sub>3</sub> 42 are also influenced by these changes, and the model indicates that taking into account the reaction 43 between CH<sub>3</sub>O<sub>2</sub> and OH leads to a 6% decrease of O<sub>3</sub>.

44

## 45 TOC Graph



## 47 Introduction

Peroxy radicals, RO<sub>2</sub>, are key species in atmospheric and low temperature combustion chemistry. 48 49 They are reactive intermediates formed during the oxidation of all hydrocarbons. In the atmosphere, 50 their subsequent fate depends on the concentration of NOx. In polluted environments at high NOx 51 they will rapidly be removed by reaction with NO to form NO<sub>2</sub>. Subsequent photolysis of NO<sub>2</sub> will 52 lead to an increase in O<sub>3</sub> concentration. With decreasing NO concentration, other reaction pathways 53 become competitive for RO<sub>2</sub>: the self- and cross-reactions with HO<sub>2</sub> or other RO<sub>2</sub> radicals, leading mostly to stable products<sup>1</sup>. Recently, a new reaction pathway has been suggested as a competitive fate 54 for RO<sub>2</sub> in clean environments, namely the reaction of RO<sub>2</sub> with OH radicals<sup>2,3</sup>. Archibald *et al.*<sup>2</sup> have 55 investigated in a modeling study the impact of including the reaction between RO<sub>2</sub> and OH on the 56 57 composition of the Marine Boundary Layer (MBL). Different scenarios were run using the BAMBO model, based on the Master Chemical Mechanism, MCM<sup>4</sup>. Different rate constants and reaction paths 58 59 were considered for RO<sub>2</sub> up to C<sub>4</sub>, and for the simplest one, CH<sub>3</sub>O<sub>2</sub>, leading to the following products:

60

 $OH + CH_3O_2 \rightarrow CH_2OO + H_2O$ (R1a)

(R1b)

(R1c)

 $OH + CH_3O_2 \rightarrow CH_3O + HO_2$ 

 $OH + CH_3O_2 \rightarrow CH_3OH + O_2$ 

63 Archibald *et al.* found for all possible reaction channels only a small, negligible effect on the mixing ratios of O<sub>3</sub>, NO<sub>x</sub>, OH and other trace gas species in the marine boundary layer. However, a substantial 64 65 increase in the mixing ratios of HCOOH was predicted (from 0.16 ppt in the base case, i.e. absence of 66 (R1), to 25.5 ppt at the scenario with the largest rate constant), if the reaction pathway were the 67 formation of the Criegee radical (R1a). A strong increase in the mixing ratio of CH<sub>3</sub>OH (from 37 ppt 68 in the base case, i.e. absence of (R1), to 294 ppt at the largest rate constant scenario) resulted if the 69 major pathway were (R1c): the impact on stable species was the smallest if (R1b) was the major path. Ab initio calculations<sup>5,6</sup> predict pathway (R1b) to be the major reaction, with possible minor 70 contributions of (R1c). Very recently, Müller et al.<sup>6</sup> have implemented the title reaction into a global 71 72 atmospheric chemistry model and have shown, that a yield of around 18% for (R1c) could explain a 73 large missing source of methanol over remote oceans. It seems therefore important to determine 74 experimentally the product distribution of (R1).

The rate constant has been measured recently for the first time for the two simplest peroxy radicals<sup>$$-10$$</sup>

76 
$$CH_3O_2 + OH \rightarrow \text{ products}$$
 (R1)  
77  $C_2H_5O_2 + OH \rightarrow \text{ products}$  (R2)

Large rate constants  $k_1 = (2.8\pm1.4)\times10^{-10}$  cm<sup>3</sup>s<sup>-1</sup> and  $k_2 = (1.2\pm0.3)\times10^{-10}$  cm<sup>3</sup>s<sup>-1</sup> have been measured by our group for CH<sub>3</sub>O<sub>2</sub><sup>7</sup> and C<sub>2</sub>H<sub>5</sub>O<sub>2</sub><sup>8</sup>, respectively, although  $k_1$  has very recently been revised by our group to a lower value<sup>9</sup>:  $k_1=(1.60\pm0.4)\times10^{-10}$  cm<sup>3</sup>s<sup>-1</sup>. The rate constant  $k_1$  has also recently been determined by Yan *et al.*<sup>10</sup> with an even lower value of  $k_1=(0.84\pm0.17)\times10^{-10}$  cm<sup>3</sup>s<sup>-1</sup> deduced from 82 fitting UV-absorption signals to a complex mechanism. Because these reactions are fast, they can 83 become competitive with other reaction paths for RO<sub>2</sub> radicals. Fittschen *et al.*<sup>3</sup> have integrated (R1) into a detailed box model utilizing the MCM<sup>11</sup> and have determined its importance as a sink for 84 CH<sub>3</sub>O<sub>2</sub> radicals in remote marine environments. Running the model using conditions found in the 85 remote tropical marine boundary layer during a field campaign at Cape Verde in 2007<sup>11</sup> have shown 86 that using the initially reported rate constant from Bossolasco et al.<sup>7</sup> results in up to 30% of all CH<sub>3</sub>O<sub>2</sub> 87 radicals being removed through  $(R1)^{3,6}$ . The revised rate constant of Assaf *et al.*<sup>9</sup> will decrease the 88 89 importance of (R1); yet, this reaction is still a substantial sink for CH<sub>3</sub>O<sub>2</sub>. However, in order to 90 evaluate the impact of this reaction on the composition of the atmosphere it is essential to identify the 91 reaction products. In the current paper we present experiments for the determination of the yields of 92 two possible products:  $HO_2$  radicals from path (R1b) and the Criegee intermediate from path (R1a). 93 The results of these experiments have been implemented into a box model utilizing the MCM, and the 94 impact on the composition of the remote atmosphere has been determined.

