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# Using a cross correlation technique to refine the accuracy of the Failure Forecast Method: Application to Soufrière Hills volcano, Montserrat

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## Abstract

Prior to many volcanic eruptions, an acceleration in seismicity has been observed, suggesting the potential for this as a forecasting tool. The Failure Forecast Method (FFM) relates an accelerating precursor to the timing of failure by an empirical power law, with failure being defined in this context as the onset of an eruption. Previous applications of the FFM have used a wide variety of accelerating time series, often generating questionable forecasts with large misfits between data and the forecast, as well as the generation of a number of different forecasts from the same data series. Here, we show an alternative approach applying the FFM in combination with a cross correlation technique which identifies seismicity from a single active source mechanism and location at depth. Isolating a single system at depth avoids additional uncertainties introduced by averaging data over a number of different accelerating phenomena, and consequently reduces the mis-

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fit between the data and the forecast. Similar seismic waveforms were identified in the precursory accelerating seismicity to dome collapses at Soufrière Hills volcano, Montserrat in June 1997, July 2003 and February 2010. These events were specifically chosen since they represent a spectrum of collapse scenarios at this volcano. The cross correlation technique generates a five-fold increase in the number of seismic events which could be identified from continuous seismic data rather than using triggered data, thus providing a more holistic understanding of the ongoing seismicity at the time. The use of similar seismicity as a forecasting tool for collapses in 1997 and 2003 greatly improved the forecasted timing of the dome collapse, as well as improving the confidence in the forecast, thereby outperforming the classical application of the FFM. We suggest that focusing on a single active seismic system at depth allows a more accurate forecast of some of the major dome collapses from the ongoing eruption at Soufrière Hills volcano, and provides a simple addition to the well used methodology of the FFM.

*Keywords:* Volcano-seismology, Failure Forecast Method, low frequency, multiplets, Eruption forecasting, Soufrière Hills volcano

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## 1. Introduction

Volcanic eruptions are often preceded by accelerating geophysical signals (McNutt, 2002), associated with the movement of magma or other fluid towards the surface. Of these precursors, seismicity is at the forefront of forecasting volcanic unrest since it is routinely observed and the change from background level can be

6 observed in real time (Chouet et al., 1994; Cornelius and Voight, 1994; Kilburn,  
7 2003; Ortiz et al., 2003). Since forecasting of volcanic eruptions relies upon the  
8 ability to forecast the timing of magma reaching the surface, low frequency seis-  
9 micity, with a spectral range of 0.2 – 5 Hz (Lahr et al., 1994), may potentially act  
10 as a forecasting tool since one of the interpretations of its low frequency content is  
11 its association with the movement of magmatic fluid at depth (Chouet et al., 1994;  
12 Neuberg et al., 2000).

13

14 The relationship between an accelerating geophysical precursor and the tim-  
15 ing of failure of the system was first considered for landslides (Fukuzono, 1985)  
16 but has since been adapted for the forecasting of volcanic eruptions (Voight, 1988,  
17 1989). The Material Failure Law or the Failure Forecast Method (FFM) as it is re-  
18 ferred to in volcanology (Cornelius and Voight, 1995), is an empirical power-law  
19 relationship based on first principles associated with failing materials, which re-  
20 lates the acceleration of a precursor ( $d^2\Omega/dt^2$ ) to the rate of that precursor ( $d\Omega/dt$ )  
21 at constant stress and temperature (Voight, 1988) by:

$$\frac{d^2\Omega}{dt^2} = K \left( \frac{d\Omega}{dt} \right)^\alpha \quad (1)$$

22 where  $K$  and  $\alpha$  are empirical constants.  $\Omega$  can represent a number of different  
23 geophysical precursors, for example low frequency seismic event rate (Hammer  
24 and Neuberg, 2009), event rate of all recorded seismicity (Kilburn and Voight,  
25 1998), or the amplitude of the seismic events (Ortiz et al., 2003). The parameter

26  $\alpha$  is thought to range between 1 and 2 in volcanic environments (Voight, 1988;  
27 Voight and Cornelius, 1991), or may even evolve from 1 towards 2 as seismic-  
28 ity proceeds (Kilburn, 2003). An infinite  $d\Omega/dt$  suggests an uncontrolled rate of  
29 change or a singularity and is associated with an impending eruption. The inverse  
30 form of  $d\Omega/dt$  is linear if  $\alpha = 2$ , and therefore the timing of failure is determined  
31 when a linear regression of inverse rate against time intersects the x-axis (Voight,  
32 1988). It is important to note that this forecasted timing of failure is associated  
33 with the potential for an eruption due to accelerated magma flux at depth, and  
34 may not necessarily result in one, since a direct pathway of magma to the surface  
35 might be buffered.

36

37 The ongoing eruption of Soufrière Hills Volcano (SHV), Montserrat began  
38 in July 1995 with a series of phreatic explosions associated with vent openings  
39 around the crater (Young et al., 1998). Since November 1995, the andesitic vol-  
40 cano has undergone a repeated cycle of dome growth and collapse, with the col-  
41 lapse phases resulting in pyroclastic flows, lahars and ash fall events (Sparks and  
42 Young, 2002; Wadge et al., 2014). The first major dome collapse at SHV occurred  
43 in June 1997, killing 19 people (Loughlin et al., 2002), and generated pulsatory  
44 block and ash flows due to the collapse of  $5 \times 10^6 \text{ m}^3$  of material, removing the  
45 top 100 m of dome material (Voight et al., 1999). A number of other major dome  
46 collapses have occurred since 1995 including: 26 December 1997 (Voight et al.,  
47 2002); 29 July 2001, which lowered the dome height by over 150 m (Matthews  
48 et al., 2002); 12 July 2003, during which 210 million  $\text{m}^3$  of material was displaced

49 (Herd et al., 2005); 20 May 2006 (Loughlin et al., 2010); and the most recent event  
50 on 11 February 2010 when 50 million m<sup>3</sup> of material was displaced (Stinton et al.,  
51 2014). As at many other volcanoes worldwide (e.g. Galeras, Colombia; Redoubt,  
52 Alaska; Mt Pinatubo, Philippines), an increase in the number of low frequency  
53 seismic events has been identified in hindsight prior to dome collapse events at  
54 SHV, in particular the event of June 1997 (Cruz and Chouet, 1997; Miller et al.,  
55 1998; Stephens and Chouet, 2001; Kilburn, 2003; Hammer and Neuberg, 2009).

56

57 White et al. (1998) first noted that low frequency earthquakes appear to occur  
58 in swarms of similar waveforms lasting from days to weeks, prior to and during  
59 unrest and extrusion periods at SHV. A swarm was originally described as a se-  
60 quence of temporally close seismic events occurring within 15 km of a volcano  
61 (Benoit and McNutt, 1996). We take the narrower definition of Voight et al. (1999)  
62 who determine a swarm as when there are more than 10 events within an hour.  
63 Often, similarity between repeating events exists within these swarms, which sug-  
64 gests that the source location and source mechanism are identical (Geller and  
65 Mueller, 1980; Caplan-Auerbach and Petersen, 2005; Petersen, 2007). Events  
66 which are statistically similar to one another are known as multiplets, and can be  
67 grouped together into a family. Many authors have shown that it is possible to  
68 classify multiplets into a number of families of highly similar waveforms using a  
69 cross correlation technique, which therefore isolates and focusses on a single sys-  
70 tem at depth. Stephens and Chouet (2001) investigated a 23 hour swarm of low  
71 frequency seismic events prior to the eruption of Redoubt volcano, Alaska in De-

72 cember 1989, finding that the events could be sorted into 3 distinct families which  
73 evolved with time, of which the majority of events were correlated (cross corre-  
74 lation coefficient  $\geq 0.68$ ) with just one distinctive family. Later analysis of the  
75 2009 Redoubt eruption by Buurman et al. (2013) also suggested the presence of  
76 multiplets, in particular prior to explosion events. Petersen (2007) suggested that  
77 a dominant family of multiplets exists within the low frequency seismic swarms at  
78 Shishaldin volcano, Alaska, although the dominant family is different within each  
79 swarm studied between 2002 and 2004. Thelen et al. (2011) suggest that the oc-  
80 currence of multiplets at Mount St. Helens, Washington and Bezymianny volcano,  
81 Russia are related at least in part to the viscosity of the magma, and are therefore  
82 more prominent during dome building eruption events. Highly correlated high fre-  
83 quency events were observed at Mt. Unzen, Japan during significant endogenous  
84 growth of a lava dome between 1993 and 1994 and were classified into over 100  
85 families (Umakoshi et al., 2003). Families of similar low frequency events have  
86 also been identified and studied at SHV in relation to tilt cycles by Voight et al.  
87 (1999) and Green and Neuberg (2006), who identified 9 multiplet families con-  
88 taining more than 45 similar events each over a time period of 6 days in June 1997,  
89 although not all of the families were active during each of the seismic swarms. In  
90 addition, Ottemöller (2008) found that 7100 hybrid events generated in the days  
91 prior to a large scale dome collapse at SHV in July 2003 all belonged to the same  
92 multiplet family.

93

94 In this paper we investigate accelerating event rates of precursory low fre-

95 quency seismicity in the days prior to a number of large dome collapses at SHV  
96 on 25 June 1997; 12 July 2003; and 11 February 2010, and its use as a forecast-  
97 ing tool. These collapses were chosen since they represent a wide spectrum of  
98 collapses at SHV: the first; the largest; and the latest collapse. We use a cross cor-  
99 relation technique in order to further classify the seismicity based on waveform  
100 similarities as well as frequency content, as has been previously adopted to find  
101 similar seismic events at SHV. Green and Neuberg (2006) have already distin-  
102 guished families of seismicity prior to the dome collapse in June 1997, however  
103 do not use these families in forecasting in any manner. Hammer and Neuberg  
104 (2009) used a single family from Green and Neuberg (2006) as a forecasting tool,  
105 however fail to detail the family used or whether any other family of seismicity  
106 produced a successful forecast. Here, we identify whether low frequency seismic  
107 families can be identified at another station at SHV compared to Green and Neu-  
108 berg (2006), and produce forecasts using each of the seismic families identified,  
109 rather than picking just one. In addition, we identify similar seismic families and  
110 use these in hindsight analysis for forecasting dome collapses in July 2003 and  
111 February 2010 at SHV. This allows analysis of the wider application of the FFM  
112 at SHV, and whether it can be used in all circumstances as a forecasting tool. In  
113 Section 2 we describe the methodology of the cross correlation technique used to  
114 identify multiplets, with Section 3 showing the identified multiplet families for  
115 June 1997, July 2003 and February 2010 respectively. In Section 4 we apply the  
116 FFM to each of the families of events, showing that a more accurate forecast is  
117 generated when using a cross correlation technique to focus on one single system



118 rather than simply using any low frequency seismicity. The implications of these  
119 results are discussed in Section 5.

## 120 **2. The Cross Correlation Technique**

### 121 *2.1. Data selection*

122 Since 1995 a continuous network of seismometers has been deployed on Montser-  
123 rat to ensure the constant and consistent monitoring of the volcano by the Montser-  
124 rat Volcano Observatory (MVO), originally installed by the USGS Volcano Dis-  
125 aster Assistance Program (Aspinall et al., 1998). At the time of the dome collapse  
126 event in June 1997, five three-component seismometers (Guralp CMG-40T with a  
127 30 second corner frequency) and three vertical component Integra LA100/F 1Hz  
128 instruments were deployed, with data being digitized at 75 Hz. In March 2005,  
129 station MBLG was upgraded to a three-component broadband instrument (Guralp  
130 CMG-40T) with data digitized at 100 Hz. Station MBLG was chosen for analysis  
131 because of its close proximity to the dome which allowed a good signal-to-noise  
132 ratio and the availability of triggered and continuous data for the entire period un-  
133 der investigation.