95

## 96 Experimental section

#### $HO_2$ yield by cw-cavity ring down spectroscopy (CRDS)

An improved version of a well-described set-up<sup>12-15</sup>, installed at the University of Lille, has been used 98 99 to quantify simultaneously OH, HO<sub>2</sub> and CH<sub>3</sub>O<sub>2</sub> radicals by two cavity ring down spectroscopy (cw-100 CRDS) absorption paths, crossing in a slow flow reactor and at a small angle to an Excimer laser 101 photolysis beam. One cavity was equipped with mirrors optimized for a detection of  $CH_3O_2$  radicals at 102 7489.16 cm<sup>-1</sup>, the other beam path was used for a sequential detection of HO<sub>2</sub> radicals<sup>16</sup> at 6638.205 cm<sup>-1</sup> and OH radicals<sup>17</sup> at 7028.831 cm<sup>-1</sup>. The sequential detection of both species, OH and HO<sub>2</sub>, is fast 103 and simple, as the mirrors cover the wavelength region for both species, OH and  $HO_2$ . The entire beam 104 105 path is fibered coupled and changing the wavelength can be accomplished in less than one minute by 106 switching the fiber connectors of the distributed feedback (DFB) lasers without re-alignment of the 107 cavity. Typical kinetic decays are obtained by acquiring ring-down events over several photolysis 108 pulses (20 to 50), and ring-down times  $\tau$  are converted to absorbance  $\alpha$  by the following equation:

109 
$$\alpha_t = [A]_t \times \sigma = \frac{R_L}{c} \left( \frac{1}{\tau_t} - \frac{1}{\tau_0} \right)$$
 [Eq. 1]

110 The quantity  $\tau_0$  and  $\tau_t$  are the ring-down times in the presence and absence of the absorbing species, in 111 our case before and after the photolysis pulse, respectively;  $\sigma$  is the absorption cross section of the 112 absorbing species;  $R_L$  is the ratio between cavity length (79 cm) and absorption length (in our case the

- overlap of photolysis laser and near IR laser which is 37.7 cm; *c* is the speed of light. For more details
- 114 on the synchronization and acquisition see Parker *et al.*<sup>14</sup>.
- 115 Radicals have been generated from the 248 nm photolysis of XeF<sub>2</sub>, which is a solid and easy to handle.
- 116 Mixtures have been prepared by introducing a few crystals (0.3-0.5 g) into a home-made Teflon bag,
- which was then filled with around 50 l helium. Evaporation of the crystals led to  $XeF_2$  concentrations
- of around 1Torr, well below its vapor pressure of ( $\approx 5 \text{ Torr}^{18}$ ). Stable flows of this mixture were
- introduced into the photolysis reactor through a Teflon needle valve: the flow rate was determined
- 120 through measurement of pressure increase into a known volume. More details on the use of  $XeF_2$  as a
- 121 photolytic source of F-atom can be found in a recent paper $^{9}$ .
- A total of four experiments has been carried out, all at a total pressure of 50 Torr and 295K. Theconcentrations are summarized in Table 1.

124 Table 1: Initial concentrations of stable species and maximum radical concentrations from cw-CRDS125 measurements

$O_2$ / $10^{18}$ cm <sup>-3</sup>	He / $10^{17}$ cm <sup>-3</sup>	$CH_4 / 10^{15} cm^{-3}$	$^{a}XeF_{2} / 10^{14} cm^{-3}$	$^{b}CH_{3}O_{2} / 10^{12} \text{ cm}^{-3}$	<sup>b</sup> OH / 10 <sup>12</sup> cm <sup>-3</sup>	$^{b}\text{HO}_{2,\text{max}}$ / 10 <sup>12</sup> cm <sup>-3</sup>
1.30	3.0	1.00	1.94	4.3	0.45	0.58
1.11	5.0	0.82	3.72	8.2	1.7	2.2
1.06	5.5	0.86	4.17	8.9	1.9	2.4
0.86	7.6	0.66	6.04	12.1	4.8	5.2

<sup>a</sup>: estimated from weighted XeF<sub>2</sub> in Teflon bag, <sup>b</sup>: as obtained from cw-CRDS measurements using the absorption cross sections given below

128

### 129 Broadband cavity-enhanced UV spectroscopy

The reaction of OH with CH<sub>3</sub>O<sub>2</sub> was also investigated by time-resolved broadband cavity-enhanced 130 UV spectroscopy to test for the possibility of production of formaldehyde oxide, CH<sub>2</sub>OO. The 131 132 experimental apparatus, which was developed at Sandia National Lab, has been described in detail previously<sup>19,20</sup>, and only a brief account is given here. The continuous-wave broadband optical buildup 133 cavity was configured to simultaneously probe the spectral range  $\lambda = 300 - 450$  nm with average 134 135 effective path length of  $\sim 40$  m and spectral resolution of 1.5 nm. The optical cavity was integrated into 136 a gas-phase flow reactor with independent control over the experimental temperature, pressure, and 137 sample mixture composition. All experiments were performed at T = 293 K and total P = 30 Torr, and 138 transient UV spectra were recorded with time resolution of 30 µs.

## 139 Results and discussion

140 This section is divided into three sections: details on the determination of HO<sub>2</sub> yield, obtained by cw-

141 CRDS in Lille, are given in the first part; the determination of an upper limit for the yield of Criegee

intermediate by UV absorption spectroscopy, obtained at Sandia National Lab, is given in the secondpart; the implementation in the MCM and the impact of the new findings onto the composition of the

part, the implementation in the MCM and the impact of the new infinings onto the composition of t

troposphere are presented in the third part.

### 145 The yield of HO<sub>2</sub> radicals as products of the reaction $CH_3O_2 + OH$

We have employed a set-up combining two cw-CRDS paths within a laser photolysis reactor to directly quantify the yield of HO<sub>2</sub> radicals,  $\phi_{1b}$ . For this purpose, the reaction has been initiated by the concurrent photolytic production of CH<sub>3</sub>O<sub>2</sub> and OH radicals through the following reaction sequence:

149	$XeF_2 + hv_{248nm} \rightarrow Xe + 2 F$	(R3)
150	$F + CH_4 \rightarrow HF + CH_3$	(R4)
151	$CH_3 + O_2 + M \rightarrow CH_3O_2 + M$	(R5)
152	$F + H_2O \rightarrow HF + OH$	(R6)

The ratio between the initial concentrations of CH<sub>3</sub> and OH radicals is given by the ratio of the rates of (R4) to (R6), i.e. the products of the rate constant and the corresponding precursor concentration (CH<sub>4</sub> or H<sub>2</sub>O). The subsequent reaction (R5) is under our conditions ( $\approx$ 30 Torr O<sub>2</sub>) very fast (k'<sub>5</sub>  $\approx$  1.3×10<sup>5</sup> cm<sup>3</sup>s<sup>-1</sup>) and it can be considered that all CH<sub>3</sub> radicals are converted to CH<sub>3</sub>O<sub>2</sub>. CH<sub>3</sub>O<sub>2</sub> radicals have been quantified on one absorption path, while HO<sub>2</sub> and OH radicals have been quantified sequentially on the second path.

The line strength of a ground state  $X^2\Pi_{1/2}$  transition of OH radicals in the near infrared region at 159 7028.831 cm<sup>-1</sup> has been very recently determined by Assaf and Fittschen<sup>17</sup> to be S =  $4.07 \times 10^{-21}$  cm, 160 with a peak absorption cross section at 50 Torr helium of  $\sigma_{OH, 50 \text{ Torr He}} = (1.54\pm0.1) \times 10^{-19} \text{ cm}^2$ . The 161 162 current work has been carried out in the presence of high O<sub>2</sub> concentrations (25 to 40 Torr, the 163 complement to 50 Torr being mostly He from the XeF<sub>2</sub> flow), and therefore the peak absorption cross 164 section will be somewhat lower due to increased pressure broadening of O2 compared to He. 165 Unfortunately, the  $O_2$  broadening coefficient  $\gamma_{O_2}$  for the line used in this work is currently not known and has been estimated. In the Hitran data base<sup>21</sup> one finds the broadening coefficient in air to be  $\gamma_{air}$  = 166 0.095 cm<sup>-1</sup>atm<sup>-1</sup>. Taking this value, i.e. considering an identical broadening efficiency for O<sub>2</sub> and N<sub>2</sub>, 167 one can calculate from the line strength<sup>17</sup> and considering a Voigt profile  $\sigma_{OH, 50 \text{ Torr air}} = 1.16 \times 10^{-19} \text{ cm}^2$ . 168 Taking into account the varying He/O<sub>2</sub> ratios, the peak absorption cross section for the different 169 experiments vary between  $(1.23 - 1.34) \times 10^{-19}$  cm<sup>2</sup>. The average value of  $\sigma_{OH} = 1.27 \times 10^{-19}$  cm<sup>2</sup> of has 170

- been used in [Eq. 1] to convert  $\alpha_{OH}$  into [OH]. An uncertainty of 20% is estimated for this absorption
- 172 cross section.
- 173 The absorption cross section for  $CH_3O_2$  has been taken from Farago *et al.*<sup>22</sup>. For  $HO_2$  it has been taken
- 174 from Thiébaud et al.<sup>23</sup>: this line has an absorption cross section nearly two times higher than
- surrounding lines and is probably the convolution of two nearly perfectly overlapping transition. As a
- 176 result the pressure broadening in Helium<sup>24</sup> or in  $O_2^{25}$  is very small. The uncertainty of  $\sigma_{HO_2}$  is estimated
- to be less than 10%.