134

135 MVO uses a STA/LTA (short-term average to long-term average) ratio trigger-  
136 ing algorithm to identify individual events from the continuous incoming seismic  
137 record and place them into a catalogue of triggered events. It consists of two slid-  
138 ing windows, one investigating the short term amplitudes and is therefore very  
139 sensitive to incoming seismic signals, and one investigating the long term ampli-

140 tude of the signal, which can provide information about the temporal amplitude  
141 of the noise at the site of the seismometer (Withers et al., 1998). Since the trigger  
142 is based on the ratio between these two windows, the algorithm is better able to  
143 record weak seismicity, compared to a simple amplitude only trigger mechanism  
144 (Trnkoczy, 2002). If a trigger (when a critical threshold of this ratio is exceeded) is  
145 found on three or more seismic stations simultaneously then an event is registered  
146 within the event count catalogue. To ensure the entirety of the earthquake signal  
147 is captured 2 seconds is added to the beginning of the trigger, and 10 seconds after  
148 the event once again drops below the triggering threshold. Using a Short-Term  
149 Averaging window length of 0.333 seconds, and a Long-Term Averaging window  
150 length of 60 seconds, with a trigger and detrieger ratio value of 4 and 2 respec-  
151 tively, a total of 1817 triggered events were placed within the catalogue from the  
152 22 - 25 June 1997; 520 events placed in the catalogue from 8 - 13 July 2003; and  
153 452 events identified from 8 - 12 February 2010. The instrument response and  
154 digitizer gains were removed from these seismograms to give velocity calibrated  
155 traces. STA/LTA analysis only identifies seismic events from the continuous seis-  
156 mic record, and does not classify them in any form. Traces were filtered with a  
157 band pass Butterworth two pole filter, with a low frequency cut-off of 0.5 Hz to  
158 reduce the influence of oceanic noise and a high frequency cut-off of 5 Hz, so as  
159 to concentrate entirely on the low frequency content of the waveform associated  
160 with magma movement.

161 2.2. *Event Classification*

162 Similarities between waveform shape can be quantified using the cross corre-  
163 lation function:

$$r_{xy}(i, i - l) = \frac{\sum_{i=1}^n (x_i - \bar{x})(y_{i-l} - \bar{y})}{\sqrt{\sum_{i=1}^n (x_i - \bar{x})^2} \sqrt{\sum_{i=1}^n (y_{i-l} - \bar{y})^2}} \quad (2)$$

164 where  $r$  is the cross correlation coefficient,  $x$  and  $y$  represent the two traces in  
165 the correlation, and therefore  $x_i$  is the  $i$ th sample of the signal  $x$  and  $y_{i-l}$  is the  
166  $i$ - $l$ th sample of the signal  $y$ . The overbar represents the mean value of the signal  
167 and  $l$  is the lag between the two signals. Identical waveforms will result in a cross  
168 correlation function of 1 or -1 dependent upon the polarity of the signal.  $r_{xy}$  is a  
169 measure of similarity in waveform shape only, since events are normalised prior  
170 to calculation. Consequently, the cross correlation function gives no information  
171 on the amplitude ratios of the events. Waveforms which are similar, and therefore  
172 from the same source location and generated by the same source mechanism, rep-  
173 resent a single active system of seismicity.

174

175 Since the triggered events varied in duration and the cross correlation tech-  
176 nique requires events of the same length, a cross correlation window of 10 seconds  
177 was chosen. This length was chosen since it allowed the entirety of the waveform  
178 to be present within the window, but ensured that only one event was captured per  
179 window. Previous attempts at cross correlating events at SHV by Green and Neu-  
180 berg (2006) used an 8 second cross correlation window, however upon inspection

181 of the events this was not sufficient to include the majority of each coda at this  
182 station. In order to determine if any similar events were present on any particular  
183 day, each 10 second event was cross correlated with every other triggered event  
184 from the same day. The maximum cross correlation coefficient of each waveform  
185 with every other waveform was determined and placed into a cross correlation ma-  
186 trix (an example of which can be seen in Figure 1 for the 24 June 1997). In theory,  
187 identical events will give a cross correlation coefficient ( $r$ ) of 1 (e.g. the autocorre-  
188 lation of events, as seen along the diagonal of Figure 1), and events which have no  
189 correlation will have a cross correlation coefficient of 0 ( $r=0$ ). Events were then  
190 classified as being significantly similar to one another if the maximum cross cor-  
191 relation coefficient was above 0.7, and are shown on a colour spectrum in Figure  
192 1. Green and Neuberg (2006) and Thelen et al. (2011) show a clear justification of  
193 using 0.7 as a correlation threshold since it is significantly above the upper limit  
194 for random correlation between waveforms and noise. Therefore it is assumed that  
195 some events with a correlation coefficient of less than 0.7 are a consequence of  
196 random noise being correlated. Visual inspection of the stacked waveforms also  
197 confirmed that 0.7 was an appropriate choice and captured the majority of simi-  
198 lar events with limited scatter of the waveforms once aligned. Figure 1 suggests  
199 distinct time periods when similar seismic events were active (coloured areas are  
200 separated by distinctly white areas), and that a highly correlated swarm of events  
201 occurred on this day (brighter and more concentrated colours).

202

203 Two different techniques, both utilizing the entire catalogue of triggered events

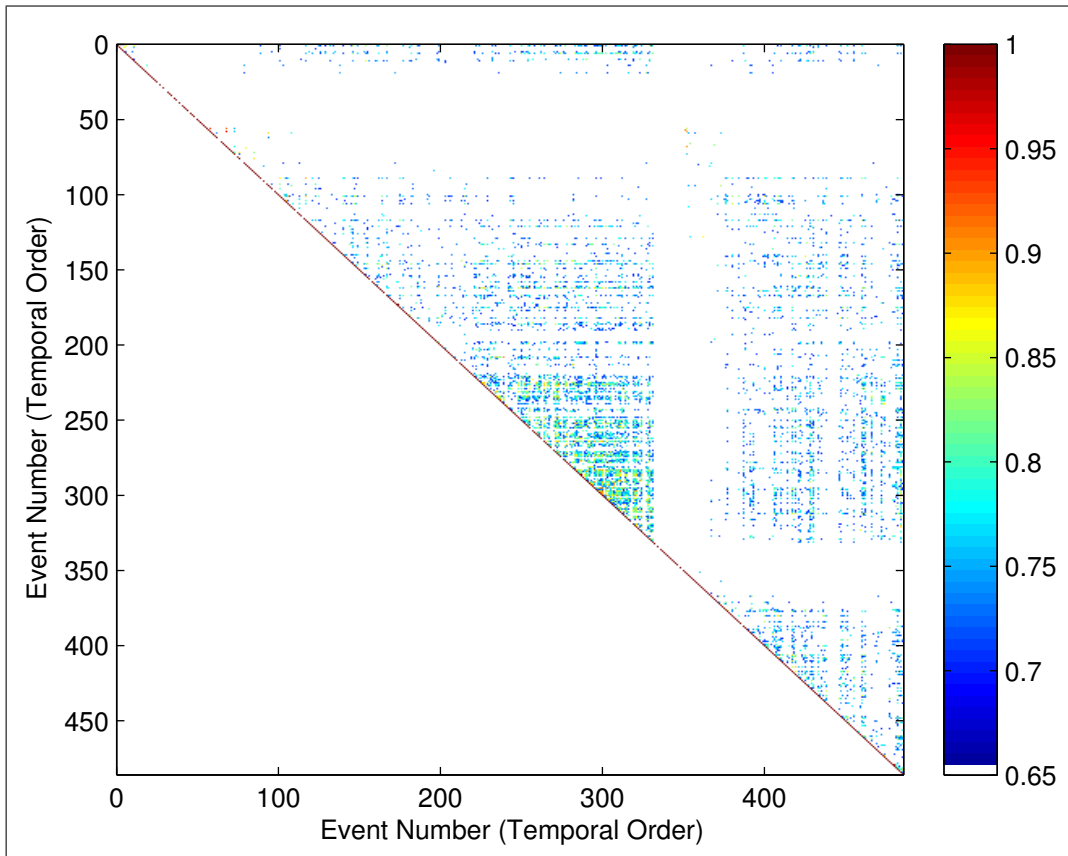


Figure 1: An example of a maximum cross correlation similarity matrix from station MBLG on 24th June 1997. A total of 486 triggered events were found within the 24 hour period and are represented from 1 to 486 along the x and y axis. Each row of the matrix therefore represents one triggered event compared to every other triggered event on that day. Only events with a cross correlation coefficient above 0.7 are shown on the colour spectrum and are deemed to be similar. The autocorrelation of each triggered event with itself (cross correlation coefficient equal to 1) is represented on the diagonal.

204 during periods of interest, were used to isolate multiplets and collate them into  
205 families. Following Petersen (2007), a dominant event for each day was identified  
206 as the event correlated with the highest number of other events from that day. The  
207 mean correlation value of each event with every other event was determined from  
208 the cross correlation matrix (Figure 1) and the event with the highest mean was  
209 taken to be the dominant event. The second technique followed Green and Neu-  
210 berg (2006) where each triggered event in turn was correlated with every other  
211 event. Events with a cross correlation coefficient above 0.7 were subsequently  
212 grouped together, labelled as a multiplet family and removed from the time series.  
213 This procedure was repeated across the entire investigated time period until all  
214 events had been classified into a number of different families. This has the advan-  
215 tage of finding all families of multiplets which may be present in the continuous  
216 data, rather than simply the dominant one, as well as finding families which may  
217 be infrequent in their repetition but still important. This procedure also allows  
218 for the identification of evolving waveforms, either by migration of their source  
219 location or change in the source process. Families which contained fewer than  
220 10 similar individual triggered waveforms were eliminated from further analysis.  
221 To avoid selection bias, the events within a single family (i.e. all had a minimum  
222 cross correlation coefficient of 0.7 with one another) were stacked, and the aver-  
223 age waveform taken (Figure 2). This average waveform is hereafter referred to as  
224 the Master Event of each family, and is used as a statistical representation of this  
225 family in terms of waveform shape.

226

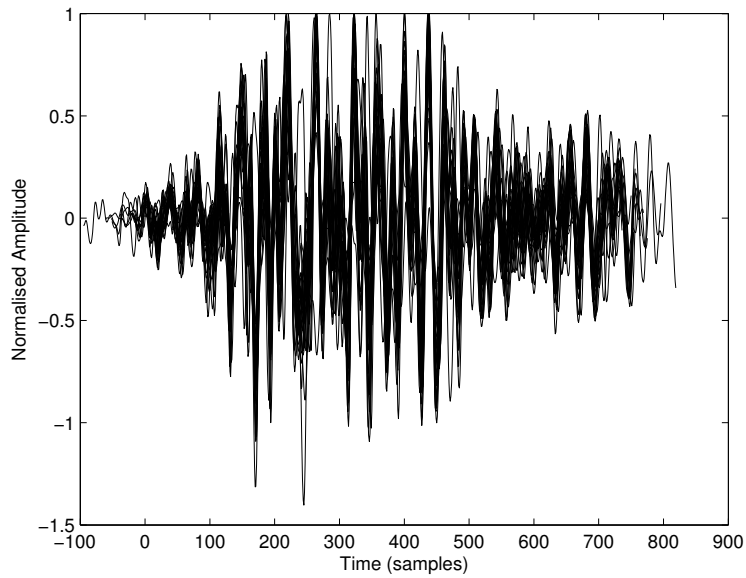
227 The master events were then cross correlated with the continuous seismic  
228 record at MBLG using a sliding window technique and multiplets identified when  
229 the cross correlation coefficient was greater than 0.7 between the master event and  
230 the continuous seismogram. The sliding window separation of 0.01 seconds al-  
231 lows the maximum number of multiplets to be identified, in particular those which  
232 are too small or are overlapping in the continuous seismic record to be identified  
233 by the triggered acquisition system at MVO (STA/LTA algorithm). The similar  
234 events were then grouped into a multiplet family.

235

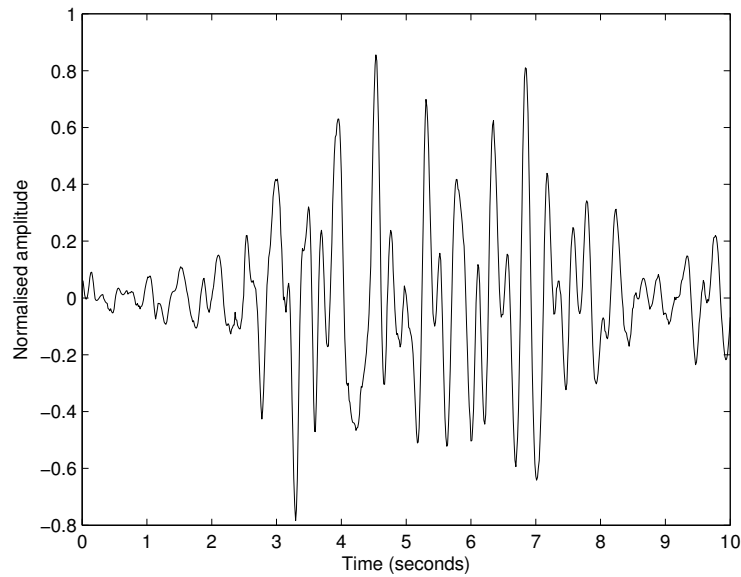
### 236 **3. Similarity of Events**

#### 237 *3.1. 22 - 25 June 1997*

238 The total number of multiplets identified using the cross correlation procedure  
239 was 7653 from 22 to 25 June 1997, in comparison to only 1435 events identified  
240 using the triggered algorithm at MVO over the same time period. This methodol-  
241 ogy therefore represents a five fold increase in the number of events which can be  
242 identified and used in further analysis. The dominant multiplet family identified  
243 using the technique by Petersen (2007) contained a total of 878 multiplets (Table  
244 1). 10 multiplet families containing over 250 multiplets each were identified us-  
245 ing the technique of Green and Neuberg (2006), although the dominant master and  
246 Master event 001 have a cross correlation coefficient of 0.93, suggesting that they  
247 belong to the same family. This emulates the conclusions of Green and Neuberg  
248 (2006), who also identified 10 waveform families during the same time period at



(a) Stack of events highly correlated with identified dominant master event on 24 June at 11:18-53s from station MBLG. A total of 26 triggered events are included in the stack. Each event has been aligned at the peak correlation coefficient.



(b) Master Event : average waveform from stack

Figure 2: Stack of highly correlated waveforms and the resulting average dominant master event from station MBLG identified on 24 June 1997



<b>Family</b>	<b>22nd</b>	<b>23rd</b>	<b>24th</b>	<b>25th</b>	<b>Total</b>
Dominant	0	133	376	369	878
Master001	0	121	336	222	679
Master010	0	256	315	32	603
Master014	2	136	302	137	577
Master100	4	71	280	193	548
Master106	0	42	169	45	256
Master121	0	100	514	400	1014
Master136	3	131	483	542	1159
Master141	0	47	276	173	496
Master210	0	20	170	349	539
Master291	0	39	390	475	904

Table 1: Number of events within each family sorted into days from the 22 - 25 June 1997.

249 station MBGA, however they did not use these families as a forecasting tool. As  
 250 these families of events were identified at a different station it is not possible to  
 251 compare directly the results of the two studies.

252

253 Clear differences between each of the families can be found in terms of the  
 254 onset timing of events and the waveform characteristics (Figure 3). In particu-  
 255 lar significant differences are noted in the expression of the waveform coda. As  
 256 expected for low frequency events, master event 210 (Figure 3a(i)) clearly de-  
 257 cays in a harmonic manner, however this is not the case for every master event  
 258 (e.g. master event 100, Figure 3a(e)). However, very little difference is seen in  
 259 the amplitude spectra (Figure 3b). All of the low frequency master events have  
 260 a dominant spectral peak at 2.1Hz, suggesting a fundamental similarity in their  
 261 resonance behaviour. A secondary spectral peak can be seen at 3.8 Hz for some

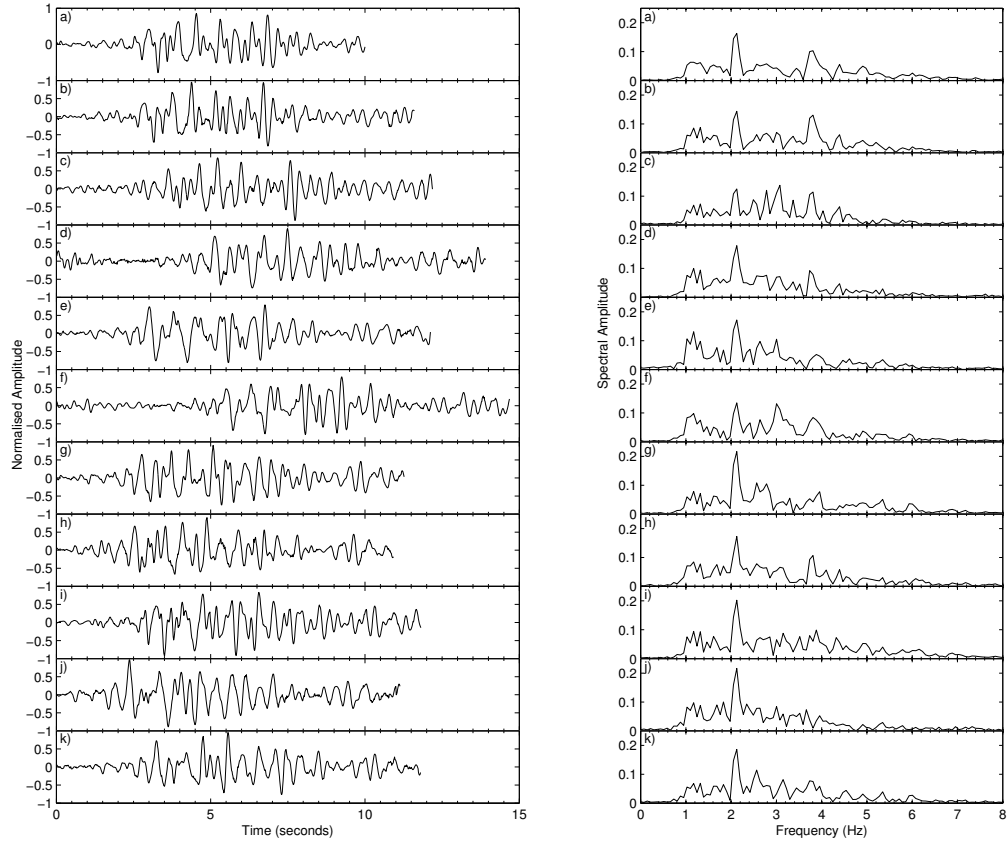
262 master events. Variations in the amount of energy distributed from 1-5 Hz varies  
263 amongst master event waveforms, although not significantly. No significant phase  
264 shifts were detected from the cross correlation analysis.

265

266 Some differences are evident in the duration and timings of swarms of each  
267 master event in relation to the dome collapse on the 25 June (Figure 4). In partic-  
268 ular, only three master events appear within the swarms identified on the 22 June,  
269 and these are very short lived. The beginning and ending of each swarm varies  
270 only slightly throughout the rest of the sequence from the 23 - 25 June. Since  
271 MBLG was destroyed by volcanic activity associated with the dome collapse at  
272 16:55 UTC (Luckett, 2005) no swarms are able to be identified past this (vertical  
273 line in Figure 4). Each of the multiplet families are persistent across each of the  
274 six swarms which occur from the 23 June onwards, suggesting that sources at the  
275 same locations are being reactivated by the same process during this time, as was  
276 also concluded by Green and Neuberg (2006). Besides master events 010, 014  
277 and 106, all other master events appear within swarms which are active right up  
278 to the dome collapse (Figure 4).

279

280 Unlike Stephens and Chouet (2001), Umakoshi et al. (2003) and Petersen  
281 (2007) the waveforms within the multiplet families observed at SHV in 1997 do  
282 not appear to significantly evolve with time, since clustering of events is not seen  
283 along the diagonal in Figure 1 (Caplan-Auerbach and Petersen, 2005) and each  
284 swarm appears to contain similar waveforms with limited evolution in the cross



(a) Dominant master and 10 other identified master event waveform signatures identified in June 1997 at station MBLG

(b) Amplitude spectra of the waveforms of the dominant master and the 10 other identified master events in June 1997 at station MBLG

Figure 3: Comparison of all master events identified from 22 to 25 June 1997 at station MBLG, by both techniques. The waveform and amplitude spectra labelled (a) is the dominant waveform identified from the cross correlation coefficient matrix (Figure 1). Waveforms and corresponding amplitude spectrum labelled (b) to (k) represent the master events 001, 010, 014, 100, 106, 121, 136, 141, 210 and 291 respectively.

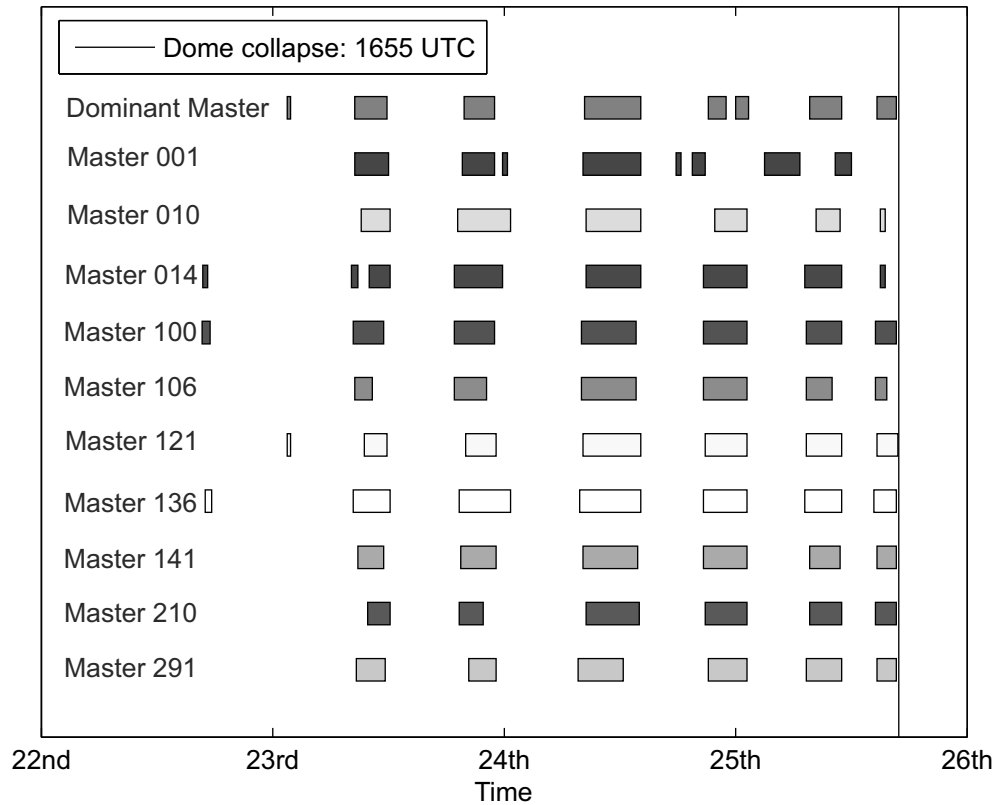


Figure 4: Comparison of the timing and duration of swarms related to each of the master events identified. The timing of the dome collapse is represented by the vertical line on the 25th June. The y axis is only an indication of each of the families present separated in space for the purpose of clarity on the plot and does not represent time or dominance, each master event is simply drawn below the last so that all can be compared. Each coloured rectangular box represents the times when the master event was active during the 22 - 25 June analysis period.