**Figure 1:** Absolute concentrations of  $CH_3O_2$  (green symbols),  $HO_2$  (red symbols) and OH (blue symbols) following the photolysis of  $[XeF_2] = 3.7 \times 10^{14} \text{ cm}^{-3}$  in the presence of  $[CH_4] = 6.8 \times 10^{14} \text{ cm}^{-3}$ ,  $[H_2O] = 7 \times 10^{14} \text{ cm}^{-3}$ ,  $[He] = 5.0 \times 10^{17} \text{ cm}^{-3}$  and  $[O_2] = 1.1 \times 10^{18} \text{ cm}^{-3}$ . Absorption cross sections of  $\sigma_{CH_3O_2} = 3.33 \times 10^{-20} \text{ cm}^2$ ,  $\sigma_{HO_2} = 2.72 \times 10^{-19} \text{ cm}^2$  and  $\sigma_{OH} = 1.27 \times 10^{-19} \text{ cm}^2$  have been used for conversion of absorbances  $\alpha$  for the three species. The black lines represent adjustment to the model in **Table 2**.

Figure 1 shows a typical experiment at 50 Torr total pressure, using the following concentrations: 185  $[XeF_2] = 3.7 \times 10^{14} \text{ cm}^{-3}, [CH_4] = 6.8 \times 10^{14} \text{ cm}^{-3}, [H_2O] = 7 \times 10^{14} \text{ cm}^{-3}, [He] = 5.0 \times 10^{17} \text{ cm}^{-3} \text{ and } [O_2] = 7 \times 10^{14} \text{ cm}^{-3}, [He] = 5.0 \times 10^{17} \text{ cm}^{-3}$ 186  $1.1 \times 10^{18}$  cm<sup>-3</sup>. The concentration time profiles of all three species are presented. It can be seen that the 187 188 initial fast decay in CH<sub>3</sub>O<sub>2</sub> concentration is on the same order of magnitude as the decay of the initial 189 OH concentration. It can also be seen that the maximum  $HO_2$  concentration is slightly higher than the 190 initial OH concentration, suggesting a high yield of  $HO_2$  radicals produced in (R1). In order to determine the HO<sub>2</sub> yield,  $\phi_{HO_2}$ , the concentration time profiles have been adjusted to a simple model, 191 192 with the reactions and rate constants shown in Table 2.

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**Table 2**: Reaction mechanism used to fit CH<sub>3</sub>O<sub>2</sub>, OH and HO<sub>2</sub> concentration time profiles

No.	Reaction		$k / \mathrm{cm}^3 \mathrm{s}^{-1}$	Reference
1a	$CH_3O_2 + OH$	$\rightarrow$ CH <sub>2</sub> OO + H <sub>2</sub> O	0	
1b		$\rightarrow HO_2 + CH_3O$	$(1.25\pm0.3) \times 10^{-10}$	This work and <sup>9</sup>

1c	$\rightarrow$ products *	$(0.35\pm0.3) \times 10^{-10}$	
5	$CH_3 + O_2 (+M) \rightarrow CH_3O_2 (+M)$	1.4 ×10 <sup>-13,**</sup>	26,27
7	$CH_3O + O_2 \rightarrow CH_2O + HO_2$	$1.92 \times 10^{-15}$	28
8a	$2 \text{ CH}_3\text{O}_2 \rightarrow 2 \text{ CH}_3\text{O} + \text{O}_2$	1.3 ×10 <sup>-13</sup>	28
8b	$2 \text{ CH}_3\text{O}_2 \rightarrow \text{ products}$	$2.2 \times 10^{-13}$	28
9	$CH_3O_2 + HO_2 \rightarrow products$	5.2 ×10 <sup>-12</sup>	28
10	$2 \text{ HO}_2 \rightarrow \text{product}$	1.7 ×10 <sup>-12</sup>	29
11	$OH + HO_2 \rightarrow product$	1.0×10 <sup>-10</sup>	29,17
12	$HO_2 \rightarrow diffusion$	$8 - 22 \text{ s}^{-1}$	This work

\*no information about the products can be obtained from the current experiments \*\*Rate constant in 50
 Torr helium

Four experiments have been carried out with changing XeF<sub>2</sub> concentrations, leading to initial total radical concentrations between  $5 - 17 \times 10^{12}$  cm<sup>-3</sup> and with ratios of CH<sub>3</sub>O<sub>2</sub> / OH between 3 and 12 (see Table 1). All four experiments can be very well reproduced with the mechanism and rate constants from **Table 2**, i.e.  $\phi_{HO_2}$ =0.8.