285 correlation coefficient with time. This suggests the waveforms are stable and per-  
286 sistent and therefore the trigger location and source process must also be. The  
287 cross correlation coefficient can vary by up to 0.25 within each swarm, however  
288 the difference between the maximum and minimum mean cross correlation coeffi-  
289 cient for each swarm as determined by the 11 master events only varies by  $\leq 0.06$ ,  
290 suggesting limited evolution in the waveforms.

291

### 292 3.2. 8 - 12 July 2003

293 A total of 520 events were identified from the continuous seismic record at  
294 station MBLG from 8 to 12 July 2003 using the MVO STA/LTA algorithm. In  
295 comparison, the total number of multiplets identified for the same time period  
296 was 2241 events, representing a four fold increase in the number of seismic events  
297 identified. However, unlike the multiplets identified in June 1997, only one fam-  
298 ily of events could be identified. A total of 79 events from this single family were  
299 stacked to create an average master event, shown in Figure 5(a), all of which had  
300 a cross correlation coefficient of above 0.7 to maintain a high signal to noise ratio  
301 and consequently pick the most similar events for use in the forecast. Ottemöller  
302 (2008) also identified one single dominant family of events during a similar time  
303 period from 00:00 on 9 July to 12:00 on 12 July, however suggests that a total of  
304 7100 events could be identified. The large difference in the number of identified  
305 events from the cross correlation technique is put down to the fact that Ottemöller  
306 (2008) used a much lower cross correlation coefficient to identify events ranging

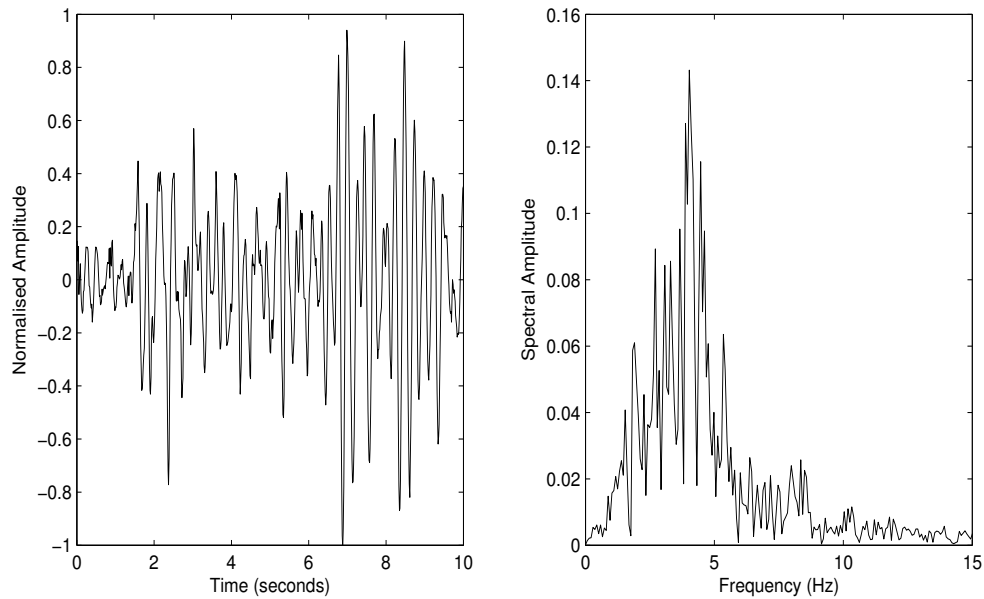
307 from 0.6 to 0.66, whereas the event identified in this study consistently used a  
308 cross correlation coefficient threshold of 0.7.

309

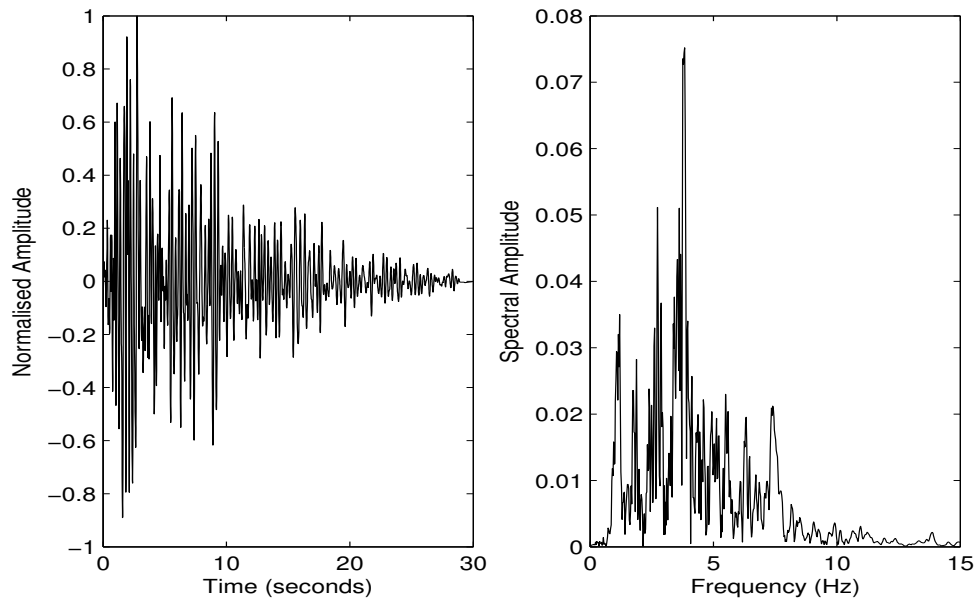
310 The dominant waveform has an emergent onset and it is difficult to pick out  
311 significant seismic phases (Figure 5(a)). Unlike the low frequency seismicity iden-  
312 tified in June 1997, the coda of the waveform does not decay in a smooth manner.  
313 It is evident that the peak energy is centred at approximately 4 Hz, in comparison  
314 to the June 1997 events which all had a dominant frequency of approximately 2.1  
315 Hz, however it is distributed across 0 Hz to 5 Hz band. Contrary to Petersen (2007)  
316 who suggested that some multiplets can be active over a number of years, there are  
317 no similarities between the events identified in June 1997 and the dominant master  
318 event identified in July 2003, pointing to an evolving system over this time period.

319

320 Figure 6 suggests a clear evolution of the cross correlation coefficient with  
321 time. One possible explanation is that this may represent a slightly migrating  
322 source location at depth. The events with a relatively lower cross correlation coef-  
323 ficient on the 9 and 10 July occurred further away from the dominant master event  
324 location, than those on 11 July. Perhaps more significantly, it should be noted  
325 that the similar seismic waveforms appear to stop in the hours prior to the dome  
326 collapse (vertical line in Figure 6), although a large amount of data is missing  
327 from this time period so it is not possible to tell the exact timing of the change  
328 from very similar events to non-similar events. This is significant because it may  
329 represent a time delay function between events occurring at depth and those at the



(a) *Left*: Dominant master waveform identified by stacking similar events (cross correlation coefficient greater than 0.7) from 8 to 12 July 2003 at station MBLG. A total of 79 events were used in the stack to create the average Master event waveform. *Right*: Single sided amplitude spectrum of the dominant master waveform.



(b) *Left*: Dominant master waveform identified by stacking similar events (cross correlation coefficient greater than 0.7) on 11 February 2010. A total of 3 events were used in the stack to create the average Master event waveform. *Right*: Single sided amplitude spectrum of the dominant master waveform.

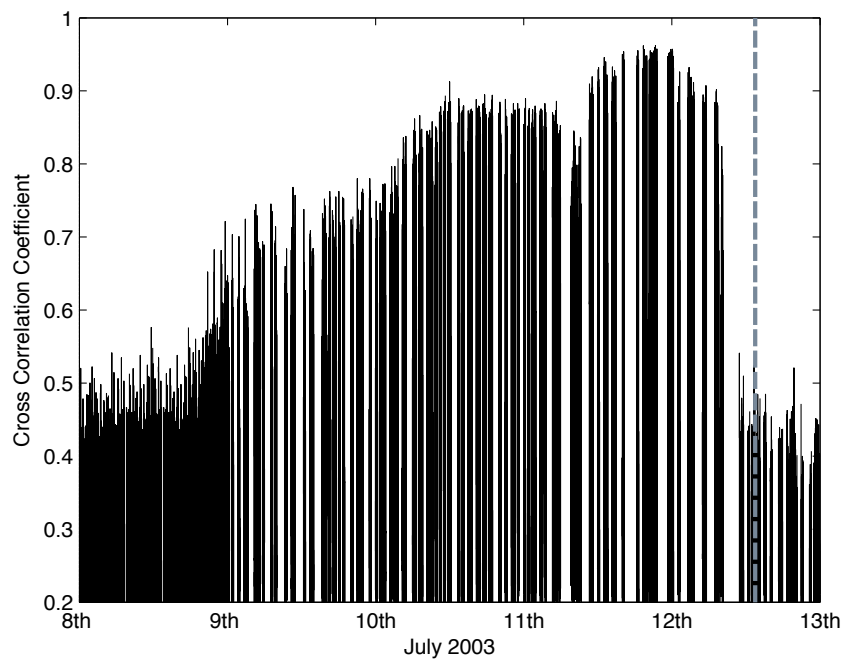


Figure 6: The evolution of the cross correlation coefficient with time: July 2003 with the Dominant Master Event, station MBLG. The dome collapse occurred at the time of the vertical line (13:30 on 12 July 2003). The gaps in the data represent gaps in the seismometer recordings rather than a dip in the cross correlation coefficient.



330 surface as first envisaged by Voight (1988).

331 *3.3. 8 - 11 February 2010*

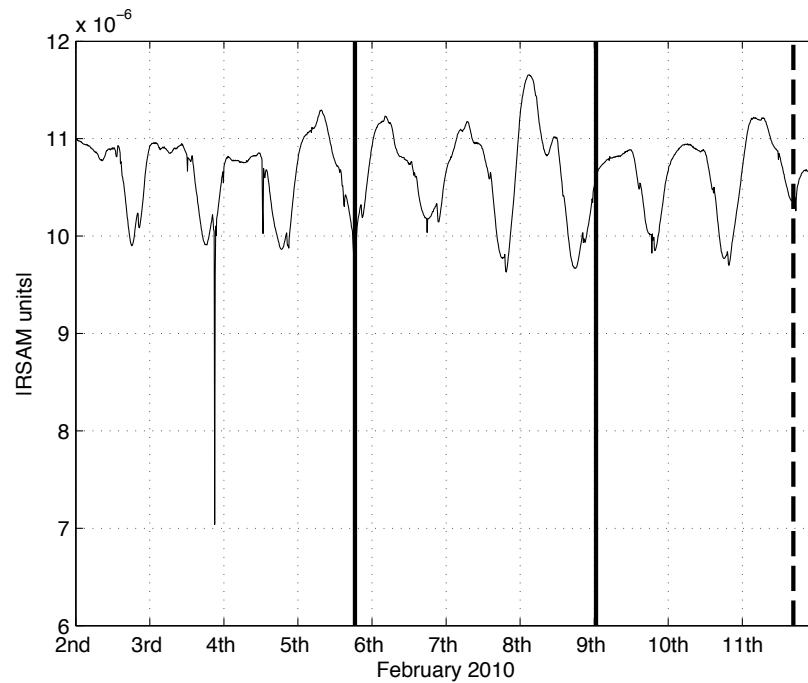


Figure 7: RSAM (10 minute averages) at Soufrière Hills volcano, 2 to 11 February 2010, station MBLG. RSAM is representative of average ground velocity in meters per second. The first two vertical lines (solid) represent two small Vulcanian explosions, on the 5 and 8 February. The final vertical line (dotted) represents the onset of the dome collapse and associated pyroclastic flows on 11 February.

332 Figure 7 suggests a clear cyclicality in seismicity in February 2010, which is  
333 unaffected by two small Vulcanian explosions in the days prior to the dome col-  
334 lapse. However, Stinton et al. (2014) reported very little precursory seismic ac-  
335 tivity before the dome collapse on 11 February 2010. Indeed, in order to identify  
336 individual seismic events from the continuous record, it was necessary to change  
337 the input STA/LTA parameters. In particular, the short term averaging window

338 length changed from 0.333 seconds (in 1997 and 2003) to 2 seconds (the maxi-  
339 mum period likely for a low frequency signal between the frequency of 0.5 and  
340 5 Hz), allowing the concentration on temporally longer events. The long term  
341 averaging window was modified from 60 to 120 seconds, again to accommodate  
342 longer events, meaning that very short events were not identified as an event. The  
343 trigger ratio value, the value at which the algorithm begins to detect an event, was  
344 moved from 4 to 8 since large amounts of noise appeared to dominate most of the  
345 signal. A total of 452 events were identified using this method between the 8 and  
346 12 February 2010.

347

348 However, of these 452 events identified in STA/LTA analysis, only 10 events  
349 were identified as true low frequency events, detected by filtering and manual in-  
350 spection. Of these 10 events, 3 were very similar to one another with a cross  
351 correlation coefficient of over 0.7. These events were subsequently stacked and a  
352 Master waveform produced (Figure 5(b)). The master waveform is much longer  
353 than the previous master events identified in 1997 and 2003, lasting  $\approx 30$  seconds.  
354 The distribution of energy is also very different, with energy existing over the  
355 ranges of 0 to 10 Hz, suggesting a hybrid nature, with a higher proportion of en-  
356 ergy concentrated above 2 Hz. However, on Montserrat, hybrid and low frequency  
357 earthquakes appear to occur on a continuum, with these two types bestowing the  
358 idealised end members (Neuberg et al., 2000). Therefore, although characteris-  
359 tically different to the master events identified prior to the 1997 and 2003 dome  
360 collapses, it is still assumed that the master event for 2010 is a low frequency type

361 earthquake related to the movement of magma at depth (Chouet, 1988; Neuberg  
362 et al., 2000).

#### 363 **4. Forecasting Volcanic Dome Collapses at SHV**

364 The FFM was first developed as a tool for identifying accelerating material  
365 creep and relating this to a slope failure; a cause and consequence of one single  
366 active system generating failure (Fukuzono, 1985). However, a volcanic system  
367 is inherently more complex, and accelerating magma ascent could be detected at  
368 several positions in the magma plumbing system with different phase delays and  
369 amplitudes. Therefore, in order for only one system to be analysed as input in  
370 the FFM it is necessary to focus only on one “family” of low frequency wave-  
371 forms which originate from the same source mechanism and location (Geller and  
372 Mueller, 1980; Neuberg et al., 2000; Thelen et al., 2011). Classification by wave-  
373 form similarity in addition to frequency content allows low frequency seismicity  
374 of a single source and depth to be exclusively analysed, meaning it is less likely  
375 that multiple sources become mixed in the forecast.

376

377 Hammer and Neuberg (2009) suggested that although individual swarm anal-  
378 ysis was not suitable for the FFM since a deceleration phase is almost always  
379 evident, consecutive swarm analysis might be at SHV. Instead of taking the event  
380 rate every 10 minutes from the continuous data, the event rate per 10 minutes is  
381 averaged across the entire duration of the seismic swarm, suggesting an overall  
382 acceleration in the event rate from swarm to swarm. This paper therefore also

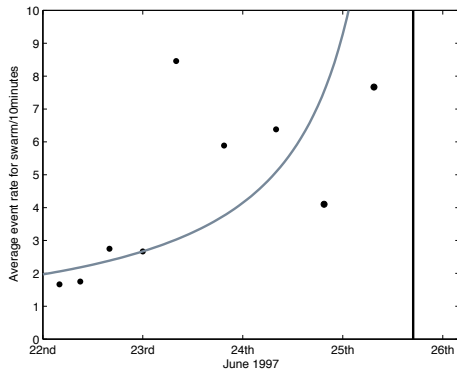
383 focuses on multiple swarm analysis for forecasting the timing of dome collapses  
384 at SHV.

#### 385 *4.1. 22 - 25 June 1997*

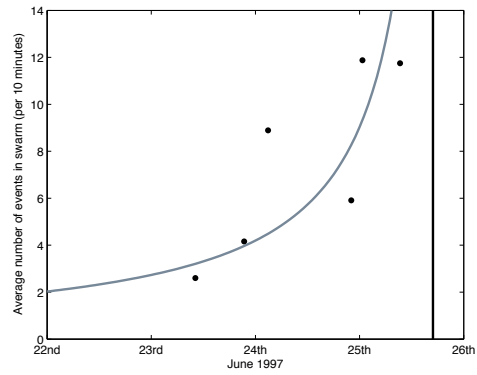
386 Using identified events from the STA/LTA algorithm, as used at MVO, Figure  
387 8a shows an initial acceleration in the average number of events per 10 minutes  
388 across swarms identified in June 1997, although the trace of the least squares fitted  
389 curve suggests a slowing of the acceleration up to the point of dome collapse (ver-  
390 tical line). Figure 8b shows the acceleration of swarms which have been identified  
391 using the dominant master event. Further classification of the low frequency seis-  
392 mic events into families appears to tighten the least squares fit and lead to a more  
393 convincing accelerating pattern of average number of events within 10 minutes of  
394 each swarm. This is further verified in Figure 8c, 8d and 8e which all show an  
395 acceleration in the average event rate with time up until the dome collapse, using  
396 master events 121, 136 and 141 respectively.

397

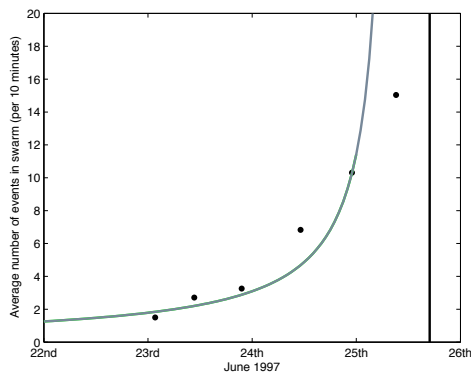
398 Figure 9 represents the application of the FFM to each of the accelerating  
399 event rates identified in Figure 8.  $\alpha = 2$  allowed graphical extrapolation to the  
400 forecasted timing of collapse as a simple linear regression. Inverse event rate  
401 trends were identified if at least three consecutive swarms formed an inverse trend.  
402 This was to try and eliminate spurious trends since a single decrease in the swarm  
403 event rate may be due to external factors, for example such as an increase in noise  
404 which obscures the number of events determined. Table 2 shows the results from



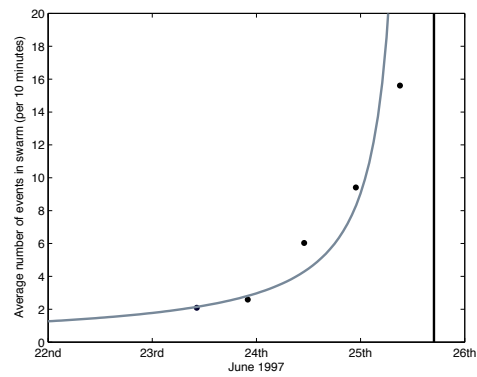
(a) All triggered low frequency seismicity



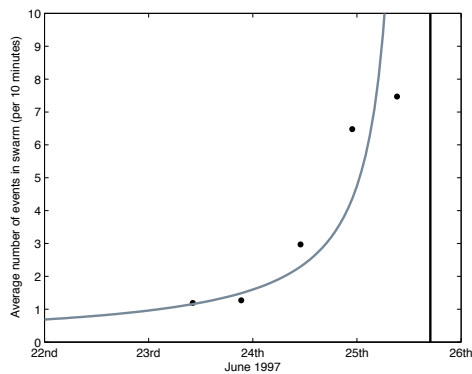
(b) Dominant master event



(c) Master event 121

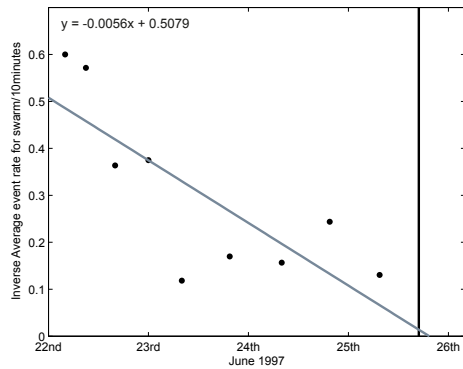


(d) Master event 136

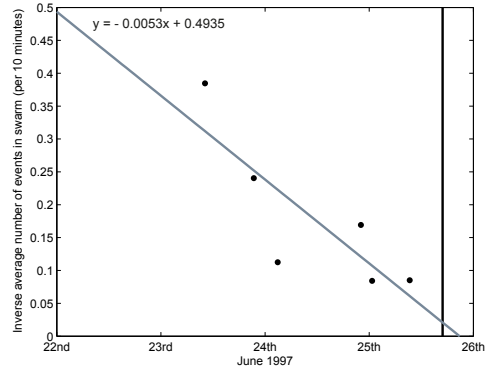


(e) Master event 141

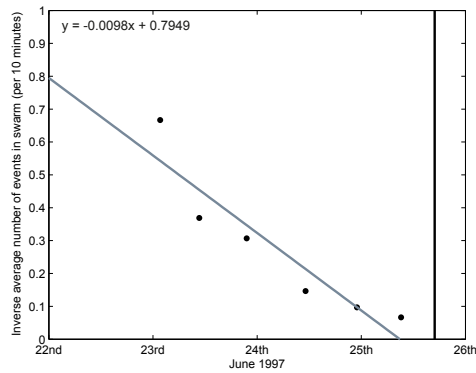
Figure 8: The average event rate per 10 minutes within swarms from 22 - 25 June 1997 at station MBLG. Each data point represents the average event rate for each individual swarm. The vertical line represents the known timing of dome collapse on the 25 June 1997 at 16:55 UTC. The acceleration of these events is depicted with the curve.



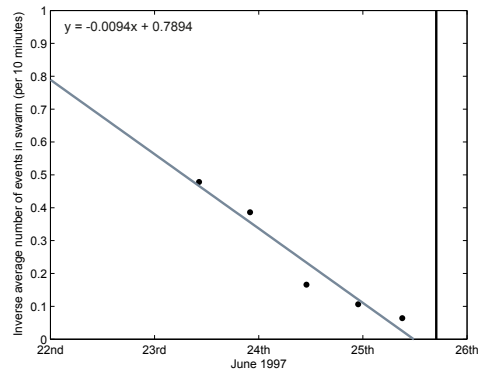
(a) All triggered low frequency seismicity



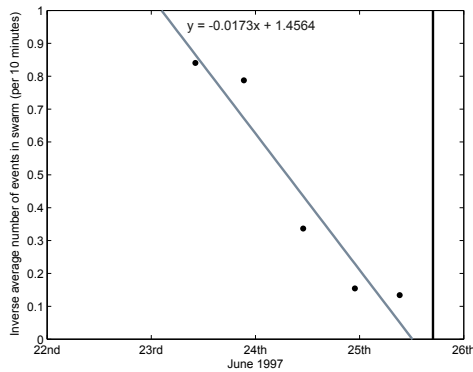
(b) Dominant master event



(c) Master event 121



(d) Master event 136



(e) Master event 141

Figure 9: Application of the FFM: the inverse average event rate per 10 minutes within swarms from 22 - 25 June 1997 at station MBLG. Each data point represents the inverse average event rate for each individual swarm. The vertical line represents the known timing of dome collapse on the 25 June 1997 at 16:55 UTC. The graphical representation of the FFM is depicted by the linear regression (it is assumed that  $\alpha = 2$ ) and the forecasted timing of failure can be read off the x-axis at the point where the linear regression crosses it.