The uncertainty of the rate constant of (R7), leading to formation of a second HO<sub>2</sub> radical, is given<sup>28</sup> 202 203 with  $\pm$ 50%. However, the O<sub>2</sub> concentration in our experiments is high, so that this reaction is not the 204 rate limiting step and an uncertainty of  $k_7$  has only a very minor impact on the HO<sub>2</sub> profile. Also, the reactions of CH<sub>3</sub>O with OH or HO<sub>2</sub> are probably not important, even though their rate constants are 205 not well known: taking an estimated rate constant  $^{30}$  of  $3 \times 10^{-11}$  cm<sup>3</sup> s<sup>-1</sup> for the reaction with OH, even 206 207 under the most unfavorable conditions (high radical and low O<sub>2</sub> concentration), only 1% of the initial 208 OH radicals would react with CH<sub>3</sub>O. The influence of (R8) and (R9) are very limited at short reaction 209 times, and therefore any uncertainty of  $k_8$  and  $k_9$  would have a negligible influence on the retrieval of 210  $\phi_{HO_2}$ . Equally, the uncertainty of  $k_{II}$  has only a limited influence on  $\phi_{HO_2}$ : while the reaction itself 211 consumes around 10% of the initial OH radicals at the experiment with the highest, initial radical 212 concentration, the estimated error of 15% on  $k_{II}$  would lead to a change of less than 3% in  $\phi_{HO_2}$ .

The major uncertainty of the retrieved value of  $\phi_{HO_2}$  comes from uncertainties in the absorption cross sections used in [Eq. 1] to convert the absorbance  $\alpha$  of OH and HO<sub>2</sub> radicals into absolute concentrations. While the absorption cross section of HO<sub>2</sub> radicals is well determined<sup>24,31,32</sup> and the uncertainty is estimated to be less than 10%,  $\sigma_{OH}$  has only been determined once<sup>17</sup>. Additionally, the value used in this work needs to take into account the increased pressure broadening by the high O<sub>2</sub> concentrations. Therefore, the uncertainty on the OH absorption cross section is estimated to be 20%. 219 The impact of this uncertainty on the retrieved  $\phi_{HO_2}$  is illustrated in the left graph of Figure 2, using the same experiment as shown in Figure 1 (zoomed on OH and HO<sub>2</sub> only, the CH<sub>3</sub>O<sub>2</sub> profile is not 220 shown). The black lines show the model using  $\sigma_{OH}=1.27\times10^{-19}$  cm<sup>2</sup>, while for the blue and red lines the 221 absorption cross section for OH has been varied by 20%. Using the lower limiting value of  $\sigma_{OH}$  in [Eq. 222 223 1] leads to higher OH concentration (red dashed line: the corresponding experimental data points, i.e. 224 the black circles multiplied by 1.2, are not plotted in Figure 2 for clarity). The model allows a good 225 reproduction of OH and HO<sub>2</sub> profiles, but in order to make up for the higher initial OH concentration, the yield of HO<sub>2</sub> radicals must be decreased ( $\phi_{HO_2} = 0.67$ ). The higher value for  $\sigma_{OH}$  (dashed blue line), 226 227 leading to lower OH concentrations, also allows a satisfactory reproduction of the HO<sub>2</sub> profile, now 228 with  $\phi_{HO_2} = 0.96$ .



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Figure 2: Concentration time profiles for OH (decaying profile) and HO<sub>2</sub> (rising profile) from the experiment shown in Figure 1, reproduced with the model utilizing the chemical mechanism from Table 2. Left graph: Influence of a 20% uncertainty of  $\sigma_{OH}$  on the HO<sub>2</sub> concentration time profile: full black line OH concentrations obtained by using  $\sigma_{OH}=1.27\times10^{-19}$  cm<sup>2</sup>, red dashed line with  $\sigma_{OH}$ decreased by 20%, blue dashed line with  $\sigma_{OH}$  increased by 20% (the corresponding experimental OH data are not plotted for clarity).  $\phi_{HO_2} = 0.67$ , 0.8 and 0.96 for the red, black and blue lines, respectively. Right graph: Impact of the rate constant  $k_1$  on both profiles: all rate constants as in Table 2, only  $k_1$ has been changed.

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The right graph of **Figure 2** shows the sensitivity of the experimental data to the rate constant  $k_1$ , 239 whereby the total value has been varied between 1 and  $2 \times 10^{-10}$  cm<sup>3</sup>s<sup>-1</sup>. In order to reproduce the 240 241 absolute HO<sub>2</sub> concentration,  $k_{1b}$  and  $k_{1c}$  have been adjusted to (9/1), (12.5/3.5) and (15/5) × 10<sup>-11</sup> cm<sup>3</sup>s<sup>-</sup> 242 <sup>1</sup> for the blue, black and red curves. The blue line represents the upper limit of the rate constant recently published by Yan *et al.*<sup>10</sup>. A change in  $k_1$  has a strong impact on the predicted HO<sub>2</sub> profile in 243 244 our experiment. Not only would the rise time of HO<sub>2</sub> slow down with a decrease in  $k_l$ , but also the 245 final, maximum HO<sub>2</sub> concentration would decrease. This is because the conversion of OH into HO<sub>2</sub> 246 through (R1) is not only governed by  $\phi_{HO_2}$ , but also by the rate constant  $k_i$ : a lower rate constant leads to a lower conversion because competing reactions gain in importance, especially the fast reaction 247 (R11). The black line shows the best fit for both species with  $k_1 = 1.6 \times 10^{-10} \text{ cm}^3 \text{s}^{-1}$ . The dotted red lines 248 correspond to the model using  $k_1 = 2 \times 10^{-10} \text{ cm}^3 \text{s}^{-1}$ : OH and HO<sub>2</sub> profiles are barely reproduced with this 249