Event	Known Timing (HH:MM)	Forecasted Timing (HH:MM)	Difference (HH:MM)	Forecasted early/late	$R^2$
Triggered Low frequency	88:55	91:24	02:29	Late	0.63
Dominant Master	88:55	92:50	03:55	Late	0.69
Master Event 001	88:55	95:12	06:17	Late	0.73
Master Event 010	88:55	n/a	n/a	n/a	n/a
Master Event 014	88:55	77:56	10:59	Early	0.60
Master Event 100	88:55	82:29	06:26	Early	0.86
Master Event 106	88:55	85:58	02:57	Early	0.59
Master Event 121	88:55	80:54	08:01	Early	0.87
Master Event 136	88:55	83:43	05:12	Early	0.94
Master Event 141	88:55	84:10	04:45	Early	0.92
Master Event 210	88:55	82:48	06:07	Early	0.84
Master Event 291	88:55	75:55	13:00	Early	0.83

Table 2: Timings of forecasted failure. Timings (known, forecasted and difference) are depicted in hours and minutes. The  $R^2$  value, as defined in the text, ranges from 0 to 1 and is the a scale of how well the model (FFM) can explain the data (event rate). Besides master events 014 and 106, each of the forecasts which were forecasted early using the FFM and a separate master event have a higher  $R^2$  value than those who were forecasted late, or when using all low frequency seismicity. Only master event 010 was unable to be used in analysis using the FFM since no acceleration in the event rate per swarm was identified. The  $R^2$  value for all triggered low frequency seismicity suggests that the FFM is inappropriate to describe the inverse event rates seen, since the linear regression does not fit well to the data.

405 the same analysis with all identified master events. When using triggered event  
406 data, although the timing of the forecasted dome collapse is within two hours of  
407 the known failure time, the fit of the linear regression (the FFM) to the data is poor.  
408 The application of the FFM to a single family of events in June 1997 has already  
409 been published in Hammer and Neuberg (2009) however, they fail to identify the  
410 family they used in analysis or the accuracy of the forecast in terms of fit to that  
411 data in a quantitative manner. Here, we show that any of the families of similar  
412 seismicity in June 1997 can be used as a forecasting tool with the FFM, with the  
413 exception of Master Event 010, which did not show an accelerating trend in the  
414 number of seismic events.

415

416 The forecasted timing of the dome collapse was never greater than 13 hours  
417 away from the known timing of collapse when using the cross correlation tech-  
418 nique first to identify similar events, and in most cases the collapse was forecasted  
419 early (Table 2). Despite increasing the difference between the known and fore-  
420 casted failure times, further classification of multiplets into families consistently  
421 allows for a better fit of the linear regression to the data, as can be seen from the  
422 high  $R^2$  values.

423

424 After Barrett (1974),  $R^2$  is defined as:

$$R^2 = \frac{\sum_{i=1}^n (y_i - \hat{y}_i)^2}{\sum_{i=1}^n (y_i - \bar{y})^2} \quad (3)$$



425 where  $y_i$  represents the observed parameter at position  $i$  (i.e. the inverse event  
426 rate at a given time),  $\hat{y}_i$  represents the predicted parameter of  $y$  at  $i$  (i.e. the FFM  
427 linear regression at this time), and  $\bar{y}$  represents the mean value of all of the  $y$  val-  
428 ues.  $R^2$  or the coefficient of determination is the proportion of variability which  
429 can be explained by the model, and ranges from a minimum value of 0 which  
430 suggests that the model does not explain any part of the data, up to a maximum of  
431 1 which suggests the model perfectly describes the data. In our case in particular,  
432 the  $R^2$  value shows how well future outcomes can be predicted by the model (the  
433 FFM), and therefore the closer the value is to one, the more confidence we can  
434 have in the forecast.  $R^2$  values of less than 0.65 are considered to represent a poor  
435 relationship between the observed data and the fitted FFM model.

436

437 Since seven out of eleven master events identified had  $R^2$  values of greater than  
438 0.7, it can be assumed that the FFM is appropriate for this data set. The  $R^2$  value  
439 for the dominant master is slightly lower at 0.69, however this still demonstrates  
440 a good fit between the model and the observations. Significantly, the  $R^2$  value for  
441 using the FFM with all triggered low frequency result gives a value of 0.63, which  
442 is deemed to not be a good fit. This is confirmed by the wide discrepancy between  
443 the observed event rates and the theoretical application of the FFM (Figure 9a).

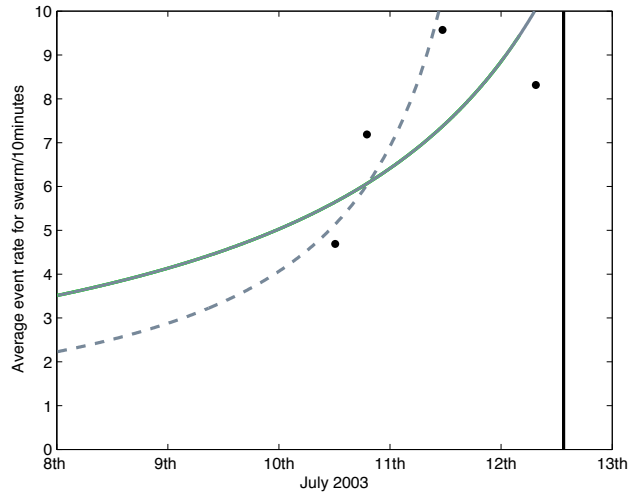
#### 444 4.2. 8 - 12 July 2003

445 Seismicity prior to the dome collapse in July 2003 did not take the form of  
446 well defined swarms, as observed in June 1997. Instead, seismicity was pulsatory

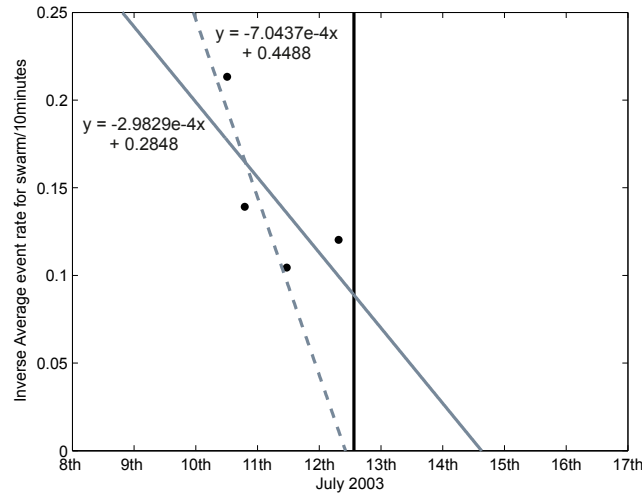
447 which was further highlighted by data gaps for MBLG. In addition, Figure 6 sug-  
448 gests that the cross correlation coefficient is constantly evolving, with highly sim-  
449 ilar events occurring throughout the precursory sequence and no evidence of an  
450 acceleration in the event rate. However “swarms” of events were identified in the  
451 triggered incoming seismicity as recorded by MVO in near real time (STA/LTA  
452 parameters unknown) and therefore the timings of these swarms were used with  
453 the FFM as a forecasting tool.

454

455 A clear acceleration can be seen in the swarms which occurred from the 10 to  
456 12 July 2003 (Figure 10(a)), followed by a slight deceleration in the event rate for  
457 the final swarm before the dome collapse (vertical line). Application of the FFM  
458 to all swarms identified (i.e. if the forecast was made on 12 July 2003 after the  
459 last swarm had ended) then the forecasted timing of collapse would be on 14 July  
460 2003 at approximately 15:00 h. However, the confidence in the forecast would be  
461 low, with an  $R^2$  value of 0.51. If only the first three swarms are used (i.e. only  
462 those swarms which exhibit an acceleration) then a forecast is made for 12 July  
463 at approximately 10:12 (dotted line in Figure 10(b)), just over 3 hours before the  
464 known timing of failure at 13:30 on 12 July. A greater amount of confidence can  
465 also be placed on this forecast, since the  $R^2$  value is 0.82, suggesting a significant  
466 relationship between the observed data points (event rate in each swarm) and the  
467 model (FFM). This is a significant improvement upon a forecast using all low  
468 frequency seismicity identified during this period, which forecasted a failure time  
469 at 17:50 h on 11 July, with an  $R^2$  value of 0.54.



(a) The average event rate per 10 minutes within swarms from 10 to 13 July 2003, station MBLG. The acceleration of all swarms is depicted with the solid curve and only the accelerating swarms by the dotted curve.



(b) Application of the FFM: the inverse average event rate per 10 minutes within swarms from 10 to 13 July 2003, station MBLG. The graphical representation of the FFM is depicted by the linear regression (it is assumed that  $\alpha = 2$ ) and the forecasted timing of failure can be read off the x-axis at the point where the linear regression crosses it. The solid regression includes data from all swarms; the dotted regression is only the swarms which were accelerating.

Figure 10: Acceleration of swarms of seismicity and application of the FFM: 10 to 13 July 2003. Each data point therefore represents the inverse average event rate for each individual swarm. The vertical line represents the known timing of dome collapse on the 12 July 2003 at 13:30.

470 *4.3. 8 - 11 February 2010*

471 Despite changing the STA/LTA parameters in order to identify seismic events  
472 during the precursory period of 8 to 11 February 2010, no additional seismicity  
473 from the continuous seismic data was identified using the cross correlation tech-  
474 nique. This not only suggests that there was very little precursory seismicity to the  
475 dome collapse, in terms of event counts or identified from accelerations in RSAM  
476 which simply appeared cyclic up to the collapse (Figure 7), but also suggests that  
477 at this time events which were detected were not similar. This is in stark contrast  
478 to both the dome collapses of 1997 and 2003 which were dominated by similar  
479 seismic events, and which showed an acceleration in seismicity prior to the col-  
480 lapse.

481

482 The fact that this collapse was not preceded by low frequency seismicity sug-  
483 gests that the collapse originated in processes unrelated to the movement of mag-  
484 matic fluid at depth in the days before the collapse or that this movement was  
485 aseismic in nature. Stinton et al. (2014) suggest that the collapse occurred due to  
486 over-steepening of the dome and talus which led to a gravitational collapse of the  
487 material. In the 4 months prior to the collapse, intensive extrusive and explosive  
488 activity had been observed, and since the collapse occurred in a piecemeal fashion  
489 over a number of hours, gravitational instability of a large dome is thought to have  
490 been a primary driving factor in collapse.

## 491 **5. Discussion**

### 492 *5.1. Characteristics of Seismicity Observed*

493 The seismicity prior to the dome collapse of June 1997 exhibited many fea-  
494 tures which have been commonly observed prior to eruptive events at SHV, and  
495 many other volcanoes worldwide. Not only is an increase in the number of seis-  
496 mic events observed in the days prior to collapse, but seismicity occurs in well de-  
497 fined swarms. The identification of multiplets within these swarms in June 1997  
498 and July 2003 echo the conclusions of Green and Neuberg (2006) and Petersen  
499 (2007) that a stable source process and location must be present in generating  
500 these events, and therefore the seismic energy must travel along similar ray paths  
501 to the seismometer. Further evidence for this comes from the fact that there is lit-  
502 tle clustering of similar events close to the diagonal of the cross correlation matrix  
503 (e.g. Figure 1). Slight clustering on the 24 June of events 250 to 350 is per-  
504 haps evident but not deemed significant. This suggests that very highly correlated  
505 events can be found over sustained periods of time (hours) and indicates very little  
506 change in the source conditions over this period (Caplan-Auerbach and Petersen,  
507 2005). Since the multiplets are repeated in swarms over a number of days, the  
508 source mechanism must be non-destructive and the trigger mechanism must be  
509 able to recharge quickly, since successive similar events occur in the continuous  
510 seismic record within seconds of one another. The identification of eleven families  
511 of multiplets each with their own waveform characteristics in June 1997 (Figure  
512 3) suggests that a number of source mechanisms and/or source locations must  
513 been active at SHV. This reflects the diversity of sources and physical processes

514 which act simultaneously at SHV. Unlike Green and Neuberg (2006), we did not  
515 find that any one multiplet family was more active than any others, suggesting that  
516 each of the source mechanisms were sustained for the duration of the precursory  
517 sequence. The identification of only one single active family in July 2003 suggests  
518 a change in source dynamics such that only one source was active during this time.

519

520 Using the quarter wavelength hypothesis of Geller and Mueller (1980), the  
521 repeating multiplets in June 1997 must occur within a maximum source distance  
522 of  $\approx 300\text{m}$ , assuming a dominant frequency of 2.1 Hz and an average P wave  
523 velocity of  $2500\text{ ms}^{-1}$  in the dome region as described by the current MVO veloc-  
524 ity model for SHV. Paulatto et al. (2010) however have suggested that the upper  
525 2.5 km of the dome at SHV could have a P wave velocity as low as  $1510\text{ ms}^{-1}$ ,  
526 therefore transforming the source volume to  $\approx 178\text{ m}$ . Furthermore, since this hy-  
527 pothesis is only really applicable for general seismic body waves, Neuberg et al.  
528 (2006) have suggested that for low frequency events to be deemed similar, the  
529 source location within a heterogeneous volcanic environment may vary by as lit-  
530 tle as one tenth of the wavelength. For our results this suggests that the each of the  
531 low frequency families in June 1997 could be located within a source volume of  
532  $\approx 120\text{ m}$  (using the MVO velocity model) or  $\approx 72\text{ m}$  (using the model of Paulatto  
533 et al. (2010)). This is in good agreement with De Angelis and Henton (2011) who  
534 located multiplet events from the same time period (22-25 June 1997) and suggest  
535 they consistently occur within a narrow depth range of 100 m to 300 m below sea  
536 level (approximately 1200 m to 1400 m beneath the volcano summit), within an

537 equally narrow longitude and latitude. Rowe et al. (2004) also relocated almost  
538 4000 similar seismic events from July 1995 to February 1996 and found that the  
539 source volume was approximately  $1km^3$ , with source dimensions of 10m to 100m  
540 in diameter. Using the dominant frequency of 4.03 Hz for the similar seismicity  
541 in July 2003, the quarter wavelength hypothesis states that the events must have  
542 occurred within a maximum distance of 155 m of one another. This is is good  
543 agreement with Ottemöller (2008) who found similar hybrid seismic events from  
544 swarms over the same time period were mostly located between 1500 and 1700 m  
545 below the dome summit, within a radius of  $\leq 150m$ .

546

547 Comparison of precursory seismicity before three large dome collapses at  
548 SHV suggests that precursory conditions are not stable or constant. This is best  
549 illustrated by the fact that 11 families of similar seismicity were identified in June  
550 1997, only one family of similar seismicity was identified in July 2003 and no  
551 families of similar seismicity were identified in February 2010. Stinton et al.  
552 (2014) has suggested that the lack of seismicity in February 2010 may be due to  
553 changes in the underlying processes of collapse, and that in 2010 gravity and over-  
554 steepening of the dome played a much larger part in the collapse event than the  
555 internal dynamics of moving magma. Forecasting dome collapse events at SHV  
556 cannot therefore solely rely upon seismicity, but must include analysis of dome  
557 size and shape in an equal capacity.

558 *5.2. The Similarity of Master Events in June 1997*

559 The similarity in frequency content between each of the master events identi-  
560 fied in June 1997 (Figure 3(b)) suggests that the master events themselves could  
561 be similar to one another. If only a shift in phase separates them from one another  
562 it could be indicative of a migrating source location at depth or a slight change  
563 in the active source process. In order to see if the master events were similar, we  
564 adopted the method of Thelen et al. (2011), who cross correlate all of the mas-  
565 ter events against one another, but use a higher cross correlation coefficient as a  
566 threshold for similarity ( $\geq 0.8$ ). This is a results of a trade off between maintain-  
567 ing an acceptable cross correlation coefficient ( $\geq 0.7$ ) for each of the individual  
568 events classified into each family, and reducing bias from the combination of mul-  
569 tiplets across several days. Table 3 shows that with a cross correlation coefficient  
570 of 0.8, the multiplets can be further organised into a number of subfamilies, for  
571 example master events 100, 106 and 210 can all be combined into a single multi-  
572 plet family with a cross correlation coefficient greater than the threshold. A lower  
573 threshold allows each of the multiplets to be placed into fewer families, however  
574 the visual similarity between the waveforms is lost. For example, with a cross  
575 correlation coefficient of 0.65 the multiplets can be sorted into just two families,  
576 however there is a much larger scatter in the waveform similarity.

577

578 Using a smaller subset of multiplets (i.e. by combining families which have a  
579 cross correlation coefficient  $\geq 0.8$ ) also provides an accurate forecast of the dome  
580 collapse event in June 1997. For example, combining master events 100, 106 and



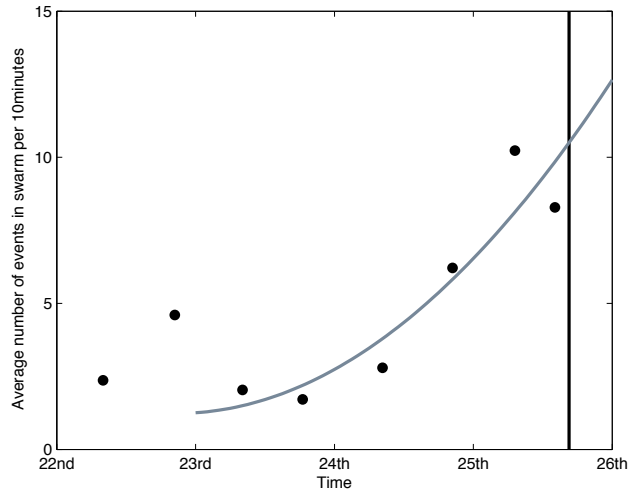
Master Event	001	010	014	100	106	121	136	141	210	291
<b>001</b>	1.000	0.778	0.567	0.590	0.669	0.727	0.699	0.700	0.453	0.682
<b>010</b>		1.000	0.694	0.389	0.462	0.590	0.736	0.359	0.468	0.599
<b>014</b>			1.000	0.776	0.614	0.699	<b>0.828</b>	0.536	0.734	0.666
<b>100</b>				1.000	<b>0.845</b>	0.783	0.656	0.794	<b>0.811</b>	0.707
<b>106</b>					1.000	0.574	0.384	0.754	0.592	0.463
<b>121</b>						1.000	0.789	0.762	0.732	<b>0.827</b>
<b>136</b>							1.000	0.573	0.776	0.798
<b>141</b>								1.000	0.668	0.701
<b>210</b>									1.000	0.794
<b>291</b>										1.000

Table 3: Cross Correlation Coefficients of each identified master event with every other master event. Values highlighted in bold represent those whose cross correlation coefficient exceeds the threshold of 0.8, and are therefore deemed to be similar. In this case, master events 100, 106 and 210 were stacked and an average waveform taken to provide a new master event. The same was done for master events 014 and 136, and 121 and 291.

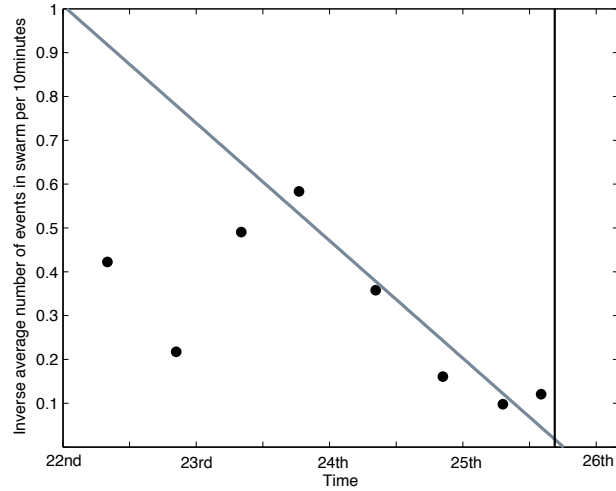
581 210 by stacking them at their maximum point of correlation and then taking an  
582 average of this stack, does not appear to produce a distinct acceleration which is  
583 tending towards a singularity (Figure 11(a)). However, upon application of the  
584 FFM we find that the timing of the dome collapse is forecasted less than 2 hours  
585 away from the known timing of the dome collapse, with a high degree of certainty  
586 between the linear regression and the inverse event rate ( $R^2 = 0.9054$ ) (Figure  
587 11), once a clear regression has been identified.

### 588 5.3. The Generation of Multiplets

589 It would be a remarkable coincidence if all of the successful forecasts of vol-  
590 canic eruptions using the FFM, whether in hindsight or real-time were chance  
591 occurrences, suggesting that some real link between the activity at depth and the



(a) The average event rate per 10 minutes within swarms from 22 - 25 June. The acceleration of these events is depicted with the curve.



(b) Application of the FFM: the inverse average event rate per 10 minutes within swarms from 22 - 25 June. The graphical representation of the FFM is depicted by the linear regression (it is assumed that  $\alpha = 2$ ) and the forecasted timing of failure can be read off the x-axis at the point where the linear regression crosses it.

Figure 11: The acceleration of swarms and application of the FFM to seismicity found in swarms which are a result of a new master event as a results of the combination of master events 100, 106 and 210. Each dot therefore represents the average event rate for each individual swarm. The vertical line represents the known timing of dome collapse on the 25 June 1997 at 16:55 UTC. Since swarms 1-3 did not produce a significant ongoing negative linear regression they were omitted from the application of the FFM. 41

592 surface must be plausible. A number of models have been proposed to explain  
593 the occurrence of low frequency seismicity in volcanic settings. The model of  
594 Iverson et al. (2006) suggests that the generation of low frequency seismicity oc-  
595 curs as a magmatic plug moves incrementally upwards within a conduit due to the  
596 movement of buoyant magmatic fluid behind the plug. In this instance it would  
597 expected for seismicity to migrate with the movement of the plug, and therefore  
598 become shallower with time. This is not observed on Montserrat, where seismic-  
599 ity consistently occurs at  $\approx 1500$  m depth below the dome summit (Aspinall et al.,  
600 1998; Rowe et al., 2004; Ottemöller, 2008; De Angelis and Henton, 2011). In  
601 addition, families of similar seismic waveforms must occur within a small spatial  
602 extent in order to maintain their similarity, and single families have been observed  
603 being sustained over a number of hours and days. These observations at Soufrière  
604 Hills volcano are inconsistent with this model, since the same family of earth-  
605 quakes would not be sustained over this period of time without major changes to  
606 the similarity of the waveforms. The evolution of cross correlation coefficients,  
607 such as that which were observed in July 2003, would require only a very small  
608 migration of the seismicity which would not be expected from the incremental  
609 movement of a volcanic plug, since it can still be classed as from the same family  
610 over the sustained period of time. This model also fails to explain the occurrence  
611 of a number of distinct families as observed in June 1997, which are repeatedly  
612 activated then deactivated.

613

614 The more recent model of Bean et al. (2014) suggests that slow rupture failure

615 within unconsolidated volcanic material in the edifice can induce low frequency  
616 seismicity, whose waveform characteristics are fundamentally dependent upon the  
617 wave propagation path. It is envisaged that families and swarms of low fre-  
618 quency seismicity develop due to slow deformation at a number of points within  
619 the upper edifice where the stress is reduced which could be induced by gas influx,  
620 gravity or magma migration. On Montserrat, the seismicity systematically occurs  
621 at the same depth such that it would require the stress drop to be maintained at  
622 the same location over a number of years. In addition, no explanation is given  
623 for the clear acceleration in seismicity which has been observed prior to the dome  
624 collapse events studied.