- 250 faster rate constant, the retrieved HO<sub>2</sub> yield would be similar with  $\phi_{HO_2} = 0.75$ . A decrease to  $k_1 =$
- $1 \times 10^{-10}$  cm<sup>3</sup>s<sup>-1</sup> results in a very poor reproduction the OH and HO<sub>2</sub> profiles: the time evolution is too 251
- slow for both species. With this rate constant,  $\phi_{HO_2}$  would rise to  $\phi_{HO_2} = 0.9$  because other loss 252
- 253 reactions, especially (R11), gain in importance.
- 254 In conclusion, our determination of the HO<sub>2</sub> yield at a total pressure of 50 Torr O<sub>2</sub> / He shows that
- 255 HO<sub>2</sub> is the major product with  $\phi_{\text{HO}_2} = 0.8 \pm 0.2$ .
- 256

#### Determination of an upper limit for the yield of Criegee intermediate 257 $CH_2OO$ 258

The Criegee intermediate CH<sub>2</sub>OO is the product of channel (R1a). Theoretical calculations predict this 259 pathway to be of minor importance<sup>5,6</sup>, and recent experiments of Yan et al.<sup>10</sup> using UV absorption 260 spectroscopy have determined an upper limit of 5% on channel (R1b). Experiments have been carried 261 262 out at Sandia National Laboratory. XeF<sub>2</sub> was not available to initiate the reaction by the same reaction 263 sequence as in Lille, also the absorption cross section would be even lower at the used wavelength. 264 Therefore, the reaction sequence was initiated by 266-nm photolysis of acetone vapor in He bath gas in the presence of 2.5 Torr  $(8.25 \times 10^{16} \text{ cm}^{-3})$  of O<sub>2</sub>. The reaction sequence begins with concurrent 265 photolytic production of methyl and acetyl radicals (at an estimated ratio of ~1.5:1) and subsequent 266 rapid reaction of these radicals with excess O<sub>2</sub> forms CH<sub>3</sub>O<sub>2</sub> and OH<sup>33-35</sup>; within less than 1 ms, as 267 summarized below: 268

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	r 1 9
~	0.0

269	$CH_3COCH_3 + hv \rightarrow CO + 2 CH_3$	~20%	(R 13a)
270	$\rightarrow$ CH <sub>3</sub> CO + CH <sub>3</sub>	~80%	(R 13b)
271	$CH_3 + O_2 (+ M) \rightarrow CH_3O_2 (+ M)$	~100%	(R 5)
272	$CH_3CO + O_2 \rightarrow OH + co-products$	10 - 50%	(R 14)

273

274 The yield of OH radicals in (R14) is pressure dependant and still controversially discussed: Carr et al.<sup>34</sup> for example report a yield of  $\phi_{OH} \approx 0.5$  for (R14) at 50 Torr, while recent publications of 275 Papadimitriou et al.<sup>36</sup> and Bouzidi et al.<sup>25</sup> report yield of 0.1 - 0.2 at 20 Torr. Therefore, CH<sub>3</sub>O<sub>2</sub> is 276 generated by this system in excess over OH (ratio  $CH_3O_2$  : OH  $\approx$  (3-15) : 1). Two different 277 278 experimental runs were performed. In the first, the initial molar fraction of acetone was adjusted to produce an estimated  $[CH_3O_2] = 4.8 \times 10^{12} \text{ cm}^{-3}$  and a lower limit for  $[OH] \ge 3.2 \times 10^{11} \text{ cm}^{-3}$ ; in the 279 second, the acetone concentration was increased to form  $[CH_3O_2] = 8.9 \times 10^{12} \text{ cm}^{-3}$  and  $[OH] \ge 6 \times 10^{11}$ 280  $\mathrm{cm}^{-3}$ . 281

The resulting transient spectra, averaged over kinetic times 0-5 ms, are presented in Figure 3. The 282 spectra reveal a single broad absorption feature with maximum intensity at the low-wavelength edge 283

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of the accessible experimental range, 300 nm, which most likely arises solely from the methyl peroxy radical. Simulations (also shown in **Figure 3**) of the predicted transient signal using the reported absorption cross-section of CH<sub>3</sub>O<sub>2</sub> at 300 nm<sup>37</sup>,  $\sigma = 3.3 \times 10^{-19}$  cm<sup>2</sup>, are in qualitative agreement with the observed spectra, which confirms our expectations of methyl peroxy radical concentration formed by reactions (R5) and (R13). The time evolution of the CH<sub>3</sub>O<sub>2</sub> absorption is included as the inset in **Figure 3**.