625

626 The generation of low frequency seismicity has also been attributed to the brittle  
627 failure of the magma itself (Webb and Dingwell, 1990; Goto, 1999) through  
628 an increase in viscosity of the melt and/or high strain rates (Lavallée et al., 2008)  
629 as the melt enters a glass transition stage. In volcanic environments, conditions  
630 which may induce this glass transition stage may include: changes in crystal  
631 and/or bubble content of the magma (Goto, 1999); an increase in the ascent rate  
632 of magma (Neuberg et al., 2006); or through the introduction of a restriction in  
633 the conduit (Thomas and Neuberg, 2012). The brittle failure model allows for  
634 the acceleration in LF seismicity observed at Soufrière Hills prior to dome col-  
635 lapses, since accelerations in magma ascent has been shown to increase the strain  
636 rate within a volcanic conduit simulation, and therefore instigate the brittle failure  
637 of the melt (Neuberg et al., 2006). A much simpler way to induce an increased

638 strain rate is to introduce a constriction within the conduit (Thomas and Neuberg,  
639 2012). Such a constriction fits with observations at Soufrière Hills that LF seis-  
640 micity consistently occurs at  $\approx 1500$  m below the dome summit (Aspinall et al.,  
641 1998; Rowe et al., 2004; Ottemöller, 2008; De Angelis and Henton, 2011), as well  
642 as the fact that multiple LF sources (i.e. families of similar seismic events) may  
643 be active at any given time, since a number of locations may exist where the strain  
644 rate threshold for brittle failure is overcome.

645

646 This model not only accounts for the acceleration in the number of multiplets  
647 that is observed and the stable source mechanism and its location, but can also  
648 account for the possible phase delay between the timing of failure at depth (i.e.  
649 when a magmatic pathway is created) and the surface expression of this failure  
650 (Voight, 1988). This could be represented by a time delay in the FFM, which  
651 forecasts the timing of failure before the known timing of the dome collapse since  
652 technically the FFM is forecasting the failure at depth related to the accelerating  
653 seismicity, rather than the surface manifestation that we observe and try to relate  
654 it to.

#### 655 *5.4. Forecasting Potential*

656 The concept of forecasting using the FFM in hindsight analysis once the erup-  
657 tion has occurred is common (e.g. Cornelius and Voight (1994); Kilburn and  
658 Voight (1998); De la Cruz-Reyna and Reyes-Dávila (2001); Ortiz et al. (2003);  
659 Hammer and Neuberg (2009); Smith and Kilburn (2010)). It is much less common

660 to employ and rely upon these tactics during developing unrest as huge responsi-  
661 bility is placed upon generating accurate forecasts which are often simply plagued  
662 with too many uncertainties. Using a cross correlation technique in conjunction  
663 with the FFM allows the isolation of single system at depth. Isolating a single sys-  
664 tem at depth avoids additional uncertainties introduced by averaging data over a  
665 number of different accelerating phenomena, and consequently reduces the misfit  
666 between the data and the forecast. On occasions when precursory seismicity could  
667 be identified prior to dome collapses at SHV in June 1997 and July 2003, use of  
668 similar seismicity and the FFM provided a more successful and more accurate  
669 forecast to the timing of collapse events than simply using the FFM in isolation  
670 with all incoming low frequency seismicity. Further investigation is required to  
671 determine whether these techniques are applicable to other volcanoes around the  
672 world, and whether it can be used to forecast volcanic phenomena other than dome  
673 collapses.

674

675 If this technique is to be used in real time forecasting, the identification of  
676 similar seismic events will also need to be undertaken in real time, i.e. a contin-  
677 uous search for similar events and defining of master events will have to become  
678 part of routine monitoring. Experience at Soufrière Hills volcano teaches us that  
679 some similar seismic events are consistent over a number of days (e.g. June 1997,  
680 July 2003) and therefore their early identification would allow simple manipu-  
681 lation of the incoming seismicity. In other circumstances, the master events for  
682 similar seismicity may be short-lived, which requires the recalculation of master

683 events more often. There are also occasions when similar seismic events may not  
684 be apparent in the seismicity at all (e.g. Soufrière Hills, February 2010) and other  
685 forecasting methods will need to be relied upon.

686

687 Other parameters used in forecasting and the identification of similarity will  
688 also need to be undertaken in real time. In this study, we have found that the use  
689 of 0.7 as a correlation threshold is appropriate for identifying different families of  
690 similar seismicity. However, Ottemöller (2008) used a cross correlation threshold  
691 of between 0.6 and 0.66 for continuous data from 9 to 12 July 2003 at SHV, allow-  
692 ing the identification of far more similar seismic events. Upon visual inspection of  
693 our data, we found these cross correlation thresholds to be too low to identify sim-  
694 ilar seismicity without noise. The determination of a cross correlation threshold  
695 is extremely important since if it is placed too low then there is a risk of placing  
696 events which are not similar into the same family, and if it is too high there may  
697 be many similar events which are not detected due to poor signal to noise ratios.  
698 Identifying this threshold in real time is likely to be another difficult parameter,  
699 and further analysis of more swarms of similar seismicity is required to determine  
700 whether it is always appropriate to use 0.7 as a threshold at SHV.

701

702 As the FFM follows a least squares regression analysis when  $\alpha$  is equal to 2,  
703 the residual error between the observed event rate and the mean event rate should  
704 follow a typical Gaussian distribution (Bell et al., 2011b). Greenhough and Main  
705 (2008) have suggested that since earthquake occurrence is a point process, the

706 rate uncertainties are best described by a Poisson distribution. This was also sug-  
707 gested by Bell et al. (2011a) who determined that the daily earthquake rates at  
708 Mauna Loa preceding the 1984 eruption were consistent with a Poisson regime,  
709 within 95 per cent confidence limits. In this instance, a generalised linear model  
710 (GLM) where  $\alpha = 1$ , rather than a least squares regression model ( $\alpha = 2$ ) may be  
711 more appropriate, since it can allow for a distribution of data that is non-Gaussian  
712 (Bell et al., 2011b). However, although using a GLM as a fitting tool does provide  
713 a higher  $R^2$  value for each of the forecasts generated from similar seismicity (al-  
714 ways  $> 0.8$ ), suggesting a GLM is a better fit to the data than a least squares linear  
715 regression, the forecasted timing of the dome collapse in 1997 was consistently  
716 late (the best forecast was still over 30 hours away from the known timing of fail-  
717 ure). Moreover, forecasting using a GLM for the July 2003 collapse generated a  
718 forecast over a week from the known timing of the collapse when using all of the  
719 available data. If only the accelerating swarms are used, the forecast is over 72  
720 hours after the known timing of failure. This is in contrast to Bell et al. (2011b)  
721 who suggest that the GLM provides more accurate forecasts for the timings of  
722 eruptions than the FFM with a linear least squares regression.

723

724 Although the GLM solves the problem of needing to use a model which can  
725 account for the appropriate error structure, its use may not be applicable in a  
726 volcanic setting (Hammer and Ohrnberger, 2012). Forecasting volcanic eruptions  
727 using the FFM and rates of temporal seismicity requires that the system has a  
728 memory, and therefore that events which have occurred before can influence the



729 outcome in the future. A poisson process is the exact opposite to this: it requires a  
730 memoryless system, in which events evolve independently. This would therefore  
731 make the use of the GLM with  $\alpha = 1$  and FFM together invalid, since one of the  
732 overriding assumptions of the FFM is that previous geophysical observables form  
733 the basis of the forecast, and therefore suggesting that the system has a memory.  
734 Analysis of a number of dome building eruptions at Mt. St. Helens in 1985 and  
735 1986, which may be comparable to Soufrière Hills volcano, suggested that an  
736 exponential model did not adequately explain the precursory trends in earthquake  
737 event rates (Bell et al., 2013), possibly implying that an exponential model is  
738 not appropriate for forecasting at andesitic dome building volcanoes, which may  
739 be one reason as to why the forecasts using a GLM in this instance were not  
740 successful.

## 741 **6. Conclusions**

742 Utilizing the cross correlation technique for the forecasting of large scale dome  
743 collapses at SHV appears to provide more consistent and more accurate forecasts  
744 than when using all precursory low frequency seismicity, since we can assume  
745 that we are forecasting using only one active system at depth. The potential for an  
746 improved forecast by isolating a single system was first proposed by Kilburn and  
747 Voight (1998) who suggested that two populations may be in effect at SHV in a  
748 qualitative manner from graphics alone (their Figure 2(c)). Here we show that this  
749 is indeed the case by a quantitative method of cross correlation for dome collapses  
750 in June 1997 and July 2003. The most recent dome collapse at SHV in February

751 2010 was not able to be forecast using the FFM since no precursory seismicity  
752 was detected. This suggests that not every dome collapse at SHV is forecastable,  
753 and suggests some events may come without warning. However, when precur-  
754 sory seismicity is detected, the cross correlation technique allows for a five-fold  
755 increase in the number of detectable events from the continuous record, and there-  
756 fore provides a more holistic picture of the ongoing seismicity. The magnitude of  
757 the events and the possible superposition of events is insignificant for detection  
758 since the continuous seismic record is normalised in near real-time and the search  
759 criteria can be set to smaller increments within the cross correlation procedure to  
760 find closely spaced events.

761

762 In June 1997, 10 families of similar waveforms were detected, signalling a  
763 number of active sources at depth occurring at the same time. Using any one of  
764 these families of seismicity in conjunction with the FFM provided more accurate  
765 forecasts to the timing of the dome collapse, with a greater degree of confidence  
766 due to high  $R^2$  values which suggest that the model (FFM) fits to the data well  
767 (Figure 9 and Table 2). In July 2003, only one family of events were identified, in  
768 agreement with Ottemöller (2008). Analysis of the cross correlation coefficients  
769 suggests a migration of the source with time (Figure 6). Significantly, and in  
770 contrast to the events of June 1997, the similar seismicity ceased hours before the  
771 dome collapse, perhaps an indication of a delay function between the seismicity at  
772 depth and the collapse at the surface as first envisaged by Voight (1988). Forecast-  
773 ing using only the accelerating swarms in July 2003 provided an accurate forecast

774 for the timing of the dome collapse, but proves the difficulty of using the FFM in  
775 real time, as the last swarm in the sequence led to a forecast a number of weeks  
776 from the known timing of collapse. Despite clear cyclic activity in RSAM (Fig-  
777 ure 7), no families of similar seismic events could be identified in the precursory  
778 seismicity of the February 2010 collapse. This echoes the conclusions of Stinton  
779 et al. (2014) who suggested that no acceleration in seismicity was observed prior  
780 to the collapse, and in fact, seismicity remained remarkably low. Further inves-  
781 tigation is required to determine whether the cross correlation technique used in  
782 conjunction with the FFM is applicable for forecasting other volcanic phenomena  
783 at other volcanoes around the world.

784

785 The overwhelming question in forecasting volcanic eruptions however still  
786 remains: how can we tell the difference between accelerating seismicity which  
787 is precursory to an eruptive event, and that which does not appear to lead to a  
788 surface expression? This is fundamental for forecasting, since the generation of  
789 false alarms can lead to deteriorating confidence in the observatory making the  
790 forecasts. False alarms are an inherent part of forecasting volcanic eruptions; the  
791 forecasts are never going to be 100% correct, 100% of the time, primarily due to  
792 the incidental nature of nature itself. However, keeping false alarms to a minimum  
793 is essential. Forecasting is complicated by precursory activity often having more  
794 than one acceleration event, and in particular at SHV having cyclic acceleration  
795 events (e.g. Figure 7). The FFM does not account well for multiple accelerations  
796 within a system, which is why searching for an overall acceleration in the precur-

797 sory activity (e.g. taking the average event rate of each swarm of activity, rather  
798 than per unit time) allows a more successful application. The use of the cross  
799 correlation technique in conjunction with the FFM appears to further enhance the  
800 success of the forecast since focus is on a single active seismic system at depth  
801 and therefore can be directly related to surface activity.

802

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