Small Criegee intermediates such as formaldehyde oxide and acetaldehyde oxide were recently 290 shown<sup>20,38-40</sup> to have very strong UV absorption bands in the 300 - 400 nm range with peak absorption 291 cross-sections of the B $\leftarrow$ X transition of  $\sim 1 \times 10^{-17}$  cm<sup>2</sup>. Thus, UV absorption is a sensitive detection 292 method for Criegee intermediates, and our experiments have a detection limit for CH2OO below 293  $\sim$ 5×10<sup>8</sup> cm<sup>-3</sup>. Transient absorption due to CH<sub>2</sub>OO is expected to decay on the timescale of a few ms, 294 primarily due to the reaction with excess acetone vapor<sup>41</sup>, yet we observe essentially no decaying 295 296 signals in our experiment. The presence of transient  $CH_3O_2$  absorption somewhat obscures our probe spectral region; nonetheless, based on the signal near the peak of CH<sub>2</sub>OO spectrum ( $\lambda \sim 360$  nm), we 297 can report a realistic upper limit for CH<sub>2</sub>OO concentration of  $<1\times10^{10}$  cm<sup>-3</sup> in our experiments. Even if 298 we adopt the most conservative possible approach and assign all of the transient absorption at 360 nm 299 to CH<sub>2</sub>OO, the upper limit for its concentration at t=1 ms is still below  $2.5 \times 10^{10}$  cm<sup>-3</sup>. Being again 300 conservative and using the lower limit for the OH-concentration this leads to an upper limit of  $\phi_{1a}$  < 301 0.05. This result is in very good agreement with theoretical calculations, but also with the recent 302 finding of Yan et al.<sup>10</sup>, who also determined an upper limit of  $\phi_{1a} < 0.05$  for the formation of the 303 304 Criegee intermediate in (R1).



**Figure 3:** Experimental transient spectra, averaged over the kinetic times 0 - 5 ms following the 248 nm photolysis of acetone in He bath gas in the presence of 2.5 Torr O<sub>2</sub> (total P = 30 Torr). Inset: transient absorption, averaged over probe  $\lambda = 300 - 340$  nm, as a function of kinetic time. Red and black spectra and kinetic traces are from run A ([CH<sub>3</sub>O<sub>2</sub>] =  $8.9 \times 10^{12}$  cm<sup>-3</sup> and [OH] =  $3 \times 10^{12}$  cm<sup>-3</sup>) and B ([CH<sub>3</sub>O<sub>2</sub>] =  $4.8 \times 10^{12}$  cm<sup>-3</sup> and [OH] =  $1.6 \times 10^{12}$  cm<sup>-3</sup>), respectively. Blue and yellow diamonds show the simulated transient absorption at 300 nm for run A and B, respectively, assuming that only CH<sub>3</sub>O<sub>2</sub> contributes to the signal.

#### 313

## 314 Modeling and Atmospheric Implication

315 A box model constrained to the Master Chemical Mechanism v3.2 has been run to test the atmospheric 316 significance of HO<sub>2</sub> production from the reaction of OH with CH<sub>3</sub>O<sub>2</sub> under clean, low NO<sub>x</sub> conditions. 317 The model was constrained with gas-phase composition field data (including the concentrations of the 318 halogen oxides IO and BrO), photolysis rates and meteorological parameters taken during the RHaMBLe project that took place in 2007<sup>42</sup> at the Cape Verde Atmospheric Observatory (CVAO) 319 320 which is situated on the island of Sao Vicente in the tropical Atlantic Ocean (23.96° S, 46.39° W). This model has been described previously and was used to calculate OH and HO<sub>2</sub> concentrations for 321 comparison with those measured at the CVAO<sup>11</sup>. Inclusion of reaction (R1b) in the model has a 322 substantial impact on the predicted CH<sub>3</sub>O<sub>2</sub> levels, decreasing the concentration of CH<sub>3</sub>O<sub>2</sub> by ~30% 323 324 during the afternoon (Figure 4). Taking into account the HO<sub>2</sub> (and CH<sub>3</sub>O) production from (R1b) 325 reported here, the model predicts approximately 20% more  $HO_2$  at CVAO (Figure 4). This increase in 326 modeled HO<sub>2</sub> further enhances the destruction of  $CH_3O_2$  (alongside the enhanced destruction due to 327 the reaction with OH) by increasing the loss of  $CH_3O_2$  via reaction with  $HO_2$  (by 15%) which is the 328 dominant CH<sub>3</sub>O<sub>2</sub> sink, accounting for 50% of the total CH<sub>3</sub>O<sub>2</sub> destruction in this environment.



**Figure 4a:** Diurnal ratios of modelled radical concentrations with reaction (R1b) included: without reaction (R1b) included in the model simulations. **b:** Diurnal concentrations of modelled OH (blue), HO<sub>2</sub> (red) and CH<sub>3</sub>O<sub>2</sub> (green) (with R1b – solid coloured lines, without R1b – dashed coloured lines) compared with measured OH and HO<sub>2</sub> (black lines). Errors bars represent the 1 $\sigma$  day-to-day variability in the averaged data.

335

Tropospheric ozone is an important greenhouse gas and in remote oceanic regions, integrating (R1b) 336 337 into a model impacts also on its formation and destruction paths. At conditions such as found at the 338 CVAO, ozone production via reaction of peroxy radicals with NO is small, and ozone is typically 339 destroyed during the day via deposition and photochemical reactions (Table 3). HO<sub>2</sub> accounts for 9% of the daily ozone destruction cycle observed at CVAO and, as the reaction of  $CH_3O_2$  with  $O_3$  is 340 significantly slower than the reaction of  $HO_2$  with  $O_3$ , overlooking an additional source of  $HO_2$  from 341 342 reaction (R1b) may influence model estimates of ozone destruction ( $dO_3$ ). The reaction of HO<sub>2</sub> with ozone proceeds  $\sim 17\%$  faster when (R1b) is included; the reactions of HO<sub>2</sub> with IO and BrO are also 343 344 enhanced by  $\sim 18\%$  and this serves to increase dO<sub>3</sub> further. Considering all the net ozone destroying 345 reactions (**Table 3**), it is estimated that (R1b) enhances  $dO_3$  by ~4%. Reaction (R1b), also serves to modify ozone production (pO<sub>3</sub>) by a reduction in the rate of CH<sub>3</sub>O<sub>2</sub>+NO and an enhancement in the 346 347 rate of HO<sub>2</sub>+NO reactions. Overall pO<sub>3</sub> is predicted to decrease modestly (by  $\sim$ 75 pptv d<sup>-1</sup>) when (R1b) 348 is considered and so taking both the changes in  $pO_3$  and  $dO_3$  into account, this study suggests that 349 tropospheric ozone in remote regions will be over-estimated by  $\sim 6\%$  by models which do not consider 350 the reaction of CH<sub>3</sub>O<sub>2</sub> with OH. Remote oceanic regions cover approximately two thirds of the Earth's 351 surface and so omission of the reaction of CH<sub>3</sub>O<sub>2</sub> with OH yielding HO<sub>2</sub> may lead to significant overpredictions in tropospheric ozone on a global scale. 352

353	Table 3: Percentage contribution of the rate limiting O <sub>3</sub> destruction steps with and without (R1b	o)
354	included in the model simulation	

O <sub>3</sub> destruction rate determining step	% Contribution to O <sub>3</sub> destruction without (R1b)	% Contribution to O <sub>3</sub> destruction with (R1b)
$O^1D + H_2O$	62.8	60.2
$OH + O_3$	5.8	5.3
$HO_2 + O_3$	8.5	9.5
$BrO + HO_2$	5.1	5.7
BrO + NO	1.1	1.1
BrO + BrO	0.2	0.2
BrO + IO	2.3	2.2
$IO + HO_2$	11.7	13.2
IO + NO	0.8	0.7
IO + IO	2.0	2.0

355

The concentration of methanol is generally underpredicted by global models compared to the 356 measured CH<sub>3</sub>OH concentration. In the recent work of Müller *et al.*<sup>6</sup> it was proposed that a significant 357 vield of CH<sub>3</sub>OH in (R1) could help to explain this discrepancy. The best estimate from their theoretical 358 359 calculations predict yields for CH<sub>3</sub>OH and for the stabilized CH<sub>3</sub>OOOH between 0.05 and 0.1, but each with a wide error margin of a factor  $\sim 3.5$ . This is in very perfect agreement with the 360 361 measurements presented in this word. However, in order to best reproduce measured CH<sub>3</sub>OH 362 concentrations, a CH<sub>3</sub>OH yield of 18%, together with a yield of 21% for stabilization of the initially 363 formed trioxide, CH<sub>3</sub>OOOH, would be needed. This scenario leaves a yield of 61% for HO<sub>2</sub> formation 364 through channel (R1b). However, Müller et al. used for their model the excessively high rate constant of Bossolasco et al.<sup>7</sup>, and hence their study overstates the overall importance of (R1). As a result the 365 366 CH<sub>3</sub>OH yield would need to be increased further (leading consequently to an even lower value for the 367 HO<sub>2</sub> yield) in order to bring into agreement the CH<sub>3</sub>OH predictions with measurements. While a yield of 61% HO<sub>2</sub> would be within the error limit of the current work, an even higher value is not in 368 369 agreement with the above presented measurements. However, it has to be noted that the current 370 measurements for the HO<sub>2</sub> yields have been carried out at low pressure (50 Torr helium  $/ O_2$ ) and that 371 the product distribution might change with pressure. Therefore, product distributions should be 372 determined at higher pressures in order to find out if the title reaction could be a significant source of 373 CH<sub>3</sub>OH in the remote troposphere. The present set-up is not very suitable for this task, because the 374 absorption lines of OH and HO<sub>2</sub> suffer from broadening, making their detection at higher pressure less 375 sensitive. Also the uncertainty of the derived yield would increase, if the appropriate broadening 376 coefficients are not well known. The better way would be the direct detection of CH<sub>3</sub>OH, but 377 unfortunately the cw-CRDS set-up is not sensitive to this species.

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